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Collision Energy Dependence of p_t Correlations in Au+Au Collisions at RHIC

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We present two-particle p_t correlations as a function of event centrality for Au+Au collisions at $\sqrt{s_{NN}} = 7.7, 11.5, 14.5, 19.6, 27, 39, 62.4,$ and 200 GeV at the Relativistic Heavy Ion Collider using the STAR detector. These results are compared to previous measurements from CERES at the Super Proton Synchrotron and from ALICE at the Large Hadron Collider. The data are compared with UrQMD model calculations and with a model based on a Boltzmann-Langevin approach incorporating effects from thermalization. The relative dynamical correlations for Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV show a power law dependence on the number of participant nucleons

and agree with the results for Pb+Pb collisions at $\sqrt{s_{NN}} = 2.76$ TeV from ALICE. As the collision energy is lowered from $\sqrt{s_{NN}} = 200$ GeV to 7.7 GeV, the centrality dependence of the relative dynamical correlations departs from the power law behavior observed at the higher collision energies. In central collisions, the relative dynamical correlations increase with collision energy up to $\sqrt{s_{NN}} = 200$ GeV in contrast to previous measurements that showed little dependence on the collision energy.

PACS numbers: 25.75.-q

The study of event-by-event correlations and fluctuations in global quantities can provide insight into the properties of the hot and dense matter created in Au+Au collisions at ultrarelativistic collision energies [1–21]. Correlations of transverse momentum, p_t , have been proposed as a measure of thermalization [15, 22, 23] and as a probe for the critical point of quantum chromodynamics (QCD) [14, 24]. A detailed study of the dependence of two-particle p_t correlations on collision energy and centrality may elucidate the effects of thermalization. If the matter produced in ultrarelativistic collisions passes through a possible QCD critical point, the fluctuations are predicted to increase with respect to a baseline of uncorrelated emission. A possible signature of the critical point could be non-monotonic behavior of two-particle correlations as a function of collision energy in central collisions.

In this paper we present an experimental study of the collision energy dependence of p_t correlations using Au+Au collisions at center of mass energies ranging from $\sqrt{s_{NN}} = 7.7$ GeV to 200 GeV, taken during the RHIC Beam Energy Scan (BES) using the Solenoidal Tracker at RHIC (STAR). The 7.7-, 11.5-, 39-, and 62.4-GeV data were taken in 2010. The 19.6-, 27-, and 200-GeV data were taken in 2011. The 14.5-GeV data were taken in 2014. The main detectors used were the Time Projection Chamber (TPC) [25] and the Time of Flight detector (TOF) [26], both located in a solenoidal magnetic field of 0.50 T. Charged tracks from the TPC with $0.2 \text{ GeV}/c \leq p_t \leq 2.0 \text{ GeV}/c$ and $|\eta| < 0.5$ were used in this analysis, where η is the pseudorapidity. Tracks in the TPC were characterized by the distance of closest approach (DCA), which is the smallest distance between the projection of the track and the measured event vertex. To suppress secondary particles from weak decays, all tracks were required to have a DCA less than 1 cm. Each track was required to have at least 15 measured points and a ratio of the number of measured points to the possible number of measured points greater than 0.52. Each event was required to have at least one track matched to a TOF hit to minimize pileup. For each collision energy, events were accepted if they originated from within 1 cm of the center of the focused beam in the plane perpendicular to the beam axis and within 30 cm of the center of STAR along the beam line to achieve uniform detector performance. The statistical errors were determined by dividing the dataset into five subsets and calculating the observables for each subset. The standard deviation of these observables divided by the square root of the number of subsets

$\sqrt{s_{NN}}$ (GeV)	Events (M)
7.7	1.43
11.5	2.46
14.5	12.0
19.6	15.4
27	28.7
39	24.8
62.4	14.9
200	22.2

TABLE I. Summary of the number of events analyzed in this analysis.

was used to calculate the error. We estimated the systematic errors of the observables by studying the effects of varying the DCA cut from 0.8 to 1.2 cm, varying the acceptance in η from $|\eta| = 0.4$ to 0.6, and by varying the lower cut for p_t from 0.18 to 0.22 GeV/ c . The average relative systematic errors related to the DCA, the η cut, and the lower cut of p_t are 1.3%, 2.7%, and 4.0% respectively.

