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### Publication Date

2014-02-01

### DOI

10.1016/j.lithos.2013.08.020

Peer reviewed



# The Dupal isotopic anomaly in the southern Paleo-Asian Ocean: Nd–Pb isotope evidence from ophiolites in Northwest China



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## ARTICLE INFO

### Article history:

Received 27 February 2013

Accepted 27 August 2013

Available online 3 September 2013

### Keywords:

Ophiolite

Paleo-Asian Ocean

Indian MORB-type mantle domain

Dupal anomaly

## ABSTRACT

It has been suggested that the Dupal isotopic anomaly in the mantle can be traced in the Paleozoic ophiolites from the Neo- and Paleo-Tethyan Ocean (275–350 Ma). The Karamaili ophiolite (KO) and Dalabute ophiolite (DO) in the eastern and western corners, respectively, of the Junggar basin in NW China represent remnants of the relatively older (>350 Ma) Paleo-Asian Ocean (PAO) crust. Thus, these ophiolites can provide additional constraints on the long-term composition and evolution of the Paleozoic suboceanic mantle. We present new major-trace element and Sr, Nd and high-precision Pb isotope data for the basalts, gabbros and a plagioclase separate from the KO and DO. Our results indicate that the PAO crust indeed has a Dupal-like isotopic signature. In detail, all samples have relatively low  $\epsilon_{\text{Nd}(t)}$  and high  $^{208}\text{Pb}/^{204}\text{Pb}(t)$  for given  $^{206}\text{Pb}/^{204}\text{Pb}(t)$  ratios (i.e., positive  $\Delta 8/4$  values), similar to the Dupal isotopic characteristics of Indian Ocean mid-ocean ridge basalts (MORB). The trace element signature of DO mafic rocks is similar to that of normal- and enriched-MORB whereas that of the KO is transitional between MORB and arc basalt. Therefore, the DO mantle domain reflects the PAO asthenosphere and the KO domain additionally shows the influence of the subduction process. Geochemical modeling using Th/Nd as well as Nd and Pb isotopic ratios indicates that up to 2% subduction component had been added to a depleted Indian MORB-type mantle to produce the bulk of KO rocks. The subduction component in the KO rocks consisted of variable proportions of  $\leq 1\%$  partial melt of unradiogenic sediment similar to modern Izu–Bonin trench sediment and hydrous fluid dehydrated from the subducted altered oceanic crust. The Devonian asthenospheric mantle beneath the southern PAO is isotopically heterogeneous, but lends support to the idea that the Dupal isotopic anomaly existed prior to the opening of the Indian Ocean. Finally, plate tectonic reconstruction indicates that the anomaly was present in the Neo- and Paleo-Tethyan oceans in the southern hemisphere and in the southern part of PAO in the northern hemisphere during the late Paleozoic.

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## 1. Introduction

Mid-ocean ridge basalts (MORB) from the Indian Ocean are isotopically distinct from Pacific–North Atlantic MORB in that they have higher  $^{208}\text{Pb}/^{204}\text{Pb}$ ,  $^{207}\text{Pb}/^{204}\text{Pb}$ ,  $^{87}\text{Sr}/^{86}\text{Sr}$  and lower  $^{143}\text{Nd}/^{144}\text{Nd}$  for a given  $^{206}\text{Pb}/^{204}\text{Pb}$  ratio (Dupre and Allegre, 1983; Hamelin and Allegre, 1985; Hart, 1984; Mahoney et al., 1992, 1998; Xu and Castillo, 2004; Xu et al., 2002). This isotopic signature of Indian MORB is typical of oceanic lavas with the so-called Dupal isotopic anomaly; these lavas are believed to originate from two large regions in the lower mantle

with atypical isotopic and seismic velocity characteristics (Castillo, 1988; Hart, 1984). One possible explanation for the occurrence of Dupal lower mantle domains is that these were produced early in the Earth's history by core–mantle–crust differentiation processes (Castillo, 1988; Hart, 1984). Such a genesis requires that the domains are long-lived regional features of the mantle (Xu and Castillo, 2004; Xu et al., 2002). Thus, understanding the long-term history and evolution of the sub-Indian mantle may lead to a better understanding of the genesis of the compositional and structural heterogeneity of the lower mantle.

Ophiolites, which are pieces of ocean floor that are occasionally preserved on land (e.g., Dilek, 2003; Hawkins, 2003), are an important source of geologic information on the formation and subsequent closure of oceanic basins (Moores and Jackson, 1974; Shervais, 2001; Shervais et al., 2005). The trace element and isotopic characteristics of ophiolites generally reflect those of the composition and evolution of their underlying mantle (e.g., Mahoney et al., 1998; Pfänder et al., 2002; Xu and

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Castillo, 2004; Xu et al., 2002; Zhang et al., 2005). For example, previous geochemical and isotopic studies of the late Paleozoic Neo- and Paleozoic ophiolites (Xu and Castillo, 2004; Xu et al., 2002; Zhang et al., 2005) showed that the bulk of Indian suboceanic mantle was most probably inherited from the Tethyan asthenosphere, which formerly occupied much of the present Indian Ocean region. Thus, these studies proposed that the Dupal isotopic anomaly existed prior to the opening of the Indian Ocean. However, what was the regional extent and how far back in time did the distinct mantle isotopic characteristics exist?

The extinct Tethys and Paleo-Asia Ocean (PAO) co-existed during the Paleozoic. Ophiolites representing relict fragments of the Tethyan and PAO crusts are preserved along the Tethys and PAO suture zones. The PAO is believed to have existed from ~1000 to ~300 Ma (Dobretsov et al., 1995; Dobretsov et al., 2003). It is slightly older than the late Paleozoic Neo- and Paleozoic Tethyan Ocean (275–350 Ma; Xu and Castillo, 2004; Xu et al., 2002; Zhang et al., 2005). The Central Asian Orogenic Belts (CAOB), which is also known as the Altai tectonic collage (Sengör and Natal'in, 1996; Sengör et al., 1993), is a subduction-accretionary complex associated with the evolution and closure of the PAO. The Karamaili ophiolite (KO) and Dalabute ophiolite (DO) are the two largest Devonian ophiolite belts in the Xinjiang region in northern China; these represent relict pieces of the crust of the southern branch of the PAO (W.J. Xiao et al., 2009) in the CAOB (Jahn, 2004). As in other typical ophiolites, most of the original rock sequences in the KO and DO are not preserved completely, but they have almost all the igneous components of a classic ophiolite (mantle peridotites, cumulate ultramafics, gabbros, and volcanics). Thus, the chemical variation of mafic rocks from the KO and DO can provide key information on the sub-PAO mantle and magma differentiation processes that formed these ophiolites. Systematic isotopic studies, especially involving combined Sr–Nb–Pb isotopes, on PAO ophiolites have not been carried out thus far.

The main objective of this paper is to constrain the geochemical and isotopic composition of the sub-PAO mantle in order to constrain the temporal and spatial extent of the Indian Ocean-type asthenospheric mantle. To attain this objective, we performed a comprehensive major-trace element and Sr, Nd and high-precision Pb isotopic investigation of KO and DO mafic rocks and combined these with existing major and trace element data (Liu et al., 2007, 2009b). Our main conclusion is that the PAO was indeed underlain by an Indian Ocean-type mantle domain as early as late Paleozoic, but caution must be taken because the subduction process had clearly modified the original composition of the mantle beneath the majority of ophiolites.

## 2. Geological background

The southern portion of CAOB is in the northern Xinjiang region of China (Fig. 1a). From north to south, it is composed of the Altai Junggar and Tienshan terranes. The Junggar terrane is traditionally further subdivided into East Junggar and West Junggar terranes and Junggar basin. The East and West Junggar are two accretionary terranes located to the east and west of Junggar basin, respectively, and thought to be the result of northward subduction of the Paleozoic PAO crust beneath the Siberian Plate (Coleman, 1989; Sengör et al., 1993; W.J. Xiao et al., 2009; Xiao et al., 2004a, 2004b; Zhang et al., 2009).

### 2.1. Karamaili ophiolite (KO)

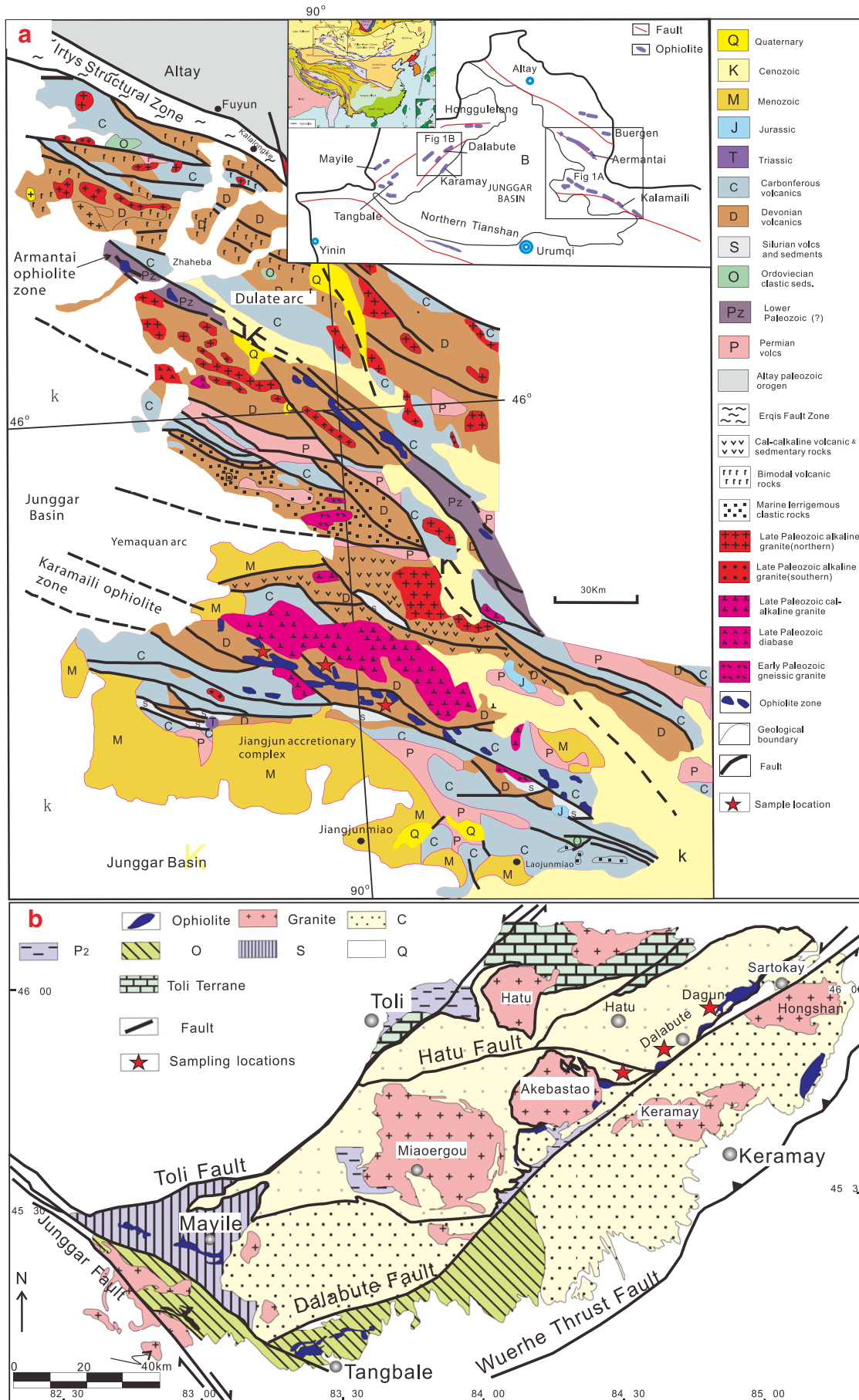
The KO is one of the two partial and dismembered ophiolite sequences in the southern part of East Junggar; the other is the Armantai

ophiolite in the central part of East Junggar (Fig. 1a). Associated with these ophiolites are two remnant arcs, the Dulate–Baytag arc in the northern side of Armantai ophiolite and the Yemaquan arc in the northern side of KO. The KO is part of a forearc accretionary complex (W.J. Xiao et al., 2009), which is similar to modern circum-Pacific forearc accretionary belts in that it contains sedimentary and volcanic rocks, ophiolitic fragments that include extrusive sequences, plutonic complex and upper mantle units, and the bulk of its accretionary prism is variably metamorphosed to greenschist facies. The overall structural geometry, abundance and types of the rock suites present, petrogenetic relationships between rock suites, and variations of deformation and metamorphism strongly resemble those of the Japan island arc orogens (W.J. Xiao et al., 2009).

