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Authors

Swanson-Hysell, NL
Avery, MS
Zhang, Y
[et al.](#)

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The paleogeography of Laurentia in its early years: new constraints from the Paleoproterozoic East-Central Minnesota batholith

Nicholas L. Swanson-Hysell¹, Margaret S. Avery¹, Yiming Zhang¹, Eben B. Hodgin¹, Robert J. Sherwood¹, Francisco E. Apen², Terrence J. Boerboom³, C. Brenhin Keller^{1,4,5}, John M. Cottle²

¹Department of Earth and Planetary Science, University of California, Berkeley, CA, USA

²Department of Earth Science, University of California, Santa Barbara, CA, USA

³Minnesota Geological Survey, St. Paul, MN, USA

⁴Berkeley Geochronology Center, 2455 Ridge Road, Berkeley, CA, USA

⁵Department of Earth Sciences, Dartmouth College, Hanover, NH, USA

Key Points:

- A new *ca.* 1780 Ma paleomagnetic pole reconstructs the Superior province of Laurentia to moderately high latitudes
- This pole establishes the coherency of the Laurentia craton following Trans-Hudson orogenesis and supports the NENA connection with Baltica
- Paleomagnetic and geologic data from Laurentia strongly support mobile-lid plate tectonics from 2.2 Ga to the present day

Corresponding author: Nicholas Swanson-Hysell, swanson-hysell@berkeley.edu

Abstract

The *ca.* 1.83 Ga Trans-Hudson orogeny resulted from collision of an upper plate consisting of the Hearne, Rae, and Slave provinces with a lower plate consisting of the Superior province. While the geologic record of *ca.* 1.83 Ga peak metamorphism within the orogen suggests that these provinces were a single amalgamated craton from this time onward, a lack of paleomagnetic poles from the Superior province following Trans-Hudson orogenesis has made this coherency difficult to test. We develop a high-quality paleomagnetic pole for northeast-trending diabase dikes of the post-Penokean orogen East-Central Minnesota Batholith (pole longitude: 265.8°; pole latitude: 20.4°; A_{95} : 4.5°; K: 45.6 N: 23) whose age we constrain to be 1779.1 ± 2.3 Ma (95% CI) with new U-Pb dates. Demagnetization and low-temperature magnetometry experiments establish dike remanence be held by low-Ti titanomagnetite. Thermochronology data constrain the intrusions to have cooled below magnetite blocking temperatures upon initial emplacement with a mild subsequent thermal history within the stable craton. The similarity of this new Superior province pole with poles from the Slave and Rae provinces establishes the coherency of Laurentia following Trans-Hudson orogenesis. This consistency supports interpretations that older discrepant 2.22 to 1.87 Ga pole positions between the provinces are the result of differential motion through mobile-lid plate tectonics. The new pole supports the NENA connection between the Laurentia and Fennoscandia cratons. The pole can be used to jointly reconstruct these cratons *ca.* 1780 Ma strengthening the paleogeographic position of these major constituents of the hypothesized late Paleoproterozoic supercontinent Nuna.

1 Introduction

In the Orosirian Period of the Paleoproterozoic Era, a series of collisional orogenies led to the amalgamation of Archean provinces to form the core of the Laurentia craton (Fig. 1; Hoffman (1988); Whitmeyer and Karlstrom (2007)). The most significant of these orogenies was the *ca.* 1850 to 1800 Ma Trans-Hudson orogeny associated with the collision between the Superior province and the Churchill province which comprised a composite of the Slave, Hearne and Rae provinces (Fig. 1; Weller and St-Onge (2017)). The length of the orogen as well as the pressure-temperature of metamorphism within it are similar to that of continent-continent collision within the Himalayan orogen (Weller & St-Onge, 2017). The terminal closure of the intervening ocean basin between the Su-

perior and composite Slave + Hearne + Rae provinces is interpreted in paleogeographic models to be associated not only with the assembly of Laurentia, but also with the conjoining of other continents into the hypothesized supercontinent Nuna (Pehrsson et al., 2015).

The rapid Paleoproterozoic amalgamation of the Laurentia craton led to the large persistent area of continental lithosphere that would grow further through accretionary orogenesis subsequently in the Paleoproterozoic Era and through the Mesoproterozoic Era (Whitmeyer & Karlstrom, 2007). This subsequent orogenesis along the southern to eastern margin of Laurentia (present-day coordinates) indicates that it was a long-lived accretionary margin (Karlstrom et al., 2001; Whitmeyer & Karlstrom, 2007). This accretionary margin has been interpreted to have extended beyond Laurentia and have continued onto Baltica and Australia (Karlstrom et al., 2001). Based on correlation of Archean provinces and Paleoproterozoic orogenic belts, Gower et al. (1990) reconstructed Baltica to Laurentia in a position known as the NENA (northern Europe and North America) configuration. This reconstruction is compatible with existing paleomagnetic constraints from *ca.* 1750 to 1270 Ma (Evans & Pisarevsky, 2008) and these conjoined cratons feature as a major component of the hypothesized Nuna supercontinent (Evans & Mitchell, 2011; Zhang et al., 2012).

Additional data that constrain the paleogeographic position of Laurentia from the time just following the Trans-Hudson orogeny can test the hypothesis of the unity of Laurentia's Archean provinces, establish the position of the newly amalgamated Laurentia, and thereby enable tests of hypothesized connections with other cratons. This study develops a new paleomagnetic pole for Laurentia from *ca.* 1780 Ma diabase dikes of the East-Central Minnesota Batholith (ECMB) that provides such constraints.

2 Geologic Setting

Coeval with collisional orogenesis between the assembling Archean provinces that formed Laurentia's core was the *ca.* 1860 to 1820 Ma accretionary Penokean orogeny along the southern margin of the Superior Province (Fig. 1; Schulz and Cannon (2007)). Penokean orogenesis resulted from island-arc and microcontinent collisions with the Superior Province that led to metamorphism of Superior Province lithologies and development of a fore-land basin (Schulz & Cannon, 2007; Holm et al., 2019). Following the Penokean orogeny,

Figure 1. Map of Laurentia showing the location of Archean provinces and younger Proterozoic crust (simplified from Whitmeyer and Karlstrom (2007)). The localities of paleomagnetic poles that constrain Laurentia’s position just after its amalgamation are shown with stars including the new pole from this study developed from the East-Central Minnesota Batholith (ECMB). The outline of the state of Minnesota around the ECMB blue star is the region for the geologic map in the right panel. This map shows interpreted Precambrian geology for the state of Minnesota (simplified from Jirsa et al. (2012)) including in regions covered by Phanerozoic sedimentary rocks where the Precambrian bedrock is inferred from geophysical data and drill cores.

there was voluminous magmatism along the southeastern margin of the west Superior Province resulting in the emplacement of the *ca.* 1780 Ma East-Central Minnesota Batholith (ECMB) and other coeval post-orogenic plutons (Fig. 1; Holm et al. (2005); Boerboom et al. (2005); Schmitz et al. (2018)). While the ECMB is dominantly comprised of felsic to intermediate plutons, mafic magmas were also generated and commingled with the more abundant felsic magmas throughout the interval of batholith generation as evidenced by mafic enclaves within some of the plutons (Boerboom et al., 2005, 2011; Schmitz et al., 2018). Mafic melt within the ECMB also led to the emplacement of a set of near-vertical northeast-trending diabase dikes (Fig. 2; Boerboom et al. (2005)). The dikes have chilled margins and are typically 1 to 3 meters wide with widths up to 8 meters (Boerboom et al., 2005). As with the granites they intrude, the dikes have primary igneous texture and no metamorphic fabric (Boerboom et al., 2005). They have experienced variable low-grade alteration such as albitization and sericitization of plagioclase and the formation

of pyrrhotite. These diabase dikes are present within all of the granitoid units of the ECMB (e.g., St. Cloud Granite, Rockville Granite, and Reformatory Granodiorite; Fig. 2) with the exception of the youngest Richmond Granite. Throughout the field area, the dikes are exposed both in glacially-polished pavement outcrops and in numerous inactive and active dimension stone granite quarries. Northeast-trending diabase dikes are present in all of the quarries in the Rockville Granite, St. Cloud Granite as well as in the Reformatory granodiorite, regardless of the size of the quarry, as well as in many natural bedrock outcrops. In many of the old inactive quarries, the north and/or south quarry walls are marked by the planar surface of a diabase dike contact, where the rock naturally separates, often resulting in elongated northeast-southwest shapes to the quarry pits. In contrast, no diabase dikes have been found in the quarries or natural exposures of the Richmond granite. Although this granite does not contain as many quarries and there are fewer natural outcrops, the lack of diabase dikes contrasts sharply with the numerous dikes present in the other nearby granites, where an equivalent exposed surface area would contain numerous diabase dikes. This absence indicates that the younger Richmond Granite post-dates the intrusion of the diabase dikes into the St. Cloud Granite, Rockville Granite, and Reformatory Granodiorite.

There are also quartz-feldspar porphyry dikes with the same northeast-trending direction as the diabase dikes found in all the granitoids also with the exception of the Richmond Granite (Boerboom et al., 2005). These porphyritic microgranite dikes have chilled, and locally flow-banded, margins. One has been observed to have intruded into a northeast-trending diabase dike and another has textures consistent with commingling of magmas between the felsic dike and adjacent diabase dike indicative of synchronous emplacement (Boerboom & Holm, 2000). The Richmond Granite has trachytoid magmatic fabric defined by aligned potassium-feldspar phenocrysts that share the same orientation with the northeast-trending dikes (Boerboom & Holm, 2000), indicating that this orientation is associated with a persistent regional stress field throughout the interval of magma emplacement and dike formation. These field relationships indicate that the quartz-feldspar porphyry and diabase dikes are comagmatic with the batholith. The diabase dikes are constrained to be younger than the St. Cloud Granite (new U-Pb date of 1781.44 ± 0.51 Ma; 2σ analytical uncertainty) which they pervasively intrude, older than the Richmond Granite (new U-Pb date of 1776.76 ± 0.49 Ma) in which they are absent, and similar in age to the quartz-feldspar porphyry dikes (new U-Pb date of 1780.78 ± 0.45 Ma).

Figure 2. Left panel: Locations of paleomagnetic sites of the northeast-trending dikes (NED) and a northwest-trending dike (NWD) within the Rockville Granite, Reformatory Granodiorite and St. Cloud Granite of the ECMB (bedrock geology from Boerboom et al. (1995) and shown in UTM zone 15N WGS 84 coordinate reference system such that each axis tick is 1 km). Within regions of the mapped St. Cloud Granite there is more complex interfingering of that granite with the Reformatory Granodiorite than is shown. The strikes of the dikes are shown as lines (black when measured on that dike; grey when using the overall mean orientation from the measured NED dikes). The location of the QP1 and ECMB6 geochronology samples are shown. The ECMB4 geochronology sample was collected ~18 km SW of the western edge of the map, in the Richmond Granite which cross-cuts the Rockville Granite and is younger than the NED dikes. Right panel: The orientations of dikes. Each individual dike orientation is the mean of multiple measurements on that dike. The mean of the poles to the NED planes is shown with a red dot and a 95% confidence ellipse on the mean calculated with Fisher statistics. This confidence ellipse intersects the equator indicating that the mean plane cannot be distinguished from vertical.

