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### Title

The Paleogeography of Laurentia in Its Early Years: New Constraints From the Paleoproterozoic East-Central Minnesota Batholith

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Peer reviewed

1    **The paleogeography of Laurentia in its early years: new**  
2        **constraints from the Paleoproterozoic East-Central**  
3            **Minnesota batholith**

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12        **Key Points:**

- 13        • A new *ca.* 1780 Ma paleomagnetic pole reconstructs the Superior province of Lau-
- 14        rentia to moderately high latitudes
- 15        • This pole establishes the coherency of the Laurentia craton following Trans-Hudson
- 16        orogenesis and supports the NENA connection with Baltica
- 17        • Paleomagnetic and geologic data from Laurentia strongly support mobile-lid plate
- 18        tectonics from 2.2 Ga to the present day

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19 **Abstract**

20 The *ca.* 1.83 Ga Trans-Hudson orogeny resulted from collision of an upper plate  
 21 consisting of the Hearne, Rae, and Slave provinces with a lower plate consisting of the  
 22 Superior province. While the geologic record of *ca.* 1.83 Ga peak metamorphism within  
 23 the orogen suggests that these provinces were a single amalgamated craton from this time  
 24 onward, a lack of paleomagnetic poles from the Superior province following Trans-Hudson  
 25 orogenesis has made this coherency difficult to test. We develop a high-quality paleo-  
 26 magnetic pole for northeast-trending diabase dikes of the post-Penokean orogen East-  
 27 Central Minnesota Batholith (pole longitude: 265.8°; pole latitude: 20.4°;  $A_{95}$ : 4.5°; K:  
 28 45.6 N: 23) whose age we constrain to be  $1779.1 \pm 2.3$  Ma (95% CI) with new U-Pb dates.  
 29 Demagnetization and low-temperature magnetometry experiments establish dike rema-  
 30 nence be held by low-Ti titanomagnetite. Thermochronology data constrain the intru-  
 31 sions to have cooled below magnetite blocking temperatures upon initial emplacement  
 32 with a mild subsequent thermal history within the stable craton. The similarity of this  
 33 new Superior province pole with poles from the Slave and Rae provinces establishes the  
 34 coherency of Laurentia following Trans-Hudson orogenesis. This consistency supports  
 35 interpretations that older discrepant 2.22 to 1.87 Ga pole positions between the provinces  
 36 are the result of differential motion through mobile-lid plate tectonics. The new pole sup-  
 37 ports the NENA connection between the Laurentia and Fennoscandia cratons. The pole  
 38 can be used to jointly reconstruct these cratons *ca.* 1780 Ma strengthening the paleo-  
 39 geographic position of these major constituents of the hypothesized late Paleoprotero-  
 40 zoic supercontinent Nuna.

41 **1 Introduction**

42 In the Orosirian Period of the Paleoproterozoic Era, a series of collisional oroge-  
 43 nies led to the amalgamation of Archean provinces to form the core of the Laurentia cra-  
 44 ton (Fig. 1; Hoffman (1988); Whitmeyer and Karlstrom (2007)). The most significant  
 45 of these orogenies was the *ca.* 1850 to 1800 Ma Trans-Hudson orogeny associated with  
 46 the collision between the Superior province and the Churchill province which comprised  
 47 a composite of the Slave, Hearne and Rae provinces (Fig. 1; Weller and St-Onge (2017)).  
 48 The length of the orogen as well as the pressure-temperature of metamorphism within  
 49 it are similar to that of continent-continent collision within the Himalayan orogen (Weller  
 50 & St-Onge, 2017). The terminal closure of the intervening ocean basin between the Su-

51 perior and composite Slave + Hearne + Rae provinces is interpreted in paleogeographic  
52 models to be associated not only with the assembly of Laurentia, but also with the con-  
53 joining of other continents into the hypothesized supercontinent Nuna (Pehrsson et al.,  
54 2015).

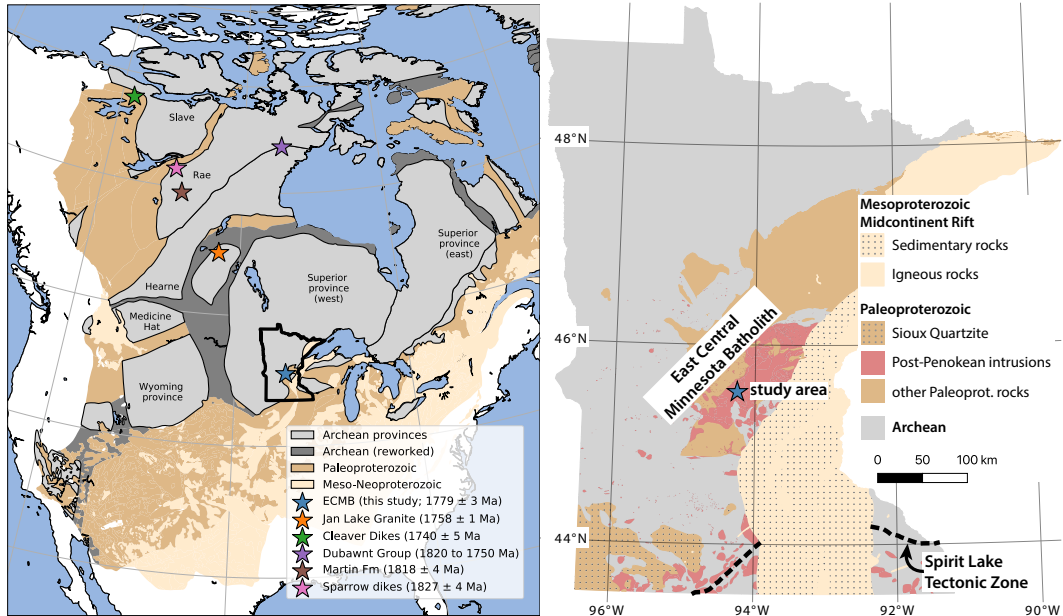
55 The rapid Paleoproterozoic amalgamation of the Laurentia craton led to the large  
56 persistent area of continental lithosphere that would grow further through accretionary  
57 orogenesis subsequently in the Paleoproterozoic Era and through the Mesoproterozoic  
58 Era (Whitmeyer & Karlstrom, 2007). This subsequent orogenesis along the southern to  
59 eastern margin of Laurentia (present-day coordinates) indicates that it was a long-lived  
60 accretionary margin (Karlstrom et al., 2001; Whitmeyer & Karlstrom, 2007). This ac-  
61 cretionary margin has been interpreted to have extended beyond Laurentia and have con-  
62 tinued onto Baltica and Australia (Karlstrom et al., 2001). Based on correlation of Archean  
63 provinces and Paleoproterozoic orogenic belts, Gower et al. (1990) reconstructed Baltica  
64 to Laurentia in a position known as the NENA (northern Europe and North America)  
65 configuration. This reconstruction is compatible with existing paleomagnetic constraints  
66 from *ca.* 1750 to 1270 Ma (Evans & Pisarevsky, 2008) and these conjoined cratons fea-  
67 ture as a major component of the hypothesized Nuna supercontinent (Evans & Mitchell,  
68 2011; Zhang et al., 2012).

69 Additional data that constrain the paleogeographic position of Laurentia from the  
70 time just following the Trans-Hudson orogeny can test the hypothesis of the unity of Lau-  
71 rentia's Archean provinces, establish the position of the newly amalgamated Laurentia,  
72 and thereby enable tests of hypothesized connections with other cratons. This study de-  
73 velops a new paleomagnetic pole for Laurentia from *ca.* 1780 Ma diabase dikes of the  
74 East-Central Minnesota Batholith (ECMB) that provides such constraints.

## 75 **2 Geologic Setting**

76 Coeval with collisional orogenesis between the assembling Archean provinces that  
77 formed Laurentia's core was the *ca.* 1860 to 1820 Ma accretionary Penokean orogeny along  
78 the southern margin of the Superior Province (Fig. 1; Schulz and Cannon (2007)). Penokean  
79 orogenesis resulted from island-arc and microcontinent collisions with the Superior Province  
80 that led to metamorphism of Superior Province lithologies and development of a fore-  
81 land basin (Schulz & Cannon, 2007; Holm et al., 2019). Following the Penokean orogeny,



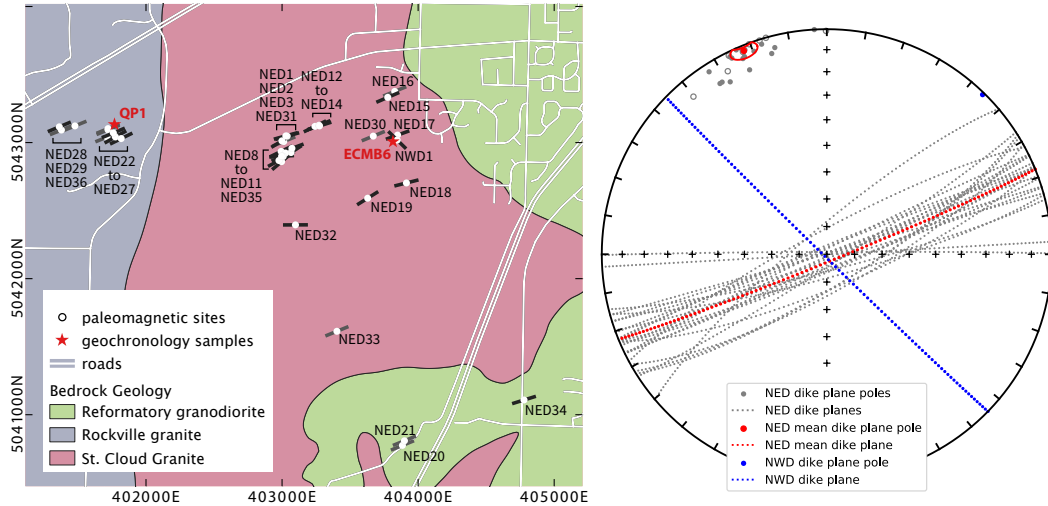


**Figure 1.** Map of Laurentia showing the location of Archean provinces and younger Proterozoic crust (simplified from Whitmeyer and Karlstrom (2007)). The localities of paleomagnetic poles that constrain Laurentia’s position just after its amalgamation are shown with stars including the new pole from this study developed from the East-Central Minnesota Batholith (ECMB). The outline of the state of Minnesota around the ECMB blue star is the region for the geologic map in the right panel. This map shows interpreted Precambrian geology for the state of Minnesota (simplified from Jirsa et al. (2012)) including in regions covered by Phanerozoic sedimentary rocks where the Precambrian bedrock is inferred from geophysical data and drill cores.

82 there was voluminous magmatism along the southeastern margin of the west Superior  
 83 Province resulting in the emplacement of the *ca.* 1780 Ma East-Central Minnesota Batholith  
 84 (ECMB) and other coeval post-orogenic plutons (Fig. 1; Holm et al. (2005); Boerboom  
 85 et al. (2005); Schmitz et al. (2018)). While the ECMB is dominantly comprised of fel-  
 86 sific to intermediate plutons, mafic magmas were also generated and commingled with the  
 87 more abundant felsic magmas throughout the interval of batholith generation as evidenced  
 88 by mafic enclaves within some of the plutons (Boerboom et al., 2005, 2011; Schmitz et  
 89 al., 2018). Mafic melt within the ECMB also led to the emplacement of a set of near-  
 90 vertical northeast-trending diabase dikes (Fig. 2; Boerboom et al. (2005)). The dikes have  
 91 chilled margins and are typically 1 to 3 meters wide with widths up to 8 meters (Boerboom  
 92 et al., 2005). As with the granites they intrude, the dikes have primary igneous texture  
 93 and no metamorphic fabric (Boerboom et al., 2005). They have experienced variable low-  
 94 grade alteration such as albitization and sericitization of plagioclase and the formation

95 of pyrrhotite. These diabase dikes are present within all of the granitoid units of the ECMB  
96 (e.g., St. Cloud Granite, Rockville Granite, and Reformatory Granodiorite; Fig. 2) with  
97 the exception of the youngest Richmond Granite. Throughout the field area, the dikes  
98 are exposed both in glacially-polished pavement outcrops and in numerous inactive and  
99 active dimension stone granite quarries. Northeast-trending diabase dikes are present  
100 in all of the quarries in the Rockville Granite, St. Cloud Granite as well as in the Re-  
101 formatory granodiorite, regardless of the size of the quarry, as well as in many natural  
102 bedrock outcrops. In many of the old inactive quarries, the north and/or south quarry  
103 walls are marked by the planar surface of a diabase dike contact, where the rock nat-  
104 urally separates, often resulting in elongated northeast-southwest shapes to the quarry  
105 pits. In contrast, no diabase dikes have been found in the quarries or natural exposures  
106 of the Richmond granite. Although this granite does not contain as many quarries and  
107 there are fewer natural outcrops, the lack of diabase dikes contrasts sharply with the nu-  
108 merous dikes present in the other nearby granites, where an equivalent exposed surface  
109 area would contain numerous diabase dikes. This absence indicates that the younger Rich-  
110 mond Granite post-dates the intrusion of the diabase dikes into the St. Cloud Granite,  
111 Rockville Granite, and Reformatory Granodiorite.

112 There are also quartz-feldspar porphyry dikes with the same northeast-trending di-  
113 rection as the diabase dikes found in all the granitoids also with the exception of the Rich-  
114 mond Granite (Boerboom et al., 2005). These porphyritic microgranite dikes have chilled,  
115 and locally flow-banded, margins. One has been observed to have intruded into a northeast-  
116 trending diabase dike and another has textures consistent with commingling of magmas  
117 between the felsic dike and adjacent diabase dike indicative of synchronous emplacement  
118 (Boerboom & Holm, 2000). The Richmond Granite has trachytoid magmatic fabric de-  
119 fined by aligned potassium-feldspar phenocrysts that share the same orientation with  
120 the northeast-trending dikes (Boerboom & Holm, 2000), indicating that this orientation  
121 is associated with a persistent regional stress field throughout the interval of magma em-  
122 placement and dike formation. These field relationships indicate that the quartz-feldspar  
123 porphyry and diabase dikes are comagmatic with the batholith. The diabase dikes are  
124 constrained to be younger than the St. Cloud Granite (new U-Pb date of  $1781.44 \pm 0.51$   
125 Ma;  $2\sigma$  analytical uncertainty) which they pervasively intrude, older than the Richmond  
126 Granite (new U-Pb date of  $1776.76 \pm 0.49$  Ma) in which they are absent, and similar  
127 in age to the quartz-feldspar porphyry dikes (new U-Pb date of  $1780.78 \pm 0.45$  Ma).



**Figure 2.** Left panel: Locations of paleomagnetic sites of the northeast-trending dikes (NED) and a northwest-trending dike (NWD) within the Rockville Granite, Reformatory Granodiorite and St. Cloud Granite of the ECMB (bedrock geology from Boerboom et al. (1995) and shown in UTM zone 15N WGS 84 coordinate reference system such that each axis tick is 1 km). Within regions of the mapped St. Cloud Granite there is more complex interfingering of that granite with the Reformatory Granodiorite than is shown. The strikes of the dikes are shown as lines (black when measured on that dike; grey when using the overall mean orientation from the measured NED dikes). The location of the QP1 and ECMB6 geochronology samples are shown. The ECMB4 geochronology sample was collected  $\sim 18$  km SW of the western edge of the map, in the Richmond Granite which cross-cuts the Rockville Granite and is younger than the NED dikes. Right panel: The orientations of dikes. Each individual dike orientation is the mean of multiple measurements on that dike. The mean of the poles to the NED planes is shown with a red dot and a 95% confidence ellipse on the mean calculated with Fisher statistics. This confidence ellipse intersects the equator indicating that the mean plane cannot be distinguished from vertical.