In detail, the KO forms a NWW-trending belt and crops out mainly along the Karamaili fault for about 150 km (Fig. 1a). It extends from Kamste in the northwest and continues southeastwards through the Moqinwula Mountains, farther east through the China–Mongolia border, and finally to the Zoolen ophiolite belt in Mongolia (Helo et al., 2006). The western extension of the KO belt is covered by Cenozoic sediments in the Junggar basin (Li et al., 2003). There are two types of outcrops in the KO belt: ophiolite sequences and associated series of volcanic rock units. Although the entire KO belt is dismembered, individual fragments of relatively intact ophiolite sequences are present. These fragments consist of large volumes of serpentinized peridotites, serpentinites, and cumulate pyroxenites; these upper mantle units, are overlain by plutonic complexes including gabbros, diabases, sparse sheeted dikes and felsic plutons, and, finally, upper extrusive sequences including basalts, red cherts and green siliceous mudstones containing Devonian and Carboniferous radiolaria (Cai, 1986; He et al., 2000; Shu and Wang, 2003). The ultramafic rocks occur as pods and lenses and are strongly serpentinized although they still contain relict olivine and pyroxene. Serpentinites are mainly metamorphosed harzburgite, dunite and minor lherzolite, and host minor pods of chromitites. The mafic rocks mainly occur as tectonic slices containing cumulate and isotropic gabbros. There are also some associated, purportedly younger volcanic rock units and these consist mainly of massive basalt flows and some pillows.

Previous reconnaissance studies indicated that the mafic rocks (gabbros and basalts) in the KO are mainly transitional in composition between MORB-like to island arc tholeiite (IAT)-like (Liu et al., 2007, 2009b); others have also reported basalts with enriched-MORB, ocean island basalt (OIB) and boninite characteristics (Bu et al., 2005; Ma, 2007; Ouyang, 2006; Zheng et al., 2007). There are also silicic rocks associated with the KO. Most of these intruded the mafic rocks and include many variably shaped bodies of interlayered diorites, quartz-diorites, tonalites, and plagiogranites in the northern side of the ophiolite (Fig. 1a). Some plagiogranite bodies intersect the gabbros discordantly or display direct contact relationship suggesting contemporaneity. The rest of the intermediate-silicic rocks consist of andesites and dacites. The age of KO has been determined by U–Pb zircon dating, and includes 336 Ma and 342 Ma for the gabbros (Jian et al., 2005), 403 Ma (Jian et al., 2005) and 373 Ma (Tang et al., 2007) for the plagiogranites and 371 Ma for the tonalites (Liu et al., 2009b). These age data indicate that our samples of the KO crust range in age from 371 Ma to 403 Ma; these are consistent with the Devonian and Early Carboniferous age of the radiolarian cherts (Cai, 1986; He et al., 2000; Shu and Wang, 2003). Finally, although the KO was definitely formed in a subduction environment, the details of its formation are still in dispute. Some argue that the KO was formed in a back-arc basin (Li, 1995; Li et al., 2003; Ma, 2007) or in a forearc region

**Fig. 1.** Geological sketch map of a) the East Junggar region and b) the west Junggar region. Inset is a simplified tectonic map of central Asia (after Jahn, 2004) and Northern Xinjiang (after Xiao et al., 2008; W. Xiao et al., 2009) showing the distribution of ophiolites. The Central Asian Orogenic Belt (CAOB) is located between Siberian craton in the north and Tarim–North China craton in the south. The northern Xinjiang region of China is composed from north to south of the Altai, Junggar and Tienshan terranes. The ophiolite belts are mainly distributed around Junggar basin. The approximate location of the northern Xinjiang region of China is indicated. Panel A is modified after Li et al. (2003) and W.J. Xiao et al. (2009). Panel B is modified after Liu et al. (2009c) and Lei et al. (2008).



(Wang et al., 2003; Xiao et al., 2004a, 2004b), but more recently it was suggested to be a section of an arc setting that experienced ridge subduction (Liu et al., 2007, 2009b).

## 2.2. Dalabute ophiolite (DO)

The DO is one of the largest of the partial and dismembered ophiolite sequences in west Junggar (Fig. 1b). The DO forms a NE–SW-trending belt and crops out mainly along the Dalabute fault for about 100 km. It extends from Miaoergou in the southwest and continues southeastwards through the Dagon, an outcrop covering an area of ~50 km<sup>2</sup>. The DO consists of tectonic slices that are mainly distributed in three locations along the Dalabute fault: Akebastao in the south, Dalabute in the middle, and Dagon in the north. Among these locations, the longest exposure of the ophiolite belt is in Dalabute, where it consists of at least 9 mafic and ultramafic outcrops. Similar to the KO, the entire ophiolite igneous sequences ranging from bottom ultramafic units to top mafic units are present. These include large volumes of serpentinized peridotites, serpentinites, cumulates, sparse diabases, sheeted dikes, and upper extrusive sequence of basalts. Above the igneous sequence are red cherts that also contain Devonian radiolarian (Xiao et al., 1992).

Previous studies indicated that the bulk composition of DO mafic rocks (gabbros and basalts) is akin to normal- and enriched-MORB and OIB (Liu et al., 2009c; Zhang et al., 2011). The DO age has been reported by many studies, and includes a Sm–Nd isochron age of 395 Ma for the gabbros (Zhang and Huang, 1992) and more recent U–Pb zircon ages of 391 Ma (Gu et al., 2009) and 332 Ma (Xu et al., 2006) for the gabbros. Another previous study reported a 302 Ma age for a gabbro, but this belongs to a later suite of mafic rocks (Liu et al., 2009c). Although quite variable, these igneous ages of the ophiolite combined with the Devonian age of the radiolarian cherts (Xiao et al., 1992) give a conclusive Early Devonian age (391–332 Ma) of the DO, although as noted earlier some younger magmas related to the ophiolite are also present.

## 3. Petrography

The igneous rocks from the different sections of KO and DO have distinctive petrographic characteristics, but in general there is no petrographic difference between KO and DO rock suites. Both whole rock sequences had been affected by oceanic hydrothermal alteration and subjected to sub-greenschist facies condition; thus, the extent of alteration is variable. Basalts, the most abundant rock in both volcanic suites, are black to gray greenish in outcrops. They are vesicular in the massive flows, amygdaloidal in the pillows, and predominantly aphyric to micropphyric with equigranular to intergranular groundmass consisting of plagioclase, pyroxene and minor anhedral titanomagnetite. Plagioclase feldspars occur as euhedral to subhedral columnar laths that are 0.03 mm to 1 mm in length. Many samples contain euhedral plagioclase with little or no alteration, but some samples contain secondary mineral replacements along the edges of grains or along twin planes within crystals; other samples experienced serious alteration and showed replacement of the plagioclase by albite, zeolite and prehnite. Pyroxenes range in size from 0.05 mm to 1 mm and are generally unzoned and variably altered to chlorite, magnetite, actinolite, and epidote.

Plutonic rocks in both ophiolites are mainly gabbros. In outcrops gabbros are massive and dark gray in color and in thin sections they are coarse- to fine-grained, isotropic and contain variable amounts of olivine, pyroxenes, plagioclase feldspar, amphibole, and titanomagnetite. The gabbros are variably altered. Some of the olivines are pseudomorphed by chlorite and clay. Pyroxenes and plagioclase are only slightly altered and are the most abundant phases in the gabbros with modes ranging from 60% to 70%.

## 4. Analytical techniques

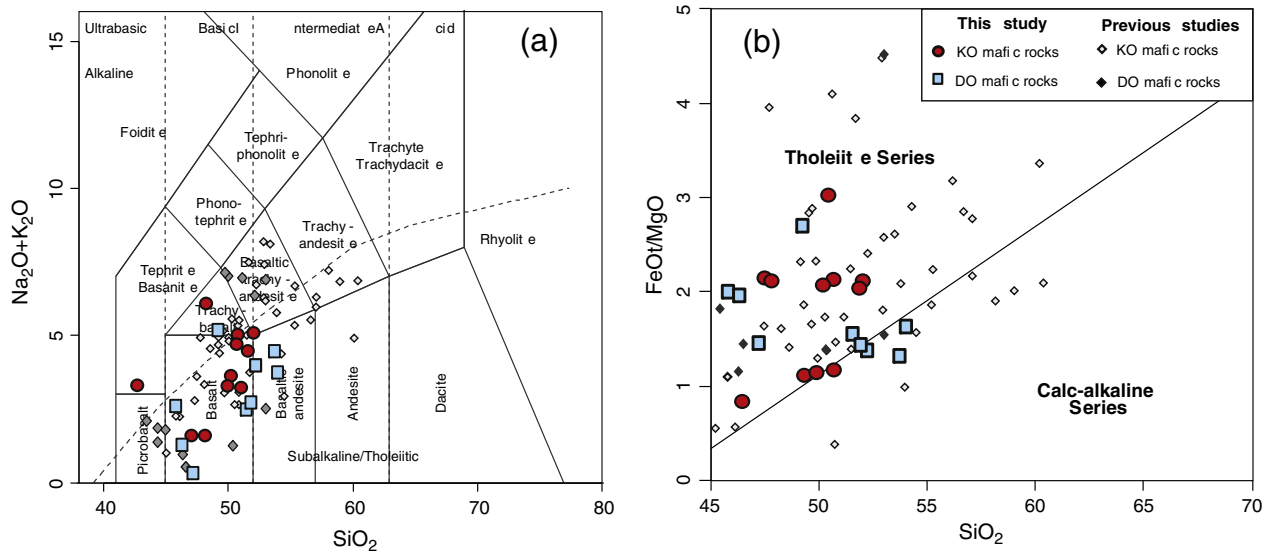
The mafic rocks analyzed include fine-grained gabbros, diabases and basalts. Prior to analysis, rock samples were crushed into small chips, highly altered chips and secondary vein materials were removed using a hand lens, and then the remaining chips were soaked in 4 N hydrochloric acid for half an hour in order to leach out alteration minerals. The rock chips were then powdered using an alumina ceramic shatterbox. Major elements were determined by X-ray fluorescence (XRF) method whereas trace element concentrations were measured using a PE Elan 6000 inductively-coupled plasma mass spectrometer (ICPMS) at the Guangzhou Institute of Geochemistry (GIG), Chinese Academy of Sciences. The analytical procedure for the ICP-MS analysis is similar to that described by Li (1997) and Chen et al. (2012). Trace element concentrations were drift corrected using the USGS standard sample BCR-1. Analytical uncertainties are generally less than 5–10%, depending upon concentrations.

Strontium and Nd isotopic compositions were measured using a Micromass Isoprobe multicollector mass spectrometer (MC-ICPMS) also at the GIG. Analytical procedures used are similar to those described by Wei et al. (2002), Liang et al. (2003) and Chen et al. (2012). Measured <sup>87</sup>Sr/<sup>86</sup>Sr were corrected to mass fractionation using <sup>86</sup>Sr/<sup>88</sup>Sr = 0.1194. Strontium standard NBS987 was used to monitor instrumental performance and yielded an average <sup>87</sup>Sr/<sup>86</sup>Sr ratio of 0.710242 ± 17 (2σ; n = 55). The measured <sup>146</sup>Nd/<sup>144</sup>Nd and fractionation-corrected <sup>143</sup>Nd/<sup>144</sup>Nd ratios of the JNdi-1 standard were 0.710288 ± 28 (2σ; n = 40) and 0.512109 ± 12 (2σ; n = 40), respectively. All measured Nd isotopic ratios were normalized to <sup>146</sup>Nd/<sup>144</sup>Nd = 0.71029. Uranium, Th and Pb elemental concentrations were measured on the same powder aliquots measured for Pb isotopes at the Scripps Institution of Oceanography (SIO), University of California, San Diego. The U, Th and Pb elemental concentrations were measured using a Finnigan Element 2 high-resolution inductively-coupled plasma mass spectrometer (HR-ICPMS) using the procedure described in Tian et al. (2008). Samples analyzed for Pb isotopic compositions were not leached because previous studies have shown that the differences in age-corrected Pb isotope ratios for leached–unleached pairs of old variably altered rocks are relatively small (Mahoney et al., 2002). Lead isotopic ratios were measured on a Micromass Sector 54 multicollector thermal ionization mass spectrometry (TIMS) also at SIO in a multicollector static mode using a <sup>207</sup>Pb–<sup>204</sup>Pb double spike. The sample preparation procedure is also similar to that described in Tian et al. (2008) and has total blank of <100 pg Pb. The Pb isotopic compositions of each sample were measured twice as spiked and unspiked aliquots, the analytical and fractionation corrected following the procedure described in Liu and Liao (2012) and Liu et al. (2013) and adjusted to the Pb standard NBS981 values of Todt et al. (1996). The precision of Pb data is taken to be <0.03% per mass unit for <sup>206</sup>Pb/<sup>204</sup>Pb and <sup>207</sup>Pb/<sup>204</sup>Pb and <0.05% for <sup>208</sup>Pb/<sup>204</sup>Pb.