3 Paleomagnetic Methods and Results

Oriented samples for paleomagnetism were collected and analyzed from 36 of the northeast-trending dikes of the ECMB and one northwest-trending dike (Fig. 2). Each sampled dike constituted a paleomagnetic site in our site naming scheme. These sites were concentrated in and around Stearns County Quarry Park near the city of St. Cloud (Fig. 2). Samples were collected from the dikes with a gas-powered drill and oriented in the field with a Pomeroy orienting fixture. The azimuthal orientations of the cores were determined either through sun or magnetic compass depending on cloud cover. Sun compass directions were preferentially used when available. When magnetic compass data were used they were corrected for local magnetic declination using the International Geomagnetic Reference Field model (Thébault et al., 2015). Specimens from the oriented

139 samples were analyzed in the UC Berkeley Paleomagnetism lab using a 2G DC-SQUID
 140 magnetometer. Samples either underwent stepwise alternating field (AF) or thermal de-
 141 magnetization. Thermal demagnetization was accomplished using an ASC thermal de-
 142 magnetizer (residual fields <10 nT). AF demagnetization was conducted with inline coils
 143 that utilize a Crest Amplifier paired with an Adwin controller to develop a well-controlled
 144 waveform. All paleomagnetic data developed in this study are available at the measure-
 145 ment level in the MagIC database ([https://www.earthref.org/](https://www.earthref.org/MagIC/) MagIC/ doi/ INSERT-
 146 DOI; UPDATE TO DOI WHEN ASSIGNED).

147 Typical behaviors of sample remanence during demagnetization are illustrated for
 148 representative specimens in Figure 3. AF demagnetization data typically reveal three
 149 components: a small low-coercivity component approximately aligned with Earth’s present
 150 local field in the study region that was typically removed below 10 mT; a medium-coercivity
 151 component that is steep and was dominantly removed between 10 and 60 mT; and a high-
 152 coercivity component that was subsequently removed incompletely as demagnetization
 153 progressed to 130 mT. These components are present to varying degrees within individ-
 154 ual specimens (Fig. 3).

155 Sister specimens from some samples underwent thermal and AF demagnetization
 156 which provides additional insight into the carriers of the components through compar-
 157 ison of the thermal and AF demagnetization spectra (such as NED2-8 in Fig. 3). These
 158 data reveal that the low-coercivity component direction is removed at the lowest unblock-
 159 ing temperatures up to 150°C. This behavior, as well as the typical direction, is most con-
 160 sistent with the component being a viscous overprint acquired in Earth’s geomagnetic
 161 field. The direction of the high-coercivity component is removed through thermal de-
 162 magnetization between 250°C and 350°C — consistent with it being held by monoclinic
 163 pyrrhotite. The direction of this magnetization held by pyrrhotite is aligned with the magnetite-
 164 held remanence within a northwest-trending dike in the region (discussed in more de-
 165 tail below) — a direction consistent with the position of Laurentia during the time pe-
 166 riod of *ca.* 1096 Ma Midcontinent Rift magmatism (Swanson-Hysell et al., 2020). We
 167 interpret this high-coercivity component held by pyrrhotite, whose presence is variable
 168 in ECMB dikes, to have formed through hydrothermal activity associated with Midcon-
 169 tinent Rift magmatism such as that represented by the emplacement of the northwest-
 170 trending dike. The pyrrhotite thereby carries a chemical remanent magnetization.

Figure 3. Paleomagnetic data from ECMB northeast-trending diabase dikes are shown in geographic coordinates on vector component plots, measurement-level equal area plots and magnetization magnitude plots (developed using PmagPy software; Tauxe et al. (2016)). Least-squares fits to the data are shown with colored arrows on the vector component plots, colored directions on the equal area plots, and as colored end-points on the magnetization magnitude plots (pink for low-coercivity; blue for medium-coercivity; yellow for high-coercivity). Specimens NED2-8a and NED2-8b are from the same core sample and were analyzed via thermal and alternating field (AF) demagnetization respectively. These data from sample NED2-8 reveal the steep medium-coercivity component to thermally unblock at temperatures characteristic of remanence held by magnetite and the high-coercivity component to thermally unblock at temperatures characteristic of pyrrhotite. Specimen NED34-6a is dominated by the steep medium-coercivity component. The medium-coercivity component is well-resolved in specimen NED12-3b which also has a substantial high-coercivity component. The high-coercivity component dominates the remanence of specimen NED36-7a such that no medium-coercivity component can be resolved.

In some specimens, the low-coercivity component is much larger and has a direction that is distinct from the present local field. It is likely that these samples acquired an isothermal remanent magnetization associated with quarrying operations or lightning strikes. This behavior can be prevalent throughout a site or can be present in just some samples from a given site. In many cases, these low-coercivity overprints can be removed through AF demagnetization and the medium-coercivity and/or high-coercivity components can be subsequently isolated. In some instances, however, these large and dominantly low-coercivity overprints extend to higher coercivities preventing the isolation of other components.

The medium-coercivity component direction is dominantly removed through thermal demagnetization between 515°C and 565°C consistent with it being held by low-Ti titanomagnetite. This direction was recovered with site mean direction uncertainty less than 8 degrees ($\alpha_{95} < 8^\circ$) for 23 sites (Table 1). We interpret this component to be a primary thermal remanence acquired at the time of dike emplacement as part of the *ca.* 1780 Ma ECMB. This interpretation gains support from the rock magnetic data, an inverse baked contact test, and thermochronology data that support an emplacement temperature well below the blocking temperature of magnetite and a mild subsequent thermal history — as discussed in more detail below.

3.1 Magnetic mineralogy constraints from low-temperature magnetometry

To gain additional insight into the magnetic mineralogy of the dikes, a Magnetic Properties Measurement System (MPMS) at the Institute for Rock Magnetism was used to conduct low-temperature remanence experiments. In the field-cooled (FC) experiments shown in Figure 4, the magnetization was measured upon warming following the specimen having cooled in an applied field of 2.5 T from 300 to 10 K. In the zero-field-cooled (ZFC) experiment, a low-temperature saturation isothermal remanence (LTSIRM) of 2.5 T was applied at 10 K after the specimen cooled in a (near-)zero field. In the room-temperature saturation isothermal remanence (RTSIRM) experiment, the sample was pulsed with a 2.5 T field at room temperature (~ 300 K) and then cooled to 10 K and warmed back to room temperature in a (near-)zero field. These experiments reveal that sister specimens to paleomagnetic specimens whose remanence is dominated by the medium-coercivity component without an appreciable high-coercivity component have strong expressions

of the ~ 120 K Verwey transition as expected for a ferromagnetic mineralogy of well-preserved low-Ti titanomagnetite (NED34-6c in Fig. 4; Verwey (1939); Feinberg et al. (2015)). In contrast, specimens from samples that have a larger contribution of the high-coercivity component have weaker saturation magnetization, minor expression of the Verwey transition, and the presence of monoclinic pyrrhotite as evidenced through the ~ 30 K Besnus transition (NED36-8c in Fig. 4; Besnus and Meyer (1964); Feinberg et al. (2015)). Samples with a smaller contribution of the high-coercivity component superimposed on the medium-coercivity component have intermediate behavior with a minor expression of the Besnus transition and a more prominent Verwey transition (NED2-8c in Fig. 4). These results support the interpretation that the medium-coercivity component is held by primary unaltered (titano)magnetite and that the high-coercivity component is the result of subsequent alteration that resulted in degradation of magnetite and formation of pyrrhotite.

Figure 4. Low-temperature remanence experiment data. The specimen from a sample whose natural remanent magnetization is dominated by the medium-coercivity component (NED34-6c) has behavior dominated by magnetite as evidenced through the response across the ~ 120 K Verwey transition. The specimen from a sample whose natural remanence is dominated by the high-coercivity component (NED36-8c) has weaker magnetization, a relatively minor expression of the Verwey transition, and expression of the ~ 32 K Besnus transition that indicates the presence of monoclinic pyrrhotite. The specimen whose natural remanence has a well-resolved medium-coercivity component with a minor high-coercivity component (NED2-8c) has intermediate behavior with a Verwey transition that is not as suppressed as in NED36-8c with a minor, but resolvable, Besnus transition (see inset). FC: field-cooled; ZFC: zero-field-cooled; LTSIRM: low-temperature saturation isothermal remanence magnetization; RTSIRM: room-temperature saturation isothermal remanence magnetization.

3.2 Baked contact test

One northwest-trending dike is exposed and was sampled within the study region as site NWD1 (Fig. 2). The magnetization direction of this dike indicates that it is associated with the main stage of the Midcontinent Rift (*ca.* 1096 Ma) as it has a normal polarity and an inclination consistent with that time interval of Midcontinent Rift volcanism (Fig. 5; Swanson-Hysell et al. (2020)). The dike cross-cuts one of the northeast-trending dikes (NED17) allowing for a baked contact test (Figs. 2 and 5). The baked contact test is positive with a distinct direction in the northeast-trending dike (corresponding to the remanence direction seen throughout the northeast-trending dikes of the ECMB) with its magnetite remanence becoming progressively overprinted by the northwest-trending dike up to the contact (Fig. 5). This positive baked contact test indicates that the northeast-trending ECMB dikes have not been overprinted since the northwest-trending dike was emplaced (*ca.* 1096 Ma). This positive baked contact test for the northwest-trending dike constitutes what is referred to as a positive “inverse” baked contact test for the northeast-trending dike remanence — it constrains the remanence to be more ancient than the *ca.* 1.1 Ga Mesoproterozoic northwest-trending dike, but does not provide a constraint back to the Paleoproterozoic time of dike emplacement. Given that the host rocks for the northeast-trending dikes are of a very similar age to the dikes themselves there is not the possibility of a Paleoproterozoic baked contact test. In contrast to the dikes, stable and consistent remanence directions were not recovered from pilot sites in ECMB granites.

The high-coercivity remanence direction held by pyrrhotite in some of the northeast-trending dikes is aligned with the remanence direction of the northwest-trending dike. While the thermal effect of the northwest-trending dike was limited to a few meters on either side of it as evidenced through the baked contact test (Fig. 5), there was more widespread hydrothermal alteration associated with Midcontinent Rift magmatic activity that led to the formation of pyrrhotite and an associated chemical remanent magnetization. In the majority of the northeast-trending dikes, the original magnetite is well-preserved (e.g., NED34 of Figs. 3 and 4) while other dikes experienced variable magnetite alteration and the formation of secondary pyrrhotite (e.g., NED36 of Figs. 3 and 4). These components can be separated through progressive demagnetization (e.g., NED12 of Figs. 3) enabling the thermal remanent magnetization held by magnetite to be used to determine site mean directions and calculate a paleomagnetic pole (Fig. 6).