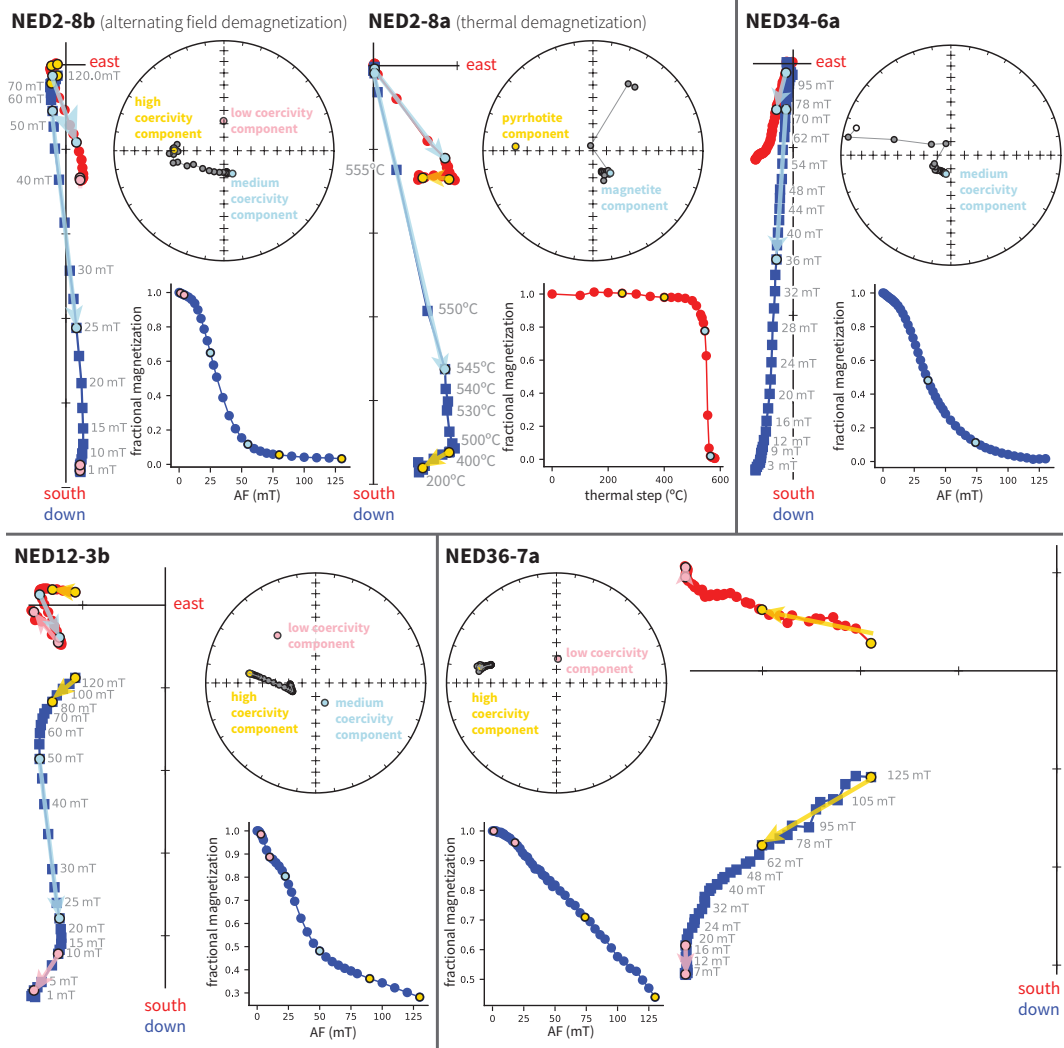
### 128 3 Paleomagnetic Methods and Results

129 Oriented samples for paleomagnetism were collected and analyzed from 36 of the  
 130 northeast-trending dikes of the ECMB and one northwest-trending dike (Fig. 2). Each  
 131 sampled dike constituted a paleomagnetic site in our site naming scheme. These sites  
 132 were concentrated in and around Stearns County Quarry Park near the city of St. Cloud  
 133 (Fig. 2). Samples were collected from the dikes with a gas-powered drill and oriented  
 134 in the field with a Pomeroy orienting fixture. The azimuthal orientations of the cores were  
 135 determined either through sun or magnetic compass depending on cloud cover. Sun com-  
 136 pass directions were preferentially used when available. When magnetic compass data  
 137 were used they were corrected for local magnetic declination using the International Ge-  
 138 omagnetic Reference Field model (Thébault et al., 2015). Specimens from the oriented

139 samples were analyzed in the UC Berkeley Paleomagnetism lab using a 2G DC-SQUID  
140 magnetometer. Samples either underwent stepwise alternating field (AF) or thermal de-  
141 magnetization. Thermal demagnetization was accomplished using an ASC thermal de-  
142 magnetizer (residual fields <10 nT). AF demagnetization was conducted with inline coils  
143 that utilize a Crest Amplifier paired with an Adwin controller to develop a well-controlled  
144 waveform. All paleomagnetic data developed in this study are available at the measure-  
145 ment level in the MagIC database ([https://www.earthref.org/ MagIC/ doi/ INSERT-](https://www.earthref.org/MagIC/doi/INSERT-DOI)  
146 DOI; UPDATE TO DOI WHEN ASSIGNED).

147 Typical behaviors of sample remanence during demagnetization are illustrated for  
148 representative specimens in Figure 3. AF demagnetization data typically reveal three  
149 components: a small low-coercivity component approximately aligned with Earth’s present  
150 local field in the study region that was typically removed below 10 mT; a medium-coercivity  
151 component that is steep and was dominantly removed between 10 and 60 mT; and a high-  
152 coercivity component that was subsequently removed incompletely as demagnetization  
153 progressed to 130 mT. These components are present to varying degrees within individ-  
154 ual specimens (Fig. 3).

155 Sister specimens from some samples underwent thermal and AF demagnetization  
156 which provides additional insight into the carriers of the components through compar-  
157 ison of the thermal and AF demagnetization spectra (such as NED2-8 in Fig. 3). These  
158 data reveal that the low-coercivity component direction is removed at the lowest unblock-  
159 ing temperatures up to 150°C. This behavior, as well as the typical direction, is most con-  
160 sistent with the component being a viscous overprint acquired in Earth’s geomagnetic  
161 field. The direction of the high-coercivity component is removed through thermal de-  
162 magnetization between 250°C and 350°C — consistent with it being held by monoclinic  
163 pyrrhotite. The direction of this magnetization held by pyrrhotite is aligned with the magnetite-  
164 held remanence within a northwest-trending dike in the region (discussed in more de-  
165 tail below) — a direction consistent with the position of Laurentia during the time pe-  
166 riod of *ca.* 1096 Ma Midcontinent Rift magmatism (Swanson-Hysell et al., 2020). We  
167 interpret this high-coercivity component held by pyrrhotite, whose presence is variable  
168 in ECMB dikes, to have formed through hydrothermal activity associated with Midcon-  
169 tinent Rift magmatism such as that represented by the emplacement of the northwest-  
170 trending dike. The pyrrhotite thereby carries a chemical remanent magnetization.



**Figure 3.** Paleomagnetic data from ECMB northeast-trending diabase dikes are shown in geographic coordinates on vector component plots, measurement-level equal area plots and magnetization magnitude plots (developed using PmagPy software; Tauxe et al. (2016)). Least-squares fits to the data are shown with colored arrows on the vector component plots, colored directions on the equal area plots, and as colored end-points on the magnetization magnitude plots (pink for low-coercivity; blue for medium-coercivity; yellow for high-coercivity). Specimens NED2-8a and NED2-8b are from the same core sample and were analyzed via thermal and alternating field (AF) demagnetization respectively. These data from sample NED2-8 reveal the steep medium-coercivity component to thermally unblock at temperatures characteristic of remanence held by magnetite and the high-coercivity component to thermally unblock at temperatures characteristic of pyrrhotite. Specimen NED34-6a is dominated by the steep medium-coercivity component. The medium-coercivity component is well-resolved in specimen NED12-3b which also has a substantial high-coercivity component. The high-coercivity component dominates the remanence of specimen NED36-7a such that no medium-coercivity component can be resolved.

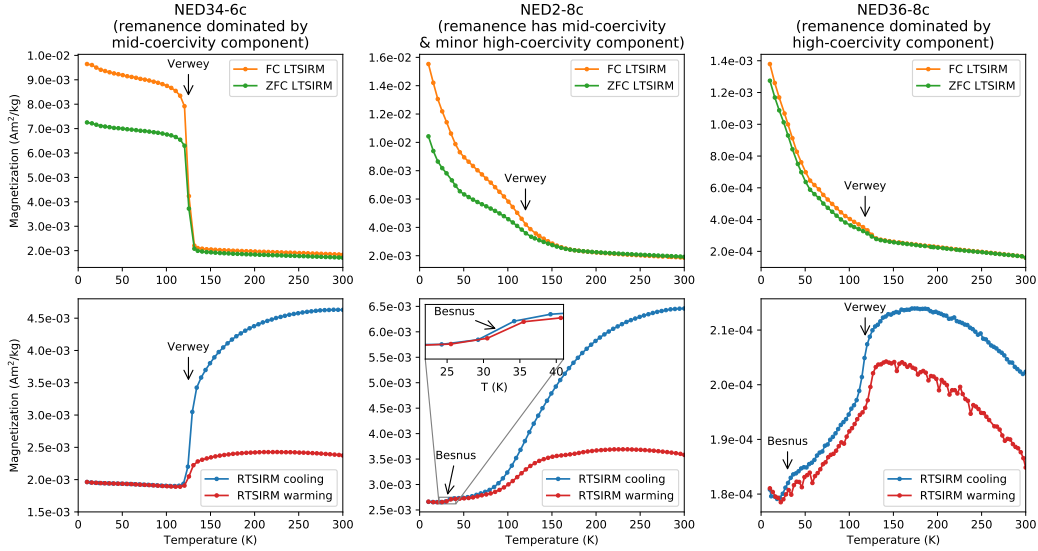
171 In some specimens, the low-coercivity component is much larger and has a direc-  
172 tion that is distinct from the present local field. It is likely that these samples acquired  
173 an isothermal remanent magnetization associated with quarrying operations or lightning  
174 strikes. This behavior can be prevalent throughout a site or can be present in just some  
175 samples from a given site. In many cases, these low-coercivity overprints can be removed  
176 through AF demagnetization and the medium-coercivity and/or high-coercivity compo-  
177 nents can be subsequently isolated. In some instances, however, these large and dom-  
178 inantly low-coercivity overprints extend to higher coercivities preventing the isolation  
179 of other components.

180 The medium-coercivity component direction is dominantly removed through ther-  
181 mal demagnetization between 515°C and 565°C consistent with it being held by low-Ti  
182 titanomagnetite. This direction was recovered with site mean direction uncertainty less  
183 than 8 degrees ( $\alpha_{95} < 8^\circ$ ) for 23 sites (Table 1). We interpret this component to be a  
184 primary thermal remanence acquired at the time of dike emplacement as part of the *ca.*  
185 1780 Ma ECMB. This interpretation gains support from the rock magnetic data, an in-  
186 verse baked contact test, and thermochronology data that support an emplacement tem-  
187 perature well below the blocking temperature of magnetite and a mild subsequent ther-  
188 mal history — as discussed in more detail below.

### 189 **3.1 Magnetic mineralogy constraints from low-temperature magnetom-** 190 **etry**

191 To gain additional insight into the magnetic mineralogy of the dikes, a Magnetic  
192 Properties Measurement System (MPMS) at the Institute for Rock Magnetism was used  
193 to conduct low-temperature remanence experiments. In the field-cooled (FC) experiments  
194 shown in Figure 4, the magnetization was measured upon warming following the spec-  
195 imen having cooled in an applied field of 2.5 T from 300 to 10 K. In the zero-field-cooled  
196 (ZFC) experiment, a low-temperature saturation isothermal remanence (LTSIRM) of 2.5  
197 T was applied at 10 K after the specimen cooled in a (near-)zero field. In the room-temperature  
198 saturation isothermal remanence (RTSIRM) experiment, the sample was pulsed with a  
199 2.5 T field at room temperature ( $\sim 300$  K) and then cooled to 10 K and warmed back  
200 to room temperature in a (near-)zero field. These experiments reveal that sister spec-  
201 imens to paleomagnetic specimens whose remanence is dominated by the medium-coercivity  
202 component without an appreciable high-coercivity component have strong expressions

203 of the  $\sim 120$  K Verwey transition as expected for a ferromagnetic mineralogy of well-preserved  
 204 low-Ti titanomagnetite (NED34-6c in Fig. 4; Verwey (1939); Feinberg et al. (2015)). In  
 205 contrast, specimens from samples that have a larger contribution of the high-coercivity  
 206 component have weaker saturation magnetization, minor expression of the Verwey transi-  
 207 tion, and the presence of monoclinic pyrrhotite as evidenced through the  $\sim 30$  K Besnus  
 208 transition (NED36-8c in Fig. 4; Besnus and Meyer (1964); Feinberg et al. (2015)). Sam-  
 209 ples with a smaller contribution of the high-coercivity component superimposed on the  
 210 medium-coercivity component have intermediate behavior with a minor expression of  
 211 the Besnus transition and a more prominent Verwey transition (NED2-8c in Fig. 4). These  
 212 results support the interpretation that the medium-coercivity component is held by pri-  
 213 mary unaltered (titano)magnetite and that the high-coercivity component is the result  
 214 of subsequent alteration that resulted in degradation of magnetite and formation of pyrrhotite.



**Figure 4.** Low-temperature remanence experiment data. The specimen from a sample whose natural remanent magnetization is dominated by the medium-coercivity component (NED34-6c) has behavior dominated by magnetite as evidenced through the response across the  $\sim 120$  K Verwey transition. The specimen from a sample whose natural remanence is dominated by the high-coercivity component (NED36-8c) has weaker magnetization, a relatively minor expression of the Verwey transition, and expression of the  $\sim 32$  K Besnus transition that indicates the presence of monoclinic pyrrhotite. The specimen whose natural remanence has a well-resolved medium-coercivity component with a minor high-coercivity component (NED2-8c) has intermediate behavior with a Verwey transition that is not as suppressed as in NED36-8c with a minor, but resolvable, Besnus transition (see inset). FC: field-cooled; ZFC: zero-field-cooled; LTSIRM: low-temperature saturation isothermal remanence magnetization; RTSIRM: room-temperature saturation isothermal remanence magnetization.

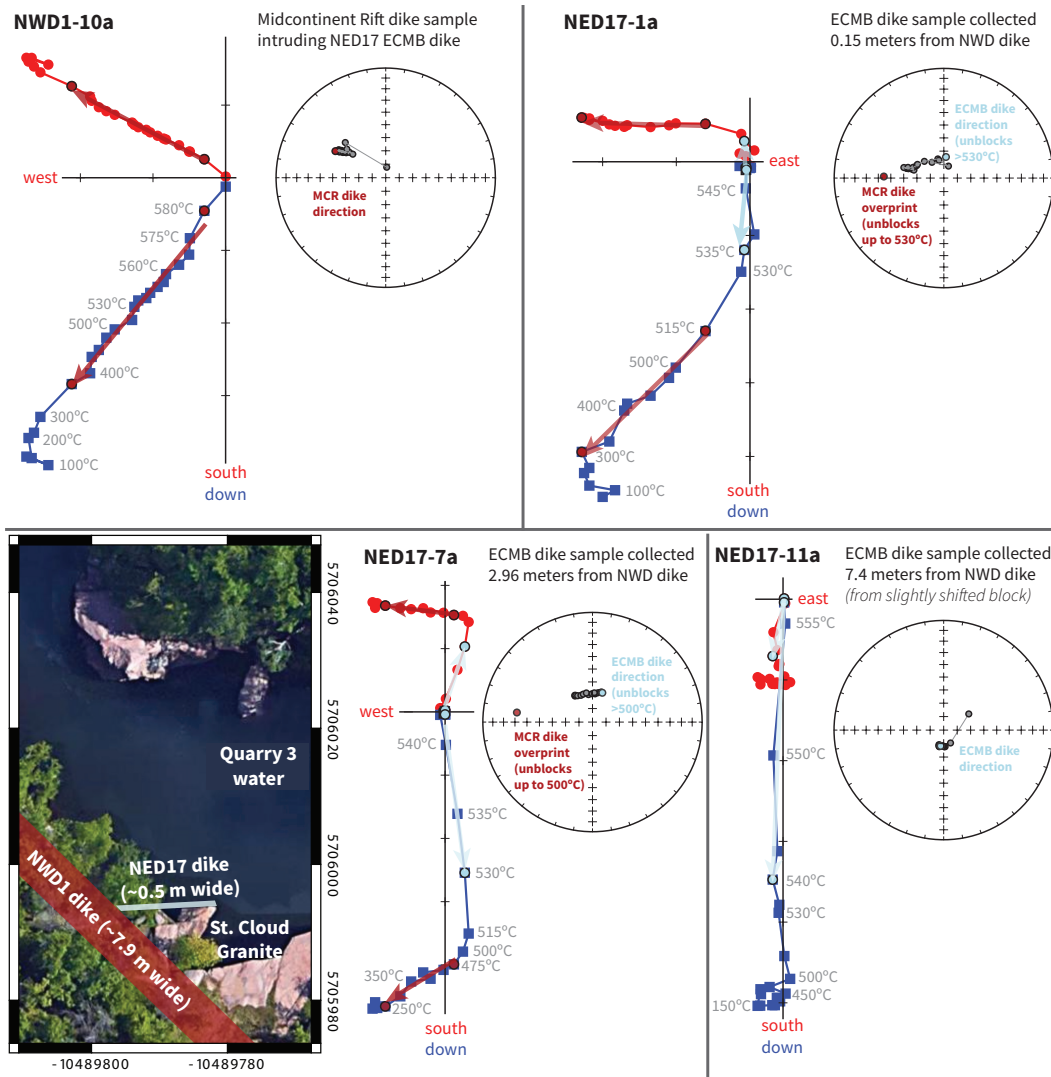


### 3.2 Baked contact test

One northwest-trending dike is exposed and was sampled within the study region as site NWD1 (Fig. 2). The magnetization direction of this dike indicates that it is associated with the main stage of the Midcontinent Rift (*ca.* 1096 Ma) as it has a normal polarity and an inclination consistent with that time interval of Midcontinent Rift volcanism (Fig. 5; Swanson-Hysell et al. (2020)). The dike cross-cuts one of the northeast-trending dikes (NED17) allowing for a baked contact test (Figs. 2 and 5). The baked contact test is positive with a distinct direction in the northeast-trending dike (corresponding to the remanence direction seen throughout the northeast-trending dikes of the ECMB) with its magnetite remanence becoming progressively overprinted by the northwest-trending dike up to the contact (Fig. 5). This positive baked contact test indicates that the northeast-trending ECMB dikes have not been overprinted since the northwest-trending dike was emplaced (*ca.* 1096 Ma). This positive baked contact test for the northwest-trending dike constitutes what is referred to as a positive “inverse” baked contact test for the northeast-trending dike remanence — it constrains the remanence to be more ancient than the *ca.* 1.1 Ga Mesoproterozoic northwest-trending dike, but does not provide a constraint back to the Paleoproterozoic time of dike emplacement. Given that the host rocks for the northeast-trending dikes are of a very similar age to the dikes themselves there is not the possibility of a Paleoproterozoic baked contact test. In contrast to the dikes, stable and consistent remanence directions were not recovered from pilot sites in ECMB granites.

The high-coercivity remanence direction held by pyrrhotite in some of the northeast-trending dikes is aligned with the remanence direction of the northwest-trending dike. While the thermal effect of the northwest-trending dike was limited to a few meters on either side of it as evidenced through the baked contact test (Fig. 5), there was more widespread hydrothermal alteration associated with Midcontinent Rift magmatic activity that led to the formation of pyrrhotite and an associated chemical remanent magnetization. In the majority of the northeast-trending dikes, the original magnetite is well-preserved (e.g., NED34 of Figs. 3 and 4) while other dikes experienced variable magnetite alteration and the formation of secondary pyrrhotite (e.g., NED36 of Figs. 3 and 4). These components can be separated through progressive demagnetization (e.g., NED12 of Figs. 3) enabling the thermal remanent magnetization held by magnetite to be used to determine site mean directions and calculate a paleomagnetic pole (Fig. 6).