## 5. Results

### 5.1. Major and trace element

Almost all the mafic volcanic and intrusive rocks from the KO (Liu et al., 2009b) and DO that we analyzed are tholeiitic basalts (or gabbros) and basaltic–andesites (Fig. 2 and Table 1). These rocks are characterized by high Al<sub>2</sub>O<sub>3</sub> and Fe<sub>2</sub>O<sub>3</sub><sup>total</sup> and variable MgO and Mg# (molar Mg / [Mg + Fe<sup>2+</sup>] \* 100). The KO mafics have low (average <1 wt.%) TiO<sub>2</sub>, similar to primitive normal-MORB (~0.6–1 wt.% TiO<sub>2</sub>; Presnall and Hoover, 1987) or typical island arc basalts (<1 wt.% TiO<sub>2</sub>; Thirlwall et al., 1994; Woodhead et al., 1993), whereas DO mafic rocks show slightly higher (average >1 wt.%) TiO<sub>2</sub>, akin to enriched-MORB.



**Fig. 2.** Classification diagrams for KO and DO igneous rocks: (a) total alkali-silica (TAS) diagram after Le Bas et al. (1986), (b) FeO\*/MgO vs SiO<sub>2</sub> after Miyashiro (1974). Shown for comparison are published data for KO mafic rocks (Liu et al., 2009b) and DO mafic rocks (Bu et al., 2005; Lei et al., 2008; Li, 1995; Ma, 2007; Ouyang, 2006; Wang et al., 2003; Zhang et al., 2009; Zheng et al., 2007).

The KO and DO mafic rocks display distinctive trace element characteristics. Although the KO mafic samples show a wide range of trace element concentrations, they can be subdivided into two distinct groups (Fig. 3). Group 1 is characterized by a light-REE depleted ( $La/Sm_N = 0.51\text{--}0.88$ ) pattern similar to that of normal-MORB. However, it differs from normal-MORB in that it has distinctive negative HFSE anomalies. Such HFSE depletion is a characteristic feature of convergent margin lavas (e.g., Pearce and Peate, 1995; Pearce et al., 1984; Tatsumi and Kogiso, 2003). The only exception to the group is samples KL-24 and KL-27, which display normal-MORB like concentration patterns. On the other hand, Group 2 shows slight enrichment in light-REE ( $La/Sm_N = 1.27\text{--}2.34$ ) similar to enriched-MORB. It also shows distinctive negative HFSE anomalies like Group 1.

The DO mafic rocks are characterized by a light-REE depleted to slightly enriched concentration pattern ( $La/Yb_N = 0.55\text{--}2.26$ ) similar to that of normal-MORB to enriched-MORB. Their incompatible trace element patterns are also similar to that of normal- to enriched-MORB. In contrast to KO mafic samples, they do not have negative HFSE anomalies. Some KO and DO samples, however, have large positive Sr anomalies, perhaps related to seawater alteration.

In summary, DO mafic rocks from the western Junggar have a normal- to enriched-MORB trace element signature whereas KO mafic rocks from the eastern Junggar have transitional between MORB- and arc basalt-like signature. Therefore, in terms of trace elements, the DO mafic rocks came from a typical MORB mantle source and were most likely produced by seafloor spreading in an open ocean or mature back arc basin whereas KO mafic rocks may have come from a mantle source similar to that beneath the DO, but was additionally influenced by a relatively recent subduction process.

## 5.2. Sr–Nd–Pb isotope characteristics

Twenty one whole rock mafic samples from the KO and DO were analyzed for Sr, Nd and Pb isotopes, and one DO plagioclase separate for Pb isotopes. One should note that the Sr and Nd isotopes of eleven KO mafic samples were previously reported (Liu et al., 2009b). The isotopic analyses and their U, Th and Pb elemental abundances are presented in Table 2. Co-variations of their initial Sr and Nd isotope ratios at a crystallization age of 371 Ma (Liu et al., 2009b) are shown in Figs. 4 and 5. The  $\epsilon_{Nd(t)}$  values of KO and DO mafic samples fall

completely within the range of age-corrected Pacific and Indian MORB fields. The  $\epsilon_{Nd(t)}$  values of the KO samples overlap although Group 1 mafic rocks extend to higher values (up to +9.7) whereas Group 2 mafic rocks extend to lower values (down to +6.2). Most DO samples display relatively lower  $\epsilon_{Nd(t)}$  values than KO samples and, thus, are more akin to Indian MORB values.

The initial Sr ratios  $^{87}Sr/^{86}Sr(t)$  of all mafic samples trend toward higher values (up to 0.7063) and, thus, are more radiogenic for given  $\epsilon_{Nd(t)}$  values relative to typical MORB. They all plot toward convergent margin arc lavas or seawater alteration. Therefore, the Sr isotopic ratios collectively are not considered a primary magmatic signature for these rocks.

The age-corrected Pb isotope compositions of KO and DO mafic samples are variable; for ease of comparison, all samples were also corrected to 371 Ma by using high-precision Th, U and Pb concentrations determined by HR-ICP-MS. In detail, the Pb isotopic ratios of KO and DO mafic rocks greatly overlap. However, several KO samples plot with age-corrected (to 371 Ma) Pacific MORB, but none of DO samples does (Fig. 5). Moreover, although there is no distinction between Group 1 and 2 KO samples, they have less radiogenic  $^{207}Pb/^{204}Pb(t)$  and  $^{208}Pb/^{204}Pb(t)$  for given  $^{206}Pb/^{204}Pb(t)$  ratios than DO mafic rocks. Combined with Nd isotopes, the Pb isotopic ratios additionally suggest a bi-modal distribution of KO and DO samples. All KO samples and two DO samples have relatively high  $\epsilon_{Nd(t)}$  for given low  $^{206}Pb/^{204}Pb(t)$  ratios whereas DO samples, except for the two above, have relatively lower  $\epsilon_{Nd(t)}$ . The former group is similar to some Tethyan samples from the Shuangou and Mianlue ophiolites (Xu and Castillo, 2004; Xu et al., 2002) whereas the latter is akin to some Tethyan samples from the Jinshajiang ophiolite (Xu and Castillo, 2004). Significantly, these Nd and Pb isotopic features of the Tethyan and PAO ophiolites are also shown by the low  $^{206}Pb/^{204}Pb$  Indian MORB (e.g., Mahoney et al., 1998, 2002).

In general, almost all KO and DO mafic samples plot above the northern hemisphere reference line (NHRL), or have high  $^{207}Pb/^{204}Pb(t)$  and  $^{208}Pb/^{204}Pb(t)$  for given  $^{206}Pb/^{204}Pb(t)$  ratios (i.e., positive  $\Delta 7/4$  and  $\Delta 8/4$  values), similar to age-corrected Indian MORB with a Dupal isotopic anomaly (Dupre and Allegre, 1983; Hart, 1984). They also plot within the fields for the 275–350 Ma Paleo-Tethyan MORB-like ophiolite lavas (Xu and Castillo, 2004; Xu et al., 2002; Zhang et al., 2005) and basalts from the Junggar basin, which had been interpreted as fragments of Paleozoic (395 to 405 Ma) PAO

**Table 1**  
Major and trace element compositions of DO rocks.

Sample	DL-137	DL-139	DL-144	DL-145	DL-150	DL-153	DL-158	DL-159	DL-161
Rock type	Gabbro	Basalt	Gabbro	Gabbro	Basalt	Basalt	Basalt	Basalt	Basalt
<i>Major elements (wt.%)</i>									
SiO <sub>2</sub>	51.52	45.74	46.34	46.73	51.55	49.05	46.11	50.32	50.32
TiO <sub>2</sub>	1.25	1.6	0.82	1.36	1.09	1.15	1.24	1.18	1.08
Al <sub>2</sub> O <sub>3</sub>	17.99	13.75	12.50	7.99	14.53	14.67	16.51	15.49	14.24
Fe <sub>2</sub> O <sub>3</sub> <sup>total</sup>	8.66	14.18	17.26	11.94	12.57	13.44	14.08	10.47	11.99
MnO	0.14	0.22	0.27	0.11	0.16	0.20	0.46	0.15	0.17
MgO	5.73	6.44	7.97	7.46	7.04	7.84	4.72	7.24	7.65
CaO	9.44	15.2	13.49	22.91	4.83	6.41	5.53	4.55	8.79
Na <sub>2</sub> O	3.12	2.03	0.81	0.27	2.51	2.28	2.50	2.93	1.91
K <sub>2</sub> O	0.8	0.53	0.45	0.00	1.06	0.05	2.36	1.26	0.72
P <sub>2</sub> O <sub>5</sub>	0.02	0.12	0.02	0.14	0.06	0.08	0.07	0.07	0.07
LOI	0.92	0.65	0.36	1.50	4.36	4.48	6.84	6.69	2.81
Total	99.6	100.45	100.29	100.41	99.74	99.66	100.43	100.34	99.74
<i>Trace element (ppm)</i>									
Sc	28.2	43.3	54.1	54.8	41.2	52.1	47.0	48.2	0.1
Ti	8278	10157	4994	8800	6242	7005	6895	7244	6256
V	208.6	320.4	394.4	478.8	314.2	337.9	323.7	258.9	300.6
Cr	209.7	197	51.54	12.21	152.7	150.2	161.8	107.2	157.4
Mn	1152	1742	2173	913.5	1160	1523	1620	1137	1286
Co	28.9	45.8	62.5	88.5	48.2	50.4	62.0	50.1	63.6
Ni	21.2	83.8	78.1	230.4	86.7	91.8	101.1	92.8	90.7
Cu	22.1	17.3	5.37	4.10	153.3	155.4	487.7	32.1	159.7
Zn	71.3	121.3	143.2	42.1	89.1	102.1	98.6	98.4	86.7
Sr	448.5	550.5	569.3	172.3	84.95	488.6	263.9	109	212.5
Rb	13.13	3.10	3.33	0.15	34.52	1.15	28.34	34.12	4.45
Y	17.76	28.35	22.22	24.47	19.07	23.33	19.58	19.62	8.05
Zr	84.93	117.7	32.66	52.57	50.59	60.39	56.56	57.57	53.74
Nb	3.55	6.64	3.02	1.30	2.43	2.82	2.59	2.78	2.58
La	5.76	6.46	1.77	3.02	2.96	3.45	3.19	3.45	2.30
Ce	14.56	17.23	4.79	8.00	7.83	8.98	8.61	8.96	7.61
Pr	2.26	2.72	0.81	1.30	1.27	1.45	1.46	1.46	1.15
Nd	10.37	12.37	4.21	6.76	6.39	7.07	7.12	7.22	5.50
Sm	2.78	3.67	1.59	2.48	2.14	2.31	2.36	2.37	1.79
Eu	1.23	1.25	0.62	1.25	0.84	0.89	0.96	0.94	0.70
Gd	3.17	4.42	2.42	3.39	2.93	3.00	3.26	3.07	2.23
Tb	0.55	0.80	0.50	0.71	0.59	0.58	0.63	0.62	0.44
Dy	3.40	4.94	3.37	4.56	3.87	3.57	3.99	3.85	2.78
Ho	0.71	1.04	0.81	1.01	0.81	0.79	0.85	0.85	0.60
Er	1.97	2.88	2.26	2.82	2.25	2.24	2.42	2.31	1.63
Tm	0.28	0.40	0.35	0.41	0.33	0.31	0.34	0.34	0.24
Yb	1.83	2.63	2.32	2.77	2.14	2.10	2.18	2.17	1.50
Lu	0.29	0.40	0.38	0.44	0.31	0.31	0.33	0.33	0.22
Hf	2.11	2.67	0.86	1.73	1.50	1.42	1.66	1.67	1.26
Ta	0.25	0.44	0.17	0.07	0.18	0.18	0.20	0.20	0.15
Th	0.94	0.50	0.09	0.17	0.27	0.26	0.29	0.29	0.01

LOI = loss on ignition. Mg# = 100 Mg / (Mg + Fe<sup>2+</sup>) assuming Fe<sub>2</sub>O<sub>3</sub> / FeO = 0.15.