Figure 5. Results from a positive baked contact test where a northwest-trending dike (site NWD1) cross-cuts a northeast-trending dike (site NED17). The NWD1 dike has a direction indicating that it formed during the *ca.* 1096 Ma main stage of rift volcanism. Close to the NWD1 dike the NED17 is nearly fully overprinted (NED17-1a). Further from NWD1 there are partial thermal overprints (NED17-7a) that by ~7 meters from the cross-cutting dike are not resolvable (NED17-11a). Note that blocks from this dike were slightly shifted due to quarrying operations which does not bear on the significance of the positive test, but has led to the exclusion of the result from the site mean compilation and overall pole.

3.3 Paleomagnetic pole

The site mean directions determined from the magnetite remanence component can be converted to virtual geomagnetic poles and then used to calculate a mean paleomagnetic pole for the ECMB diabase dikes (pole longitude: 265.8°; pole latitude: 20.4°; A_{95} : 4.5°; K: 45.6 N: 23; Fig. 6). These 23 virtual geomagnetic poles have a distribution con-

sistent with a Fisher distribution as determined through the Fisher et al. (1987) goodness-of-fit method. The A_{95} uncertainty on the mean pole position of 4.5° is within the bounds of reliability proposed by Deenen et al. (2011). It is well below the A_{95} -max value proposed to establish a well-determined mean for 23 sites (11.4°) and above the A_{95} -min value (3.4°) consistent with the site directions having sufficiently sampled secular variation of the geomagnetic field.

In a massive host rock such as the ECMB plutons without preferential bedding or foliation, it is expected that lithospheric stresses will lead to the emplacement of near vertical dikes. Dike plane orientations were measured on each dike for which there was sufficient three-dimensional exposure. Multiple measurements were made for each dike to constrain their orientation. The mean strike calculated from 17 dike orientations is 067° and the mean dip is 88° (Fig. 2). The α_{95} uncertainty associated with the Fisher mean for the poles to these dike orientation planes (i.e. the lines perpendicular to the planes) is 4.7° which means that the overall orientation of the planes is statistically indistinguishable from vertical (Fig. 2). Due to this verticality, we interpret the exhumation of the ECMB plutons to have not resulted in appreciable tilting since dike emplacement and do not apply a tilt correction to the paleomagnetic data.

4 Geochronology Methods and Results

The field relationships show the diabase dikes to be younger than the Rockville Granite, Reformatory Granodiorite and the St. Cloud Granite which they pervasively intrude and to be older than the Richmond Granite where they are absent (Boerboom et al., 2005). Holm et al. (2005) developed U-Pb dates calculated as concordia intercept dates from these intrusions. The dates reported by Holm et al. (2005) for granites intruded by the dikes are 1783 ± 11 Ma for the Reformatory Granodiorite, 1780 ± 7 Ma for the Rockville Granite and 1779 ± 5 Ma for the St. Cloud Granite. The younger cross-cutting Richmond Granite has a date of 1772 ± 3 Ma (Holm et al., 2005). An age of 1774 ± 7 Ma for one of the quartz-feldspar porphyry dikes developed by Holm et al. (2005) is consistent with this interpretation of these dikes being older than the Richmond Granite (and younger than the granites they intrude). While the dates published in Holm et al. (2005) are valuable constraints and are consistent with the field relationships, they are of lower precision than what is possible with modern analytical approaches and therefore lead to overlapping uncertainties. Higher precision constraints resulting from methods that

Figure 6. A) Site mean directions for the magnetite remanence of the northeast-trending (NED) ECMB diabase dikes with $\alpha_{95} < 8^\circ$. B) Virtual geomagnetic poles (VGPs) calculated from these site means and the overall mean paleomagnetic pole for the ECMB dikes. C) Comparison between the new ECMB paleomagnetic pole and other *ca.* 1780 to 1740 Ma poles for Laurentia. D) Comparison of poles from Laurentia's provinces from 1830 to 1740 Ma from Evans et al. (2021) as well as poles from the Trans-Hudson orogen (THO; grey; Symons and Harris (2005)) with the new ECMB pole.

apply ion-exchange separation with low blank analyses of chemically-abraded single zircon grains can test the field relationship interpretations and provide more confidence in the overall age constraints.

To further constrain the age of the northeast-trending diabase dikes, we developed new isotope dilution-thermal ionization mass spectrometry (ID-TIMS) U-Pb zircon dates

Figure 7. U-Pb dates for ECMB samples. Weighted mean dates (horizontal lines) are calculated from individual zircon dates (vertical bars). The NED diabase dikes intrude the St. Cloud Granite such that the ECMB6 weighted mean date of 1781.44 ± 0.51 Ma is a maximum age. The quartz-feldspar porphyry dikes (one of which was sampled as QP1) also intrude the St. Cloud Granite and are parallel to the NED diabase dikes. Neither the quartz-porphyry dikes nor the diabase dikes intrude the younger cross-cutting Richmond Granite such that the ECMB4 weighted mean date of 1776.76 ± 0.49 Ma provides a minimum age for the dikes. Details for the weighted mean dates are given in Table 2 and individual zircon data are in the Supporting Information.

from the St. Cloud Granite that host the dikes (sample ECMB6), the Richmond Granite from which the dikes are absent (sample ECMB4), and from a northeast-trending quartz-feldspar porphyry dike (sample QP1) that is likely coeval with the diabase dikes (Figs. 2 and 7). Zircon crystals were chemically abraded prior to analysis of single zircon grains by ID-TIMS at the Boise State Isotope Geology Laboratory (detailed geochronology methods are provided in the Supporting Information). Weighted mean dates were calculated from multiple single zircon dates (Fig. 7; Table 2). While chemical abrasion served to reduce Pb-loss and resulted in concordant analyses, some grains have persistent Pb-loss and are discordant (Fig. S1). As a result, we calculate weighted mean $^{207}\text{Pb}/^{206}\text{Pb}$ dates rather than $^{206}\text{Pb}/^{238}\text{U}$ dates (Fig. 7; Table 2). These $^{207}\text{Pb}/^{206}\text{Pb}$ dates are 1781.44 ± 0.51 Ma (2σ analytical uncertainty; MSWD = 1.24; $n=8$) for the St. Cloud Granite (ECMB6) and 1776.76 ± 0.49 Ma (MSWD = 1.15; $n=7$) for the Richmond Granite (ECMB4; Fig. 7). The date for the sampled quartz-feldspar porphyry dike (QP1) of 1780.78 ± 0.45 Ma (MSWD = 0.53; $n=8$) is between these two dates as expected on the basis of field relationships. Taking into account the analytical uncertainty on the maximum and minimum age constraints, the diabase dikes are younger than the 1781.44 ± 0.51 Ma St. Cloud

Granite and older than the 1776.76 ± 0.49 Ma Richmond Granite. If one assumes a uniform probability of diabase emplacement timing between the maximum and minimum age constraints that have normally distributed uncertainties, the 95% confidence interval can be estimated through Monte Carlo simulation. Applying this approach gives a mean age of 1779.1 Ma with 95% confidence interval (CI) bounds of 1776.8 to 1781.4 Ma. We can succinctly state the age of the northeast-trending ECMB diabase dikes as being 1779.1 ± 2.3 Ma (95% CI).

5 Thermochronology Methods and Results

While the U-Pb zircon dates constrain the crystallization ages of the ECMB intrusions, additional insight into the thermal history of the batholith can help with interpretation of the paleomagnetic data given that the thermal remanent magnetization of magnetite will be blocked at temperatures below 580°C. As discussed below, existing Ar-Ar dates on hornblende and biotite from the ECMB provide valuable constraints in this regard (Fig. 8). In this study, we also develop new U-Pb apatite dates from three ECMB granites (ECMB1, the Isle Granite; ECMB3, the Rockville Granite; ECMB4, the Richmond Granite). In contrast to zircon, for which the temperatures of appreciable Pb diffusion exceed the liquidus of granite (Cherniak & Watson, 2001), the temperature window for closure of the U-Pb system in apatite is much lower (~ 510 to 460°C ; Smye et al. (2018)). As a result, U-Pb dates of apatite serve as a thermochronometer at moderately-high temperatures (Chamberlain & Bowring, 2001; Schoene & Bowring, 2007; Cochrane et al., 2014). These temperatures are of particular relevance to the interpretation of the paleomagnetic data as they are lower than, or correspond with, the blocking temperature of low-Ti titanomagnetite. If a pluton was emplaced at depths where temperatures exceed the closure temperature of apatite, or if it experienced prolonged reheating, the U-Pb apatite dates would be appreciably younger than the U-Pb zircon crystallization dates.

U-Pb data were developed from apatite grains separated from ECMB granites through laser ablation inductively coupled plasma mass spectrometry (LA-ICP-MS) at UC Santa Barbara (method details are provided in the supporting information). In contrast to zircon, apatite incorporates significant Pb at the time of crystallization. As a result of this elevated common Pb, U-Pb dates were determined through the calculation of Tera-Wasserburg concordia lower intercept dates where the upper intercept corresponds to the ratio of ini-

338 tial $^{207}\text{Pb}/^{206}\text{Pb}$ and the lower intercept is the $^{206}\text{Pb}/^{238}\text{U}$ date (following the method
 339 of Ludwig (1998) as implemented in the IsoplotR software of Vermeesch (2018); Fig. S2).
 340 Sample ECMB1 is from the Isle Granite which has a U-Pb zircon date of 1779 ± 26 Ma
 341 (Holm et al., 2005). The U-Pb apatite date for ECMB1 is $1800.3 \pm 33.4/65.2$ Ma where
 342 the first uncertainty is the 2σ analytical uncertainty and the second uncertainty is 95%
 343 confidence interval for the date that incorporates overdispersion (this uncertainty scheme
 344 will be used for all the presented apatite dates). This U-Pb apatite date is therefore in-
 345 distinguishable from the U-Pb crystallization date (Fig. 8). Sample ECMB3 is from the
 346 Rockville Granite which has a U-Pb zircon date of 1780 ± 7 Ma (Holm et al., 2005). The
 347 U-Pb apatite date for ECMB3 is $1810.5 \pm 16.2/23.0$ Ma. The youngest end of the overdis-
 348 persion uncertainty range is very similar to (albeit just slightly older than) the U-Pb zir-
 349 con date. It is not geologically reasonable for the U-Pb apatite date to be older than the
 350 U-Pb zircon date. This result therefore suggests that the U-Pb apatite date uncertainty
 351 is a slight underestimate with the U-Pb apatite system having closed just after the time
 352 of zircon crystallization as constrained through the U-Pb zircon date. Sample ECMB4
 353 is from the Richmond Granite for which we have developed a new ID-TIMS U-Pb zir-
 354 con date of 1776.76 ± 0.49 Ma. The U-Pb apatite date for ECMB4 is $1751.7 \pm 17.8/36.6$
 355 Ma which is indistinguishable from the U-Pb zircon date (Fig. 8).

356 Pb closure temperatures (T_c) for the analyzed apatite grains can be estimated with
 357 the Dodson (1973) approach assuming a cylindrical geometry with half-widths as the char-
 358 acteristic diffusion length. Apatite grain sizes are similar across the three dated spec-
 359 imens; they typically are 100-200 μm long and 50-75 μm wide. Using the Pb diffusiv-
 360 ity values of Cherniak et al. (1991), a cooling rate of $20^\circ\text{C}/\text{Myr}$ results in closure tem-
 361 peratures of 463°C to 473°C for these grain sizes. A more rapid cooling rate is likely for
 362 the batholith given the similarity of the U-Pb apatite dates with the U-Pb zircon dates.
 363 A cooling rate of $100^\circ\text{C}/\text{Myr}$ from crystallization to apatite closure temperatures results
 364 in closure temperatures of 493°C to 505°C .