**Figure 5.** Results from a positive baked contact test where a northwest-trending dike (site NWD1) cross-cuts a northeast-trending dike (site NED17). The NWD1 dike has a direction indicating that it formed during the *ca.* 1096 Ma main stage of rift volcanism. Close to the NWD1 dike the NED17 is nearly fully overprinted (NED17-1a). Further from NWD1 there are partial thermal overprints (NED17-7a) that by  $\sim 7$  meters from the cross-cutting dike are not resolvable (NED17-11a). Note that blocks from this dike were slightly shifted due to quarrying operations which does not bear on the significance of the positive test, but has led to the exclusion of the result from the site mean compilation and overall pole.

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### 3.3 Paleomagnetic pole

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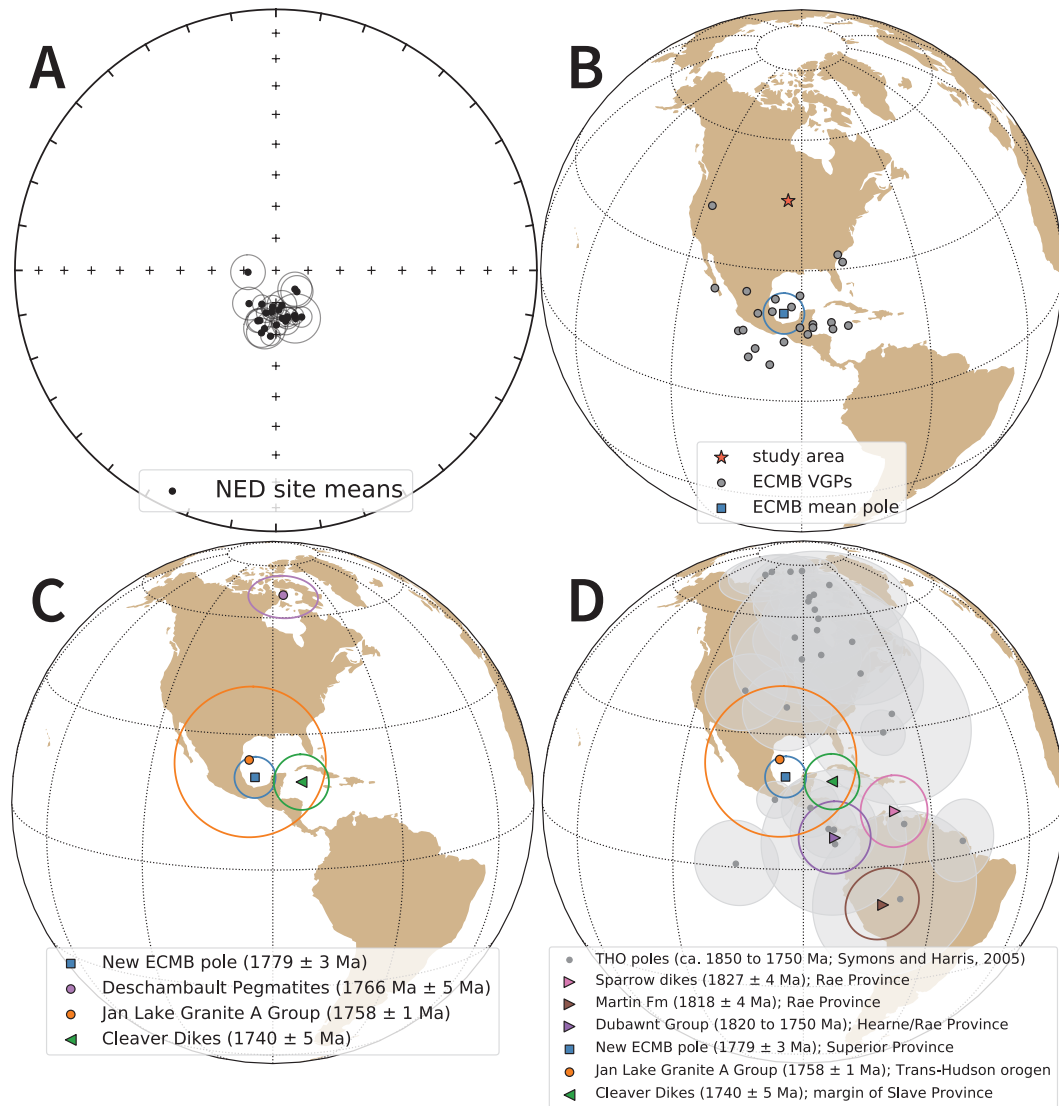
The site mean directions determined from the magnetite remanence component can be converted to virtual geomagnetic poles and then used to calculate a mean paleomagnetic pole for the ECMB diabase dikes (pole longitude:  $265.8^\circ$ ; pole latitude:  $20.4^\circ$ ;  $A_{95}$ :  $4.5^\circ$ ; K: 45.6 N: 23; Fig. 6). These 23 virtual geomagnetic poles have a distribution con-

253 consistent with a Fisher distribution as determined through the Fisher et al. (1987) goodness-  
 254 of-fit method. The  $A_{95}$  uncertainty on the mean pole position of  $4.5^\circ$  is within the bounds  
 255 of reliability proposed by Deenen et al. (2011). It is well below the  $A_{95}$ -max value pro-  
 256 posed to establish a well-determined mean for 23 sites ( $11.4^\circ$ ) and above the  $A_{95}$ -min  
 257 value ( $3.4^\circ$ ) consistent with the site directions having sufficiently sampled secular vari-  
 258 ation of the geomagnetic field.

259 In a massive host rock such as the ECMB plutons without preferential bedding or  
 260 foliation, it is expected that lithospheric stresses will lead to the emplacement of near  
 261 vertical dikes. Dike plane orientations were measured on each dike for which there was  
 262 sufficient three-dimensional exposure. Multiple measurements were made for each dike  
 263 to constrain their orientation. The mean strike calculated from 17 dike orientations is  
 264  $067^\circ$  and the mean dip is  $88^\circ$  (Fig. 2). The  $\alpha_{95}$  uncertainty associated with the Fisher  
 265 mean for the poles to these dike orientation planes (i.e. the lines perpendicular to the  
 266 planes) is  $4.7^\circ$  which means that the overall orientation of the planes is statistically in-  
 267 distinguishable from vertical (Fig. 2). Due to this verticality, we interpret the exhuma-  
 268 tion of the ECMB plutons to have not resulted in appreciable tilting since dike emplace-  
 269 ment and do not apply a tilt correction to the paleomagnetic data.

#### 270 **4 Geochronology Methods and Results**

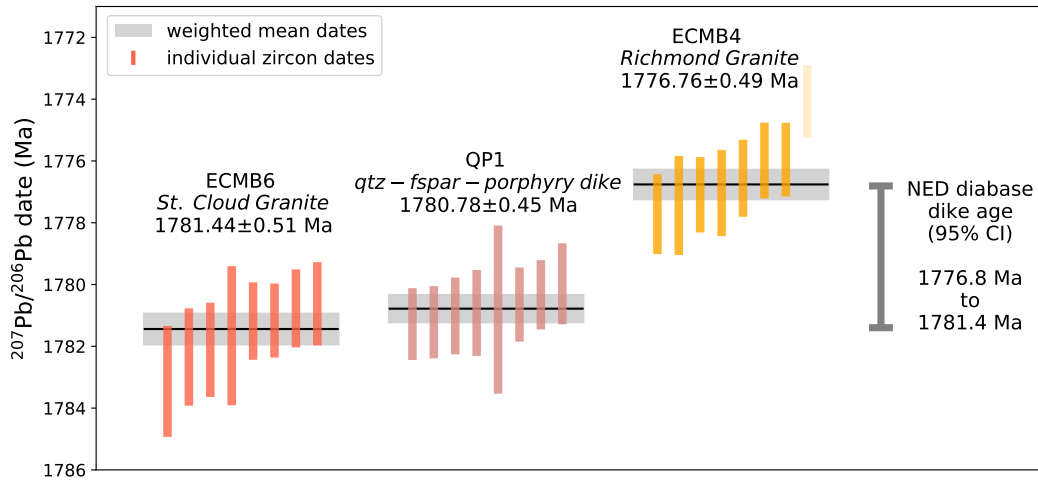
271 The field relationships show the diabase dikes to be younger than the Rockville Gran-  
 272 ite, Reformatory Granodiorite and the St. Cloud Granite which they pervasively intrude  
 273 and to be older than the Richmond Granite where they are absent (Boerboom et al., 2005).  
 274 Holm et al. (2005) developed U-Pb dates calculated as concordia intercept dates from  
 275 these intrusions. The dates reported by Holm et al. (2005) for granites intruded by the  
 276 dikes are  $1783 \pm 11$  Ma for the Reformatory Granodiorite,  $1780 \pm 7$  Ma for the Rockville  
 277 Granite and  $1779 \pm 5$  Ma for the St. Cloud Granite. The younger cross-cutting Rich-  
 278 mond Granite has a date of  $1772 \pm 3$  Ma (Holm et al., 2005). An age of  $1774 \pm 7$  Ma  
 279 for one of the quartz-feldspar porphyry dikes developed by Holm et al. (2005) is consis-  
 280 tent with this interpretation of these dikes being older than the Richmond Granite (and  
 281 younger than the granites they intrude). While the dates published in Holm et al. (2005)  
 282 are valuable constraints and are consistent with the field relationships, they are of lower  
 283 precision than what is possible with modern analytical approaches and therefore lead  
 284 to overlapping uncertainties. Higher precision constraints resulting from methods that



**Figure 6.** A) Site mean directions for the magnetite remanence of the northeast-trending (NED) ECMB diabase dikes with  $\alpha_{95} < 8^\circ$ . B) Virtual geomagnetic poles (VGPs) calculated from these site means and the overall mean paleomagnetic pole for the ECMB dikes. C) Comparison between the new ECMB paleomagnetic pole and other *ca.* 1780 to 1740 Ma poles for Laurentia. D) Comparison of poles from Laurentia's provinces from 1830 to 1740 Ma from Evans et al. (2021) as well as poles from the Trans-Hudson orogen (THO; grey; Symons and Harris (2005)) with the new ECMB pole.

285 apply ion-exchange separation with low blank analyses of chemically-abraded single zir-  
 286 con grains can test the field relationship interpretations and provide more confidence in  
 287 the overall age constraints.

288 To further constrain the age of the northeast-trending diabase dikes, we developed  
 289 new isotope dilution-thermal ionization mass spectrometry (ID-TIMS) U-Pb zircon dates



**Figure 7.** U-Pb dates for ECMB samples. Weighted mean dates (horizontal lines) are calculated from individual zircon dates (vertical bars). The NED diabase dikes intrude the St. Cloud Granite such that the ECMB6 weighted mean date of  $1781.44 \pm 0.51$  Ma is a maximum age. The quartz-feldspar porphyry dikes (one of which was sampled as QP1) also intrude the St. Cloud Granite and are parallel to the NED diabase dikes. Neither the quartz-porphyry dikes nor the diabase dikes intrude the younger cross-cutting Richmond Granite such that the ECMB4 weighted mean date of  $1776.76 \pm 0.49$  Ma provides a minimum age for the dikes. Details for the weighted mean dates are given in Table 2 and individual zircon data are in the Supporting Information.

290 from the St. Cloud Granite that host the dikes (sample ECMB6), the Richmond Gran-  
 291 ite from which the dikes are absent (sample ECMB4), and from a northeast-trending quartz-  
 292 feldspar porphyry dike (sample QP1) that is likely coeval with the diabase dikes (Figs.  
 293 2 and 7). Zircon crystals were chemically abraded prior to analysis of single zircon grains  
 294 by ID-TIMS at the Boise State Isotope Geology Laboratory (detailed geochronology meth-  
 295 ods are provided in the Supporting Information). Weighted mean dates were calculated  
 296 from multiple single zircon dates (Fig. 7; Table 2). While chemical abrasion served to  
 297 reduce Pb-loss and resulted in concordant analyses, some grains have persistent Pb-loss  
 298 and are discordant (Fig. S1). As a result, we calculate weighted mean  $^{207}\text{Pb}/^{206}\text{Pb}$  dates  
 299 rather than  $^{206}\text{Pb}/^{238}\text{U}$  dates (Fig. 7; Table 2). These  $^{207}\text{Pb}/^{206}\text{Pb}$  dates are  $1781.44$   
 300  $\pm 0.51$  Ma ( $2\sigma$  analytical uncertainty; MSWD = 1.24;  $n=8$ ) for the St. Cloud Granite  
 301 (ECMB6) and  $1776.76 \pm 0.49$  Ma (MSWD = 1.15;  $n=7$ ) for the Richmond Granite (ECMB4;  
 302 Fig. 7). The date for the sampled quartz-feldspar porphyry dike (QP1) of  $1780.78 \pm 0.45$   
 303 Ma (MSWD = 0.53;  $n=8$ ) is between these two dates as expected on the basis of field  
 304 relationships. Taking into account the analytical uncertainty on the maximum and min-  
 305 imum age constraints, the diabase dikes are younger than the  $1781.44 \pm 0.51$  Ma St. Cloud

306 Granite and older than the  $1776.76 \pm 0.49$  Ma Richmond Granite. If one assumes a uni-  
 307 form probability of diabase emplacement timing between the maximum and minimum  
 308 age constraints that have normally distributed uncertainties, the 95% confidence inter-  
 309 val can be estimated through Monte Carlo simulation. Applying this approach gives a  
 310 mean age of 1779.1 Ma with 95% confidence interval (CI) bounds of 1776.8 to 1781.4 Ma.  
 311 We can succinctly state the age of the northeast-trending ECMB diabase dikes as be-  
 312 ing  $1779.1 \pm 2.3$  Ma (95% CI).

## 313 5 Thermochronology Methods and Results

314 While the U-Pb zircon dates constrain the crystallization ages of the ECMB in-  
 315 trusions, additional insight into the thermal history of the batholith can help with in-  
 316 terpretation of the paleomagnetic data given that the thermal remanent magnetization  
 317 of magnetite will be blocked at temperatures below  $580^{\circ}\text{C}$ . As discussed below, existing  
 318 Ar-Ar dates on hornblende and biotite from the ECMB provide valuable constraints in  
 319 this regard (Fig. 8). In this study, we also develop new U-Pb apatite dates from three  
 320 ECMB granites (ECMB1, the Isle Granite; ECMB3, the Rockville Granite; ECMB4, the  
 321 Richmond Granite). In contrast to zircon, for which the temperatures of appreciable Pb  
 322 diffusion exceed the liquidus of granite (Cherniak & Watson, 2001), the temperature win-  
 323 dows for closure of the U-Pb system in apatite is much lower ( $\sim 510$  to  $460^{\circ}\text{C}$ ; Smye et  
 324 al. (2018)). As a result, U-Pb dates of apatite serve as a thermochronometer at moderately-  
 325 high temperatures (Chamberlain & Bowring, 2001; Schoene & Bowring, 2007; Cochrane  
 326 et al., 2014). These temperatures are of particular relevance to the interpretation of the  
 327 paleomagnetic data as they are lower than, or correspond with, the blocking tempera-  
 328 ture of low-Ti titanomagnetite. If a pluton was emplaced at depths where temperatures  
 329 exceed the closure temperature of apatite, or if it experienced prolonged reheating, the  
 330 U-Pb apatite dates would be appreciably younger than the U-Pb zircon crystallization  
 331 dates.

332 U-Pb data were developed from apatite grains separated from ECMB granites through  
 333 laser ablation inductively coupled plasma mass spectrometry (LA-ICP-MS) at UC Santa  
 334 Barbara (method details are provided in the supporting information). In contrast to zir-  
 335 con, apatite incorporates significant Pb at the time of crystallization. As a result of this  
 336 elevated common Pb, U-Pb dates were determined through the calculation of Tera-Wasserburg  
 337 concordia lower intercept dates where the upper intercept corresponds to the ratio of ini-

338 tial  $^{207}\text{Pb}/^{206}\text{Pb}$  and the lower intercept is the  $^{206}\text{Pb}/^{238}\text{U}$  date (following the method  
 339 of Ludwig (1998) as implemented in the IsoplotR software of Vermeesch (2018); Fig. S2).  
 340 Sample ECMB1 is from the Isle Granite which has a U-Pb zircon date of  $1779 \pm 26$  Ma  
 341 (Holm et al., 2005). The U-Pb apatite date for ECMB1 is  $1800.3 \pm 33.4/65.2$  Ma where  
 342 the first uncertainty is the  $2\sigma$  analytical uncertainty and the second uncertainty is 95%  
 343 confidence interval for the date that incorporates overdispersion (this uncertainty scheme  
 344 will be used for all the presented apatite dates). This U-Pb apatite date is therefore in-  
 345 distinguishable from the U-Pb crystallization date (Fig. 8). Sample ECMB3 is from the  
 346 Rockville Granite which has a U-Pb zircon date of  $1780 \pm 7$  Ma (Holm et al., 2005). The  
 347 U-Pb apatite date for ECMB3 is  $1810.5 \pm 16.2/23.0$  Ma. The youngest end of the overdis-  
 348 persion uncertainty range is very similar to (albeit just slightly older than) the U-Pb zir-  
 349 con date. It is not geologically reasonable for the U-Pb apatite date to be older than the  
 350 U-Pb zircon date. This result therefore suggests that the U-Pb apatite date uncertainty  
 351 is a slight underestimate with the U-Pb apatite system having closed just after the time  
 352 of zircon crystallization as constrained through the U-Pb zircon date. Sample ECMB4  
 353 is from the Richmond Granite for which we have developed a new ID-TIMS U-Pb zir-  
 354 con date of  $1776.76 \pm 0.49$  Ma. The U-Pb apatite date for ECMB4 is  $1751.7 \pm 17.8/36.6$   
 355 Ma which is indistinguishable from the U-Pb zircon date (Fig. 8).

356 Pb closure temperatures ( $T_c$ ) for the analyzed apatite grains can be estimated with  
 357 the Dodson (1973) approach assuming a cylindrical geometry with half-widths as the char-  
 358 acteristic diffusion length. Apatite grain sizes are similar across the three dated spec-  
 359 imens; they typically are 100-200  $\mu\text{m}$  long and 50-75  $\mu\text{m}$  wide. Using the Pb diffusiv-  
 360 ity values of Cherniak et al. (1991), a cooling rate of  $20^\circ\text{C}/\text{Myr}$  results in closure tem-  
 361 peratures of  $463^\circ\text{C}$  to  $473^\circ\text{C}$  for these grain sizes. A more rapid cooling rate is likely for  
 362 the batholith given the similarity of the U-Pb apatite dates with the U-Pb zircon dates.  
 363 A cooling rate of  $100^\circ\text{C}/\text{Myr}$  from crystallization to apatite closure temperatures results  
 364 in closure temperatures of  $493^\circ\text{C}$  to  $505^\circ\text{C}$ .