All the data shown are from minimum bias triggers, which were defined as a coincidence of the signals from the Zero Degree Calorimeters (ZDC) [27], the Vertex Position Detectors (VPD) [28], and/or the Beam-Beam Counters (BBC) [29]. Table I shows the number of events analyzed at each collision energy. The centrality bins were defined in terms of a reference multiplicity, which was defined as the number of detected charged particles within an acceptance of $0.5 < |\eta| < 1.0$. For the 200 GeV data, this quantity was corrected for the luminosity dependence and the position of the event vertex along the beam axis. This centrality was defined so that the particles used to determine the event centrality did not include the particles used to calculate the p_t correlations. The centrality bins used in this analysis were defined in terms of the fraction of total inelastic cross section. Specifically the bins were 0-5% (most central collisions), 5-10%, 10-20%, 20-30%, 30-40%, 40-50%, 50-60%, 60-70%, and 70-80% (most peripheral). The average number of participating nucleons, N_{part} , was calculated for each centrality bin at each collision energy using a Monte Carlo Glauber model [30, 31].

The results are compared with calculations using the UrQMD model [32, 33]. Version 3.3 of UrQMD with default parameters was used for Au+Au collisions at RHIC energies and version 3.4 was used for Pb+Pb collisions at $\sqrt{s_{NN}} = 2.76$ TeV. UrQMD is a hadronic transport

model that does not incorporate effects from a deconfined system of quarks and gluons. For comparison to the STAR results, the STAR acceptance and tracking efficiency were applied with a dependence on particle type, p_t , collision energy, and centrality. The detector efficiencies were first obtained from simulations and then applied to the UrQMD results. For comparison to the ALICE results, the ALICE acceptance was applied to the UrQMD calculations but no efficiency effects were considered.

To characterize the p_t correlations, we used the two-particle p_t correlation defined as the covariance given by:

$$\langle \Delta p_{t,i}, \Delta p_{t,j} \rangle = \frac{1}{N_{\text{events}}} \sum_{k=1}^{N_{\text{events}}} \frac{C_k}{N_k (N_k - 1)}, \quad (1)$$

where

$$C_k = \sum_{i=1}^{N_k} \sum_{j=1, j \neq i}^{N_k} (p_{t,i} - \langle p_t \rangle) (p_{t,j} - \langle p_t \rangle). \quad (2)$$

N_{events} is the number of events, N_k is the number of tracks in the k th event, and $p_{t,i}$ is the transverse momentum of the i th track in the given event. The event-averaged p_t is defined as

$$\langle p_t \rangle = \frac{\sum_{k=1}^{N_{\text{events}}} \langle p_t \rangle_k}{N_{\text{events}}}, \quad (3)$$

where $\langle p_t \rangle_k$ is the average p_t of the k th event defined as

$$\langle p_t \rangle_k = \frac{\sum_{i=1}^{N_k} p_{t,i}}{N_k}. \quad (4)$$

The quantities $\langle p_t \rangle$ and $\langle \Delta p_{t,i}, \Delta p_{t,j} \rangle$ were calculated as a function of the reference multiplicity and then averaged over the centrality bin to remove any dependence on the size of the centrality bins [34].

To characterize two-particle p_t correlations, we present the relative dynamical correlation, $\sqrt{\langle \Delta p_{t,i}, \Delta p_{t,j} \rangle} / \langle p_t \rangle$. The relative dynamical correlation represents the magnitude of the dynamic fluctuations of the average transverse momentum in units of $\langle p_t \rangle$ and can be compared directly to the observables used by CERES [35] and ALICE [36].