365 Overall, these data indicate that the samples cooled through the $\sim 500^\circ\text{C}$ closure
 366 temperatures of the U-Pb apatite system near the time of zircon crystallization consis-
 367 tent with rapid cooling rates of the plutons (Fig. 8). Additionally, there has not been
 368 significant diffusion due to later tectonothermal events. As discussed below, this result
 369 is consistent with Ar-Ar hornblende dates from the ECMB granites and supports the mag-
 370 netite remanence being a primary thermal remanent magnetization.

6 Discussion

6.1 Thermal history of the ECMB and a primary interpretation of the ECMB dike pole

Prior to the emplacement of the ECMB, Paleoproterozoic host rocks were metamorphosed to amphibolite facies during the Penokean orogeny (Holm & Selverstone, 1990). Emplacement of the ECMB has been hypothesized to be post-orogenic and associated with an interval of extensional collapse of the orogen (Holm & Lux, 1996; Boerboom & Holm, 2000). The Al-in-hornblende igneous barometer was applied to the St. Cloud and Isle Granites of the ECMB by Holm, Darrah, and Lux (1998). This barometer has varying published calibrations. Applying the pressure calibration of Mutch et al. (2016) to the data in Holm, Darrah, and Lux (1998) and assuming a 2.7 g/cm^3 overburden gives an estimated emplacement depth of $10.8 \pm 1.7 \text{ km}$ for the Freedhem Granodiorite, $\sim 10.4 \pm 1.7 \text{ km}$ for the Isle Granite and $13.4 \pm 2.1 \text{ km}$ for the St. Cloud Granite. The calibration of Ague (1997) leads to slightly higher calculated pressures implying depths that are $\sim 2.3 \text{ km}$ deeper.

Thermochronology data give additional insight into emplacement temperatures (and thereby depth). Both the Ar-Ar hornblende dates published by Holm et al. (2005) and the U-Pb apatite dates developed in this study from ECMB lithologies are indistinguishable from the crystallization ages of the intrusions (Fig. 8). The closure temperature for Ar in hornblende is ~ 580 to 490°C (Harrison, 1982). The closure of the U-Pb system in the dated apatite grains is ~ 500 to 460°C . The consistency between the U-Pb zircon crystallization and the U-Pb apatite and Ar-Ar hornblende cooling dates indicates that the present-day erosion level of the ECMB was at a shallow enough depth that the crustal temperatures were lower than these closure temperatures at the time of emplacement of the plutons. Geothermal gradients in continental arc settings are typically between 25 to 45°C/km – potentially higher at 1.8 Ga (Rothstein & Manning, 2003). Taking a geothermal gradient of 30°C/km and the closure temperature constraints indicates that the plutons were emplaced at 15 km or shallower in the continental lithosphere. This emplacement depth is consistent with the Al-in-hornblende paleobarometry estimates.

Ar-Ar biotite dates provide insight into even lower temperatures as the system blocks at $\sim 330^\circ\text{C}$ (Grove & Harrison, 1996), well below the blocking temperature of magnetite magnetization in the dikes. Ar-Ar biotite dates from ECMB plutons range from over-

Figure 8. Summary of U-Pb zircon dates, U-Pb apatite dates, Ar-Ar hornblende, and Ar-Ar biotite dates from the East-Central Minnesota batholith (ECMB) from this study, Holm and Lux (1996) and Holm et al. (2005). Approximate closure temperatures associated with the thermochronometers are labeled next to the relevant data. That the U-Pb apatite and Ar-Ar hornblende dates are indistinguishable from the U-Pb zircon crystallization dates indicates that the plutons were emplaced at upper middle to upper crustal levels. The Ar-Ar biotite dates from both the ECMB plutons and older host rock lithologies indicate exhumation within ~ 20 million years to shallower depths and a lack of regional tectonothermal activity over the following 1.75 billion years.

lapping the crystallization dates to being younger by ~ 20 million years (Fig. 8). Ar-Ar biotite dates from older host lithologies to the ECMB are either older than the age of the batholith itself or the same age as the Ar-Ar biotite dates from the batholith (Fig. 8). These data suggest that the batholith was emplaced near the upper depth range estimates from the Al-in-hornblende barometry (~ 10 km) and underwent exhumation to below the $\sim 330^\circ\text{C}$ closure temperature of the K-Ar biotite system soon after emplacement of the plutons. These data also indicate that there has not been reheating or pervasive fluid flow that would have perturbed the Ar-Ar thermochronometers in the granites in the time since initial cooling. The magnetization used to develop the paleomagnetic pole comes from remanence held by low-Ti magnetite that dominantly unblocks between 540 and 560°C (Figs. 3 and 5). The thermochronology results constrain the rocks

to have cooled through the magnetite blocking temperatures at the time that the dikes were emplaced within the batholith and to not have experienced reheating that would have perturbed the thermochronometers. These data support an interpretation that the magnetite remanence within the ECMB dikes is a primary thermal remanent magnetization.

These thermochronology data also demonstrate that following the emplacement of the ECMB there were not regional tectonothermal events with the potential to have thermally modified the magnetization of the magnetite within the ECMB diabase dikes. Subsequent Paleoproterozoic Yavapai and Mazatzal accretion occurred to the southeast of the Spirit Lake Tectonic zone on the other side of the Minnesota River Valley promontory (~ 160 km south of the study region; Fig. 1; Holm et al. (2007)). In contrast to the northeast-trending dikes in the ECMB, northeast-trending dikes in Wisconsin within *ca.* 1840 Ma plutons were metamorphosed to amphibolite facies associated with such accretionary orogenesis (Holm et al., 2019). Mazatzal orogeny deformation and metamorphism occurred *ca.* 1650 to 1630 Ma within the juvenile accreted island arc of the Wisconsin Magmatic Terrane (Holm, Schneider, & Coath, 1998). The region of the ECMB was not affected by these tectonothermal events (Fig. 1, Holm et al. (2005)). Holm et al. (2005) proposed that the voluminous ECMB batholith stabilized the continental lithosphere and prevented the region from being modified during subsequent collisions along the margin. This lack of deformation in the region of the ECMB is further supported by the nearly horizontal bedding of *ca.* 1.63 Ga siliciclastic sedimentary rocks on either side of the batholith (Holm, Schneider, & Coath, 1998; Medaris et al., 2021). In southwestern Minnesota, plutons coeval with the ECMB are overlain by the subhorizontal Sioux Quartzite (Fig. 1) with the correlative Barron Quartzite of northwestern Wisconsin also being undeformed (Southwick et al., 1986). This lack of deformation contrasts with correlative Baraboo quartzite south of the Spirit Lake Tectonic Zone (~ 400 km from the ECMB) that underwent compressional deformation during subsequent orogenesis (Holm, Schneider, & Coath, 1998; Medaris et al., 2021). The Yavapai and Mazatzal terranes were intruded by *ca.* 1470 to 1430 Ma granites of the Eastern Granite Rhyolite Province and there is a sizeable pluton of this age within accreted Penokean rocks in northern Wisconsin (the *ca.* 1470 Ma Wolf River batholith; Dewane and Van Schmus (2007)). However, the thermal effects of the Wolf River batholith (~ 370 km east of the ECMB study area) were limited to a 10-15 km wide contact zone surrounding the intrusion (Holm et al., 2019).

The one major subsequent tectonothermal event in the region for which there is localized evidence in the ECMB is the development of the Midcontinent Rift that initiated *ca.* 1109 Ma and in which magmatic activity continued to *ca.* 1084 Ma (Fig. 1; Fairchild et al. (2017); Swanson-Hysell et al. (2019)). While the main rift axis can be inferred from gravity and aeromagnetic anomaly data to be located ~ 75 km southeast of the study region (Fig. 1), the studied northwest-trending dike has a magnetization direction that implies that it was emplaced during Midcontinent Rift development *ca.* 1096 Ma (Fig. 5). The baked contact test between that dike (NWD1) and the northeast-trending dike that it cross-cuts (NED17), indicates that the thermal effect of the dike emplacement and the Midcontinent Rift in general was localized within the immediate vicinity of that dike (a few meters; Fig. 5). However, this Midcontinent Rift magmatic activity did result in local hydrothermal alteration as evidenced by magnetization held by monoclinic pyrrhotite that is variably present through the ECMB dikes and is in the same direction as the magnetization of the northwest-trending dike (Fig. 3). This chemical remanent magnetization held by monoclinic pyrrhotite likely formed at relatively low temperatures. Phase relationships in the Fe-S system developed through hydrothermal recrystallization experiments show monoclinic pyrrhotite to form at temperatures below 250°C and likely above 75°C (Kissin & Scott, 1982). While in some sites, this pyrrhotite-forming alteration obscured the primary thermal remanence held by magnetite (e.g., NED36 in Fig. 3), in the majority of sites the magnetite remanence direction can be well-resolved (e.g., NED2, NED12 and NED34 in Fig. 3). As a result, the paleomagnetic directions used to calculate the paleomagnetic pole shown in Figure 6 are held by (titano)magnetite that recorded a thermal remanent magnetization upon cooling of the diabase dikes. This evidence for variable late Mesoproterozoic hydrothermal alteration of the dikes provides an explanation for Ar-Ar data developed from two northeast-trending diabase dikes that were reported in Boerboom and Holm (2000). These Ar-Ar data did not yield a plateau age, but give whole rock total gas dates that are late Mesoproterozoic in age. An interpretation that these whole rock total gas ages correspond with the age of emplacement is difficult to reconcile with the cogenetic relationship between the diabase dikes, the quartz-feldspar porphyry dikes and the ECMB granites. K-Ar whole-rock ages of 1570 to 1280 Ma from the dikes reported in Hanson (1968) and discussed in Horan et al. (1987) are attributed to partial resetting. The evidence for fluid flow that led to the formation of pyrrhotite *ca.* 1096 Ma supports the hypothesis put forward by Horan et al. (1987), as

well as by Boerboom and Holm (2000), that there was Mesoproterozoic disruption of the K-Ar isotopic system in the dikes such that the Mesoproterozoic Ar-Ar dates are the result of alteration of dikes which are Paleoproterozoic in age. The field relationships indicating that the northeast-trending diabase dikes are comagmatic with ECMB granites is also consistent with whole rock Pb isotope data that reveal very similar arrays implying a *ca.* 1.8 Ga isochron age for both lithologies (Horan et al., 1987).