365 Overall, these data indicate that the samples cooled through the  $\sim 500^\circ\text{C}$  closure  
 366 temperatures of the U-Pb apatite system near the time of zircon crystallization consis-  
 367 tent with rapid cooling rates of the plutons (Fig. 8). Additionally, there has not been  
 368 significant diffusion due to later tectonothermal events. As discussed below, this result  
 369 is consistent with Ar-Ar hornblende dates from the ECMB granites and supports the mag-  
 370 netite remanence being a primary thermal remanent magnetization.

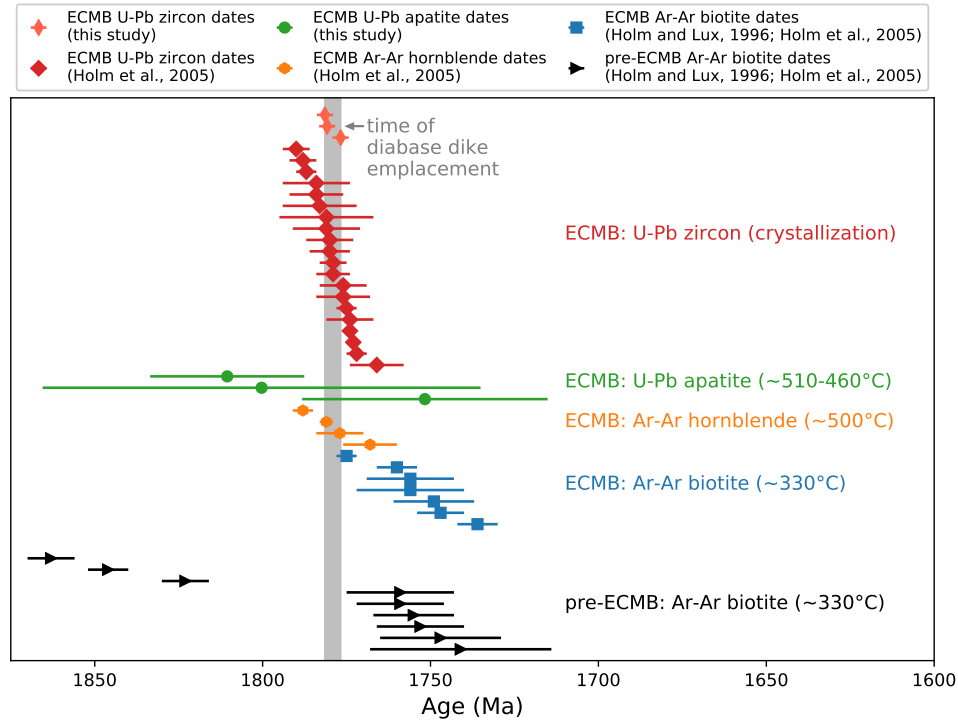
## 6 Discussion

### 6.1 Thermal history of the ECMB and a primary interpretation of the ECMB dike pole

Prior to the emplacement of the ECMB, Paleoproterozoic host rocks were metamorphosed to amphibolite facies during the Penokean orogeny (Holm & Selverstone, 1990). Emplacement of the ECMB has been hypothesized to be post-orogenic and associated with an interval of extensional collapse of the orogen (Holm & Lux, 1996; Boerboom & Holm, 2000). The Al-in-hornblende igneous barometer was applied to the St. Cloud and Isle Granites of the ECMB by Holm, Darrah, and Lux (1998). This barometer has varying published calibrations. Applying the pressure calibration of Mutch et al. (2016) to the data in Holm, Darrah, and Lux (1998) and assuming a  $2.7 \text{ g/cm}^3$  overburden gives an estimated emplacement depth of  $10.8 \pm 1.7 \text{ km}$  for the Freedhem Granodiorite,  $\sim 10.4 \pm 1.7 \text{ km}$  for the Isle Granite and  $13.4 \pm 2.1 \text{ km}$  for the St. Cloud Granite. The calibration of Ague (1997) leads to slightly higher calculated pressures implying depths that are  $\sim 2.3 \text{ km}$  deeper.

Thermochronology data give additional insight into emplacement temperatures (and thereby depth). Both the Ar-Ar hornblende dates published by Holm et al. (2005) and the U-Pb apatite dates developed in this study from ECMB lithologies are indistinguishable from the crystallization ages of the intrusions (Fig. 8). The closure temperature for Ar in hornblende is  $\sim 580$  to  $490^\circ\text{C}$  (Harrison, 1982). The closure of the U-Pb system in the dated apatite grains is  $\sim 500$  to  $460^\circ\text{C}$ . The consistency between the U-Pb zircon crystallization and the U-Pb apatite and Ar-Ar hornblende cooling dates indicates that the present-day erosion level of the ECMB was at a shallow enough depth that the crustal temperatures were lower than these closure temperatures at the time of emplacement of the plutons. Geothermal gradients in continental arc settings are typically between  $25$  to  $45^\circ\text{C/km}$  – potentially higher at  $1.8 \text{ Ga}$  (Rothstein & Manning, 2003). Taking a geothermal gradient of  $30^\circ\text{C/km}$  and the closure temperature constraints indicates that the plutons were emplaced at  $15 \text{ km}$  or shallower in the continental lithosphere. This emplacement depth is consistent with the Al-in-hornblende paleobarometry estimates.

Ar-Ar biotite dates provide insight into even lower temperatures as the system blocks at  $\sim 330^\circ\text{C}$  (Grove & Harrison, 1996), well below the blocking temperature of magnetite magnetization in the dikes. Ar-Ar biotite dates from ECMB plutons range from over-



**Figure 8.** Summary of U-Pb zircon dates, U-Pb apatite dates, Ar-Ar hornblende, and Ar-Ar biotite dates from the East-Central Minnesota batholith (ECMB) from this study, Holm and Lux (1996) and Holm et al. (2005). Approximate closure temperatures associated with the thermochronometers are labeled next to the relevant data. That the U-Pb apatite and Ar-Ar hornblende dates are indistinguishable from the U-Pb zircon crystallization dates indicates that the plutons were emplaced at upper middle to upper crustal levels. The Ar-Ar biotite dates from both the ECMB plutons and older host rock lithologies indicate exhumation within  $\sim 20$  million years to shallower depths and a lack of regional tectonothermal activity over the following 1.75 billion years.

lapping the crystallization dates to being younger by  $\sim 20$  million years (Fig. 8). Ar-Ar biotite dates from older host lithologies to the ECMB are either older than the age of the batholith itself or the same age as the Ar-Ar biotite dates from the batholith (Fig. 8). These data suggest that the batholith was emplaced near the upper depth range estimates from the Al-in-hornblende barometry ( $\sim 10$  km) and underwent exhumation to below the  $\sim 330^\circ\text{C}$  closure temperature of the K-Ar biotite system soon after emplacement of the plutons. These data also indicate that there has not been reheating or pervasive fluid flow that would have perturbed the Ar-Ar thermochronometers in the granites in the time since initial cooling. The magnetization used to develop the paleomagnetic pole comes from remanence held by low-Ti magnetite that dominantly unblocks between  $540$  and  $560^\circ\text{C}$  (Figs. 3 and 5). The thermochronology results constrain the rocks



414 to have cooled through the magnetite blocking temperatures at the time that the dikes  
415 were emplaced within the batholith and to not have experienced reheating that would  
416 have perturbed the thermochronometers. These data support an interpretation that the  
417 magnetite remanence within the ECMB dikes is a primary thermal remanent magneti-  
418 zation.

419 These thermochronology data also demonstrate that following the emplacement of  
420 the ECMB there were not regional tectonothermal events with the potential to have ther-  
421 mally modified the magnetization of the magnetite within the ECMB diabase dikes. Sub-  
422 sequent Paleoproterozoic Yavapai and Mazatzal accretion occurred to the southeast of  
423 the Spirit Lake Tectonic zone on the other side of the Minnesota River Valley promon-  
424 tory (~160 km south of the study region; Fig. 1; Holm et al. (2007)). In contrast to the  
425 northeast-trending dikes in the ECMB, northeast-trending dikes in Wisconsin within *ca.*  
426 1840 Ma plutons were metamorphosed to amphibolite facies associated with such accre-  
427 tionary orogenesis (Holm et al., 2019). Mazatzal orogeny deformation and metamorphism  
428 occurred *ca.* 1650 to 1630 Ma within the juvenile accreted island arc of the Wisconsin  
429 Magmatic Terrane (Holm, Schneider, & Coath, 1998). The region of the ECMB was not  
430 affected by these tectonothermal events (Fig.1, Holm et al. (2005)). Holm et al. (2005)  
431 proposed that the voluminous ECMB batholith stabilized the continental lithosphere and  
432 prevented the region from being modified during subsequent collisions along the mar-  
433 gin. This lack of deformation in the region of the ECMB is further supported by the nearly  
434 horizontal bedding of *ca.* 1.63 Ga siliciclastic sedimentary rocks on either side of the batholith  
435 (Holm, Schneider, & Coath, 1998; Medaris et al., 2021). In southwestern Minnesota, plu-  
436 tons coeval with the ECMB are overlain by the subhorizontal Sioux Quartzite (Fig. 1)  
437 with the correlative Barron Quartzite of northwestern Wisconsin also being undeformed  
438 (Southwick et al., 1986). This lack of deformation contrasts with correlative Baraboo  
439 quartzite south of the Spirit Lake Tectonic Zone (~400 km from the ECMB) that un-  
440 derwent compressional deformation during subsequent orogenesis (Holm, Schneider, &  
441 Coath, 1998; Medaris et al., 2021). The Yavapai and Mazatzal terranes were intruded  
442 by *ca.* 1470 to 1430 Ma granites of the Eastern Granite Rhyolite Province and there is  
443 a sizeable pluton of this age within accreted Penokean rocks in northern Wisconsin (the  
444 *ca.* 1470 Ma Wolf River batholith; Dewane and Van Schmus (2007)). However, the ther-  
445 mal effects of the Wolf River batholith (~370 km east of the ECMB study area) were  
446 limited to a 10-15 km wide contact zone surrounding the intrusion (Holm et al., 2019).

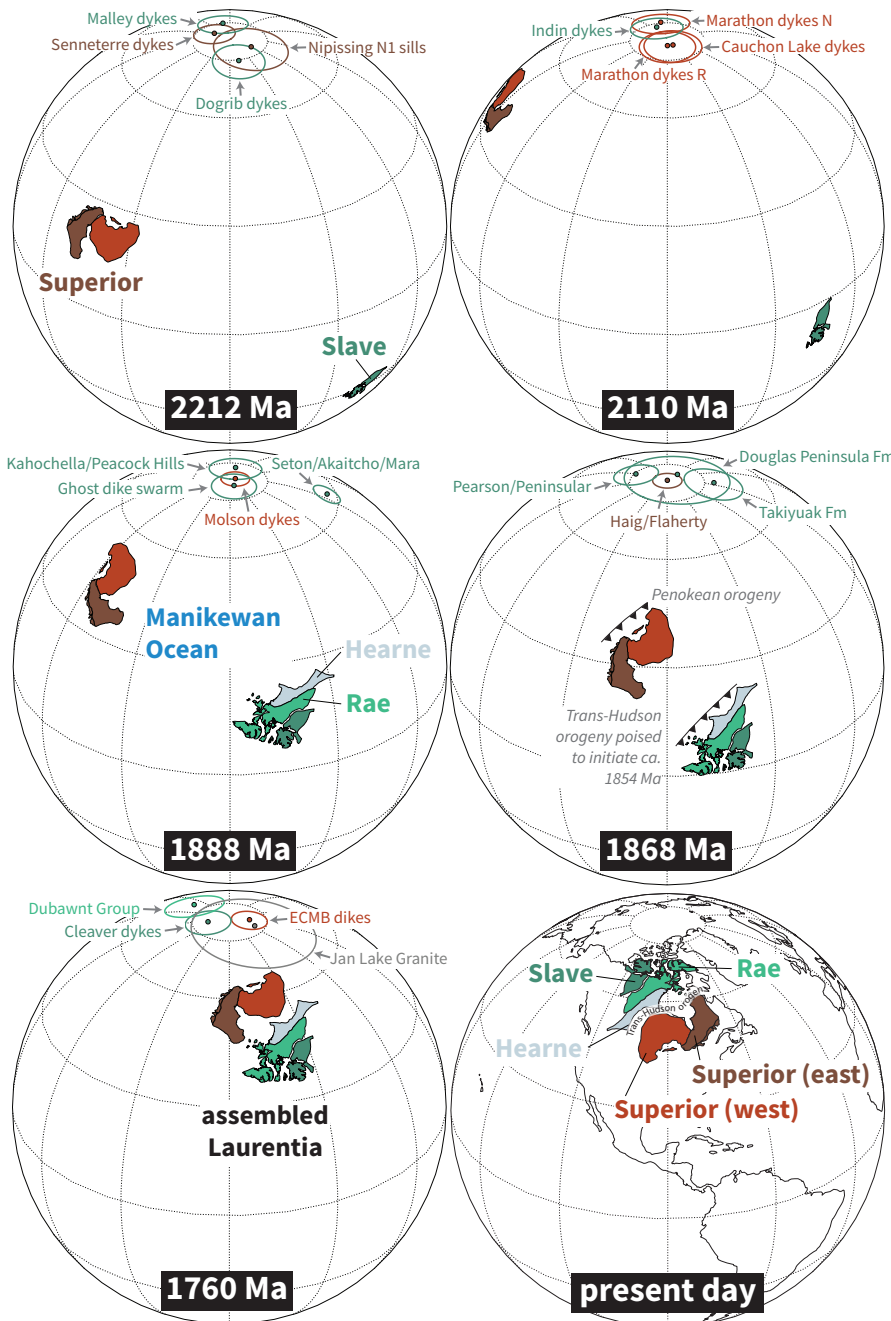
447 The one major subsequent tectonothermal event in the region for which there is  
448 localized evidence in the ECMB is the development of the Midcontinent Rift that ini-  
449 tiated *ca.* 1109 Ma and in which magmatic activity continued to *ca.* 1084 Ma (Fig. 1;  
450 Fairchild et al. (2017); Swanson-Hysell et al. (2019)). While the main rift axis can be  
451 inferred from gravity and aeromagnetic anomaly data to be located  $\sim 75$  km southeast  
452 of the study region (Fig. 1), the studied northwest-trending dike has a magnetization  
453 direction that implies that it was emplaced during Midcontinent Rift development *ca.*  
454 1096 Ma (Fig. 5). The baked contact test between that dike (NWD1) and the northeast-  
455 trending dike that it cross-cuts (NED17), indicates that the thermal effect of the dike  
456 emplacement and the Midcontinent Rift in general was localized within the immediate  
457 vicinity of that dike (a few meters; Fig. 5). However, this Midcontinent Rift magmatic  
458 activity did result in local hydrothermal alteration as evidenced by magnetization held  
459 by monoclinic pyrrhotite that is variably present through the ECMB dikes and is in the  
460 same direction as the magnetization of the northwest-trending dike (Fig. 3). This chem-  
461 ical remanent magnetization held by monoclinic pyrrhotite likely formed at relatively low  
462 temperatures. Phase relationships in the Fe-S system developed through hydrothermal  
463 recrystallization experiments show monoclinic pyrrhotite to form at temperatures be-  
464 low  $250^{\circ}\text{C}$  and likely above  $75^{\circ}\text{C}$  (Kissin & Scott, 1982). While in some sites, this pyrrhotite-  
465 forming alteration obscured the primary thermal remanence held by magnetite (e.g., NED36  
466 in Fig. 3), in the majority of sites the magnetite remanence direction can be well-resolved  
467 (e.g., NED2, NED12 and NED34 in Fig. 3). As a result, the paleomagnetic directions  
468 used to calculate the paleomagnetic pole shown in Figure 6 are held by (titano)magnetite  
469 that recorded a thermal remanent magnetization upon cooling of the diabase dikes. This  
470 evidence for variable late Mesoproterozoic hydrothermal alteration of the dikes provides  
471 an explanation for Ar-Ar data developed from two northeast-trending diabase dikes that  
472 were reported in Boerboom and Holm (2000). These Ar-Ar data did not yield a plateau  
473 age, but give whole rock total gas dates that are late Mesoproterozoic in age. An inter-  
474 pretation that these whole rock total gas ages correspond with the age of emplacement  
475 is difficult to reconcile with the cogenetic relationship between the diabase dikes, the quartz-  
476 feldspar porphyry dikes and the ECMB granites. K-Ar whole-rock ages of 1570 to 1280  
477 Ma from the dikes reported in Hanson (1968) and discussed in Horan et al. (1987) are  
478 attributed to partial resetting. The evidence for fluid flow that led to the formation of  
479 pyrrhotite *ca.* 1096 Ma supports the hypothesis put forward by Horan et al. (1987), as

480 well as by Boerboom and Holm (2000), that there was Mesoproterozoic disruption of the  
 481 K-Ar isotopic system in the dikes such that the Mesoproterozoic Ar-Ar dates are the re-  
 482 sult of alteration of dikes which are Paleoproterozoic in age. The field relationships in-  
 483 dicated that the northeast-trending diabase dikes are comagmatic with ECMB gran-  
 484 ites is also consistent with whole rock Pb isotope data that reveal very similar arrays im-  
 485 plying a *ca.* 1.8 Ga isochron age for both lithologies (Horan et al., 1987).