Figure 1 shows the relative dynamical correlation $\sqrt{\langle \Delta p_{t,i}, \Delta p_{t,j} \rangle} / \langle p_t \rangle$ as a function of centrality for eight collision energies. Also shown in this figure are the UrQMD calculations. The measured relative dynamical correlations for Au+Au collisions at 200 GeV are well reproduced by a power law given by $22.32\% / \sqrt{N_{\text{part}}}$. This power law distribution is also shown for the other seven collision energies. The relative dynamical correlation distributions deviate from this power law with decreasing collision energy. Figure 1 also shows the ratio of the measured relative dynamical correlation to the power

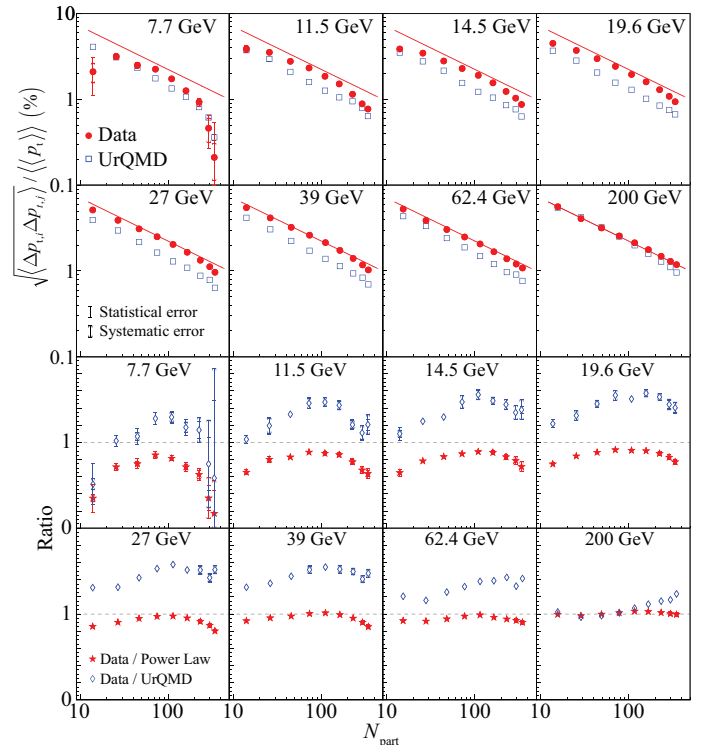


FIG. 1. The relative dynamical correlation $\sqrt{\langle \Delta p_{t,i}, \Delta p_{t,j} \rangle} / \langle p_t \rangle$ as a function of N_{part} for eight collision energies and UrQMD calculations. Statistical and systematic errors are shown. The solid straight lines represent a power law given by $22.32\% / \sqrt{N_{\text{part}}}$. Also shown are the ratios of the measured data to the power law and to UrQMD calculations.

law distribution observed at 200 GeV and the ratio of the measured values to the UrQMD calculations at each collision energy.

The previous STAR measurements of the relative dynamical correlation at 19.6, 62.4, 130, and 200 GeV [16] used different acceptance cuts including $0.15 \text{ GeV}/c \leq p_t \leq 2.0 \text{ GeV}/c$ and $|\eta| < 1.0$ as well as a different centrality definition using detected charged particles with $|\eta| < 0.5$. The previous data at 19.6, 62.4, and 200 GeV are consistent with the current data.

Figure 2 shows the UrQMD results for the relative dynamical correlation for three cases. The first case is the direct output from the model. The second case is UrQMD in which the effect of an 80% constant tracking efficiency was introduced. The third case is the method used in this paper in which the UrQMD calculations are obtained by introducing the effect of the STAR tracking efficiency, which depends on the particle type, the particle p_t , the collision energy, and the collision centrality. These calculations show that the relative dynamical correlation is not sensitive to the efficiency, which allows for the presentation of the experimental results without correction for tracking efficiency.