Overall, the constraints requiring that the ECMB granites were emplaced at depths where the ambient temperature was below the closure of U-Pb apatite and Ar-Ar hornblende systems indicate that the comagmatic diabase dikes would have acquired their magnetization at the time of emplacement. The lack of significant thermal events that could have reset the magnetite magnetization is indicated by the geologic setting, the thermochronology data (including the Paleoproterozoic Ar-Ar biotite dates), and the positive inverse baked contact test. We therefore interpret the pole calculated from the magnetite remanence of the ECMB diabase dikes as a high-quality constraint on the paleogeographic position of Laurentia at the time the dikes intruded (1779.1 ± 2.3 Ma). The ECMB diabase dikes pole meets six of the seven criteria for the quality criteria Van der Voo (1990) and the Meert et al. (2020) reliability criteria with the only one not satisfied being due to the lack of dual polarity directions. This single polarity normal polarity is consistent with the polarity of the Cleaver Dykes and the proposal of Irving (2004) that there was a normal geomagnetic superchron that followed the Trans-Hudson orogeny.

6.2 Laurentia's paleomagnetic poles following the Trans-Hudson orogeny

The Trans-Hudson orogeny is a major event in the formation of Laurentia resulting from collision between the Superior province and the Churchill plate consisting of the composite Slave + Rae + Hearne provinces (Hoffman, 1988; Corrigan et al., 2009; Weller & St-Onge, 2017). Geologic data on the timing of Trans-Hudson orogenesis include a $^{206}\text{Pb}/^{238}\text{U}$ date of 1854.2 ± 1.6 Ma from the base of a foredeep sedimentary succession on the northern margin of the East Superior province that constrains flexural subsidence associated with Trans-Hudson orogenesis to have initiated at that time (Hodgskiss et al., 2019). This timing of orogenesis is consistent with $^{207}\text{Pb}/^{206}\text{Pb}$ dates of monazite within garnet of Trans-Hudson orogen eclogites for which a mean date of 1831 ± 5 Ma has been interpreted to record peak metamorphism (Weller & St-Onge, 2017). A similar timing of *ca.* 1860 to 1820 Ma peak Trans-Hudson metamorphism resulting from col-

Figure 9. Paleogeographic reconstructions at five different times in the Paleoproterozoic and the position of the provinces at present. Paleomagnetic poles within 20 Myr of the given time (10 Myr for 1888 and 1868 Ma) are shown from the compilation of Evans et al. (2021) as well as the new ECMB pole. These data illustrate differential plate motion between the Superior and Slave Provinces that is required by the data leading up to the closure of the Manikewan Ocean and the assembly of Laurentia during the Trans-Hudson orogeny. The ECMB pole is consistent with an assembled Laurentia following the Trans-Hudson orogeny which contrasts with the disparate orientations and paleolatitudes between Laurentia's constituent provinces prior to the orogeny.

lisional orogenesis has been interpreted from U-Pb zircon rim and monazite dates from the orogen in Baffin Island, Nunavut, Canada (Skipton et al., 2016). These geologic data strongly support that the Superior Province was conjoined with the Slave + Rae + Hearne provinces prior to 1800 Ma in their present-day relative positions (Fig. 1). There are high-quality paleomagnetic poles for the Superior province during the time interval when the Manikewan Ocean was closing leading up to the Trans-Hudson orogeny (the *ca.* 1880 Molson dykes pole and the *ca.* 1870 Ma Haig/Flaherty pole; Fig. 9). There are no Paleoproterozoic paleomagnetic poles for the Superior Province during or after the Trans-Hudson orogeny that can test Laurentia’s coherency with paleomagnetic data. The new 1779.1 ± 2.3 Ma ECMB paleomagnetic pole fills this gap.

While abundant paleomagnetic data have been developed from rocks within the Trans-Hudson orogen (Symons & Harris, 2005), both the primary nature of the remanence directions as well as the age of remanence acquisition has been challenging to establish. As a result, relatively few poles from this interval during and following Laurentia’s amalgamation have been included in curated pole compilations such as that compiled by the Nordic paleogeography workshops (Evans et al., 2021). The best constrained of these poles comes from the post-orogenic 1740 ± 5 Ma Cleaver Dykes of the Great Bear Magmatic Arc on the western margin of the Slave Province (Fig. 1; Irving (2004)). The age of these dikes is constrained by a U-Pb baddeleyite date and there is a positive baked contact test supporting the interpretation that the pole is primary. As can be seen in Figure 6, the position of this *ca.* 1740 Ma Cleaver Dykes pole for Laurentia (Irving, 2004) is similar to the new pole for the *ca.* 1780 Ma ECMB diabase dikes (Fig. 6). They do not share a common mean (as determined through Watson and bootstrap common mean tests; Tauxe et al. (2016)), but are within 11° of one another (less when considering uncertainty). The similarity in these pole positions provides an independent test of the coherency of the Laurentia craton at that time.

An additional pole for comparison comes from hematite remanence of volcanics and sediments of the Baker Lake Group of the Dubawnt Supergroup which were deposited in a basin atop the suture between the Rae and Hearne provinces (Figs. 1 and 6; Park et al. (1973)). Both the depositional age of the succession and the age of hematite remanence are roughly constrained leading to a wide assigned age range of 1820 to 1750 Ma in Evans et al. (2021) for the pole. While the loose age constraints hinder firm comparisons, the broad similarity of this pole with the Cleaver Dykes pole of the Slave Province

as well as that for Martin Formation (1818 ± 4 Ma) and the Sparrow dikes (1827 ± 4 Ma) of the Rae Province establishes a largely consistent position of the Churchill plate from *ca.* 1820 to 1740 Ma.

A less robust pole for comparison with the new ECMB pole comes from the post-orogenic Jan Lake Granite within the Trans-Hudson orogen in southeast Saskatchewan developed in Gala et al. (1995). A U-Pb zircon date of 1758 ± 1 Ma for this granite was reported in Bickford et al. (2005). Directional data from the Jan Lake Granite falls into two groups. Based on thermal demagnetization behavior, Irving (2004) interpreted the ‘A’ grouping to be a TRM held by magnetite that was acquired at the time of the emplacement of the intrusion *ca.* 1758 Ma. The ECMB dikes pole shares a common mean with this Jan Lake Granite pole with overlapping A_{95} confidence circles (Fig. 6). This result is consistent with both the Jan Lake Granite and ECMB being post-orogenic magmatic events that occurred following the amalgamation of Laurentia. This similarity suggests that despite the large uncertainty on the Jan Lake Granite pole and the ambiguity resulting from multiple directional groups that the pole does constrain the position of Laurentia *ca.* 1758 Ma.

A pole that does not hold up to such comparative scrutiny is that developed for the Deschambault Pegmatites from within the Trans-Hudson orogen (Fig. 6C; Symons et al. (2000)). This pole has been interpreted to constrain the position of Laurentia *ca.* 1766 Ma — an age based on U-Pb monazite and allanite dates of other pegmatites in the region. As noted in D’Agrella-Filho et al. (2020), there are no field tests for this pole and the remanence directions from which it is calculated are quite close to the modern geomagnetic field. This pole was included in the curated Nordic paleogeography workshop compilation of Evans et al. (2021) less because of the quality of the individual pole, but rather because there are a number of poles of similar position to this one from the Trans-Hudson orogen (Fig. 6). Many of these poles have individual VGPs that are streaked between directions similar to the Jan Lake Granite and that of the present-local field. The direction of the Deschambault pole is far from the new *ca.* 1780 Ma ECMB dikes pole. In contrast, the similarity of pole position between the new ECMB dikes pole and the *ca.* 1758 Ma Jan Lake Granite pole as well as the *ca.* 1740 Ma Cleaver Dykes pole supports that these poles, rather than the Deschambault Pegmatites pole, constrain Laurentia’s position during this interval. This Deschambault Pegmatites pole played a role, in conjunction with other poles from the Trans-Hudson orogen, such as that from the

Wapisi gneiss and the Deschambault Post pluton, in an interpretation that Laurentia's pole path was at a standstill in the vicinity of the Deschambault Pegmatites pole from *ca.* 1800 Ma through to *ca.* 1766 Ma (Symons et al., 2000; Symons & Harris, 2005). As reviewed in Raub (2008), there are numerous difficulties in interpreting these data from the Trans-Hudson orogen as useful constraints including: 1) a lack of field tests; 2) uncertainty in the timescale of cooling and the timing of the acquisition of magnetization in these slowly cooled units; 3) poorly constrained tilt corrections and 4) large secondary viscous remanent magnetizations that are prevalent due to the coarse grain-size of the igneous lithologies. The preferred interpretation of Raub (2008), which is echoed in D'Agrella-Filho et al. (2020), is that there is unresolved component mixing between primary directions (which would be in the vicinity of the Jan Lake Granite A Group pole and our new ECMB pole) and the present-day north pole as the result of unresolved viscous overprints (Fig. 6). This component mixing leads to streaked site mean directions in individual studies as well as the database of Trans-Hudson orogen poles including the Deschambault Pegmatites pole (Fig. 6; Raub (2008)). The new ECMB pole significantly strengthens this interpretation by demonstrating that a proposed northerly apparent polar wander path to satisfy the Deschambault Pegmatites pole and other Trans-Hudson orogen poles streaked between the Jan Lake Granite and the modern-day pole is indeed fictitious (Fig. 6). Instead, the ECMB pole and the Cleaver Dykes pole establish the paleogeographic position of Laurentia to have been consistent *ca.* 1780 to 1740 Ma (Fig. 9).

6.3 The paleogeography of Laurentia

As is expected by the geologic record of Trans-Hudson orogenesis, the similarity in pole positions from the southeastern margin of the Superior Provinces (the new ECMB pole) and the northwestern margin of the Slave Province (the Cleaver Dykes pole; Fig. 6) indicate a coherent assembled Laurentia following 1.8 Ga (Fig. 9). The coherency of the record of high-quality paleomagnetic poles at this time when the geologic record indicates a recently assembled Laurentia increases confidence that differing pole positions between Laurentia's Archean provinces earlier in the Paleoproterozoic are indeed a record of differential plate tectonic motion (Fig. 9). There is a particularly rich record of paleomagnetic poles from the Archean Superior and Slave provinces that can be paired between 2.23 and 1.89 Ga that constrain the provinces to not be in their modern relative

orientation and to be undergoing differential motion (Mitchell et al., 2014; Buchan et al., 2016; Swanson-Hysell, 2021). These poles result in reconstructions where prior to the Trans-Hudson orogeny there was an ocean basin between the Superior province and the Hearne + Rae + Slave provinces known as the Manikewan Ocean (Fig. 9; Stauffer (1984)). The poles are consistent with the Superior Province approaching the joint Slave + Hearne + Rae provinces prior to the onset of the Trans-Hudson orogeny (Fig. 9). These data provide strong evidence for mobile lid plate tectonics from 2.23 Ga onward (Mitchell et al., 2014; Buchan et al., 2016; Swanson-Hysell, 2021).

The orogenesis associated with Laurentia’s assembly is hypothesized to have resulted in the formation of the supercontinent, or semi-supercontinent, Nuna (Hoffman, 1997; Evans & Mitchell, 2011; Evans et al., 2016). Given that Laurentia is the largest craton hypothesized to have been part of this supercontinent, its paleogeographic position is key to reconstructions of Nuna. The new ECMB pole provides higher confidence in the paleogeographic position of Laurentia in the time just following its formation from the collision of constituent Archean provinces (Fig. 9). This new pole can be used to evaluate hypothesized connections between Laurentia and other cratons. There is an increasingly rich global database of paleomagnetic poles *ca.* 1780 Ma including poles from the Amazonia, Baltica, India, Rio de la Plata, São Francisco and North China cratons (Zhang et al., 2012; Xu et al., 2014; Bispo-Santos et al., 2014; Shankar et al., 2018; D’Agrella-Filho et al., 2020).