486 Overall, the constraints requiring that the ECMB granites were emplaced at depths  
 487 where the ambient temperature was below the closure of U-Pb apatite and Ar-Ar horn-  
 488 blende systems indicate that the comagmatic diabase dikes would have acquired their  
 489 magnetization at the time of emplacement. The lack of significant thermal events that  
 490 could have reset the magnetite magnetization is indicated by the geologic setting, the  
 491 thermochronology data (including the Paleoproterozoic Ar-Ar biotite dates), and the pos-  
 492 itive inverse baked contact test. We therefore interpret the pole calculated from the mag-  
 493 netite remanence of the ECMB diabase dikes as a high-quality constraint on the pale-  
 494 ogeographic position of Laurentia at the time the dikes intruded ( $1779.1 \pm 2.3$  Ma). The  
 495 ECMB diabase dikes pole meets six of the seven criteria for the quality criteria Van der  
 496 Voo (1990) and the Meert et al. (2020) reliability criteria with the only one not satis-  
 497 fied being due to the lack of dual polarity directions. This single polarity normal polar-  
 498 ity is consistent with the polarity of the Cleaver Dykes and the proposal of Irving (2004)  
 499 that there was a normal geomagnetic superchron that followed the Trans-Hudson orogeny.

## 500 **6.2 Laurentia’s paleomagnetic poles following the Trans-Hudson orogeny**

501 The Trans-Hudson orogeny is a major event in the formation of Laurentia result-  
 502 ing from collision between the Superior province and the Churchill plate consisting of  
 503 the composite Slave + Rae + Hearne provinces (Hoffman, 1988; Corrigan et al., 2009;  
 504 Weller & St-Onge, 2017). Geologic data on the timing of Trans-Hudson orogenesis in-  
 505 clude a  $^{206}\text{Pb}/^{238}\text{U}$  date of  $1854.2 \pm 1.6$  Ma from the base of a foredeep sedimentary suc-  
 506 cession on the northern margin of the East Superior province that constrains flexural sub-  
 507 sidence associated with Trans-Hudson orogenesis to have initiated at that time (Hodgkiss  
 508 et al., 2019). This timing of orogenesis is consistent with  $^{207}\text{Pb}/^{206}\text{Pb}$  dates of monazite  
 509 within garnet of Trans-Hudson orogen eclogites for which a mean date of  $1831 \pm 5$  Ma  
 510 has been interpreted to record peak metamorphism (Weller & St-Onge, 2017). A sim-  
 511 ilar timing of *ca.* 1860 to 1820 Ma peak Trans-Hudson metamorphism resulting from col-



**Figure 9.** Paleogeographic reconstructions at five different times in the Paleoproterozoic and the position of the provinces at present. Paleomagnetic poles within 20 Myr of the given time (10 Myr for 1888 and 1868 Ma) are shown from the compilation of Evans et al. (2021) as well as the new ECMB pole. These data illustrate differential plate motion between the Superior and Slave Provinces that is required by the data leading up to the closure of the Manikewan Ocean and the assembly of Laurentia during the Trans-Hudson orogeny. The ECMB pole is consistent with an assembled Laurentia following the Trans-Hudson orogeny which contrasts with the disparate orientations and paleolatitudes between Laurentia’s constituent provinces prior to the orogeny.

512 lisional orogenesis has been interpreted from U-Pb zircon rim and monazite dates from  
513 the orogen in Baffin Island, Nunavut, Canada (Skipton et al., 2016). These geologic data  
514 strongly support that the Superior Province was conjoined with the Slave + Rae + Hearne  
515 provinces prior to 1800 Ma in their present-day relative positions (Fig. 1). There are high-  
516 quality paleomagnetic poles for the Superior province during the time interval when the  
517 Manikewan Ocean was closing leading up to the Trans-Hudson orogeny (the *ca.* 1880 Mol-  
518 lison dykes pole and the *ca.* 1870 Ma Haig/Flaherty pole; Fig. 9). There are no Paleo-  
519 proterozoic paleomagnetic poles for the Superior Province during or after the Trans-Hudson  
520 orogeny that can test Laurentia’s coherency with paleomagnetic data. The new  $1779.1$   
521  $\pm 2.3$  Ma ECMB paleomagnetic pole fills this gap.

522 While abundant paleomagnetic data have been developed from rocks within the  
523 Trans-Hudson orogen (Symons & Harris, 2005), both the primary nature of the rema-  
524 nence directions as well as the age of remanence acquisition has been challenging to es-  
525 tablish. As a result, relatively few poles from this interval during and following Laurentia’s  
526 amalgamation have been included in curated pole compilations such as that com-  
527 piled by the Nordic paleogeography workshops (Evans et al., 2021). The best constrained  
528 of these poles comes from the post-orogenic  $1740 \pm 5$  Ma Cleaver Dykes of the Great  
529 Bear Magmatic Arc on the western margin of the Slave Province (Fig. 1; Irving (2004)).  
530 The age of these dikes is constrained by a U-Pb baddeleyite date and there is a positive  
531 baked contact test supporting the interpretation that the pole is primary. As can be seen  
532 in Figure 6, the position of this *ca.* 1740 Ma Cleaver Dykes pole for Laurentia (Irving,  
533 2004) is similar to the new pole for the *ca.* 1780 Ma ECMB diabase dikes (Fig. 6). They  
534 do not share a common mean (as determined through Watson and bootstrap common  
535 mean tests; Tauxe et al. (2016)), but are within  $11^\circ$  of one another (less when consid-  
536 ering uncertainty). The similarity in these pole positions provides an independent test  
537 of the coherency of the Laurentia craton at that time.

538 An additional pole for comparison comes from hematite remanence of volcanics and  
539 sediments of the Baker Lake Group of the Dubawnt Supergroup which were deposited  
540 in a basin atop the suture between the Rae and Hearne provinces (Figs. 1 and 6; Park  
541 et al. (1973)). Both the depositional age of the succession and the age of hematite re-  
542 manence are roughly constrained leading to a wide assigned age range of 1820 to 1750  
543 Ma in Evans et al. (2021) for the pole. While the loose age constraints hinder firm com-  
544 parisons, the broad similarity of this pole with the Cleaver Dykes pole of the Slave Province

545 as well as that for Martin Formation ( $1818 \pm 4$  Ma) and the Sparrow dikes ( $1827 \pm 4$   
546 Ma) of the Rae Province establishes a largely consistent position of the Churchill plate  
547 from *ca.* 1820 to 1740 Ma.

548 A less robust pole for comparison with the new ECMB pole comes from the post-  
549 orogenic Jan Lake Granite within the Trans-Hudson orogen in southeast Saskatchewan  
550 developed in Gala et al. (1995). A U-Pb zircon date of  $1758 \pm 1$  Ma for this granite was  
551 reported in Bickford et al. (2005). Directional data from the Jan Lake Granite falls into  
552 two groups. Based on thermal demagnetization behavior, Irving (2004) interpreted the  
553 ‘A’ grouping to be a TRM held by magnetite that was acquired at the time of the em-  
554 placement of the intrusion *ca.* 1758 Ma. The ECMB dikes pole shares a common mean  
555 with this Jan Lake Granite pole with overlapping  $A_{95}$  confidence circles (Fig. 6). This  
556 result is consistent with both the Jan Lake Granite and ECMB being post-orogenic mag-  
557 matic events that occurred following the amalgamation of Laurentia. This similarity sug-  
558 gests that despite the large uncertainty on the Jan Lake Granite pole and the ambigu-  
559 ity resulting from multiple directional groups that the pole does constrain the position  
560 of Laurentia *ca.* 1758 Ma.

561 A pole that does not hold up to such comparative scrutiny is that developed for  
562 the Deschambault Pegmatites from within the Trans-Hudson orogen (Fig. 6C; Symons  
563 et al. (2000)). This pole has been interpreted to constrain the position of Laurentia *ca.*  
564 1766 Ma — an age based on U-Pb monazite and allanite dates of other pegmatites in  
565 the region. As noted in D’Agrella-Filho et al. (2020), there are no field tests for this pole  
566 and the remanence directions from which it is calculated are quite close to the modern  
567 geomagnetic field. This pole was included in the curated Nordic paleogeography work-  
568 shop compilation of Evans et al. (2021) less because of the quality of the individual pole,  
569 but rather because there are a number of poles of similar position to this one from the Trans-  
570 Hudson orogen (Fig. 6). Many of these poles have individual VGPs that are streaked  
571 between directions similar to the Jan Lake Granite and that of the present-local field.  
572 The direction of the Deschambault pole is far from the new *ca.* 1780 Ma ECMB dikes  
573 pole. In contrast, the similarity of pole position between the new ECMB dikes pole and  
574 the *ca.* 1758 Ma Jan Lake Granite pole as well as the *ca.* 1740 Ma Cleaver Dykes pole  
575 supports that these poles, rather than the Deschambault Pegmatites pole, constrain Lau-  
576 rentia’s position during this interval. This Deschambault Pegmatites pole played a role,  
577 in conjunction with other poles from the Trans-Hudson orogen, such as that from the

578 Wapisi gneiss and the Deschambault Post pluton, in an interpretation that Laurentia's  
579 pole path was at a standstill in the vicinity of the Deschambault Pegmatites pole from  
580 *ca.* 1800 Ma through to *ca.* 1766 Ma (Symons et al., 2000; Symons & Harris, 2005). As  
581 reviewed in Raub (2008), there are numerous difficulties in interpreting these data from  
582 the Trans-Hudson orogen as useful constraints including: 1) a lack of field tests; 2) un-  
583 certainty in the timescale of cooling and the timing of the acquisition of magnetization  
584 in these slowly cooled units; 3) poorly constrained tilt corrections and 4) large secondary  
585 viscous remanent magnetizations that are prevalent due to the coarse grain-size of the  
586 igneous lithologies. The preferred interpretation of Raub (2008), which is echoed in D'Agrella-  
587 Filho et al. (2020), is that there is unresolved component mixing between primary di-  
588 rections (which would be in the vicinity of the Jan Lake Granite A Group pole and our  
589 new ECMB pole) and the present-day north pole as the result of unresolved viscous over-  
590 prints (Fig. 6). This component mixing leads to streaked site mean directions in indi-  
591 vidual studies as well as the database of Trans-Hudson orogen poles including the De-  
592 schambault Pegmatites pole (Fig. 6; Raub (2008)). The new ECMB pole significantly  
593 strengthens this interpretation by demonstrating that a proposed northerly apparent po-  
594 lar wander path to satisfy the Deschambault Pegmatites pole and other Trans-Hudson  
595 orogen poles streaked between the Jan Lake Granite and the modern-day pole is indeed  
596 fictitious (Fig. 6). Instead, the ECMB pole and the Cleaver Dykes pole establish the pa-  
597 leogeographic position of Laurentia to have been consistent *ca.* 1780 to 1740 Ma (Fig.  
598 9).

### 599 **6.3 The paleogeography of Laurentia**

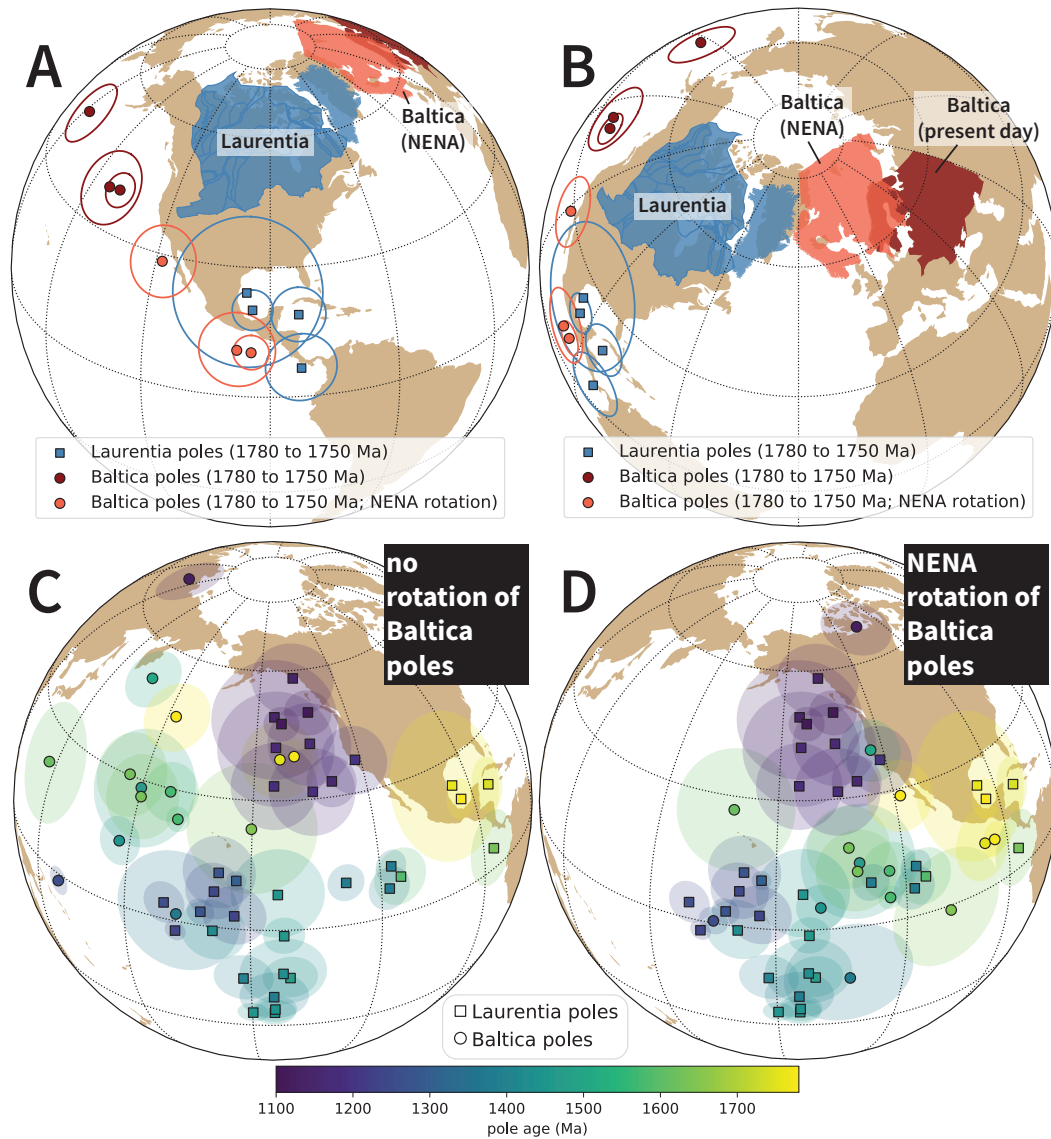
600 As is expected by the geologic record of Trans-Hudson orogenesis, the similarity  
601 in pole positions from the southeastern margin of the Superior Provinces (the new ECMB  
602 pole) and the northwestern margin of the Slave Province (the Cleaver Dykes pole; Fig.  
603 6) indicate a coherent assembled Laurentia following 1.8 Ga (Fig. 9). The coherency of  
604 the record of high-quality paleomagnetic poles at this time when the geologic record in-  
605 dicates a recently assembled Laurentia increases confidence that differing pole positions  
606 between Laurentia's Archean provinces earlier in the Paleoproterozoic are indeed a record  
607 of differential plate tectonic motion (Fig. 9). There is a particularly rich record of pa-  
608 leomagnetic poles from the Archean Superior and Slave provinces that can be paired be-  
609 tween 2.23 and 1.89 Ga that constrain the provinces to not be in their modern relative

610 orientation and to be undergoing differential motion (Mitchell et al., 2014; Buchan et  
611 al., 2016; Swanson-Hysell, 2021). These poles result in reconstructions where prior to the  
612 Trans-Hudson orogeny there was an ocean basin between the Superior province and the  
613 Hearne + Rae + Slave provinces known as the Manikewan Ocean (Fig. 9; Stauffer (1984)).  
614 The poles are consistent with the Superior Province approaching the joint Slave + Hearne  
615 + Rae provinces prior to the onset of the Trans-Hudson orogeny (Fig. 9). These data  
616 provide strong evidence for mobile lid plate tectonics from 2.23 Ga onward (Mitchell et  
617 al., 2014; Buchan et al., 2016; Swanson-Hysell, 2021).

618 The orogenesis associated with Laurentia’s assembly is hypothesized to have re-  
619 sulted in the formation of the supercontinent, or semi-supercontinent, Nuna (Hoffman,  
620 1997; Evans & Mitchell, 2011; Evans et al., 2016). Given that Laurentia is the largest  
621 craton hypothesized to have been part of this supercontinent, its paleogeographic po-  
622 sition is key to reconstructions of Nuna. The new ECMB pole provides higher confidence  
623 in the paleogeographic position of Laurentia in the time just following its formation from  
624 the collision of constituent Archean provinces (Fig. 9). This new pole can be used to eval-  
625 uate hypothesized connections between Laurentia and other cratons. There is an increas-  
626 ingly rich global database of paleomagnetic poles *ca.* 1780 Ma including poles from the  
627 Amazonia, Baltica, India, Rio de la Plata, São Francisco and North China cratons (Zhang  
628 et al., 2012; Xu et al., 2014; Bispo-Santos et al., 2014; Shankar et al., 2018; D’Agrella-  
629 Filho et al., 2020).

630 One hypothesized connection of particular interest is that with Baltica. The two  
631 cratons have been hypothesized to have been conjoined such that they shared a margin  
632 with a long-lived history of accretionary orogenesis (Hoffman, 1997; Karlstrom et al., 2001).  
633 The proposed NENA (northern Europe and North America) configuration between Lau-  
634 rentia and Baltica allows for such a shared margin (Gower et al., 1990; Buchan et al.,  
635 2000; Evans & Pisarevsky, 2008). The increased concordance between *ca.* 1780 to 1750  
636 paleomagnetic poles from Laurentia and Baltica upon the NENA rotation of Baltica can  
637 be seen in Figure 10. Paleomagnetic poles support a continued NENA connection un-  
638 til at least 1260 Ma (supported by poles from Baltica’s *ca.* 1258 Ma post-Jotnian intru-  
639 sions and Laurentia’s *ca.* 1267 Ma Mackenzie dikes) and perhaps to 1120 Ma (where the  
640 paleomagnetic comparison is reliant on the *ca.* 1122 Ma Salla dike of Baltica developed  
641 from a single cooling unit; Salminen et al. (2009)). This connection supports the long-  
642 lived active margin where Laurentia grew through the rest of the Paleoproterozoic and





**Figure 10.** A) Paleomagnetic poles for Laurentia and Baltica between 1780 and 1750 Ma with the Baltica poles shown with and without the NENA rotation (Euler pole for Baltica of  $[47.5^\circ, 001.5^\circ, +49.0^\circ]$  as in Evans and Pisarevsky (2008)). Baltica is shown in its present day location in dark red and shown in the NENA position in light red. B) Same data as in panel A shown with a different center of projection which allows for easier visualization of the reconstructed position. C,D) Comparison between poles between 1780 and 1110 Ma between Baltica and Laurentia without (C) and with (D) the NENA rotation. The poles that are shown are those from the Nordic compilation with 'A' and 'B' grades as well as the new ECMB pole from this study.