Figure 3 shows the comparison of a Boltzmann-

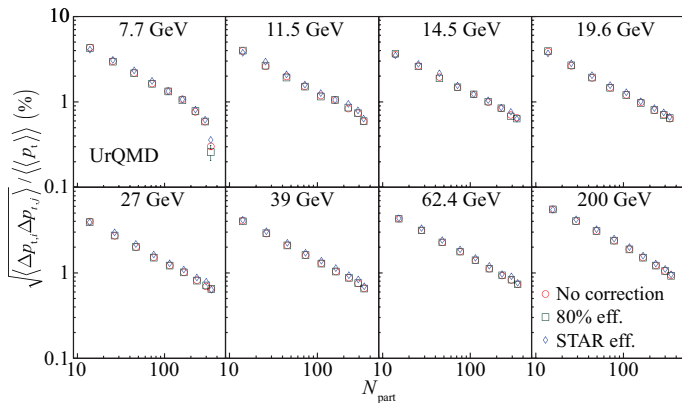


FIG. 2. UrQMD calculations for the relative dynamical correlation $\sqrt{\langle \Delta p_{t,i} \Delta p_{t,j} \rangle} / \langle \langle p_t \rangle \rangle$ as a function of N_{part} for eight collision energies. Three cases are shown; UrQMD uncorrected, which has no efficiency corrections, UrQMD 80% eff., which uses a fixed efficiency of 80%, and UrQMD, which uses a tracking efficiency that depends on particle type, p_t , centrality, and collision energy.

Langevin approach to the study of equilibration and thermalization effects on two-particle p_t correlations [23]. The results for local equilibrium flow and partial thermalization are shown compared with current results for 19.6 and 200 GeV collisions as well as the results from Pb+Pb collisions at 2.76 TeV [36]. The local equilibrium flow predictions are realized using a blast wave model including the fluctuation of thermalized flow while a time dependent relaxation time is used to obtain the partial thermalization results. The authors of Ref. [23] point out that these comparisons suggest incomplete thermalization in peripheral collisions because they disagree with a local equilibrium flow model. The agreement in Fig. 3 of the model calculations for partial thermalization with the measured two-particle p_t correlations at all centralities at these three widely-spaced collision energies lends support to this model.

Figure 4 shows the relative dynamical correlation for Au+Au collisions at 7.7 and 200 GeV compared with similar results from Pb+Pb collisions at 2.76 TeV [36]. The ALICE collaboration determined the relative dynamical correlation using tracks with $0.15 \text{ GeV}/c \leq p_t \leq 2.0 \text{ GeV}/c$ and $|\eta| < 0.8$. The results for Au+Au collisions at 200 GeV agree well with the results for Pb+Pb collisions at 2.76 TeV. The dashed line represents a power law fit to the STAR Au+Au data at 200 GeV of the form $22.32\% / \sqrt{N_{\text{part}}}$. This fit also reproduces the ALICE Pb+Pb results at 2.76 TeV except for the most central collisions. Not only does the relative dynamical correlation scale as a power law, but it scales as $1/\sqrt{N_{\text{part}}}$, adding credence to the idea that the observed particle production comes from uncorrelated sources. As the collision energy is lowered, the relative dynamical correlation as a function of N_{part} shows a breakdown in this power law scaling as demonstrated by the results for

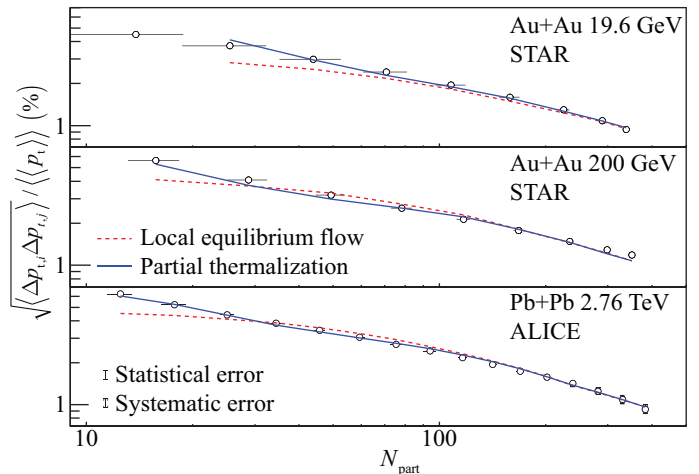


FIG. 3. Comparison of a model [23] incorporating a Boltzmann-Langevin approach to the calculation of thermalization effects for the relative dynamical correlation from Au+Au collisions at $\sqrt{s_{\text{NN}}} = 19.6$ and 200 GeV. Also shown are model comparisons to results from Pb+Pb collisions at $\sqrt{s_{\text{NN}}} = 2.76$ TeV [36].