Major element analyzed by XRF at the Guangzhou Institute of Geochemistry (GIG), Chinese Academy of Sciences. Trace element concentration are measured by ICP-MS also at GIG, the analytical uncertainties are generally less than 5–10%, depending upon concentrations.

crust (Zheng et al., 2007). Junggar basin samples on average, however, have lower <sup>207</sup>Pb/<sup>204</sup>Pb<sub>(t)</sub> but higher <sup>208</sup>Pb/<sup>204</sup>Pb<sub>(t)</sub> for given <sup>206</sup>Pb/<sup>204</sup>Pb<sub>(t)</sub> than KO and DO mafic samples.

## 6. Discussion

### 6.1. Mantle source composition

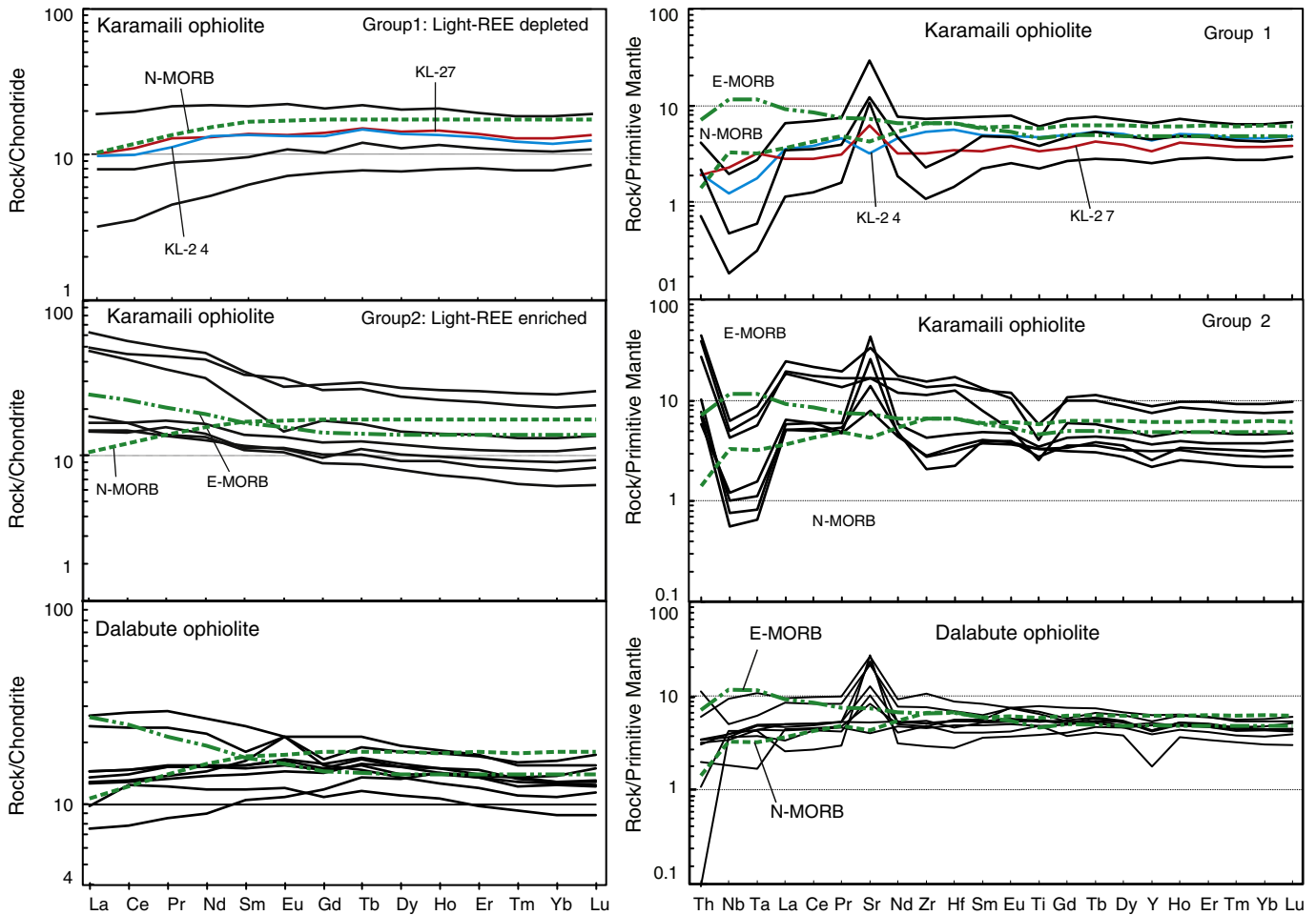
#### 6.1.1. Trace element constraints

Ratios of incompatible trace elements, especially those with similar incompatibilities with respect to mantle rocks, have been used to constrain mantle sources (Condie, 2003). Almost all of our KO mafic rocks display HFSE/LILE depletion and LILE/REE enrichment. Specifically, many samples display slight Th enrichment relative to La as well as Nb depletion relative to Th compared to normal-MORB (Fig. 6a). Combined with previous results, however, KO igneous rocks display a wide range of Th/La and Nb/Th ratios, overlapping with the complete spectrum of magmas from forearcs to ocean ridges and oceanic islands. Such chemical variations indicate that the KO actually

consists of both subduction-related as well as other oceanic igneous rocks. In other words, the bulk of our KO samples have the trace element signature of lavas derived from a depleted mantle source (or sources) contaminated by subduction-derived fluid(s) or mixing of magmas from a depleted mantle and recycled crustal materials (e.g., Condie, 2003; Godard et al., 2003, 2006; Peate and Pearce, 1998).

In comparison, the composition of DO mafic samples is similar to those of modern MORB and mature backarc basin basalts. Their HFSE/LILE and LILE/REE ratios are close to 1; together with previous results, these indicate that the source of DO mafic rocks is most likely a depleted, but slightly heterogeneous asthenospheric mantle.

The difference in the trace element composition of the mantle sources of KO and DO samples can be analyzed further using Th/Yb and Ta/Yb ratios (Fig. 6b). These ratios are excellent tracers of mantle source composition because they are largely independent of variations caused by degree of partial melting and crystal fractionation (Pearce and Peate, 1995). The element Th in marine sediments is more than 2 orders of magnitude larger than in the mantle (Plank and Langmuir, 1998) and highly soluble in melts but insoluble in fluids



**Fig. 3.** Chondrite-normalized rare earth element (REE) concentrations and primitive mantle-normalized incompatible element concentrations of KO and DO mafic rocks. Chondrite and primitive mantle values are from Sun and McDonough (1989). The KO mafic rocks can be subdivided into two groups according to light-REE depleted ( $La/Sm_N < 1$ ) and enriched ( $La/Sm_N > 1$ ) concentration patterns. Values of N-MORB and E-MORB used are from Sun and McDonough (1989).

(Turner and Hawkesworth, 1997), Ta and Yb, unlike Th, are not enrichment in marine sediments and less effectively transferred from subducting slab to melting column (Pearce and Peate, 1995). Thus, Ta/Yb represents the degree of mantle enrichment or depletion of the mantle source (Pearce and Peate, 1995) whereas increasing Th/Yb ratio represents increasing sediment melt contribution (Class et al., 2000; Elliott et al., 1997; Singer et al., 2007). In other words, addition of sediment melt would result in addition of Th, but not Ta, to the mantle source and would make a vertical vector in Fig. 6b.

The KO igneous rocks have wide ranges of Th/Yb and Ta/Yb ratios, but most of them have higher Th/Yb for given Ta/Yb than MORB. Such relatively high Th/Yb ratios, plotting above the mantle array, indicate addition of melt from subducted sediment. In comparison, the bulk of DO samples show relatively low Th/Yb for given Ta/Yb ratios. Combined with published data, DO samples plot within the depleted- to enriched-MORB mantle array, suggesting a depleted to slightly enriched mantle source. Therefore, as above, the DO mafic rocks most likely originated from a slightly heterogeneous asthenospheric mantle whereas the mantle source of KO is more heterogeneous as it has the signature of a depleted mantle source that had been contaminated by subduction component.

#### 6.1.2. Sr–Nd–Pb isotope constraints

Strontium is a fluid mobile element and is relatively abundant (8 ppm) in seawater. The exchange of Sr between rock and seawater during hydrothermal alteration of the oceanic crust raises the  $^{87}Sr/^{86}Sr$  ratios of altered submarine rocks (e.g., Godard et al., 2006; Kawahata

et al., 2001; McCulloch et al., 1981). Thus, the high  $^{87}Sr/^{86}Sr(t)$  ratios of KO and DO rocks are most likely due to seawater alteration and do not represent the primary magmatic composition of these rocks. In contrast, Nd is relatively immobile in fluid and the Nd isotopic values of altered submarine rocks remain unaffected until after the water/rock ratio reaches  $10^5$  (Faure and Mensing, 2004; McCulloch et al., 1981). Another source of concern is the mobility of U (when oxidized to  $U^{+VI}$ ) and Pb during hydrothermal alteration as hydrothermal activity can significantly affect the Pb isotope systematics, especially in old samples (Faure and Mensing, 2004; Hauff et al., 2003). Consequently, we analyzed the Pb isotopic composition of the whole rock sample (DL-137) and its plagioclase separate (DL-137-pl) to see the extent of alteration in the Pb isotopes. Results show that their modern day Pb isotopic ratios are different, owing to differences in the respective parent (U, Th) and daughter (Pb) concentrations (Table 2). However, the age-corrected Pb isotopic values for the plagioclase ( $^{206}Pb/^{204}Pb(t) = 17.945$ ,  $^{207}Pb/^{204}Pb(t) = 15.533$  and  $^{208}Pb/^{204}Pb(t) = 37.994$ ) are only slightly outside the analytical errors for those of the whole rock ( $^{206}Pb/^{204}Pb(t) = 17.931$ ,  $^{207}Pb/^{204}Pb(t) = 15.569$  and  $^{208}Pb/^{204}Pb(t) = 37.852$ ) and, thus, cluster with other whole rock samples in Pb–Pb covariation diagrams (Fig. 5). Therefore, the Pb isotopic ratios of our whole rock samples appear not to have been severely affected by alteration and can be used to constrain the Pb isotopic composition of their mantle source(s).

As noted earlier, DO samples and plagioclase separate have low  $\epsilon_{Nd(t)}$  and high  $^{207}Pb/^{204}Pb(t)$  and  $^{208}Pb/^{204}Pb(t)$  for given  $^{206}Pb/^{204}Pb(t)$  ratios (i.e., positive  $\Delta 7/4$  and  $\Delta 8/4$  values), similar to those for MORB