643 through the Mesoproterozoic until the *ca.* 1.08 Ga continent-continent collision of the  
 644 Grenvillian orogeny (Whitmeyer & Karlstrom, 2007).

## 7 Conclusions

The East-Central Minnesota Batholith was emplaced following Penokean orogenesis on the southeast margin of the Superior Province. While the southeast margin of Laurentia experienced subsequent intervals of accretionary orogenesis, thermochronology data constrain the batholith to have a straight-forward history of post-emplacment rapid exhumation without substantial reheating. Subsequent orogenesis occurred well southeast of the batholith — consistent with the batholith having played a role in stabilizing Laurentian lithosphere. Comagmatic diabase dikes of the East-Central Minnesota Batholith can be constrained through U-Pb geochronology on the felsic units to have been emplaced at  $1779.1 \pm 2.3$  Ma. A new paleomagnetic pole developed from the magnetite remanence of these dikes provides a high-quality constraint on the position of Laurentia following Trans-Hudson orogenesis. This pole confirms the coherency of an amalgamated Laurentia at the time and supports the NENA connection with Baltica. This paleomagnetic coherency further strengthens the case that previously disparate pole positions between the Superior and Slave provinces are the result of *ca.* 2.2 to 1.8 Ga mobile-lid plate tectonics. The geologic and paleomagnetic record of Laurentia is inconsistent with a stagnant-lid regime anytime over the past 2.2 billion years.

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677 is also archived on Zenodo (<https://doi.org/10.5281/zenodo.4625041>). This repository  
 678 also contains Python code related to calculations, visualizations and statistical tests dis-  
 679 cussed herein.

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**Table 1.** Summary of site level paleomagnetic data

site	site lat	site lon	n	dec	inc	k	R	$\alpha_{95}$	VGP lat	VGP lon
<i>northeast-trending dike magnetite-component site means</i>										
NED1	45.53423	265.75804	8	157.6	74.7	380	7.98	2.8	18.4	276.9
NED2	45.53421	265.75816	8	172.3	74.9	420	7.98	2.7	17.4	269.6
NED5	45.53309	265.75803	6	170.3	79.0	213	5.98	4.6	24.5	269.6
NED6	45.53299	265.75773	9	197.8	73.6	183	8.96	3.8	16.1	256.5
NED7	45.53288	265.75767	5	266.3	81.3	204	4.98	5.4	42.0	242.6
NED8	45.53286	265.75782	7	202.6	78.6	423	6.99	2.9	24.8	256.6
NED9	45.53314	265.75855	7	191.1	71.5	90	6.93	6.4	12.2	259.5
NED10	45.53259	265.75742	6	199.7	73.3	324	5.98	3.7	15.8	255.4
NED11	45.53252	265.75768	7	166.1	75.1	391	6.98	3.1	18.1	272.6
NED12	45.53489	265.76076	6	179.5	73.0	374	5.99	3.5	14.1	266.0
NED13	45.53497	265.76113	7	169.1	73.9	199	6.97	4.3	15.9	271.4
NED14	45.53492	265.76119	6	193.0	70.0	217	5.98	4.6	10.1	258.0
NED15	45.53688	265.76758	8	175.8	77.7	604	7.99	2.3	22.0	267.6
NED16	45.53728	265.76822	5	185.1	78.6	449	4.99	3.6	23.6	263.7
NED18	45.53124	265.76945	6	134.7	81.7	159	5.97	5.3	33.2	279.5
NED23	45.53398	265.74200	7	193.0	76.4	411	6.99	3.0	20.2	259.7
NED25	45.53396	265.74119	4	135.4	80.7	334	3.99	5.0	31.5	280.6
NED26	45.53445	265.74129	5	151.1	73.4	572	4.99	3.2	17.4	280.8
NED28	45.53467	265.73817	4	157.8	73.8	145	3.98	7.7	16.9	277.2
NED29	45.53438	265.73690	8	219.0	76.7	119	7.94	5.1	24.4	248.6
NED31	45.53385	265.75785	3	184.7	69.4	1730	3.00	3.0	8.7	262.9
NED34	45.51700	265.78083	8	185.9	77.0	362	7.98	2.9	20.8	263.1
NED35	45.53320	265.75761	8	166.4	74.7	261	7.97	3.4	17.4	272.6
<b>mean pole: pole longitude: 265.8; pole latitude: 20.4; <math>A_{95}</math>: 4.5; K: 45.6 N: 23</b>										
<i>northwest-trending dike magnetite-component site mean</i>										
NWD1	45.53407	265.76852	9	293.4	41.6	66	8.88	6.4	32.9	177.5

Notes: site lat–latitude of site ( $^{\circ}$ ; WGS84); site lon–longitude of site ( $^{\circ}$ ; WGS84) n–number of samples analyzed and included in the site mean; dec–tilt-correction mean declination for the site; inc–tilt-correction mean inclination for the site; k–Fisher precision parameter; R–resultant vector length;  $\alpha_{95}$ –95% confidence limit in degrees; VGP lat–latitude of the virtual geomagnetic pole for the site; VGP lon–longitude of the virtual geomagnetic pole for the site.

**Table 2.** Summary of ID-TIMS  $^{207}\text{Pb}/^{206}\text{Pb}$  East-Central Minnesota Batholith zircon dates

Sample	Unit	Latitude	$^{207}\text{Pb}/^{206}\text{Pb}$ date (Ma)	Uncertainty ( $2\sigma$ )		MSWD	n/N
		Longitude		X	Z		
ECMB6	St. Cloud Granite	45.53396° N	1781.44	0.51	2.4	1.24	8/8
		94.23187° W					
QP1	quartz-feldspar	45.53481° N	1780.78	0.45	2.4	0.53	8/8
	porphyry dike	94.25811° W					
ECMB4	Richmond Granite	45.44343° N	1776.76	0.49	2.4	1.15	7/8
		94.48360° W					

Notes: X is  $2\sigma$  analytical uncertainty; Z is  $2\sigma$  uncertainty including decay constant uncertainty. This Z uncertainty needs to be utilized when comparing to dates using other decay systems (e.g.,  $^{40}\text{Ar}/^{39}\text{Ar}$ ,  $^{187}\text{Re}$ - $^{187}\text{Os}$ ); MSWD is the mean squared weighted deviation; n is the number of individual zircon dates included in the calculated sample mean date; N is the number of individual zircons analyzed for the sample.

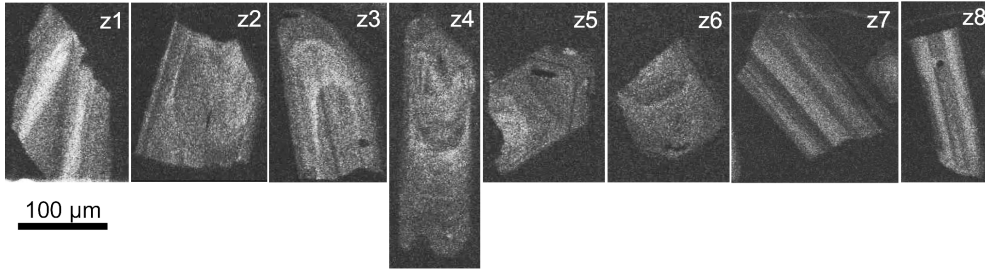
## Supporting Information for “The paleogeography of Laurentia in its early years: new constraints from the Paleoproterozoic East Central Minnesota batholith”

Swanson-Hysell, N. L., Avery, M. S., Zhang, Y., Hodgin, E. B., Sherwood, R. J., Apen, F. E., Boerboom, T. J., Keller, C. B., and Cottle, J. M. (2021), The paleogeography of Laurentia in its early years: new constraints from the Paleoproterozoic East Central Minnesota batholith *Tectonics*.

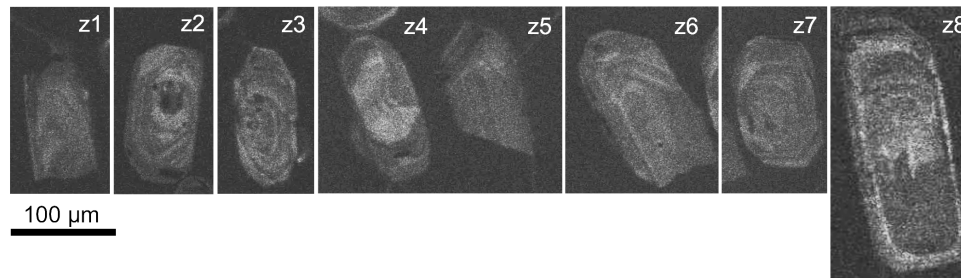
### ID-TIMS U-Pb zircon geochronology methods

U-Pb dates were obtained by chemical abrasion isotope dilution thermal ionization mass spectrometry (ID-TIMS) in the Boise State University (BSU) Isotope Geology Laboratory (Table DR1; Fig. SI2). Chemical abrasion of single zircon grains was modified after Mattinson (2005). Zircons were separated from rocks using standard techniques, annealed in a muffle furnace at 900°C for 60 hours in quartz crucibles, and imaged by cathodoluminescence in grain mounts (Fig. SI1).

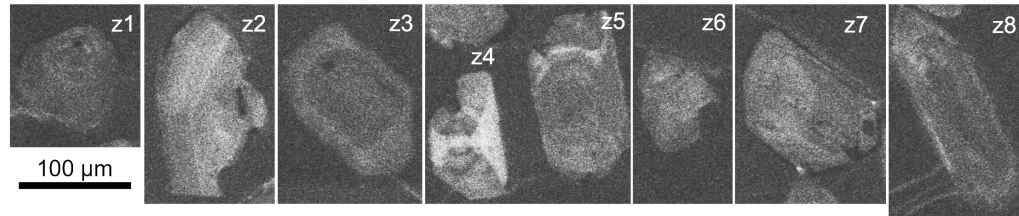
#### QP1 – Quartz Porphyry Dike



#### ECMB4 – Richmond Granite



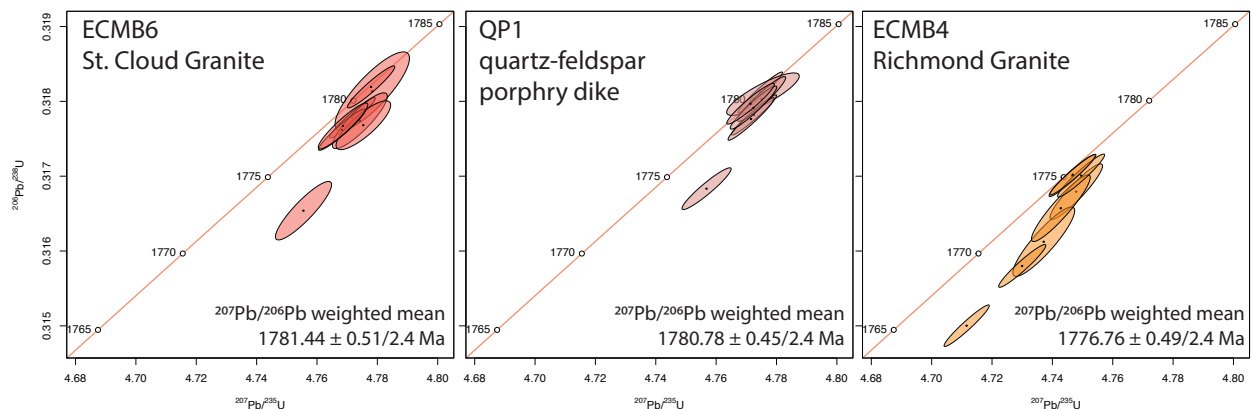
#### ECMB6 – St. Cloud Granite



**Figure SI1.** Cathodoluminescence (CL) images of the zircons dated by ID-TIMS. The 100 μm scale bars applies for all imaged grains in a given sample.

Individual zircons were removed from grain mounts and chemically abraded. Chemical abrasion was carried out by transferring zircons to 3 ml Teflon Perfluoroalkoxy alkane (PFA) beakers in which they were rinsed in 3.5 M HNO<sub>3</sub> and ultrapure H<sub>2</sub>O prior to loading into 300  $\mu$ l Teflon PFA microcapsules. Fifteen microcapsules were placed in a large-capacity Parr vessel and the zircon partially dissolved in 120  $\mu$ l of 29 M HF for 12 hours at 190°C. Zircons were returned to 3 ml Teflon PFA beakers, HF was removed, and zircons were immersed in 3.5 M HNO<sub>3</sub>, ultrasonically cleaned for an hour, and fluxed on a hotplate at 80°C for an hour. The HNO<sub>3</sub> was removed and zircon was rinsed twice in ultrapure H<sub>2</sub>O before being reloaded into the 300  $\mu$ l Teflon PFA microcapsules (rinsed and fluxed in 6 M HCl during sonication and washing of the zircons) and spiked with the <sup>233</sup>U-<sup>235</sup>U-<sup>205</sup>Pb BSU tracer solution (BSU1B). Zircons were dissolved in Parr vessels in 120  $\mu$ l of 29 M HF at 220°C for 48 hours, dried to fluorides, and re-dissolved in 6 M HCl at 180°C overnight. Pb and U were separated from the zircon matrix using an HCl-based anion-exchange chromatographic procedure (Krogh, 1973), eluted together and dried with 2  $\mu$ l of 0.05 N H<sub>3</sub>PO<sub>4</sub>.

Pb and U were loaded on a single outgassed Re filament in 5  $\mu$ l of a silica-gel/phosphoric acid mixture (Gerstenberger and Haase, 1997), and Pb and U isotopic measurements made on a GV Isoprobe-T multicollector thermal ionization mass spectrometer equipped with an ion-counting Daly detector. Pb isotopes were measured by peak-jumping all isotopes on the Daly detector for 190 cycles with a mass bias correction of  $0.16 \pm 0.03\%$ /a.m.u. ( $1\sigma$ ). Transitory isobaric interferences due to high-molecular weight organics, particularly on <sup>204</sup>Pb and <sup>207</sup>Pb, disappeared within 30-45 cycles, while ionization efficiency averaged 104 cps/pg of each Pb isotope. Linearity (to  $\geq 1.4 \times 10^6$  cps) and the associated deadtime correction of the Daly detector were determined by analysis of NBS982. Uranium was analyzed as UO<sub>2</sub><sup>+</sup> ions in static Faraday mode on 10<sup>12</sup> ohm resistors for up to 300 cycles, and corrected for isobaric interference of <sup>233</sup>U<sup>18</sup>O<sup>16</sup>O on <sup>235</sup>U<sup>16</sup>O<sup>16</sup>O with an <sup>18</sup>O/<sup>16</sup>O of 0.00206. Ionization efficiency averaged 20 mV/ng of each U isotope. U mass fractionation was corrected using the <sup>233</sup>U/<sup>235</sup>U ratio of the BSU1B tracer.



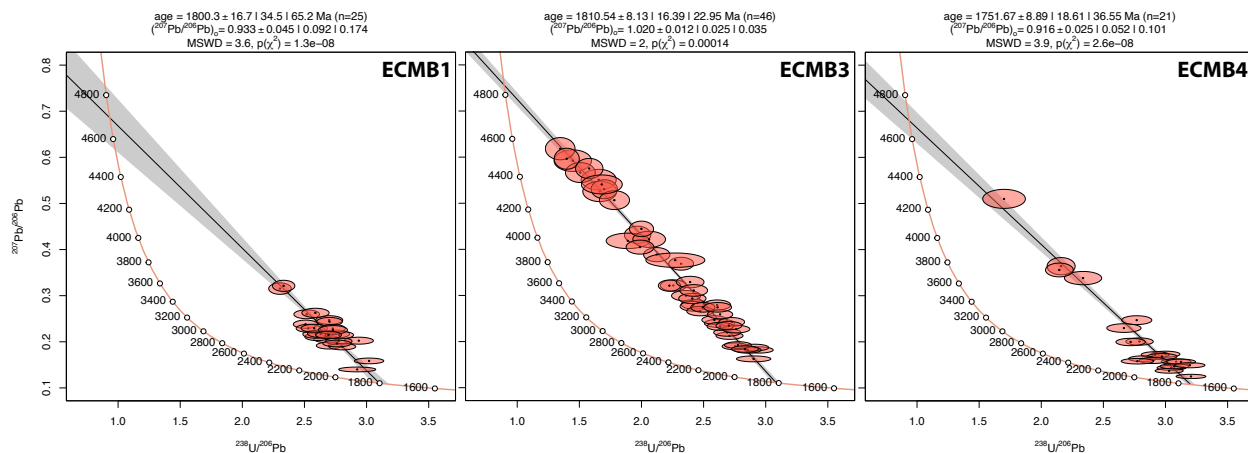
**Figure SI2.** U-Pb concordia plots for the new zircon dates. Ellipses represent  $2\sigma$  analytical uncertainty on individual zircon dates. The weighted mean dates are shown with X/Y uncertainty where X is  $2\sigma$  analytical uncertainty and Z is  $2\sigma$  uncertainty including decay constant uncertainty.