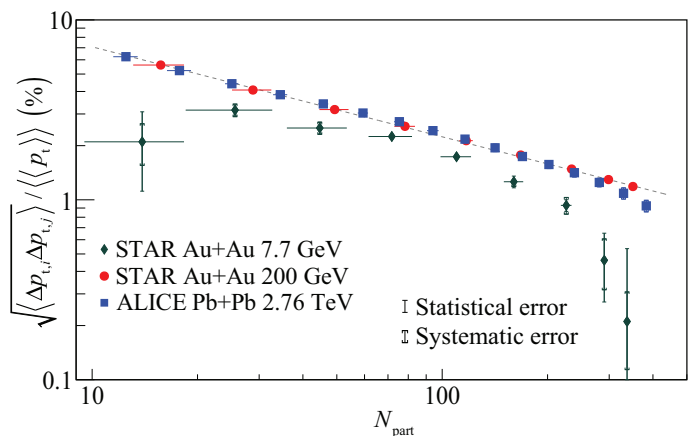


FIG. 4. The relative dynamical correlation $\sqrt{\langle \Delta p_{t,i} \Delta p_{t,j} \rangle} / \langle \langle p_t \rangle \rangle$ for $\sqrt{s_{\text{NN}}} = 7.7$ GeV and 200 GeV Au+Au collisions compared with similar results from Pb+Pb collisions at $\sqrt{s_{\text{NN}}} = 2.76$ TeV [36]. The dashed line represents a fit to the data at $\sqrt{s_{\text{NN}}} = 200$ GeV given by $22.32\% / \sqrt{N_{\text{part}}}$. Statistical and systematic errors are shown.

7.7 GeV in Fig. 4.

Figure 5 shows the relative dynamical correlation $\sqrt{\langle \Delta p_{t,i} \Delta p_{t,j} \rangle} / \langle \langle p_t \rangle \rangle$ as a function of $\sqrt{s_{\text{NN}}}$ for the most central bin (0-5%). Also shown are the results for Pb+Pb collisions from ALICE [36] and Pb+Pb collisions from CERES [35]. UrQMD calculations are shown as described above.

The data from CERES [35] in Fig. 5 are from Pb+Pb collisions at $\sqrt{s_{\text{NN}}} = 8.7, 12.3$ and 17.3 GeV. The CERES results were published using an observable (Σp_t), which is mathematically identical to $\sqrt{\langle \Delta p_{t,i} \Delta p_{t,j} \rangle} / \langle \langle p_t \rangle \rangle$. STAR had shown previously [16] that the CERES results

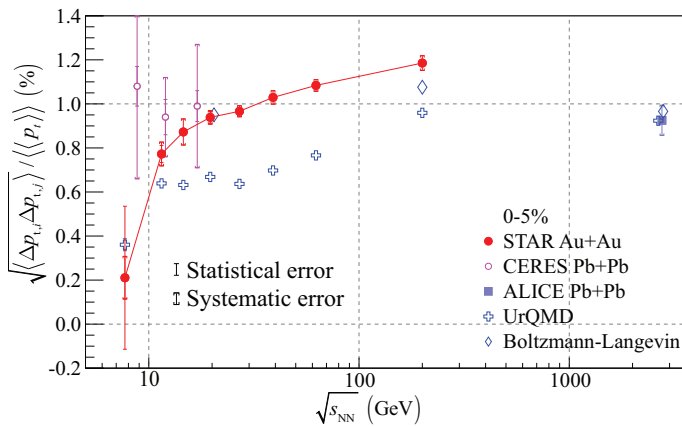


FIG. 5. The relative dynamical correlation $\sqrt{\langle \Delta p_{t,i} \Delta p_{t,j} \rangle} / \langle \langle p_t \rangle \rangle$ for Au+Au collisions as a function of collision energy for the 0-5% centrality bin along with results for Pb+Pb from CERES [35] and results for Pb+Pb from ALICE [36] along with UrQMD calculations and results from Boltzmann-Langevin model calculations [23]. The solid line is drawn to guide the eye. Statistical and systematic errors are shown for the data points.

taken together with STAR results at 19.6, 62.4, 130, and 200 GeV for 0-5% centrality indicated that the relative dynamical correlation was constant with collision energy. The present results for 7.7 GeV to 200 GeV, although in reasonable agreement with the CERES data, show that the relative dynamical correlation decreases at lower collision energy.