**Table 2**  
Pb, Nd and Sr isotope ratios and concentrations of KO and DO mafic rocks.

Karamaili mafic rocks												
Sample	KL-01	KL-03	KL-04	KL-12	KL-14	KL-15	KL-16	KL-17	KL-21	KL-24	KL-27	KL-06
Rock type	Basalt	Basalt	Basalt	Diabase	Gabbro	Gabbro	Gabbro	Gabbro	Gabbro	Basalt	Basalt	Basal andesite
Pb (ppm)	0.61	2.50	5.63	1.66	2.64	2.93	1.09	0.68	1.03	0.34	0.46	1.60
U (ppm)	0.15	0.15	0.65	0.14	0.80	2.43	0.58	0.14	0.19	0.07	0.06	1.17
Th (ppm)	0.24	0.04	1.86	0.04	0.29	0.76	0.59	0.15	0.26	0.01	0.16	3.42
$^{238}\text{U}/^{204}\text{Pb}$	15.4	3.8	7.2	5.4	19.6	54.9	34.7	13.1	11.5	12.6	6.2	48.4
$^{232}\text{Th}/^{204}\text{Pb}$	25.9	1.2	21.5	1.4	7.3	17.8	35.8	13.9	16.2	1.5	12.0	146.2
$^{206}\text{Pb}/^{204}\text{Pb}$	18.669	18.046	18.252	18.025	19.691	21.347	19.895	18.080	18.519	18.290	18.437	20.084
$^{207}\text{Pb}/^{204}\text{Pb}$	15.558	15.469	15.516	15.465	15.616	15.718	15.593	15.410	15.485	15.391	15.545	15.628
$^{208}\text{Pb}/^{204}\text{Pb}$	38.204	37.850	38.107	37.807	38.422	38.453	38.375	37.617	37.908	37.746	38.394	39.960
$^{206}\text{Pb}/^{204}\text{Pb}_{(t)}$	17.756	17.818	17.824	17.705	18.531	18.097	17.841	17.306	17.837	17.543	17.983	17.215
$^{207}\text{Pb}/^{204}\text{Pb}_{(t)}$	15.509	15.457	15.493	15.448	15.554	15.542	15.482	15.369	15.448	15.350	15.521	15.473
$^{208}\text{Pb}/^{204}\text{Pb}_{(t)}$	37.723	37.828	37.709	37.782	38.287	38.124	37.711	37.360	37.607	37.719	37.966	37.252
$\Delta 7/4$	9.4	3.5	7.1	3.8	5.5	9.1	5.8	0.3	2.5	-4.1	7.6	11.6
$\Delta 8/4$	62.9	65.9	53.3	75.0	25.6	61.8	51.4	81.0	41.5	88.2	69.5	81.2
Nd (ppm)	10.46	16.16	21.64	2.53	6.62	6.28	5.96	6.38	7.76	6.27	4.39	-
Sm (ppm)	3.36	3.49	5.54	0.99	1.73	1.68	1.73	2.13	2.13	2.18	1.50	-
$^{147}\text{Sm}/^{144}\text{Nd}$	0.1957	0.1313	0.1559	0.2377	0.1591	0.1629	0.1766	0.2031	0.1669	0.2120	0.2082	-
$^{143}\text{Nd}/^{144}\text{Nd}$	0.513132	0.513060	0.512932	0.513212	0.512922	0.512871	0.512928	0.513065	0.513023	0.513149	0.513029	-
$^{143}\text{Nd}/^{144}\text{Nd}_{(t)}$	0.512676	0.512741	0.512572	0.512658	0.512551	0.512553	0.512499	0.512475	0.512634	0.512654	0.512544	-
$\epsilon_{\text{Nd}(t)}$	9.7	11.3	7.7	9.3	7.4	6.2	6.6	8.0	8.9	9.2	7.1	-
Sr (ppm)	615.2	349.2	349.5	224.8	291.9	543	167.1	258.6	912.8	67.39	131.2	-
Rb (ppm)	3.03	21.82	32.11	0.08	0.90	4.28	43.80	7.72	10.89	0.25	3.32	-
$^{87}\text{Rb}/^{86}\text{Sr}$	0.0143	0.1809	0.2660	0.0010	0.0089	0.0228	0.7590	0.0864	0.0345	0.0109	0.0732	-
$^{87}\text{Sr}/^{86}\text{Sr}$	0.703917	0.705040	0.705386	0.705000	0.705417	0.704809	0.707937	0.706083	0.704265	0.704356	0.705443	-
$^{87}\text{Sr}/^{86}\text{Sr}_{(t)}$	0.703844	0.704084	0.705627	0.704995	0.704862	0.703981	0.703928	0.704688	0.704088	0.704300	0.705068	-

Dalabute mafic rocks

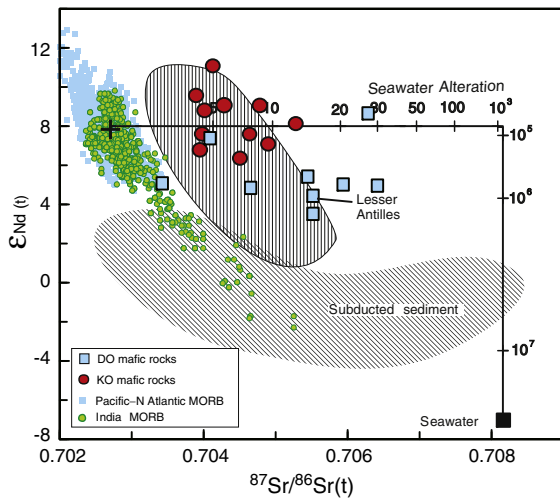
Sample	DL-137	DL-137-pl	DL-139	DL-144	DL-145	DL-150	DL-153	DL-158	DL-159	DL-161
Rock type	Gabbro	Plagioclase	Basalt	Gabbro	Gabbro	Basalt	Basalt	Basalt	Basalt	Basalt
Pb (ppm)	2.40	4.52	2.40	0.63	0.23	0.88	0.31	0.26	0.42	0.33
U (ppm)	0.19	0.30	0.16	0.07	0.18	0.06	0.08	0.06	0.11	0.07
Th (ppm)	1.10	0.28	0.80	0.23	0.19	0.21	0.22	0.24	0.21	0.22
<sup>238</sup> U/ <sup>204</sup> Pb	4.893	4.177	4.158	7.169	52.600	3.959	15.952	13.527	17.279	13.028
<sup>232</sup> Th/ <sup>204</sup> Pb	29.949	4.028	21.810	23.392	56.706	15.407	47.045	60.535	32.278	42.977
<sup>206</sup> Pb/ <sup>204</sup> Pb	18.221	18.194	18.184	18.383	21.078	18.122	18.514	18.387	18.449	18.750
<sup>207</sup> Pb/ <sup>204</sup> Pb	15.585	15.546	15.566	15.622	15.716	15.560	15.593	15.599	15.591	15.722
<sup>208</sup> Pb/ <sup>204</sup> Pb	38.407	38.069	38.314	38.477	39.013	38.258	38.559	38.837	38.538	39.154
<sup>206</sup> Pb/ <sup>204</sup> Pb <sub>(t)</sub>	17.932	17.946	17.938	17.958	17.962	17.888	17.569	17.585	17.425	17.978
<sup>207</sup> Pb/ <sup>204</sup> Pb <sub>(t)</sub>	15.569	15.533	15.553	15.599	15.548	15.547	15.542	15.555	15.536	15.680
<sup>208</sup> Pb/ <sup>204</sup> Pb <sub>(t)</sub>	37.852	37.995	37.910	38.044	37.962	37.972	37.687	37.716	37.940	38.358
Δ7/4	13.5	9.7	11.9	16.2	11.1	11.8	14.7	15.9	15.7	24.1
Δ8/4	54.6	67.1	59.7	70.6	61.9	71.9	82.0	82.8	124.7	99.6
Nd (ppm)	10.37		12.37	4.206	6.76	6.394	7.066	7.124	7.223	5.502
Sm (ppm)	2.777		3.665	1.587	2.48	2.143	2.305	2.362	2.365	1.789
<sup>147</sup> Sm/ <sup>144</sup> Nd	0.1630		0.1803	0.2296	0.2229	0.2040	0.1985	0.2018	0.1993	0.1979
<sup>143</sup> Nd/ <sup>144</sup> Nd	0.512940		0.512881	0.512904	0.513153	0.512921	0.512905	0.512883	0.512900	0.512900
<sup>143</sup> Nd/ <sup>144</sup> Nd <sub>(t)</sub>	0.512544		0.512443	0.512346	0.512611	0.512425	0.512422	0.512393	0.512415	0.512419
ε <sub>Nd(t)</sub>	7.5		5.5	3.6	8.8	5.2	5.1	4.5	5.0	5.1
Sr (ppm)	448.5		550.5	569.3	172.3	84.95	488.6	263.9	109	212.5
Rb (ppm)	13.13		3.1	3.325	0.15	34.52	1.152	28.34	34.12	4.449
<sup>87</sup> Rb/ <sup>86</sup> Sr	0.0848		0.0163	0.0169	0.0024	1.1767	0.0068	0.3110	0.9064	0.0606
<sup>87</sup> Sr/ <sup>86</sup> Sr	0.704508		0.705503	0.705578	0.706272	0.709621	0.705953	0.707136	0.709410	0.706709
<sup>87</sup> Sr/ <sup>86</sup> Sr <sub>(t)</sub>	0.704060		0.705417	0.705489	0.706259	0.703406	0.705917	0.705494	0.704622	0.706389

Lead isotopic ratios were measured on a Micromass Sector 54 multicollector TIMS at the Scripps Institution of Oceanography (SIO) in a multicollector static mode using a <sup>207</sup>Pb–<sup>204</sup>Pb double spike. The procedure has total blank of <100 pg Pb. The Pb isotopic compositions of each sample were adjusted to the Pb standard NBS981 values of [Todd et al. \(1996\)](#). The precision of Pb data is taken to be <0.03% per mass unit for <sup>206</sup>Pb/<sup>204</sup>Pb and <sup>207</sup>Pb/<sup>204</sup>Pb and <0.05% for <sup>208</sup>Pb/<sup>204</sup>Pb. U, Th and Pb elemental concentrations were measured on the same powder aliquots measured for Pb isotopes also do at the SIO by using a Finnigan Element 2 high-resolution inductively-coupled plasma mass spectrometer (HR-ICPMS). <sup>238</sup>U/<sup>204</sup>Pb and <sup>232</sup>Th/<sup>204</sup>Pb ratios are calculated from U, Th and Pb concentrations.

Strontium and Nd isotopic ratios were measured using a MC-ICPMS at the Guangzhou Institute of Geochemistry (GIG), Chinese Academy of Sciences. Measured <sup>87</sup>Sr/<sup>86</sup>Sr were corrected to mass fractionation using <sup>86</sup>Sr/<sup>88</sup>Sr = 0.1194. Strontium standard NBS987 was used to monitor instrumental performance and yielded an average <sup>87</sup>Sr/<sup>86</sup>Sr ratio of 0.710242 ± 17 (2σ; n = 55). The measured <sup>146</sup>Nd/<sup>144</sup>Nd and fractionation-corrected <sup>143</sup>Nd/<sup>144</sup>Nd ratios of the JNdi-1 standard were 0.710288 ± 28 (2σ; n = 40) and 0.512109 ± 12 (2σ; n = 40), respectively. All measured Nd isotopic ratios were normalized to <sup>146</sup>Nd/<sup>144</sup>Nd = 0.71029. Rb, Sr, Sm and Nd were also measured at GIG. <sup>147</sup>Sm/<sup>144</sup>Nd and <sup>87</sup>Rb/<sup>86</sup>Sr ratios are calculated from Rb, Sr, Sm and Nd concentrations.

Initial <sup>206</sup>Pb/<sup>204</sup>Pb<sub>(t)</sub>, <sup>208</sup>Pb/<sup>204</sup>Pb<sub>(t)</sub>, <sup>143</sup>Nd/<sup>144</sup>Nd<sub>(t)</sub> and <sup>87</sup>Sr/<sup>86</sup>Sr<sub>(t)</sub> are calculated using <sup>238</sup>U/<sup>204</sup>Pb = 5 and <sup>232</sup>Th/<sup>238</sup>U = 2.3 ([White, 1993](#)) and <sup>87</sup>Rb/<sup>86</sup>Sr = 0.02 and <sup>147</sup>Sm/<sup>144</sup>Nd = 0.24 ([Mahoney et al., 1998](#)). ε<sub>Nd(t)</sub> is calculated using <sup>143</sup>Nd/<sup>144</sup>Nd = 0.512638 and <sup>147</sup>Sm/<sup>144</sup>Nd = 0.1967 for present day CHUR.

Δ7/4 Pb and Δ8/4 Pb are calculated as described by [Hart \(1984\)](#).