ID-TIMS U-Pb dates and uncertainties were calculated using the algorithms of Schmitz and Schoene (2007), BSU1B tracer solution with calibration of  $^{235}\text{U}/^{205}\text{Pb} = 77.93$  and  $^{233}\text{U}/^{235}\text{U} = 1.007066$ , and U decay constants recommended by Hiess et al. (2012), including  $^{238}\text{U}/^{235}\text{U}$  of

137.818.  $^{206}\text{Pb}/^{238}\text{U}$  ratios and dates were corrected for initial  $^{230}\text{Th}$  disequilibrium using  $\text{DTh}/\text{U} = 0.20 \pm 0.05$  ( $1\sigma$ ). All common Pb in analyses was attributed to laboratory blank and subtracted based on the measured laboratory Pb isotopic composition and associated uncertainty. U blanks are estimated at 0.013 pg. ID-TIMS weighted mean  $^{207}\text{Pb}/^{206}\text{Pb}$  dates were calculated from equivalent dates ( $\text{pof} > 0.05$ ) using Isoplot 3.0 (Ludwig, 2003). Errors on the weighted mean dates are given as  $\pm x / z$ , where  $x$  is the internal error based on analytical uncertainties including counting statistics, subtraction of tracer solution, and blank and initial common Pb subtraction;  $z$  also includes the U decay constant uncertainties propagated in quadrature. Dates from individual zircon fractions and weighted mean dates are reported at  $2\sigma$ .

### LA-ICP-MS U-Pb apatite geochronology methods

Apatite U-Pb geochronology data were developed via laser ablation inductively coupled mass spectrometry (LA-ICP-MS) at UC Santa Barbara. U-Pb isotopes were analyzed with a Cetac Teledyne 193 nm excimer Analyte laser with a HelEx ablation cell coupled to a Nu Instruments Plasma 3D multi-collector (MC) ICP-MS. On the Plasma 3D,  $^{204}(\text{Pb}+\text{Hg})$ ,  $^{206}\text{Pb}$ ,  $^{207}\text{Pb}$ , and  $^{208}\text{Pb}$  were measured on Daly detectors and  $^{238}\text{U}$  and  $^{232}\text{Th}$  were measured on Faraday collectors. Apatite was ablated with a 40  $\mu\text{m}$  diameter laser spot for 60 pulses fired at a 4 Hz repetition rate and 50% of 5 mJ laser power. Each ablation sequence consisted of 2 cleaning shots, followed by 25 secs of monitored washout and 15 secs of ablation. The Iolite v. 2.5 program (Paton et al., 2011) in the Igor Pro software environment was used to correct the raw U-Pb ratios for baselines, laser- and plasma-induced element fractionation, and instrument drift.



**Figure SI3.** U-Pb Tera-Wasserburg plots for the new apatite dates. Ellipses represent  $2\sigma$  analytical uncertainty on individual apatite dates. The

Multiple apatite reference materials (RMs) were analyzed throughout the analytical session to monitor data quality. Apatite RM Madagascar ( $478.4 \pm 6.1$  Ma, ID-TIMS age; Schmitz, 2020 personal communication) served as the primary bracketing standard with RMs McClure ( $523.5 \pm 1.5$  Ma, ID-TIMS age; Schoene and Bowring, 2006), 401 ( $530.3 \pm 1.5$  Ma, ID-ICP-MS age; Thompson et al., 2016) and OD306 ( $1596.7 \pm 7.1$  Ma, ID-ICP-MS age; Thompson et al., 2016) analyzed as secondary standards to assess precision and accuracy. All of the secondary standards form mixing arrays between common-Pb and an age intercept. The excess variances of



$^{206}\text{Pb}/^{238}\text{U}$  and  $^{207}\text{Pb}/^{206}\text{Pb}$  required for secondary standards to conform to a statistically single mixing line are 2.1% and 2.0% ( $2\sigma$ ), respectively; these values were added in quadrature to the internal uncertainties of each U-Pb datum (Horstwood et al., 2016). The ages of the secondary standards with these designated uncertainties are  $547.4 \pm 30.4/35.9$  Ma (McClure),  $512.8 \pm 9.8/10.5$  Ma (401),  $1586.5 \pm 14.9/15.7$  Ma (OD306) (uncertainties following the same format used in main text), within uncertainty of their published ages. Although mass-204 was measured on the MC-ICP-MS, isobaric interferences with  $^{204}\text{Hg}$  in the He carrier gas preclude the use of the  $^{204}\text{Pb}$  method for common-Pb corrections. The U-Pb apatite data are reported in Table S2.

Table S1: Zircon U-Th-Pb isotopic data

Sample	Compositional Parameters					Radiogenic Isotope Ratios							Isotopic Ages (Ma)				Weighted Mean (Ma)			
	Th/U	<sup>206</sup> Pb*/ <sup>206</sup> Pb	mol % <sup>206</sup> Pb*	Pb <sub>c</sub> */Pb <sub>c</sub>	Pb <sub>c</sub> (pg)	<sup>206</sup> Pb/ <sup>204</sup> Pb	<sup>208</sup> Pb/ <sup>206</sup> Pb	<sup>207</sup> Pb/ <sup>206</sup> Pb	% err	<sup>207</sup> Pb/ <sup>235</sup> U	% err	<sup>206</sup> Pb/ <sup>238</sup> U	% err	corr. coef.	<sup>207</sup> Pb/ <sup>206</sup> Pb ±	<sup>207</sup> Pb/ <sup>235</sup> U ±	<sup>206</sup> Pb/ <sup>238</sup> U ±	<sup>206</sup> Pb/ <sup>206</sup> Pb ±	<sup>207</sup> Pb/ <sup>206</sup> Pb	
(a)	(b)	(c)	(c)	(c)	(d)	(e)	(e)	(f)	(e)	(f)	(e)	(f)	(f)	(g)	(f)	(g)	(f)	(g)	(f)	(h)
<b>ECMB4</b>																				
z1	0.459	1.7560	99.84%	189	0.24	11080	0.134	0.108748	0.071	4.747907	0.150	0.316793	0.096	0.927	1777.72	1.29	1775.74	1.26	1774.05	1.50
z2	0.538	2.9136	99.88%	258	0.29	14816	0.157	0.108642	0.065	4.746500	0.131	0.317007	0.072	0.961	1775.96	1.19	1775.49	1.10	1775.09	1.12
z3	0.482	0.8672	99.78%	138	0.16	8032	0.141	0.108731	0.088	4.737140	0.177	0.316124	0.120	0.896	1777.44	1.60	1773.83	1.49	1770.77	1.85
z4	0.468	1.1223	99.79%	144	0.20	8395	0.137	0.108707	0.076	4.742804	0.168	0.316572	0.115	0.921	1777.04	1.39	1774.84	1.41	1772.97	1.78
z5	0.572	2.6740	99.84%	201	0.35	11478	0.167	0.108678	0.068	4.730009	0.135	0.315801	0.074	0.952	1776.56	1.25	1772.57	1.13	1769.19	1.14
z6	0.467	3.2913	99.82%	173	0.49	10119	0.136	0.108710	0.067	4.749514	0.134	0.317009	0.074	0.956	1777.10	1.22	1776.02	1.12	1775.11	1.15
z7	0.446	3.8518	99.92%	381	0.26	22347	0.130	0.108530	0.064	4.711602	0.130	0.315002	0.072	0.959	1774.07	1.17	1769.30	1.09	1765.27	1.11
z8	0.506	5.8883	99.75%	119	1.28	6606	0.148	0.108644	0.067	4.746793	0.132	0.317021	0.070	0.959	1775.98	1.23	1775.54	1.10	1775.17	1.09
																			<b>1776.76 ± 0.49 [2.40]</b>	
																			MSWD = 1.15	POF
																			= 0.33	
<b>QPI</b>																				
z1	0.359	9.6958	99.96%	768	0.32	45979	0.105	0.108903	0.061	4.774622	0.126	0.318120	0.070	0.967	1780.33	1.12	1780.45	1.06	1780.54	1.08
z2	0.401	4.9637	98.54%	21	6.10	1236	0.117	0.108932	0.149	4.775515	0.206	0.318096	0.073	0.847	1780.81	2.72	1780.60	1.73	1780.43	1.14
z3	0.423	6.4130	99.89%	269	0.61	15865	0.123	0.108960	0.064	4.772557	0.128	0.317818	0.069	0.967	1781.28	1.16	1780.08	1.07	1779.06	1.07
z4	0.336	2.7906	99.91%	325	0.21	19550	0.098	0.108956	0.064	4.771584	0.131	0.317764	0.073	0.963	1781.22	1.17	1779.91	1.10	1778.80	1.13
z5	0.366	2.5302	99.85%	198	0.32	11825	0.107	0.108944	0.068	4.775083	0.136	0.318032	0.077	0.944	1781.02	1.24	1780.53	1.14	1780.11	1.19
z6	0.360	2.8906	99.92%	401	0.18	23996	0.105	0.108922	0.066	4.772373	0.132	0.317916	0.074	0.951	1780.65	1.20	1780.05	1.11	1779.54	1.15
z7	0.451	5.0856	99.69%	100	1.30	5857	0.131	0.108882	0.072	4.771395	0.135	0.317968	0.070	0.952	1779.98	1.31	1779.88	1.13	1779.80	1.08
z8	0.455	2.5040	99.63%	84	0.76	4943	0.133	0.108938	0.076	4.756832	0.140	0.316833	0.072	0.942	1780.92	1.39	1777.31	1.18	1774.25	1.12
																			<b>1780.78 ± 0.45 [2.39]</b>	
																			MSWD = 0.53	POF
																			= 0.82	
<b>ECMB6</b>																				
z1	0.284	1.213	0.990	30.870	0.97	1899.098	0.083	0.108982	0.123	4.778311	0.211	0.318136	0.135	0.834	1781.66	2.25	1781.10	1.77	1780.62	2.11
z2	0.335	4.276	0.999	226.896	0.47	13671.762	0.098	0.108953	0.066	4.771689	0.130	0.317780	0.070	0.958	1781.17	1.20	1779.93	1.09	1778.88	1.10
z3	0.376	2.430	0.998	155.473	0.39	9279.850	0.110	0.109024	0.086	4.774201	0.156	0.317741	0.096	0.872	1782.35	1.57	1780.37	1.31	1778.69	1.48
z4	0.476	1.918	0.998	191.621	0.26	11163.360	0.139	0.109071	0.098	4.775361	0.155	0.317681	0.084	0.822	1783.14	1.80	1780.58	1.30	1778.39	1.30
z5	0.315	2.478	0.999	228.261	0.27	13819.538	0.092	0.108954	0.069	4.777931	0.134	0.318193	0.072	0.953	1781.18	1.25	1781.03	1.13	1780.90	1.13
z6	0.332	1.064	0.998	155.064	0.17	9355.845	0.097	0.109010	0.084	4.755533	0.160	0.316539	0.101	0.893	1782.11	1.52	1777.08	1.35	1772.81	1.57
z7	0.342	2.990	0.999	215.887	0.35	12986.019	0.100	0.108930	0.069	4.768270	0.133	0.317621	0.071	0.946	1780.77	1.26	1779.33	1.11	1778.10	1.11
z8	0.368	2.149	0.998	126.819	0.43	7588.570	0.107	0.108921	0.074	4.768637	0.141	0.317671	0.079	0.931	1780.62	1.35	1779.39	1.18	1778.35	1.22
																			<b>1781.44 ± 0.51 [2.40]</b>	
																			MSWD = 1.24	POF
																			= 0.28	

(a) z1, z2 etc. are labels for single zircon grains or fragments annealed and chemically abraded after Mattinson (2005). Bold indicates z fraction included in weighted mean.  
 (b) Model Th/U ratio iteratively calculated from the radiogenic <sup>208</sup>Pb/<sup>206</sup>Pb ratio and <sup>206</sup>Pb/<sup>238</sup>U age.  
 (c) Pb\* and Pb<sub>c</sub> represent radiogenic and common Pb, respectively; mol % <sup>206</sup>Pb\* with respect to radiogenic, blank and initial common Pb.  
 (d) Measured ratio corrected for spike and fractionation only. Fractionation estimated at 0.16 ± 0.03 ‰/a.m.u. (1 sigma) for Daly analyses, based on analyses of EARTHTIME 202-205 trace solution run recently.  
 (e) Corrected for fractionation, spike, and common Pb; all common Pb was assumed to be procedural blank: <sup>206</sup>Pb/<sup>204</sup>Pb = 18.042 ± 0.61%; <sup>207</sup>Pb/<sup>206</sup>Pb = 15.537 ± 0.52%; <sup>208</sup>Pb/<sup>206</sup>Pb = 37.686 ± 0.63% (all uncertainties 1-sigma).  
 (f) Errors are 2 sigma, propagated using the algorithms of Schmitz and Schoene (2007).  
 (g) Calculations are based on the decay constants of Jaffey et al. (1971) and Hiess et al. (2012). <sup>206</sup>Pb/<sup>238</sup>U and <sup>207</sup>Pb/<sup>206</sup>Pb ages corrected for initial disequilibrium in <sup>230</sup>Th/<sup>232</sup>U using DTh/U [magma] = 0.20 ± 0.05 (1 sigma).  
 (h) Weighted mean ± 2s internal uncertainty [± 2s internal + decay constant uncertainties]. MSWD = Mean Standard Weighted Deviation. POF = Probability of Fit