Figure 5 also shows the relative dynamical correlation for the 5% most central collisions from Pb+Pb collisions at 2.76 TeV from ALICE [36]. This result seems to show that the relative dynamical correlation plateaus above 200 GeV. The relative dynamical correlation at 2.76 TeV is somewhat lower than the value at 200 GeV. This difference could be partially due to the fact that the 0-5% centrality bin for Pb+Pb collisions at 2.76 TeV is associated with a somewhat larger value of N_{part} than the value for 200 GeV Au+Au collisions, leading to a lower value of the relative dynamical correlation assuming a $1/\sqrt{N_{\text{part}}}$ scaling.

The UrQMD calculations agree with the measured relative dynamical correlation for Au+Au central collisions at 7.7 GeV and with the relative dynamical correlation for Pb+Pb collisions at 2.76 TeV. However, the measured relative dynamical correlation increases more than the calculated values from UrQMD as the collision energy is increased from 7.7 GeV to 200 GeV. Also shown in Fig. 5 are the predictions of the Boltzmann-Langevin calculations [23] for central collisions of Au+Au at 19.6 and 200 GeV and central collisions of Pb+Pb at 2.76 TeV. These results show little dependence on the collision energy and agree with the measured results at 19.6 GeV and 2.76 TeV but slightly under-predict the results at 200 GeV.

Figure 6 shows the relative dynamical correlation for

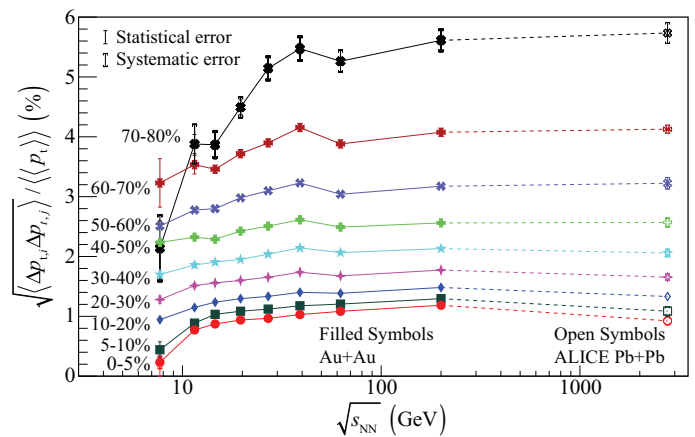


FIG. 6. The relative dynamical correlation $\sqrt{\langle \Delta p_{t,i} \Delta p_{t,j} \rangle} / \langle \langle p_t \rangle \rangle$ for Au+Au collisions grouped by centrality, compared to Pb+Pb collisions at 2.76 TeV from ALICE [36]. The solid and dashed lines are drawn to guide the eye. The filled symbols refer to the present Au+Au results while the open symbols refer to Pb+Pb results from ALICE [36]. Statistical and systematic errors are shown.

Au+Au collisions along with Pb+Pb collisions from ALICE [36] grouped by the collision centrality. The results for the most central collisions (0-5%) increase as the collision energy is increased and tend to plateau at the highest collision energy. The mid-central centrality bins show less dependence on the collision energy. In the most peripheral bin, the relative dynamical correlations again increase with collision energy. For all centralities, the relative dynamical correlations seem to plateau at the highest collision energy.

In conclusion, we observe a power law scaling of the form $1/\sqrt{N_{\text{part}}}$ for the relative dynamical correlation in Au+Au collisions at 200 GeV. A similar power law scaling had been previously observed in Pb+Pb collisions at 2.76 TeV [36] except in the most central collisions. As the collision energy for Au+Au collisions is decreased to 7.7 GeV, the power law scaling observed at 200 GeV breaks down. For the most central Au+Au collisions, the relative dynamical correlations increase with collision energy up to 200 GeV and seem to reach a plateau from 200 GeV to 2.76 TeV. We observe no non-monotonic scaling of the measured relative dynamical correlation in central collisions. We observe that two-particle p_t correlations show evidence of incomplete thermalization when compared with the Boltzmann-Langevin model in Ref. [23] in the most peripheral collisions. New calculations from this model at collision energies below 19.6 GeV would be of interest to better determine the extent of thermalization.