One hypothesized connection of particular interest is that with Baltica. The two cratons have been hypothesized to have been conjoined such that they shared a margin with a long-lived history of accretionary orogenesis (Hoffman, 1997; Karlstrom et al., 2001). The proposed NENA (northern Europe and North America) configuration between Laurentia and Baltica allows for such a shared margin (Gower et al., 1990; Buchan et al., 2000; Evans & Pisarevsky, 2008). The increased concordance between *ca.* 1780 to 1750 paleomagnetic poles from Laurentia and Baltica upon the NENA rotation of Baltica can be seen in Figure 10. Paleomagnetic poles support a continued NENA connection until at least 1260 Ma (supported by poles from Baltica’s *ca.* 1258 Ma post-Jotnian intrusions and Laurentia’s *ca.* 1267 Ma Mackenzie dikes) and perhaps to 1120 Ma (where the paleomagnetic comparison is reliant on the *ca.* 1122 Ma Salla dike of Baltica developed from a single cooling unit; Salminen et al. (2009)). This connection supports the long-lived active margin where Laurentia grew through the rest of the Paleoproterozoic and

Figure 10. A) Paleomagnetic poles for Laurentia and Baltica between 1780 and 1750 Ma with the Baltica poles shown with and without the NENA rotation (Euler pole for Baltica of $[47.5^\circ, 001.5^\circ, +49.0^\circ]$ as in Evans and Pisarevsky (2008)). Baltica is shown in its present day location in dark red and shown in the NENA position in light red. B) Same data as in panel A shown with a different center of projection which allows for easier visualization of the reconstructed position. C,D) Comparison between poles between 1780 and 1110 Ma between Baltica and Laurentia without (C) and with (D) the NENA rotation. The poles that are shown are those from the Nordic compilation with 'A' and 'B' grades as well as the new ECMB pole from this study.

643 through the Mesoproterozoic until the *ca.* 1.08 Ga continent-continent collision of the
 644 Grenvillian orogeny (Whitmeyer & Karlstrom, 2007).

7 Conclusions

The East-Central Minnesota Batholith was emplaced following Penokean orogenesis on the southeast margin of the Superior Province. While the southeast margin of Laurentia experienced subsequent intervals of accretionary orogenesis, thermochronology data constrain the batholith to have a straight-forward history of post-emplacment rapid exhumation without substantial reheating. Subsequent orogenesis occurred well southeast of the batholith — consistent with the batholith having played a role in stabilizing Laurentian lithosphere. Comagmatic diabase dikes of the East-Central Minnesota Batholith can be constrained through U-Pb geochronology on the felsic units to have been emplaced at 1779.1 ± 2.3 Ma. A new paleomagnetic pole developed from the magnetite remanence of these dikes provides a high-quality constraint on the position of Laurentia following Trans-Hudson orogenesis. This pole confirms the coherency of an amalgamated Laurentia at the time and supports the NENA connection with Baltica. This paleomagnetic coherency further strengthens the case that previously disparate pole positions between the Superior and Slave provinces are the result of *ca.* 2.2 to 1.8 Ga mobile-lid plate tectonics. The geologic and paleomagnetic record of Laurentia is inconsistent with a stagnant-lid regime anytime over the past 2.2 billion years.

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is also archived on Zenodo (<https://doi.org/10.5281/zenodo.4625041>). This repository also contains Python code related to calculations, visualizations and statistical tests discussed herein.

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Table 1. Summary of site level paleomagnetic data

site	site lat	site lon	n	dec	inc	k	R	α_{95}	VGP lat	VGP lon
<i>northeast-trending dike magnetite-component site means</i>										
NED1	45.53423	265.75804	8	157.6	74.7	380	7.98	2.8	18.4	276.9
NED2	45.53421	265.75816	8	172.3	74.9	420	7.98	2.7	17.4	269.6
NED5	45.53309	265.75803	6	170.3	79.0	213	5.98	4.6	24.5	269.6
NED6	45.53299	265.75773	9	197.8	73.6	183	8.96	3.8	16.1	256.5
NED7	45.53288	265.75767	5	266.3	81.3	204	4.98	5.4	42.0	242.6
NED8	45.53286	265.75782	7	202.6	78.6	423	6.99	2.9	24.8	256.6
NED9	45.53314	265.75855	7	191.1	71.5	90	6.93	6.4	12.2	259.5
NED10	45.53259	265.75742	6	199.7	73.3	324	5.98	3.7	15.8	255.4
NED11	45.53252	265.75768	7	166.1	75.1	391	6.98	3.1	18.1	272.6
NED12	45.53489	265.76076	6	179.5	73.0	374	5.99	3.5	14.1	266.0
NED13	45.53497	265.76113	7	169.1	73.9	199	6.97	4.3	15.9	271.4
NED14	45.53492	265.76119	6	193.0	70.0	217	5.98	4.6	10.1	258.0
NED15	45.53688	265.76758	8	175.8	77.7	604	7.99	2.3	22.0	267.6
NED16	45.53728	265.76822	5	185.1	78.6	449	4.99	3.6	23.6	263.7
NED18	45.53124	265.76945	6	134.7	81.7	159	5.97	5.3	33.2	279.5
NED23	45.53398	265.74200	7	193.0	76.4	411	6.99	3.0	20.2	259.7
NED25	45.53396	265.74119	4	135.4	80.7	334	3.99	5.0	31.5	280.6
NED26	45.53445	265.74129	5	151.1	73.4	572	4.99	3.2	17.4	280.8
NED28	45.53467	265.73817	4	157.8	73.8	145	3.98	7.7	16.9	277.2
NED29	45.53438	265.73690	8	219.0	76.7	119	7.94	5.1	24.4	248.6
NED31	45.53385	265.75785	3	184.7	69.4	1730	3.00	3.0	8.7	262.9
NED34	45.51700	265.78083	8	185.9	77.0	362	7.98	2.9	20.8	263.1
NED35	45.53320	265.75761	8	166.4	74.7	261	7.97	3.4	17.4	272.6
mean pole: pole longitude: 265.8; pole latitude: 20.4; A_{95}: 4.5; K: 45.6 N: 23										
<i>northwest-trending dike magnetite-component site mean</i>										
NWD1	45.53407	265.76852	9	293.4	41.6	66	8.88	6.4	32.9	177.5

Notes: site lat–latitude of site ($^{\circ}$; WGS84); site lon–longitude of site ($^{\circ}$; WGS84) n–number of samples analyzed and included in the site mean; dec–tilt-correction mean declination for the site; inc–tilt-correction mean inclination for the site; k–Fisher precision parameter; R–resultant vector length; α_{95} –95% confidence limit in degrees; VGP lat–latitude of the virtual geomagnetic pole for the site; VGP lon–longitude of the virtual geomagnetic pole for the site.

Table 2. Summary of ID-TIMS $^{207}\text{Pb}/^{206}\text{Pb}$ East-Central Minnesota Batholith zircon dates

Sample	Unit	Latitude	$^{207}\text{Pb}/^{206}\text{Pb}$	Uncertainty (2σ)		MSWD	n/N
		Longitude	date (Ma)	X	Z		
ECMB6	St. Cloud Granite	45.53396° N	1781.44	0.51	2.4	1.24	8/8
		94.23187° W					
QP1	quartz-feldspar	45.53481° N	1780.78	0.45	2.4	0.53	8/8
	porphyry dike	94.25811° W					
ECMB4	Richmond Gran-	45.44343° N	1776.76	0.49	2.4	1.15	7/8
	ite	94.48360° W					

Notes: X is 2σ analytical uncertainty; Z is 2σ uncertainty including decay constant uncertainty. This Z uncertainty needs to be utilized when comparing to dates using other decay systems (e.g., $^{40}\text{Ar}/^{39}\text{Ar}$, ^{187}Re - ^{187}Os); MSWD is the mean squared weighted deviation; n is the number of individual zircon dates included in the calculated sample mean date; N is the number of individual zircons analyzed for the sample.

Supporting Information for “The paleogeography of Laurentia in its early years: new constraints from the Paleoproterozoic East Central Minnesota batholith”

Swanson-Hysell, N. L., Avery, M. S., Zhang, Y., Hodgins, E. B., Sherwood, R. J., Apen, F. E., Boerboom, T. J., Keller, C. B., and Cottle, J. M. (2021), The paleogeography of Laurentia in its early years: new constraints from the Paleoproterozoic East Central Minnesota batholith *Tectonics*.

ID-TIMS U-Pb zircon geochronology methods

U-Pb dates were obtained by chemical abrasion isotope dilution thermal ionization mass spectrometry (ID-TIMS) in the Boise State University (BSU) Isotope Geology Laboratory (Table DR1; Fig. SI2). Chemical abrasion of single zircon grains was modified after Mattinson (2005). Zircons were separated from rocks using standard techniques, annealed in a muffle furnace at 900°C for 60 hours in quartz crucibles, and imaged by cathodoluminescence in grain mounts (Fig. SI1).

QP1 – Quartz Porphyry Dike

ECMB4 – Richmond Granite

ECMB6 – St. Cloud Granite

Figure SI1. Cathodoluminescence (CL) images of the zircons dated by ID-TIMS. The 100 μm scale bars applies for all imaged grains in a given sample.

Individual zircons were removed from grain mounts and chemically abraded. Chemical abrasion was carried out by transferring zircons to 3 ml Teflon Perfluoroalkoxy alkane (PFA) beakers in which they were rinsed in 3.5 M HNO_3 and ultrapure H_2O prior to loading into 300 μl Teflon PFA microcapsules. Fifteen microcapsules were placed in a large-capacity Parr vessel and the zircon partially dissolved in 120 μl of 29 M HF for 12 hours at 190°C. Zircons were returned to 3 ml Teflon PFA beakers, HF was removed, and zircons were immersed in 3.5 M HNO_3 , ultrasonically cleaned for an hour, and fluxed on a hotplate at 80°C for an hour. The HNO_3 was removed and zircon was rinsed twice in ultrapure H_2O before being reloaded into the 300 μl Teflon PFA microcapsules (rinsed and fluxed in 6 M HCl during sonication and washing of the zircons) and spiked with the ^{233}U - ^{235}U - ^{205}Pb BSU tracer solution (BSU1B). Zircons were dissolved in Parr vessels in 120 μl of 29 M HF at 220°C for 48 hours, dried to fluorides, and re-dissolved in 6 M HCl at 180°C overnight. Pb and U were separated from the zircon matrix using an HCl-based anion-exchange chromatographic procedure (Krogh, 1973), eluted together and dried with 2 μl of 0.05 N H_3PO_4 .