Table S2: Apatite U-Pb data

Spot name	U (ppm)	Th (ppm)	<sup>207</sup> Pb/ <sup>235</sup> U	2 SE (%)	<sup>206</sup> Pb/ <sup>238</sup> U	2 SE (%)	rho	<sup>238</sup> U/ <sup>206</sup> Pb	2 SE (%)	<sup>207</sup> Pb/ <sup>206</sup> Pb	2 SE (%)
<b>ECMB1</b>											
ECMB1_1	49.6	28.3	10.6010	4.4	0.3734	3.1	0.71	2.6781	3.1	0.2060	3.1
ECMB1_2	54.4	24.3	9.7938	4.4	0.3707	3.1	0.70	2.6976	3.1	0.1917	3.1
ECMB1_3	33.1	28.4	12.6190	4.5	0.3704	3.3	0.73	2.6998	3.3	0.2472	3.1
ECMB1_4	44.1	12.5	10.8580	4.6	0.3634	3.3	0.71	2.7518	3.3	0.2168	3.2
ECMB1_5	53.9	12.7	9.4952	4.5	0.3404	3.3	0.73	2.9377	3.3	0.2024	3.1
ECMB1_7	36.7	21.4	12.4303	4.4	0.3701	3.1	0.72	2.7020	3.1	0.2437	3.0
ECMB1_8	35.0	11.3	13.0744	4.4	0.3986	3.2	0.72	2.5088	3.2	0.2380	3.1
ECMB1_9	18.9	4.0	18.8677	4.6	0.4342	3.2	0.71	2.3031	3.2	0.3153	3.2
ECMB1_10	43.1	17.7	11.1589	4.5	0.3841	3.1	0.70	2.6035	3.1	0.2108	3.2
ECMB1_11	35.2	92.0	11.5084	4.7	0.3857	3.3	0.71	2.5927	3.3	0.2165	3.3
ECMB1_12	34.6	15.8	12.4772	4.4	0.3950	3.2	0.72	2.5316	3.2	0.2292	3.1
ECMB1_13	48.9	16.5	9.9917	4.4	0.3601	3.1	0.72	2.7770	3.1	0.2013	3.0
ECMB1_14	56.7	21.5	9.3120	4.4	0.3565	3.1	0.72	2.8050	3.1	0.1895	3.0
ECMB1_15	35.4	15.2	11.3952	4.5	0.3801	3.3	0.73	2.6309	3.3	0.2175	3.0
ECMB1_16	56.8	23.2	9.7936	4.5	0.3621	3.4	0.74	2.7617	3.4	0.1963	3.0
ECMB1_17	97.9	67.0	7.2326	4.5	0.3311	3.2	0.71	3.0202	3.2	0.1585	3.2
ECMB1_19	28.9	6.7	14.2073	4.5	0.3971	3.4	0.74	2.5183	3.4	0.2596	3.0
ECMB1_20	41.0	13.5	10.6404	4.6	0.3596	3.4	0.73	2.7809	3.4	0.2147	3.1
ECMB1_21	27.3	4.0	13.9859	4.7	0.3863	3.5	0.74	2.5887	3.5	0.2627	3.1
ECMB1_22	41.1	19.6	12.2605	4.6	0.3878	3.4	0.74	2.5786	3.4	0.2294	3.1
ECMB1_23	18.8	18.3	18.9832	4.4	0.4287	3.2	0.72	2.3326	3.2	0.3213	3.1
ECMB1_24	39.2	13.3	11.5599	4.4	0.3666	3.1	0.69	2.7278	3.1	0.2288	3.2
ECMB1_25	38.3	14.2	11.4021	4.9	0.3664	3.5	0.72	2.7293	3.5	0.2258	3.4
ECMB1_26	164.9	3840.0	6.5949	5.2	0.3419	4.2	0.81	2.9248	4.2	0.1400	3.0
ECMB1_27	38.6	11.9	10.9887	4.4	0.3712	3.1	0.72	2.6940	3.1	0.2148	3.1
<del>ECMB1_6</del>	<del>18.4</del>	<del>3.7</del>	<del>18.7929</del>	<del>4.6</del>	<del>0.5173</del>	<del>3.4</del>	<del>0.73</del>	<del>1.9331</del>	<del>3.4</del>	<del>0.2636</del>	<del>3.1</del>
<del>ECMB1_18</del>	<del>30.5</del>	<del>6.4</del>	<del>15.0260</del>	<del>4.8</del>	<del>0.4509</del>	<del>3.6</del>	<del>0.76</del>	<del>2.2178</del>	<del>3.6</del>	<del>0.2418</del>	<del>3.2</del>
<del>ECMB1_28</del>	<del>38.1</del>	<del>24.4</del>	<del>14.3511</del>	<del>4.8</del>	<del>0.4319</del>	<del>3.7</del>	<del>0.77</del>	<del>2.3154</del>	<del>3.7</del>	<del>0.2411</del>	<del>3.1</del>
<del>ECMB1_29</del>	<del>48.7</del>	<del>29.8</del>	<del>11.2605</del>	<del>4.5</del>	<del>0.4015</del>	<del>3.2</del>	<del>0.72</del>	<del>2.4907</del>	<del>3.2</del>	<del>0.2035</del>	<del>3.1</del>
<del>ECMB1_30</del>	<del>34.2</del>	<del>10.6</del>	<del>15.0964</del>	<del>4.8</del>	<del>0.4970</del>	<del>3.8</del>	<del>0.78</del>	<del>2.0121</del>	<del>3.8</del>	<del>0.2204</del>	<del>3.1</del>
<b>ECMB3</b>											
ECMB3_1	70.1	179.1	8.5024	4.4	0.3394	3.2	0.72	2.9464	3.2	0.1818	3.0
ECMB3_2	33.3	121.6	12.5593	4.9	0.3805	3.8	0.78	2.6281	3.8	0.2395	3.1
ECMB3_3	28.5	72.2	13.1515	4.4	0.3865	3.2	0.72	2.5873	3.2	0.2469	3.1
ECMB3_4	22.0	64.8	15.9828	4.6	0.4130	3.4	0.74	2.4213	3.4	0.2808	3.1
ECMB3_5	22.3	60.6	14.7595	4.7	0.3833	3.4	0.73	2.6089	3.4	0.2794	3.2
ECMB3_6	22.5	72.0	15.7819	4.7	0.4140	3.5	0.75	2.4155	3.5	0.2766	3.1
ECMB3_7	15.9	42.4	19.6392	4.9	0.4431	3.8	0.77	2.2568	3.8	0.3216	3.1
ECMB3_8	4.8	6.3	45.8747	7.7	0.6040	7.0	0.90	1.6556	7.0	0.5511	3.3
ECMB3_9	3.8	6.7	56.4061	9.1	0.6910	8.5	0.93	1.4472	8.5	0.5923	3.2
ECMB3_10	59.0	158.7	8.8574	4.5	0.3435	3.2	0.71	2.9112	3.2	0.1871	3.1
ECMB3_11	48.1	147.1	10.8222	4.5	0.3692	3.3	0.73	2.7086	3.3	0.2127	3.0
ECMB3_12	21.0	53.0	14.8251	4.8	0.4044	3.6	0.75	2.4728	3.6	0.2660	3.1
ECMB3_13	15.9	49.1	19.8998	4.5	0.4494	3.3	0.73	2.2252	3.3	0.3213	3.1
ECMB3_14	15.4	33.1	18.9633	4.9	0.4181	3.8	0.78	2.3918	3.8	0.3291	3.1
ECMB3_15	17.0	50.8	17.4913	5.1	0.4239	4.1	0.79	2.3590	4.1	0.2994	3.1
ECMB3_16	42.6	65.7	8.7769	4.4	0.3503	3.1	0.70	2.8547	3.1	0.1818	3.1
ECMB3_17	42.6	118.3	11.2602	4.5	0.3717	3.3	0.74	2.6903	3.3	0.2198	3.0
ECMB3_18	4.9	11.6	43.5264	7.5	0.6010	6.7	0.89	1.6639	6.7	0.5255	3.4
ECMB3_19	8.0	13.0	30.2100	5.7	0.5080	4.4	0.78	1.9685	4.4	0.4315	3.5
ECMB3_20	47.8	96.7	9.4929	4.4	0.3631	3.1	0.70	2.7541	3.1	0.1897	3.2
ECMB3_21	12.5	32.3	21.9116	4.7	0.4311	3.5	0.75	2.3196	3.5	0.3688	3.1
ECMB3_22	4.2	7.6	51.7630	7.2	0.6630	6.5	0.90	1.5083	6.5	0.5665	3.1
ECMB3_23	5.2	7.7	42.9820	6.0	0.5880	5.1	0.85	1.7007	5.1	0.5304	3.2
ECMB3_24	3.4	5.5	63.3132	7.9	0.7430	7.3	0.92	1.3459	7.3	0.6183	3.2
ECMB3_26	20.3	51.3	16.7809	4.9	0.4150	3.7	0.77	2.4096	3.7	0.2934	3.1
ECMB3_27	8.1	18.1	30.5038	8.2	0.5290	7.6	0.92	1.8904	7.6	0.4184	3.3
ECMB3_28	6.0	10.9	39.1760	6.5	0.5610	5.6	0.86	1.7825	5.6	0.5067	3.3
ECMB3_29	60.4	174.2	9.5511	4.5	0.3602	3.2	0.72	2.7762	3.2	0.1924	3.1
ECMB3_30	28.5	54.8	12.2504	4.5	0.3667	3.1	0.69	2.7270	3.1	0.2424	3.3
ECMB3_31	112.7	387.0	7.7152	5.0	0.3445	3.9	0.77	2.9028	3.9	0.1625	3.2
ECMB3_32	10.4	28.3	25.2184	5.2	0.4710	4.1	0.78	2.1231	4.1	0.3885	3.2
ECMB3_33	8.6	20.1	28.1872	6.2	0.4850	5.2	0.84	2.0619	5.2	0.4217	3.4
ECMB3_34	10.0	31.6	22.8309	9.2	0.4400	8.5	0.93	2.2727	8.5	0.3765	3.4
ECMB3_35	22.9	62.7	15.1554	4.4	0.4009	3.2	0.71	2.4944	3.2	0.2743	3.1
ECMB3_36	40.7	128.6	11.3333	4.5	0.3622	3.3	0.74	2.7609	3.3	0.2270	3.0
ECMB3_38	45.8	70.3	8.9418	4.4	0.3530	3.1	0.71	2.8329	3.1	0.1838	3.1
ECMB3_39	25.8	85.4	13.5569	4.4	0.3798	3.2	0.72	2.6330	3.2	0.2590	3.1
ECMB3_40	36.3	128.5	12.1928	4.5	0.3797	3.3	0.73	2.6337	3.3	0.2330	3.0
ECMB3_42	4.2	6.3	58.7790	6.7	0.7150	5.9	0.88	1.3986	5.9	0.5965	3.1
ECMB3_43	33.8	90.3	11.9596	4.4	0.3699	3.1	0.71	2.7034	3.1	0.2346	3.1
ECMB3_44	23.1	72.5	14.4674	4.7	0.3827	3.5	0.75	2.6130	3.5	0.2743	3.1
ECMB3_45	8.6	11.8	28.0617	5.5	0.5030	4.6	0.83	1.9881	4.6	0.4048	3.1
ECMB3_46	4.3	6.7	50.1885	6.5	0.6330	5.7	0.88	1.5798	5.7	0.5753	3.1
ECMB3_47	7.6	9.7	30.6232	5.0	0.5000	4.0	0.79	2.0000	4.0	0.4444	3.1
ECMB3_48	16.9	32.0	17.7004	5.0	0.4131	3.8	0.75	2.4207	3.8	0.3109	3.3
ECMB3_49	5.2	9.3	44.3465	8.6	0.5950	8.0	0.93	1.6807	8.0	0.5408	3.1
<del>ECMB3_25</del>	<del>9.4</del>	<del>17.7</del>	<del>29.5612</del>	<del>6.0</del>	<del>0.5490</del>	<del>5.0</del>	<del>0.83</del>	<del>1.8215</del>	<del>5.0</del>	<del>0.3907</del>	<del>3.4</del>
<del>ECMB3_37</del>	<del>18.1</del>	<del>33.5</del>	<del>19.5908</del>	<del>4.7</del>	<del>0.4972</del>	<del>3.6</del>	<del>0.76</del>	<del>2.0113</del>	<del>3.6</del>	<del>0.2859</del>	<del>3.1</del>
<del>ECMB3_41</del>	<del>42.7</del>	<del>97.2</del>	<del>11.2835</del>	<del>4.5</del>	<del>0.3959</del>	<del>3.3</del>	<del>0.73</del>	<del>2.5259</del>	<del>3.3</del>	<del>0.2068</del>	<del>3.1</del>

<b>ECMB4</b>											
ECMB4_1	42.2	132.0	7.9502	4.4	0.3331	3.1	0.71	3.0021	3.1	0.1732	3.1
ECMB4_2	20.5	32.1	9.9057	4.6	0.3583	3.2	0.71	2.7910	3.2	0.2006	3.2
ECMB4_3	86.8	161.3	6.4967	4.3	0.3282	3.1	0.71	3.0469	3.1	0.1436	3.0
ECMB4_4	23.4	75.2	10.1149	4.8	0.3677	3.6	0.74	2.7196	3.6	0.1996	3.3
ECMB4_5	6.8	16.9	23.2267	6.0	0.4630	4.3	0.71	2.1598	4.3	0.3640	4.2
ECMB4_6	54.4	168.0	7.0172	4.5	0.3241	3.2	0.71	3.0855	3.2	0.1571	3.1
ECMB4_7	42.2	92.3	8.1260	4.4	0.3428	3.2	0.71	2.9172	3.2	0.1720	3.1
ECMB4_8	43.1	47.8	7.8004	4.5	0.3365	3.3	0.73	2.9718	3.3	0.1682	3.1
ECMB4_9	77.4	218.3	6.9050	4.3	0.3200	3.1	0.71	3.1250	3.1	0.1566	3.1
ECMB4_10	243.4	362.4	5.3710	4.3	0.3120	3.1	0.71	3.2051	3.1	0.1249	3.0
ECMB4_11	128.9	247.7	6.2358	4.3	0.3301	3.1	0.71	3.0294	3.1	0.1371	3.1
ECMB4_12	7.9	25.3	19.9491	6.3	0.4280	5.2	0.83	2.3364	5.2	0.3382	3.5
ECMB4_13	2.6	4.4	41.3289	8.9	0.5880	8.3	0.93	1.7007	8.3	0.5100	3.3
ECMB4_14	7.8	16.4	22.8506	5.5	0.4660	4.3	0.78	2.1459	4.3	0.3558	3.4
ECMB4_15	42.9	96.1	7.9985	4.5	0.3554	3.2	0.72	2.8137	3.2	0.1633	3.1
ECMB4_16	71.2	137.0	6.5874	4.3	0.3256	3.1	0.71	3.0713	3.1	0.1468	3.0
ECMB4_17	67.9	117.3	6.4340	4.3	0.3129	3.1	0.72	3.1959	3.1	0.1492	3.0
ECMB4_18	43.6	44.7	7.5746	4.6	0.3448	3.2	0.69	2.9002	3.2	0.1594	3.3
ECMB4_19	49.5	147.3	7.8430	4.9	0.3605	3.8	0.78	2.7739	3.8	0.1579	3.1
ECMB4_20	12.3	20.5	11.8828	5.4	0.3752	4.2	0.78	2.6652	4.2	0.2298	3.3
ECMB4_21	15.2	37.2	12.2956	4.8	0.3612	3.6	0.75	2.7685	3.6	0.2470	3.2
<b>Standards</b>											
MAD_1	28.4	713.3	1.5713	3.3	0.0860	2.3	0.69	11.6306	2.3	0.1326	2.4
MAD_2	28.0	709.2	1.5796	3.2	0.0855	2.2	0.68	11.7000	2.2	0.1341	2.4
MAD_3	27.8	700.3	1.5342	3.4	0.0846	2.2	0.65	11.8217	2.2	0.1316	2.6
MAD_4	27.8	690.6	1.5090	3.3	0.0831	2.2	0.66	12.0279	2.2	0.1317	2.5
MAD_5	27.9	696.4	1.5248	3.3	0.0863	2.4	0.71	11.5875	2.4	0.1282	2.4
MAD_6	28.1	693.6	1.5438	3.3	0.0855	2.3	0.70	11.6945	2.3	0.1310	2.3
MAD_7	27.1	673.7	1.4123	3.2	0.0839	2.2	0.67	11.9246	2.2	0.1222	2.4
MAD_8	27.6	692.1	1.5598	3.3	0.0860	2.3	0.67	11.6279	2.3	0.1316	2.5
MAD_9	28.2	704.0	1.5330	3.5	0.0846	2.3	0.66	11.8217	2.3	0.1315	2.6
MAD_10	28.1	697.2	1.5045	3.3	0.0847	2.1	0.65	11.8078	2.1	0.1289	2.5
MAD_11	28.0	691.9	1.5490	3.3	0.0852	2.2	0.65	11.7440	2.2	0.1320	2.5
MAD_12	28.0	702.9	1.5404	3.2	0.0836	2.2	0.70	11.9617	2.2	0.1337	2.3
MAD_13	28.3	712.3	1.4711	3.4	0.0834	2.2	0.66	11.9919	2.2	0.1280	2.5
MAD_14	28.0	701.2	1.5407	3.4	0.0843	2.3	0.67	11.8610	2.3	0.1326	2.6
MAD_15	28.1	704.5	1.5307	3.4	0.0847	2.2	0.66	11.8036	2.2	0.1311	2.5
MAD_16	28.1	700.1	1.5184	3.5	0.0851	2.1	0.61	11.7454	2.1	0.1294	2.8
MAD_17	27.2	680.8	1.4555	3.3	0.0847	2.2	0.66	11.8078	2.2	0.1247	2.5
MAD_18	28.7	725.3	1.6133	3.2	0.0858	2.2	0.68	11.6605	2.2	0.1365	2.3
McClure_1	31.5	65.0	2.5769	3.4	0.1009	2.2	0.64	9.9157	2.2	0.1854	2.6
McClure_2	16.5	34.0	3.5750	3.5	0.1089	2.4	0.69	9.1827	2.4	0.2382	2.5
McClure_3	15.2	33.9	3.7798	3.7	0.1101	2.6	0.71	9.0827	2.6	0.2491	2.6
McClure_4	13.6	29.4	3.8994	3.5	0.1125	2.6	0.72	8.8889	2.6	0.2515	2.4
McClure_5	14.6	31.3	3.7320	3.4	0.1099	2.4	0.71	9.0992	2.4	0.2464	2.4
McClure_6	15.0	31.0	3.7084	4.2	0.1085	2.4	0.57	9.2166	2.4	0.2480	3.4
McClure_7	14.5	30.3	3.5285	3.7	0.1078	2.7	0.72	9.2764	2.7	0.2375	2.5
McClure_8	14.7	32.0	4.0311	3.6	0.1105	2.5	0.68	9.0498	2.5	0.2647	2.7
McClure_9	15.3	31.7	3.3715	3.8	0.1099	2.5	0.65	9.0992	2.5	0.2226	2.9
OD306_1	24.3	69.8	5.1614	6.1	0.2861	2.4	0.39	3.4953	2.4	0.1309	5.6
OD306_2	24.8	67.1	4.5124	3.4	0.2825	2.3	0.66	3.5398	2.3	0.1159	2.6
OD306_3	14.0	56.0	4.9799	6.3	0.2893	2.6	0.42	3.4566	2.6	0.1249	5.7
OD306_4	22.9	67.7	4.9913	4.1	0.2845	2.4	0.59	3.5149	2.4	0.1273	3.3
OD306_5	23.8	68.6	4.1846	3.1	0.2813	2.1	0.70	3.5549	2.1	0.1079	2.2
OD306_6	18.3	36.9	4.4793	3.3	0.2889	2.3	0.71	3.4614	2.3	0.1125	2.3
OD306_7	20.3	48.6	4.3296	3.3	0.2843	2.3	0.69	3.5174	2.3	0.1105	2.4
OD306_8	19.8	40.4	4.4902	3.4	0.2909	2.3	0.68	3.4376	2.3	0.1120	2.5
OD306_9	23.9	62.4	5.2166	3.8	0.2861	2.2	0.58	3.4953	2.2	0.1323	3.1
OD306_10	19.7	39.8	4.2460	3.7	0.2811	2.4	0.64	3.5575	2.4	0.1096	2.8
OD306_11	22.9	45.4	4.3426	3.2	0.2849	2.2	0.69	3.5100	2.2	0.1106	2.3
OD306_12	24.0	73.8	4.6749	4.0	0.2815	2.2	0.57	3.5524	2.2	0.1205	3.3
OD306_13	11.3	25.1	4.2567	3.4	0.2831	2.5	0.72	3.5323	2.5	0.1091	2.4
OD306_14	24.0	65.6	4.8882	3.6	0.2784	2.4	0.67	3.5920	2.4	0.1274	2.7
OD306_15	26.1	80.7	4.4955	3.1	0.2812	2.1	0.69	3.5562	2.1	0.1160	2.3
OD306_16	19.0	38.6	5.6322	5.0	0.2880	2.4	0.48	3.4722	2.4	0.1419	4.4
OD306_17	9.5	20.9	9.2240	6.4	0.3210	4.0	0.62	3.1153	4.0	0.2085	5.1
OD306_18	24.1	67.8	5.1256	3.5	0.2901	2.2	0.62	3.4471	2.2	0.1282	2.7
401_1	19.3	139.5	1.0279	4.2	0.0890	2.5	0.59	11.2360	2.5	0.0838	3.4
401_2	19.0	144.6	1.2461	4.4	0.0917	2.4	0.55	10.9051	2.4	0.0986	3.6
401_3	19.1	140.0	0.9271	4.1	0.0884	2.4	0.58	11.3122	2.4	0.0761	3.3
401_4	18.3	137.1	0.9784	4.8	0.0883	2.4	0.49	11.3250	2.4	0.0804	4.2
401_5	19.0	132.4	0.8781	3.6	0.0861	2.3	0.62	11.6144	2.3	0.0740	2.8
401_6	19.7	138.0	0.8524	3.8	0.0859	2.6	0.68	11.6414	2.6	0.0720	2.8
401_7	19.6	137.3	0.8867	4.5	0.0871	2.2	0.49	11.4863	2.2	0.0739	3.9
401_8	19.2	140.3	1.0044	4.6	0.0892	2.8	0.60	11.2108	2.8	0.0817	3.7
401_9	19.1	139.2	0.8821	3.6	0.0872	2.2	0.61	11.4679	2.2	0.0734	2.9
401_10	19.8	143.1	0.8847	3.8	0.0855	2.2	0.58	11.6986	2.2	0.0751	3.1
401_11	19.6	141.1	0.8927	3.6	0.0873	2.4	0.67	11.4548	2.4	0.0742	2.7
401_12	20.1	149.2	0.8859	3.7	0.0852	2.2	0.59	11.7302	2.2	0.0754	3.0
401_13	19.1	139.5	0.9152	4.3	0.0877	2.2	0.52	11.3999	2.2	0.0757	3.6
401_14	19.4	140.3	0.8451	4.1	0.0853	2.3	0.55	11.7247	2.3	0.0719	3.4
401_15	19.6	140.1	0.8696	3.8	0.0855	2.3	0.60	11.6959	2.3	0.0738	3.1
401_16	19.6	140.9	0.9218	4.3	0.0851	2.5	0.58	11.7509	2.5	0.0786	3.5

Strikethrough indicates data discarded because final integrations were not flat

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