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- [1] H. Heiselberg and A. D. Jackson, Phys. Rev. C **63**, 064904 (2001).
- [2] Z.-w. Lin and C. M. Ko, Phys. Rev. C **64**, 041901 (2001).
- [3] S. Jeon and V. Koch, Phys. Rev. Lett. **85**, 2076 (2000).
- [4] S. A. Bass, P. Danielewicz, and S. Pratt, Phys. Rev. Lett. **85**, 2689 (2000).
- [5] S. A. Voloshin, V. Koch, and H. G. Ritter, Phys. Rev. C **60**, 024901 (1999).
- [6] M. Stephanov, K. Rajagopal, and E. Shuryak, Phys. Rev. D **60**, 114028 (1999).
- [7] H. Heiselberg, Phys. Rept. **351**, 161 (2001).
- [8] M. Asakawa, U. Heinz, and B. Müller, Phys. Rev. Lett. **85**, 2072 (2000).
- [9] Q. Liu and T. A. Trainor, Phys. Lett. B **567**, 184 (2003).
- [10] E. V. Shuryak and M. A. Stephanov, Phys. Rev. C **63**, 064903 (2001).
- [11] C. Pruneau, S. Gavin, and S. Voloshin, Phys. Rev. C **66**, 044904 (2002).
- [12] S. A. Bass, M. Gyulassy, H. Stöcker, and W. Greiner, J. Phys. G **25**, R1 (1999).
- [13] M. Stephanov, K. Rajagopal, and E. Shuryak, Phys. Rev. Lett. **81**, 4816 (1998).
- [14] M. Stephanov, Phys. Rev. D **65**, 096008 (2002).
- [15] S. Gavin, Phys. Rev. Lett. **92**, 162301 (2004).
- [16] J. Adams *et al.* (STAR Collaboration), Phys. Rev. C **72**, 044902 (2005).
- [17] G. Agakishiev *et al.* (STAR Collaboration), Phys. Lett. B **704**, 467 (2011).
- [18] J. Adam *et al.* (ALICE Collaboration), Phys. Rev. Lett. **118**, 162302 (2017).
- [19] L. Adamczyk *et al.* (STAR Collaboration), Phys. Rev. Lett. **112**, 032302 (2014).
- [20] L. Adamczyk *et al.* (STAR Collaboration), Phys. Rev. Lett. **113**, 092301 (2014).
- [21] L. Adamczyk *et al.* (STAR Collaboration), Phys. Lett. B **785**, 551 (2018).
- [22] S. Gavin and G. Moschelli, Phys. Rev. C **85**, 014905 (2012).
- [23] S. Gavin, G. Moschelli, and C. Zin, Phys. Rev. C **95**, 064901 (2017).
- [24] M. A. Stephanov, Phys. Rev. Lett. **102**, 032301 (2009).
- [25] K. H. Ackermann *et al.* (STAR Collaboration), Nucl. Instrum. Methods A **499**, 624 (2003).
- [26] W. J. Llope, Nucl. Instrum. Methods A **661**, S110 (2012).
- [27] C. Adler, A. Denisov, E. Garcia, M. Murray, H. Strobele, and S. White, Nucl. Instrum. Methods A **499**, 433 (2003).
- [28] W. J. Llope *et al.*, Nucl. Instrum. Methods A **759**, 23 (2014).
- [29] C. A. Whitten (STAR Collaboration), AIP Conference Proceedings **980**, 390 (2008).
- [30] M. L. Miller, K. Reygers, S. J. Sanders, and P. Steinberg, Ann. Rev. Nucl. Part. Sci. **57**, 205 (2007).
- [31] B. I. Abelev *et al.* (STAR Collaboration), Phys. Rev. C **79**, 034909 (2009).
- [32] M. Bleicher *et al.*, J. Phys. G **25**, 1859 (1999).
- [33] S. Bass *et al.*, Prog. Part. Nucl. Phys. **41**, 255 (1998).
- [34] X. Luo, J. Xu, B. Mohanty, and N. Xu, J. Phys. G **40**, 105104 (2013).
- [35] D. Adamova *et al.* (CERES Collaboration), Nucl. Phys. A **727**, 97 (2003).
- [36] B. Abelev *et al.* (ALICE Collaboration), Eur. Phys. J. C **74**, 3077 (2014).