Pb and U were loaded on a single outgassed Re filament in 5 μl of a silica-gel/phosphoric acid mixture (Gerstenberger and Haase, 1997), and Pb and U isotopic measurements made on a GV Isoprobe-T multicollector thermal ionization mass spectrometer equipped with an ion-counting Daly detector. Pb isotopes were measured by peak-jumping all isotopes on the Daly detector for 190 cycles with a mass bias correction of $0.16 \pm 0.03\%$ /a.m.u. (1σ). Transitory isobaric interferences due to high-molecular weight organics, particularly on ^{204}Pb and ^{207}Pb , disappeared within 30-45 cycles, while ionization efficiency averaged 104 cps/pg of each Pb isotope. Linearity (to $\geq 1.4 \times 10^6$ cps) and the associated deadtime correction of the Daly detector were determined by analysis of NBS982. Uranium was analyzed as UO_2^+ ions in static Faraday mode on 10^{12} ohm resistors for up to 300 cycles, and corrected for isobaric interference of $^{233}\text{U}^{18}\text{O}^{16}\text{O}$ on $^{235}\text{U}^{16}\text{O}^{16}\text{O}$ with an $^{18}\text{O}/^{16}\text{O}$ of 0.00206. Ionization efficiency averaged 20 mV/ng of each U isotope. U mass fractionation was corrected using the $^{233}\text{U}/^{235}\text{U}$ ratio of the BSU1B tracer.

Figure SI2. U-Pb concordia plots for the new zircon dates. Ellipses represent 2σ analytical uncertainty on individual zircon dates. The weighted mean dates are shown with X/Y uncertainty where X is 2σ analytical uncertainty and Z is 2σ uncertainty including decay constant uncertainty.

ID-TIMS U-Pb dates and uncertainties were calculated using the algorithms of Schmitz and Schoene (2007), BSU1B tracer solution with calibration of $^{235}\text{U}/^{205}\text{Pb} = 77.93$ and $^{233}\text{U}/^{235}\text{U} = 1.007066$, and U decay constants recommended by Hiess et al. (2012), including $^{238}\text{U}/^{235}\text{U}$ of

137.818. $^{206}\text{Pb}/^{238}\text{U}$ ratios and dates were corrected for initial ^{230}Th disequilibrium using $\text{DTh}/\text{U} = 0.20 \pm 0.05$ (1σ). All common Pb in analyses was attributed to laboratory blank and subtracted based on the measured laboratory Pb isotopic composition and associated uncertainty. U blanks are estimated at 0.013 pg. ID-TIMS weighted mean $^{207}\text{Pb}/^{206}\text{Pb}$ dates were calculated from equivalent dates ($\text{pof} > 0.05$) using Isoplot 3.0 (Ludwig, 2003). Errors on the weighted mean dates are given as $\pm x / z$, where x is the internal error based on analytical uncertainties including counting statistics, subtraction of tracer solution, and blank and initial common Pb subtraction; z also includes the U decay constant uncertainties propagated in quadrature. Dates from individual zircon fractions and weighted mean dates are reported at 2σ .

LA-ICP-MS U-Pb apatite geochronology methods

Apatite U-Pb geochronology data were developed via laser ablation inductively coupled mass spectrometry (LA-ICP-MS) at UC Santa Barbara. U-Pb isotopes were analyzed with a Cetac Teledyne 193 nm excimer Analyte laser with a HelEx ablation cell coupled to a Nu Instruments Plasma 3D multi-collector (MC) ICP-MS. On the Plasma 3D, $^{204}(\text{Pb}+\text{Hg})$, ^{206}Pb , ^{207}Pb , and ^{208}Pb were measured on Daly detectors and ^{238}U and ^{232}Th were measured on Faraday collectors. Apatite was ablated with a 40 μm diameter laser spot for 60 pulses fired at a 4 Hz repetition rate and 50% of 5 mJ laser power. Each ablation sequence consisted of 2 cleaning shots, followed by 25 secs of monitored washout and 15 secs of ablation. The Iolite v. 2.5 program (Paton et al., 2011) in the Igor Pro software environment was used to correct the raw U-Pb ratios for baselines, laser- and plasma-induced element fractionation, and instrument drift.

Figure SI3. U-Pb Tera-Wasserburg plots for the new apatite dates. Ellipses represent 2σ analytical uncertainty on individual apatite dates. The

Multiple apatite reference materials (RMs) were analyzed throughout the analytical session to monitor data quality. Apatite RM Madagascar (478.4 ± 6.1 Ma, ID-TIMS age; Schmitz, 2020 personal communication) served as the primary bracketing standard with RMs McClure (523.5 ± 1.5 Ma, ID-TIMS age; Schoene and Bowring, 2006), 401 (530.3 ± 1.5 Ma, ID-ICP-MS age; Thompson et al., 2016) and OD306 (1596.7 ± 7.1 Ma, ID-ICP-MS age; Thompson et al., 2016) analyzed as secondary standards to assess precision and accuracy. All of the secondary standards form mixing arrays between common-Pb and an age intercept. The excess variances of

$^{206}\text{Pb}/^{238}\text{U}$ and $^{207}\text{Pb}/^{206}\text{Pb}$ required for secondary standards to conform to a statistically single mixing line are 2.1% and 2.0% (2σ), respectively; these values were added in quadrature to the internal uncertainties of each U-Pb datum (Horstwood et al., 2016). The ages of the secondary standards with these designated uncertainties are $547.4 \pm 30.4/35.9$ Ma (McClure), $512.8 \pm 9.8/10.5$ Ma (401), $1586.5 \pm 14.9/15.7$ Ma (OD306) (uncertainties following the same format used in main text), within uncertainty of their published ages. Although mass-204 was measured on the MC-ICP-MS, isobaric interferences with ^{204}Hg in the He carrier gas preclude the use of the ^{204}Pb method for common-Pb corrections. The U-Pb apatite data are reported in Table S2.

Table S1: Zircon U-Th-Pb isotopic data

Sample (a)	Compositional Parameters					Radiogenic Isotope Ratios								Isotopic Ages (Ma)						Weighted Mean (Ma)	
	$\frac{^{238}\text{Th}}{\text{U}}$	$\frac{^{206}\text{Pb}^*}{^{206}\text{Pb}}$	mol %	$\frac{^{207}\text{Pb}^*}{^{207}\text{Pb}}$	$\frac{^{206}\text{Pb}}{(\text{pg})}$	$\frac{^{208}\text{Pb}}{^{208}\text{Pb}}$	$\frac{^{207}\text{Pb}}{^{207}\text{Pb}}$	$\frac{^{206}\text{Pb}}{^{206}\text{Pb}}$	% err	$\frac{^{207}\text{Pb}}{^{235}\text{U}}$	% err	$\frac{^{206}\text{Pb}}{^{238}\text{U}}$	% err	corr. coef.	$\frac{^{207}\text{Pb}}{^{206}\text{Pb}}$	$\pm$	$\frac{^{206}\text{Pb}}{^{235}\text{U}}$	$\pm$	$\frac{^{206}\text{Pb}}{^{238}\text{U}}$	$\pm$	$\frac{^{207}\text{Pb}}{^{206}\text{Pb}}$
	(b)	(c)	(c)	(c)	(c)	(d)	(e)	(e)	(f)	(e)	(f)	(e)	(f)		(g)	(f)	(g)	(f)	(g)	(f)	(h)
ECMB4																					
z1	0.459	1.7560	99.84%	189	0.24	11080	0.134	0.108748	0.071	4.747907	0.150	0.316793	0.096	0.927	1777.72	1.29	1775.74	1.26	1774.05	1.50	1776.76 ± 0.49 [2.40] MSWD = 1.15 POF = 0.33
z2	0.538	2.9136	99.88%	258	0.29	14816	0.157	0.108642	0.065	4.746500	0.131	0.317007	0.072	0.961	1775.96	1.19	1775.49	1.10	1775.09	1.12	
z3	0.482	0.8672	99.78%	138	0.16	8032	0.141	0.108731	0.088	4.737140	0.177	0.316124	0.120	0.896	1777.44	1.60	1773.83	1.49	1770.77	1.85	
z4	0.468	1.1223	99.79%	144	0.20	8395	0.137	0.108707	0.076	4.742804	0.168	0.316572	0.115	0.921	1777.04	1.39	1774.84	1.41	1772.97	1.78	
z5	0.572	2.6740	99.84%	201	0.35	11478	0.167	0.108678	0.068	4.730009	0.135	0.315801	0.074	0.952	1776.56	1.25	1772.57	1.13	1769.19	1.14	
z6	0.467	3.2913	99.82%	173	0.49	10119	0.136	0.108710	0.067	4.749514	0.134	0.317009	0.074	0.956	1777.10	1.22	1776.02	1.12	1775.11	1.15	
z7	0.446	3.8518	99.92%	381	0.26	22347	0.130	0.108530	0.064	4.711602	0.130	0.315002	0.072	0.959	1774.07	1.17	1769.30	1.09	1765.27	1.11	
z8	0.506	5.8883	99.75%	119	1.28	6606	0.148	0.108644	0.067	4.746793	0.132	0.317021	0.070	0.959	1775.98	1.23	1775.54	1.10	1775.17	1.09	
QPI																					
z1	0.359	9.6958	99.96%	768	0.32	45979	0.105	0.108903	0.061	4.774622	0.126	0.318120	0.070	0.967	1780.33	1.12	1780.45	1.06	1780.54	1.08	1780.78 ± 0.45 [2.39] MSWD = 0.53 POF = 0.82
z2	0.401	4.9637	98.54%	21	6.10	1236	0.117	0.108932	0.149	4.775515	0.206	0.318096	0.073	0.847	1780.81	2.72	1780.60	1.73	1780.43	1.14	
z3	0.423	6.4130	99.89%	269	0.61	15865	0.123	0.108960	0.064	4.772557	0.128	0.317818	0.069	0.967	1781.28	1.16	1780.08	1.07	1779.06	1.07	
z4	0.336	2.7906	99.91%	325	0.21	19550	0.098	0.108956	0.064	4.771584	0.131	0.317764	0.073	0.963	1781.22	1.17	1779.91	1.10	1778.80	1.13	
z5	0.366	2.5302	99.85%	198	0.32	11825	0.107	0.108944	0.068	4.775083	0.136	0.318032	0.077	0.944	1781.02	1.24	1780.53	1.14	1780.11	1.19	
z6	0.360	2.8906	99.92%	401	0.18	23996	0.105	0.108922	0.066	4.772373	0.132	0.317916	0.074	0.951	1780.65	1.20	1780.05	1.11	1779.54	1.15	
z7	0.451	5.0856	99.69%	100	1.30	5857	0.131	0.108882	0.072	4.771395	0.135	0.317968	0.070	0.952	1779.98	1.31	1779.88	1.13	1779.80	1.08	
z8	0.455	2.5040	99.63%	84	0.76	4943	0.133	0.108938	0.076	4.756832	0.140	0.316833	0.072	0.942	1780.92	1.39	1777.31	1.18	1774.25	1.12	
ECMB6																					
z1	0.284	1.213	0.990	30.870	0.97	1899.098	0.083	0.108982	0.123	4.778311	0.211	0.318136	0.135	0.834	1781.66	2.25	1781.10	1.77	1780.62	2.11	1781.44 ± 0.51 [2.40] MSWD = 1.24 POF = 0.28
z2	0.335	4.276	0.999	226.896	0.47	13671.762	0.098	0.108953	0.066	4.771689	0.130	0.317780	0.070	0.958	1781.17	1.20	1779.93	1.09	1778.88	1.10	
z3	0.376	2.430	0.998	155.473	0.39	9279.850	0.110	0.109024	0.086	4.774201	0.156	0.317741	0.096	0.872	1782.35	1.57	1780.37	1.31	1778.69	1.48	
z4	0.476	1.918	0.998	191.621	0.26	11163.360	0.139	0.109071	0.098	4.775361	0.155	0.317681	0.084	0.822	1783.14	1.80	1780.58	1.30	1778.39	1.30	
z5	0.315	2.478	0.999	228.261	0.27	13819.538	0.092	0.108954	0.069	4.777931	0.134	0.318193	0.072	0.953	1781.18	1.25	1781.03	1.13	1780.90	1.13	
z6	0.332	1.064	0.998	155.064	0.17	9355.845	0.097	0.109010	0.084	4.755533	0.160	0.316539	0.101	0.893	1782.11	1.52	1777.08	1.35	1772.81	1.57	
z7	0.342	2.990	0.999	215.887	0.35	12986.019	0.100	0.108930	0.069	4.768270	0.133	0.317621	0.071	0.946	1780.77	1.26	1779.33	1.11	1778.10	1.11	
z8	0.368	2.149	0.998	126.819	0.43	7588.570	0.107	0.108921	0.074	4.768637	0.141	0.317671	0.079	0.931	1780.62	1.35	1779.39	1.18	1778.35	1.22	