**Fig. 4.** Age-corrected ( $t = 371$  Ma)  $\epsilon_{Nd(t)}$  versus  $^{87}Sr/^{86}Sr(t)$  ratios for KO (circles) and DO (square) mafic rocks compared to Pacific–North Atlantic and Indian MORB and convergent margin arc lavas. These are age-corrected to 371 Ma by assuming that the upper mantle MORB source has an average  $^{87}Rb/^{86}Sr = 0.02$  and  $^{147}Sm/^{144}Nd = 0.24$  (Mahoney et al., 1998). The solid line shows the effect of alteration on the Sr and Nd isotope composition of rocks for different water/rock ratios (crosses). Notice the shift of  $^{87}Sr/^{86}Sr$  ratios towards more radiogenic values but leaving the Nd values unaffected until water/rock ratio is  $> 10^5$ . The calculation equation used is by McCulloch et al. (1981) that is modified by Faure and Mensing (2004).

$$\epsilon_r^f = \frac{\epsilon_r^i + \epsilon_w^i X_w (Water/Rock)}{X_r + X_w (Water/Rock)}$$

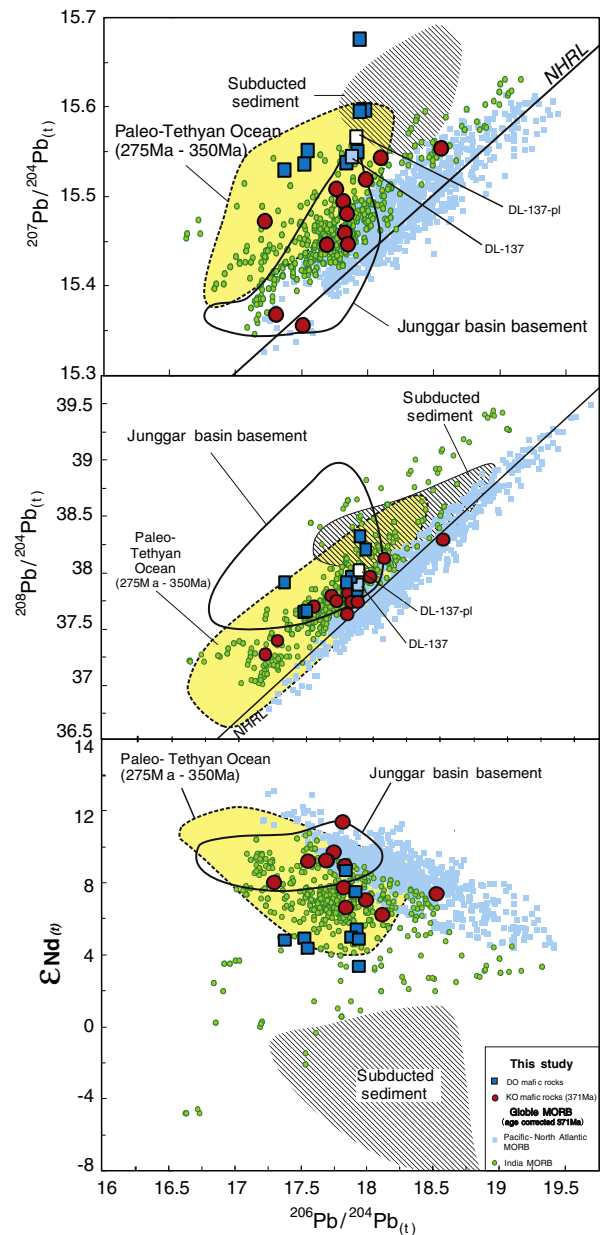
The  $\epsilon_r^i$ ,  $\epsilon_r^f$  and  $\epsilon_w^i$ ,  $\epsilon_w^f$  are isotope ratios in the initial and final stages of rocks and water, respectively.  $X_r$  and  $X_w$  are concentrations of element X in the rocks and water. The Sr parameter used in the calculation are  $\epsilon_r^i Sr = 0.70816$  (recalculated to 371 Ma from modern seawater),  $\epsilon_w^i Sr = 0.709211$  (Veizer, 1989) and  $X_w Sr = 8$  ppm; the rock value  $\epsilon_r^f Sr = 0.70270$  was recalculated to 371 Ma from the MORB value of  $\epsilon_r^f Sr = 0.70300$  (Hauff et al., 2003) and  $X_r Sr = 120$  ppm. The Nd parameters used are  $\epsilon_w^i Nd = -7$  and  $X_w Nd = 2.6 \times 10^{-6}$  ppm (McCulloch et al., 1981) and  $\epsilon_r^f Nd = 8$  and  $X_r Nd = 10$  ppm (Hauff et al., 2003).

Data for present-day Pacific–North Atlantic and Indian MORB are from the PetDB database (<http://www.petdb.org/>) and the compilation of Stracke et al. (2003). Data for convergent margin arc lavas are from the GEOROC database (<http://georoc.mpch-mainz.gwdg.de/georoc/>).

from the Indian Ocean. Since the trace element data indicate that DO mafic rocks came from a depleted to slightly enriched asthenospheric mantle, the combined isotopic and trace element data indicate that the DO mafic rocks most likely came from an Indian Ocean-type asthenospheric mantle with a Dupal isotopic anomaly.

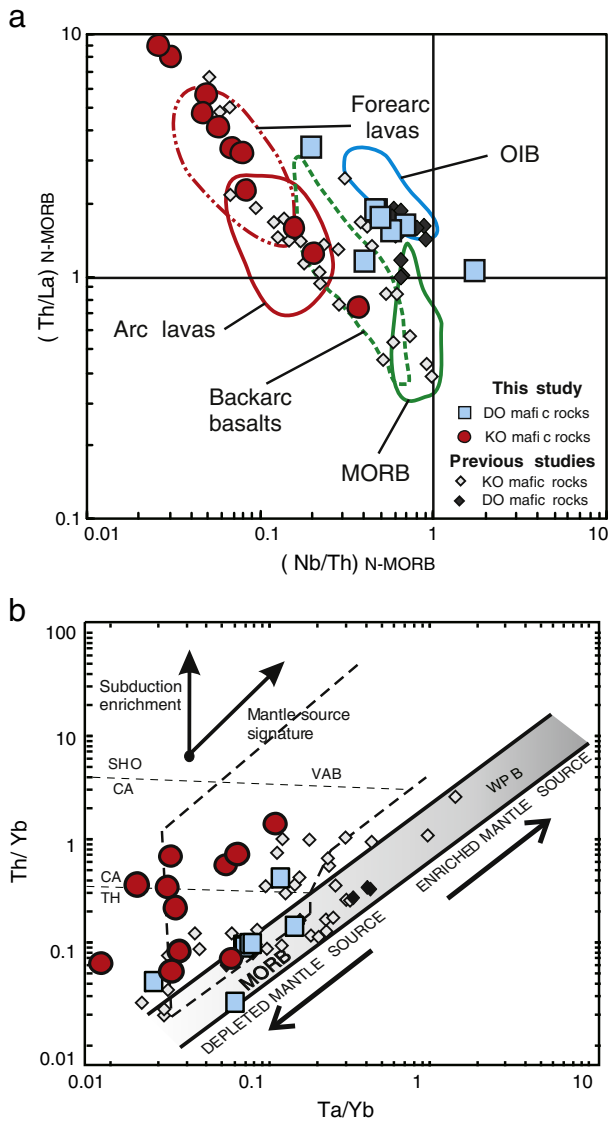
The KO samples, except for sample KL-14, have a limited range of  $^{206}Pb/^{204}Pb$  ratios, which additionally argues for minimal alteration effect. Sample KL-14 has a relatively high  $^{206}Pb/^{204}Pb$  value most likely as a result of high degree of alteration (LOI = 10.99%). Such a high LOI value may have caused mobility of U that resulted in variation in the  $^{238}U/^{204}Pb$  or  $\mu$  ratio, which in turn increases the  $^{206}Pb/^{204}Pb$  ratio over time. On the other hand, although U can be mobilized, the low abundance of  $^{235}U$  ( $^{238}U/^{235}U = 137.88$ ) makes the  $^{207}Pb/^{204}Pb$  ratio insensitive to alteration and can only be affected when substantial U is added relative to Pb (Faure and Mensing, 2004; Staudigel et al., 1996).

Similar to DO samples, KO samples have elevated  $^{207}Pb/^{204}Pb$  and  $^{208}Pb/^{204}Pb$  for given  $^{206}Pb/^{204}Pb$  ratios, a characteristic isotopic feature of Indian MORB. The Indian MORB-type Pb isotopic signature of KO mafic rocks, however, must be treated with caution because trace element data indicate that KO mafic rocks originated from a mantle source that was contaminated by a subduction component. Interestingly, the Pb isotopic signature of Indian MORB can also be acquired from subducted, ancient marine sediments (e.g., Rehkämper and Hofmann, 1997). Addition of Pb from subducted sediment can be easily detected



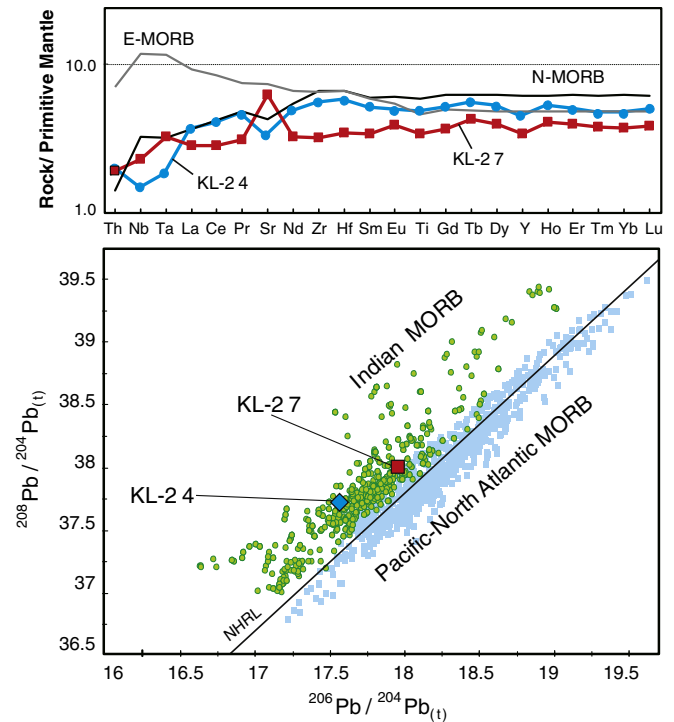
**Fig. 5.** Age-corrected ( $t = 371$  Ma)  $^{207}Pb/^{204}Pb(t)$ ,  $^{208}Pb/^{204}Pb(t)$  and  $\epsilon_{Nd(t)}$  vs  $^{206}Pb/^{204}Pb(t)$  for KO and DO mafic rocks compared to age-corrected Indian and Pacific–North Atlantic MORB. Data source for MORB are the same as in Fig. 4. Age-corrected MORB field at 371 Ma was calculated by assuming  $^{238}U/^{204}Pb = 5$  and  $^{232}Th/^{238}U = 2.3$  in the depleted MORB source (White, 1993). Also shown are fields for Paleo-Tethyan basalts from ophiolites (Xu and Castillo, 2004; Xu et al., 2002; Zhang et al., 2005), volcanic rocks from the Junggar basin basement (Zheng et al., 2007) and subducted sediment (Plank and Langmuir, 1998). NHRL (Northern Hemisphere Reference Line) is from Hart (1984).

because of the high  $^{207}Pb/^{204}Pb$  of sediment, producing a steep trend of rapidly increasing  $^{207}Pb/^{204}Pb$  with increasing  $^{206}Pb/^{204}Pb$  ratios. In other words, an “excess”  $^{207}Pb$  indicates mixing of Pb between, e.g., a MORB-type, depleted mantle source and subducted sediment, which is in the form of melt and/or hydrous fluids from the subducted slab (Elliott, 2003; Elliott et al., 1997; Godard et al., 2006; Hickey-Vargas et al., 2008). Specific examples of such additions of excess  $^{207}Pb$  are shown by magmas from the 500-Ma-old Devil River arc system in New Zealand (Münker, 2000) and for 570 Ma Agardagh Tex-Chem ophiolite in Central Asia (Pfänder et al., 2002); both examples suggest that the excess  $^{207}Pb$  was produced by subduction of sediment derived from an Archaean source (Münker, 2000) or small amounts of old crustal material (Pfänder et al., 2002).



**Fig. 6.** Normal-MORB (Sun and McDonough, 1989) normalized (a) Th/La versus Nb/Th of KO and DO mafic rocks. The composition of bulk KO and DO mafic rocks (this and previous studies) are compared to basalts from oceanic spreading centers, seamounts, backarc, and forearc (MORB, OIB, backarc, arc and forearc data are from Godard et al. (2006, 2003). (b) Th/Yb versus Ta/Yb (original diagram from Pearce, 1982). Previous data for KO and DO mafic rocks are from Lei et al. (2008), Ma (2007), Ouyang (2006), Wang et al. (2003) and Zheng et al. (2007).

On the other hand, Th is immobile during alteration so that the parent/daughter (Th/Pb) ratio is generally not be affected during alteration; thus, initial  $^{208}\text{Pb}/^{204}\text{Pb}$  isotope ratios of the igneous rocks should reflect those of their mantle source (Hauff et al., 2003). The well-defined correlation between  $^{208}\text{Pb}/^{204}\text{Pb}$  and  $^{206}\text{Pb}/^{204}\text{Pb}$  values of KO mafic rocks (Fig. 5) suggests that the KO Pb isotopic signature was most likely that of the original mantle source(s) before the subduction component addition, which is Indian MORB-like. More importantly, the Pb isotopic signature of samples KL-24 and KL-27, which have minimal influence of subduction, also indicates that the mantle source of KO mafic rocks indeed originally possessed an Indian MORB isotopic signature with a Dupal isotopic anomaly (Fig. 7), consistent with the Pb isotopic signature of DO mafic rocks. In fact, other normal-MORB-like samples from nearby PAO ophiolites also have an Indian MORB-type isotopic characteristics (Lei et al., 2008; Liu et al., 2009a, 2010). Therefore, combined with the subduction origin of the KO (e.g., Liu et al., 2007, 2009a), we propose that the KO mafic rocks were derived from an Indian



**Fig. 7.** Primitive mantle-normalized incompatible element concentrations and  $^{208}\text{Pb}/^{204}\text{Pb}(t)$  vs  $^{206}\text{Pb}/^{204}\text{Pb}(t)$  for KO mafic sample KL-24, 27. Primitive mantle values are from Sun and McDonough (1989).

MORB-type mantle that was additionally and ~recently modified by a subduction component.

To briefly summarize, our results indicate the Nd and Pb isotopic composition of PAO ophiolite fragments is similar to that of Indian Ocean MORB and this, in turn, suggests the existence of an Indian MORB-type mantle domain at least, beneath the southern part of the PAO since the Early Devonian. An Indian Ocean-type mantle also existed beneath the nearby Paleo-Tethyan Ocean at 275–350 Ma, and perhaps earlier (Xu and Castillo, 2004; Xu et al., 2002; Zhang et al., 2005). Thus, the Indian MORB-like characteristics of the 371 Ma KO rocks and 390 Ma DO rocks lend support to the idea that the mantle source with this isotopic signature existed prior to the opening of the present day Indian Ocean (Hickey-Vargas et al., 2008; Xu and Castillo, 2004; Xu et al., 2002; Zhang et al., 2005).

The present-day Dupal isotopic anomaly is centered in the southern hemisphere (Hart, 1984), but it also includes some parts of the Indian and Atlantic oceans in the northern hemisphere (Castillo, 1988; Regelous et al., 2009). Plate tectonic reconstruction (Scotese, 2001) showed that the Neo- and Paleo-Tethyan oceans were mainly centered in the southern hemisphere although a part of the Paleo-Tethys extended from the equator to the northern hemisphere, akin to modern Indian and Atlantic oceans. Moreover, the PAO was formed during the break-up of the Rodinia Supercontinent between 1000 and 300 Ma in the northern hemisphere (Dobretsov et al., 1995, 2003). In particular, the Junggar Ocean, which represents the southern branch of PAO, was located from 0° N to 40° N (Klootwijk, 2013; Windley et al., 2007; Xiao et al., 2010). If these are all true, then our isotopic data combined with previous results (e.g., Xu and Castillo, 2004; Xu et al., 2002) indicate that the Dupal isotopic anomaly was present in the sub-oceanic mantle beneath the Neo- and Paleo-Tethys in the southern hemisphere and in the Junggar Ocean in the northern hemisphere.

Goldstein et al. (2008) summarized all possible origins of the Dupal anomaly, including upwelling of the lower mantle through plumes (Castillo, 1988) followed by contamination of the asthenosphere by such mantle plumes (e.g., Class and le Roex, 2011; Weis and Frey,

**Table 3**  
Model parameter used for mixing model calculation.

	Partition coefficients				
	$D_{Th}$	$D_{Nd}$	$D_{Pb}$		
Sediment/melt <sup>a</sup>	0.89	1.53	1.29		
Sediment/fluid <sup>a</sup>	–	3.26	0.64		
AOC/fluid <sup>b</sup>	–	51	0.1		
	Trace elements (ppm)				
	Th	Nd	Pb		
DIMM <sup>c</sup>	0.004	0.483	0.014		
Sediment <sup>d</sup>	6.91	27	19.9		
Sediment melt <sup>e</sup>	7.72	17.96	15.6		
LAOC (KL-14) <sup>f</sup>	–	6.62	2.64		
Sediment fluid <sup>g</sup>	–	8.24	30.92		
LAOC fluid <sup>h</sup>	–	0.13	24.22		
Slab fluid <sup>i</sup>	–	1.75	25.56		
	Isotopic compositions				
0 Ma	$^{143}Nd/^{144}Nd_{(0)}$	$^{206}Pb/^{204}Pb_{(0)}$	$^{207}Pb/^{204}Pb_{(0)}$	$^{208}Pb/^{204}Pb_{(0)}$	
DIMM <sup>j</sup>	0.51326	17.90	15.46	37.77	
Sediment (GLOSS) <sup>k</sup>	0.51218	–	–	–	
Sediment melt (GLOSS) <sup>k</sup>	0.51218	–	–	–	
Sediment (unradiogenic Pb) <sup>l</sup>	0.51231	18.82	15.65	38.92	
Sediment melt (unradiogenic Pb) <sup>l</sup>	0.51231	18.82	15.65	38.92	
Age-corrected to 371 Ma	$^{143}Nd/^{144}Nd_{(t)}$	$^{206}Pb/^{204}Pb_{(t)}$	$^{207}Pb/^{204}Pb_{(t)}$	$^{208}Pb/^{204}Pb_{(t)}$	$\epsilon_{Nd(t)}$
DIMM <sup>j</sup>	0.51267	17.18	15.35	37.22	10.0
Sediment (GLOSS) <sup>k</sup>	0.51186	–	–	–	–
Sediment melt (GLOSS) <sup>k</sup>	0.51186	–	–	–	–
Sediment (unradiogenic Pb) <sup>l</sup>	0.51196	18.15	15.61	38.48	–3.9
Sediment melt (unradiogenic Pb) <sup>l</sup>	0.51196	18.15	15.61	38.48	–3.9
Slab fluid <sup>m</sup>	0.51255	18.53	15.55	38.29	7.3

<sup>a</sup> Sediment/melt and sediment/fluid partition coefficients from Johnson and Plank (1999).

<sup>b</sup> AOC/fluid partition coefficients calculated from Brenan et al. (1995a, 1995b) by assuming subducted oceanic crust has garnet:cpx = 60:40.

<sup>c</sup> DIMM trace element concentration values are from Workman and Hart (2005).

<sup>d</sup> Sediment trace element concentration values are from GLOSS (Plank and Langmuir, 1998).

<sup>e</sup> Calculated sediment melt trace element composition assuming 5% of melting of GLOSS (Plank and Langmuir, 1998) and partition coefficient in note a.

<sup>f</sup> LAOC trace element values are represented by highly altered sample KL-14 (LOI = 10.99%) from Liu et al. (2007).

<sup>g</sup> Calculated trace element composition of fluid released during dehydration of subducted sediment (GLOSS) (Plank and Langmuir, 1998). This assumes 1% of fluid in equilibrium with sediment and partition coefficient in note a.

<sup>h</sup> Calculated trace element composition of fluid released from LAOC as noted f. This assumes 1% of fluid in equilibrium with LAOC and partition coefficient in note b.

<sup>i</sup> Calculated trace element composition of slab fluid assumes a 80:20 mixture of LAOC:sediment-derived fluids.

<sup>j</sup> Neodymium isotopic values of DIMM are from Workman and Hart (2005). Pb isotopic values are from Indian MORB compiled from Stracke et al. (2003) and PetDB database (website: <http://www.petdb.org/>). DIMM age-corrected isotope values are calculated by assuming  $^{87}Rb/^{86}Sr = 0.02$  and  $^{147}Sm/^{144}Nd = 0.24$  (Mahoney et al., 1998), and  $^{238}U/^{204}Pb = 5$  and  $^{232}Th/^{238}U = 2.3$  in the depleted MORB source (White, 1993).  $\epsilon_{Nd(t)}$  is calculated using  $^{143}Nd/^{144}Nd = 0.512638$  and  $^{147}Sm/^{144}Nd = 0.1967$  for present day CHUR.

<sup>k</sup> Sediment (GLOSS) Nd isotopic values are from Plank and Langmuir (1998), age-corrected to 371 Ma are calculated by using trace element concentration of GLOSS: Sm = 5.78 ppm and Nd = 27 ppm.

<sup>l</sup> Sediment (unradiogenic Pb) Nd and Pb isotopic values are from Izu–Bonin sediment (Plank and Langmuir, 1998; Plank et al., 2007). Age-corrected to 371 Ma are calculated by using trace element concentration of average Izu–Bonin sediment: Sm = 6.35 ppm, Nd = 28.7 ppm, Pb = 6.82 ppm, U = 1.19 ppm and Th = 2.43 ppm (Plank and Langmuir, 1998).  $\epsilon_{Nd(t)}$  calculated as note j.

<sup>m</sup> Slab fluid Nd and Pb isotopic values are represented by highly altered sample KL-14 (LOI = 10.99%) from Liu et al. (2007).

1996) and mantle modification either through focused subduction (Staudigel et al., 1996) of altered oceanic crust and sediment (Rehkämper and Hofmann, 1997) or through recycling of arc crust (Kempton et al., 2002) or detached sub-continental lithospheric mantle (SCLM) during continental break-up (Chung et al., 2001; Hanan et al., 2004; Mahoney et al., 1998). Therefore, the exact origin of the anomaly is still a contentious issue. Our new results cannot exclude any of the competing models for the exact origin of the isotopically anomalous regions in the lower mantle. However, our results reinforce the idea that the distinct Indian MORB-type (i.e., Dupal) isotopic signature is an asthenospheric feature and it is produced through contamination by upwelling plumes carrying materials from the deep mantle (e.g., Weis and Frey, 1996).

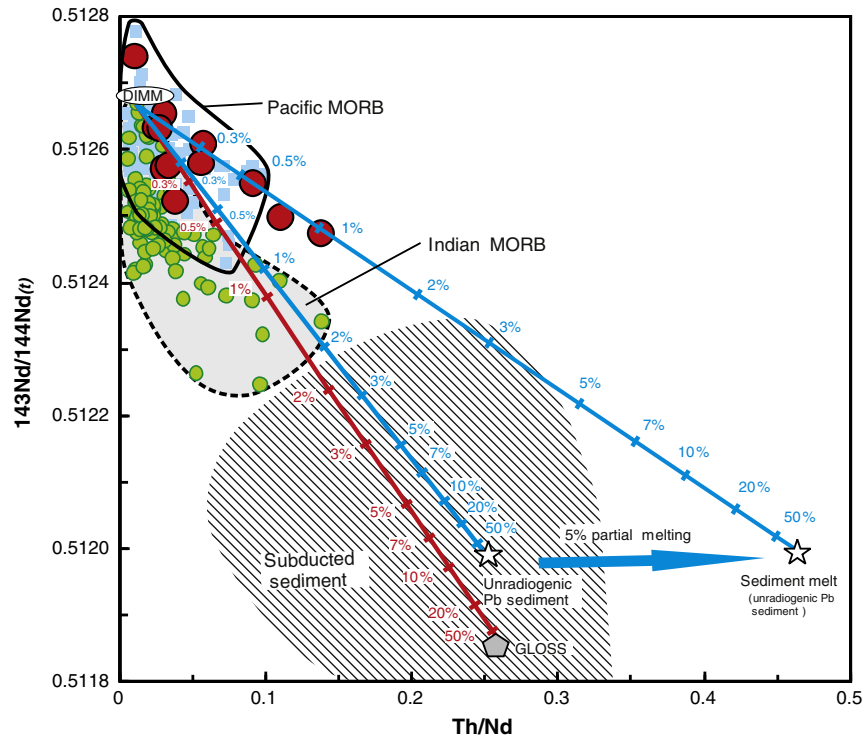
In summary, we propose that the long-lived Dupal isotopic anomaly of the Indian Ocean-type mantle domain was ultimately derived from the lower mantle although some of its geochemical features may have been modified by relatively more recent, shallow mantle geological

processes such as lithospheric delamination or subduction. Our data, however, cannot constrain the exact origin of the isotopically anomalous lower mantle. More investigations, particularly of ancient (>390 Ma) ophiolites, are needed in order to resolve this important problem.

## 6.2. The subduction component

### 6.2.1. Subducted sediment

As discussed above, trace element and isotopic data indicate that the DO mafic rocks most likely came from an Indian MORB-type asthenospheric mantle whereas the KO mafic rocks came from a similar source that had been additionally contaminated during a relatively ~recent subduction event (e.g., Liu et al., 2007, 2009a). Here we further discuss the nature of the subduction component and quantify its influence on the Indian MORB-type mantle source of the KO mafic samples. To constrain its nature, we first investigate the behavior of Th/Nd ratios against  $^{143}Nd/^{144}Nd$  ratio of KO mafic samples to evaluate sediment melt



**Fig. 8.** Plot of Th/Nd versus  $^{143}\text{Nd}/^{144}\text{Nd}(t)$  for KO mafic rocks. KO mafic rocks form a binary linear array between mantle wedge and sediment components. The mantle wedge end-member was assumed to be a very depleted-Indian MORB mantle (DIMM; Stracke et al., 2003) with high  $^{143}\text{Nd}/^{144}\text{Nd}(t)$  but low Th/Nd ratios. The low  $^{143}\text{Nd}/^{144}\text{Nd}(t)$  and high Th/Nd ratio end-member is interpreted to be bulk sediment or sediment melt calculated in Table 3. Note that the global subducted sediment (GLOSS; Plank and Langmuir, 1998) and its melts were not appropriate for explaining the entire compositional spectrum of KO mafic rocks. Thus, we assumed that the PAO sediment would have unradiogenic Pb isotopes, similar to present day sediment in the Izu–Bonin trench (Plank and Langmuir, 1998; Plank et al., 2007). Results of calculation indicate that addition of  $\leq 1\%$  of this sediment partial melt to DIMM can account for the subduction component signature of KO mafic rocks. Tick marks on the calculated DIMM–sediment (melt) mixing lines are labeled with wt.% of sediment (melt). Data source are same as in Fig. 4. Partition coefficients of the sediment melt are in Table 3 and explained in the text.

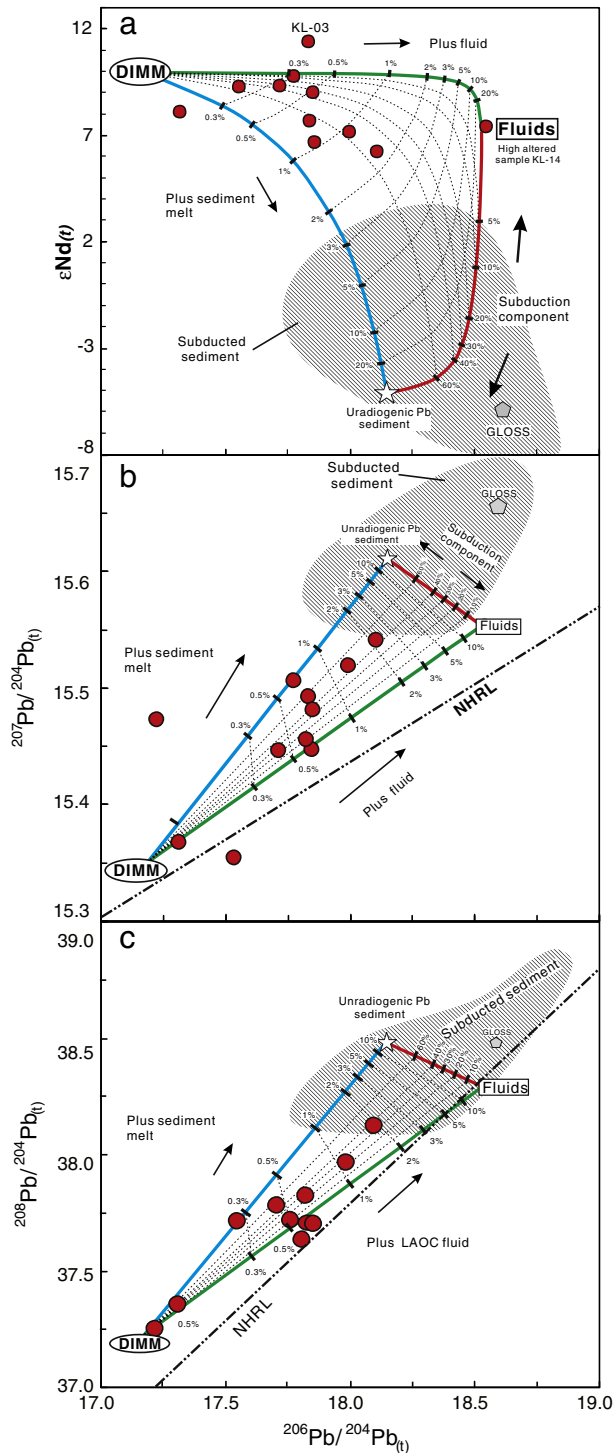
contribution to a depleted mantle source. We use Th and Nd because these are immobile in fluid dehydrated from subducted oceanic crust and sediment, but are relatively mobile in sediment melt (Class et al., 2000; Elliott et al., 1997; Johnson and Plank, 1999). Moreover, Th is more incompatible ( $D_{\text{sediment/melt}}^{\text{Th}} = 0.89$ ) than Nd ( $D_{\text{sediment/melt}}^{\text{Nd}} = 1.53$ ) during melting (Johnson and Plank, 1999) and, thus, a sediment partial melt could fractionate the Th/Nd ratio from the bulk sediment. We assume a depleted-Indian MORB mantle (DIMM; compiled from Stracke et al., 2003) as the mantle end-member. For the sediment end-member, we first use the global subducted sediment (GLOSS; Plank and Langmuir, 1998), and results of all our calculations, presented below, show that GLOSS does constitute an appropriate sediment end-member. Previous studies have suggested that the PAO was surrounded by active margins and many pieces of PAO crust similar to KO were most probably exposed along these margins (Dobretsov et al., 1995; Jahn, 2004; Sengör et al., 1993; W.J. Xiao et al., 2009). This is especially evident in the southern CAOB as it is mainly comprised of juvenile mantle-derived materials (Jahn, 2004; Sun et al., 2008). Consequently, sediments derived from these margins would have unradiogenic Pb but radiogenic Nd isotopes. Among present-day subducted sediments (Plank and Langmuir, 1998; Plank et al., 2007), sediment from the Izu–Bonin trench has the least radiogenic Pb and radiogenic Nd isotopes, and incidentally its provenance is connected to an Indian MORB type mantle (Tollstrup et al., 2010). Thus, the Izu–Bonin sediment represents a realistic sediment end-member.

We use a simple two-component mixing model and the parameters in Table 3 to calculate the relative proportions of the depleted mantle and bulk sediment end-members in the source of KO mafic samples. Results of our calculations (Fig. 8) show that, except for one sample, mixing of bulk Izu–Bonin sediment and DIMM simply cannot produce the high Th/Nd for given  $^{143}\text{Nd}/^{144}\text{Nd}$  of the bulk of the mafic samples.

In order to explain the spectrum of Th/Nd for given  $^{143}\text{Nd}/^{144}\text{Nd}$  values of the bulk of the samples, we have to elevate the Th/Nd ratio of the bulk sediment first by partial melting (see also, Class et al., 2000; Singer et al., 2007). By iteration, we found that the sediment needs to be partially melted by  $\sim 5$  wt.% in order to cover the range of Th/Nd and  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios of the bulk of KO mafic rocks. Addition of  $\leq 1\%$  of this sediment partial melt to a highly depleted mantle source can account for the subduction component signature of KO mafic rocks.

### 6.2.2. Dehydration fluid

Many investigators believe that the subduction component consists primarily of hydrous fluid dehydrated from subducted altered oceanic crust and only a lesser amount of melt from subducted sediment (e.g., Plank and Langmuir, 1998; Turner and Hawkesworth, 1997). Thus, it is reasonable to assume that dehydration fluid, in addition to sediment melt discussed above, was also a part of the subduction component that influenced the source of KO mafic rocks. Thus, to further constrain the nature of subduction component in the mantle source of KO mafic rocks, we use both Nd and Pb isotopes to evaluate the importance of the dehydrated fluid in the slab component. Neodymium and Pb isotopes provide powerful constraints because they combine the relatively hydrous fluid-immobile character of Nd with hydrous fluid mobile character of Pb. For the model calculations, we use the same depleted mantle and sediment melt end-members as in the previous sediment melt model (Table 3). For the fluid end-member, we use the isotopic composition of the highly altered basalt sample KL-14 (LOI = 10.99%) of local altered oceanic crust (LAOC); for trace element composition, we mix 80% of fluid released from LAOC sample KL-14 with 20% fluid released from sediment (see Table 3).



**Fig. 9.** Diagrams of  $\epsilon_{Nd(t)}$ ,  $^{207}Pb/^{204}Pb(t)$  and  $^{208}Pb/^{204}Pb(t)$  versus  $^{206}Pb/^{204}Pb(t)$  showing the amounts of Nd and Pb in the subduction component that affected the KO sub-mantle. Partition coefficients and compositions of fluid from local altered oceanic crust (LAOC), sediment, and sediment melt are listed in Table 3 and explained in the text. Values for the subducted sediment and DIMM are the same as Fig. 8. Mixing lines between DIMM and fluid from LAOC and subducted sediment, DIMM and sediment melt, and especially between fluids and sediment melt indicate the isotopic composition of the subduction component (SC). Mixing calculation results suggest that the compositional array of the sub-KO mantle cannot be explained by mixing a SC that is either purely sediment or purely fluid; instead, the SC is a variable mixture of sediment and fluid; addition  $\leq 2\%$  of such a SC is consistent with the data. Symbols as in Fig. 8. Tick marks on the calculated mixing lines are labeled in wt.% of individual end-members.

Our model calculation results show that the bulk of KO mafic rocks are inside the area defined by the depleted mantle, sediments melt and fluid (Fig. 9). None of the samples has a subduction component that is either purely sediment or purely fluid; instead, the subduction component is a variable mixture of the two. Moreover, the subduction component flux to the depleted mantle wedge is variable because the amount of sediment subducted, the extent of partial melting of such sediment, amount of slab dehydration, and other factors such as temperature, pressure, water content, and oxidation state in the subduction zone (Kelley and Cottrell, 2009; Kelley et al., 2006) are also variable. In any case, the total amount of subduction component that mixed with the depleted mantle to generate the range of mantle sources of the parental melts of KO mafic samples in all model calculations is consistently  $\leq 2\%$ . Note that this combined fluid and sediment contribution to the subduction component is higher than sediment melt contribution alone (Fig. 8). Moreover, this is an upper limit of subduction component contribution as some of the KO mafic rocks (i.e., group 2 samples) may have been derived from a slightly enriched mantle source to begin with. The main point of these simple exercises using refractory elements and Nd and Pb isotopes is simply to show that subducted sediment and fluid dehydrated from altered oceanic crust are needed to produce the subduction component that influenced the mantle source of KO parental magmas.

## 7. Conclusions

Results of our combined major-trace element and Nd and Pb isotopic investigation suggest a common, Indian MORB-type mantle source for DO and KO mafic rocks. Thus, the isotopically Indian MORB-type upper mantle with a Dupal isotopic anomaly had been in existence since at least  $\sim 390$  Ma ago beneath the PAO. This lends support to the idea that the Dupal isotopic anomaly existed prior to the opening of the Indian Ocean. However, the composition of portions of the KO mantle was additionally influenced by a relatively  $\sim$ recent subduction process, which, importantly, can also produce a Dupal-like isotopic signature. Thus, caution must be taken when interpreting Dupal isotopic signals coming from ancient ophiolites. Model calculations for KO mafic rocks indicate that  $\leq 2\%$  of subduction component consisting of a variable mixture of  $\sim 5\%$  partial melt of sediment with unradiogenic Pb sediment and fluid dehydrated from altered oceanic crust was added to the depleted Indian mantle to produce KO parental magmas.

We conclude that the Dupal isotopic signature is a long-lived asthenospheric feature that existed beneath the Neo- and Paleo-Tethyan Ocean (275–350 Ma) in the southern hemisphere and in the southern section of PAO (371–390 Ma) in the northern hemisphere. The Dupal isotopic signature of the asthenosphere was most likely produced through contamination by upwelling plumes of an isotopically unusual deep mantle component generated through ancient subduction or delamination of lithospheric materials.

## Acknowledgments

We are grateful for constructive reviews by M. Sun, another one anonymous referee and editors that led to clarification and improvement of the manuscript. We also thank Liyan Tian and Chris MacIsaac for their help in the isotopic analysis. This study was jointly supported by NSFC funds (No. 41302041; 40930316; 41230207; 40425003), Guangxi National Natural Science Foundation of China (No. 2012GXNSFCA053007), China Postdoctoral Science Foundation Grant (2013M530440), open research grants of Guangxi Key Laboratory of Hidden Metallic Ore Deposits Exploration (No. 13-A-01-01) and the State Key Laboratory for Mineral Deposits Research (Nanjing University) (18-13-6), Bagui Scholar Innovation Project of Guangxi Province (to Xu Jifeng), research grants of GUT and China Scholarship Council Fund Ph.D. Study Abroad Program (2008). Part of this study was done by X. L. as a

visiting graduate student at the Scripps Institution of Oceanography, University of California, San Diego. This is a contribution to IGCP 592.

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