(a) z1, z2 etc. are labels for single zircon grains or fragments annealed and chemically abraded after Mattinson (2005). Bold indicates z fraction included in weighted mean.

(b) Model Th/U ratio iteratively calculated from the radiogenic $^{208}\text{Pb}/^{206}\text{Pb}$ ratio and $^{206}\text{Pb}/^{238}\text{U}$ age.(c) Pb* and Pb_c represent radiogenic and common Pb, respectively; mol % $^{206}\text{Pb}^*$ with respect to radiogenic, blank and initial common Pb.(d) Measured ratio corrected for spike and fractionation only. Fractionation estimated at 0.16 ± 0.03 ‰ a.m.u. (1 sigma) for Daly analyses, based on analyses of EARTHTIME 202-205 trace solution run recently.(e) Corrected for fractionation, spike, and common Pb; all common Pb was assumed to be procedural blank: $^{206}\text{Pb}/^{204}\text{Pb} = 18.042 \pm 0.61\%$; $^{207}\text{Pb}/^{204}\text{Pb} = 15.537 \pm 0.52\%$; $^{208}\text{Pb}/^{204}\text{Pb} = 37.686 \pm 0.63\%$ (all uncertainties 1-sigma).

(f) Errors are 2 sigma, propagated using the algorithms of Schmitz and Schoene (2007).

(g) Calculations are based on the decay constants of Jaffey et al. (1971) and Hiess et al. (2012). $^{206}\text{Pb}/^{238}\text{U}$ and $^{207}\text{Pb}/^{206}\text{Pb}$ ages corrected for initial disequilibrium in $^{230}\text{Th}/^{232}\text{Th}$ using DTh/U [magma] = 0.20 ± 0.05 (1 sigma).(h) Weighted mean $\pm 2\sigma$ internal uncertainty [$\pm 2\sigma$ internal + decay constant uncertainties]. MSWD = Mean Standard Weighted Deviation. POF = Probability of Fit

Table S2: Apatite U-Pb data

Spot name	U (ppm)	Th (ppm)	²⁰⁷ Pb ²³⁵ U	2 SE (%)	²⁰⁶ Pb ²³⁸ U	2 SE (%)	rho	²³⁸ U ²⁰⁶ Pb	2 SE (%)	²⁰⁷ Pb ²⁰⁶ Pb	2 SE (%)
ECMB1											
ECMB1_1	49.6	28.3	10.6010	4.4	0.3734	3.1	0.71	2.6781	3.1	0.2060	3.1
ECMB1_2	54.4	24.3	9.7938	4.4	0.3707	3.1	0.70	2.6976	3.1	0.1917	3.1
ECMB1_3	33.1	28.4	12.6190	4.5	0.3704	3.3	0.73	2.6998	3.3	0.2472	3.1
ECMB1_4	44.1	12.5	10.8580	4.6	0.3634	3.3	0.71	2.7518	3.3	0.2168	3.2
ECMB1_5	53.9	12.7	9.4952	4.5	0.3404	3.3	0.73	2.9377	3.3	0.2024	3.1
ECMB1_7	36.7	21.4	12.4303	4.4	0.3701	3.1	0.72	2.7020	3.1	0.2437	3.0
ECMB1_8	35.0	11.3	13.0744	4.4	0.3986	3.2	0.72	2.5088	3.2	0.2380	3.1
ECMB1_9	18.9	4.0	18.8677	4.6	0.4342	3.2	0.71	2.3031	3.2	0.3153	3.2
ECMB1_10	43.1	17.7	11.1589	4.5	0.3841	3.1	0.70	2.6035	3.1	0.2108	3.2
ECMB1_11	35.2	92.0	11.5084	4.7	0.3857	3.3	0.71	2.5927	3.3	0.2165	3.3
ECMB1_12	34.6	15.8	12.4772	4.4	0.3950	3.2	0.72	2.5316	3.2	0.2292	3.1
ECMB1_13	48.9	16.5	9.9917	4.4	0.3601	3.1	0.72	2.7770	3.1	0.2013	3.0
ECMB1_14	56.7	21.5	9.3120	4.4	0.3565	3.1	0.72	2.8050	3.1	0.1895	3.0
ECMB1_15	35.4	15.2	11.3952	4.5	0.3801	3.3	0.73	2.6309	3.3	0.2175	3.0
ECMB1_16	56.8	23.2	9.7936	4.5	0.3621	3.4	0.74	2.7617	3.4	0.1963	3.0
ECMB1_17	97.9	67.0	7.2326	4.5	0.3311	3.2	0.71	3.0202	3.2	0.1585	3.2
ECMB1_19	28.9	6.7	14.2073	4.5	0.3971	3.4	0.74	2.5183	3.4	0.2596	3.0
ECMB1_20	41.0	13.5	10.6404	4.6	0.3596	3.4	0.73	2.7809	3.4	0.2147	3.1
ECMB1_21	27.3	4.0	13.9859	4.7	0.3863	3.5	0.74	2.5887	3.5	0.2627	3.1
ECMB1_22	41.1	19.6	12.2605	4.6	0.3878	3.4	0.74	2.5786	3.4	0.2294	3.1
ECMB1_23	18.8	18.3	18.9832	4.4	0.4287	3.2	0.72	2.3326	3.2	0.3213	3.1
ECMB1_24	39.2	13.3	11.5599	4.4	0.3666	3.1	0.69	2.7278	3.1	0.2288	3.2
ECMB1_25	38.3	14.2	11.4021	4.9	0.3664	3.5	0.72	2.7293	3.5	0.2258	3.4
ECMB1_26	164.9	3840.0	6.5949	5.2	0.3419	4.2	0.81	2.9248	4.2	0.1400	3.0
ECMB1_27	38.6	11.9	10.9887	4.4	0.3712	3.1	0.72	2.6940	3.1	0.2148	3.1
ECMB1_28	38.4	3.7	18.7929	4.6	0.5173	3.4	0.73	1.9331	3.4	0.2636	3.1
ECMB1_29	30.5	6.4	15.0260	4.8	0.4509	3.6	0.76	2.2178	3.6	0.2418	3.2
ECMB1_30	38.1	24.4	14.3511	4.8	0.4319	3.7	0.77	2.3154	3.7	0.2411	3.1
ECMB1_31	48.7	29.8	11.2605	4.5	0.4015	3.2	0.72	2.4907	3.2	0.2035	3.1
ECMB1_32	34.2	10.6	15.0964	4.8	0.4970	3.8	0.78	2.0121	3.8	0.2204	3.1
ECMB3											
ECMB3_1	70.1	179.1	8.5024	4.4	0.3394	3.2	0.72	2.9464	3.2	0.1818	3.0
ECMB3_2	33.3	121.6	12.5593	4.9	0.3805	3.8	0.78	2.6281	3.8	0.2395	3.1
ECMB3_3	28.5	72.2	13.1515	4.4	0.3865	3.2	0.72	2.5873	3.2	0.2469	3.1
ECMB3_4	22.0	64.8	15.9828	4.6	0.4130	3.4	0.74	2.4213	3.4	0.2808	3.1
ECMB3_5	22.3	60.6	14.7595	4.7	0.3833	3.4	0.73	2.6089	3.4	0.2794	3.2
ECMB3_6	22.5	72.0	15.7819	4.7	0.4140	3.5	0.75	2.4155	3.5	0.2766	3.1
ECMB3_7	15.9	42.4	19.6392	4.9	0.4431	3.8	0.77	2.2568	3.8	0.3216	3.1
ECMB3_8	4.8	6.3	45.8747	7.7	0.6040	7.0	0.90	1.6556	7.0	0.5511	3.3
ECMB3_9	3.8	6.7	56.4061	9.1	0.6910	8.5	0.93	1.4472	8.5	0.5923	3.2
ECMB3_10	59.0	158.7	8.8574	4.5	0.3435	3.2	0.71	2.9112	3.2	0.1871	3.1
ECMB3_11	48.1	147.1	10.8222	4.5	0.3692	3.3	0.73	2.7086	3.3	0.2127	3.0
ECMB3_12	21.0	53.0	14.8251	4.8	0.4044	3.6	0.75	2.4728	3.6	0.2660	3.1
ECMB3_13	15.9	49.1	19.8998	4.5	0.4494	3.3	0.73	2.2252	3.3	0.3213	3.1
ECMB3_14	15.4	33.1	18.9633	4.9	0.4181	3.8	0.78	2.3918	3.8	0.3291	3.1
ECMB3_15	17.0	50.8	17.4913	5.1	0.4239	4.1	0.79	2.3590	4.1	0.2994	3.1
ECMB3_16	42.6	65.7	8.7769	4.4	0.3503	3.1	0.70	2.8547	3.1	0.1818	3.1
ECMB3_17	42.6	118.3	11.2602	4.5	0.3717	3.3	0.74	2.6903	3.3	0.2198	3.0
ECMB3_18	4.9	11.6	43.5264	7.5	0.6010	6.7	0.89	1.6639	6.7	0.5255	3.4
ECMB3_19	8.0	13.0	30.2100	5.7	0.5080	4.4	0.78	1.9685	4.4	0.4315	3.5
ECMB3_20	47.8	96.7	9.4929	4.4	0.3631	3.1	0.70	2.7541	3.1	0.1897	3.2
ECMB3_21	12.5	32.3	21.9116	4.7	0.4311	3.5	0.75	2.3196	3.5	0.3688	3.1
ECMB3_22	4.2	7.6	51.7630	7.2	0.6630	6.5	0.90	1.5083	6.5	0.5665	3.1
ECMB3_23	5.2	7.7	42.9820	6.0	0.5880	5.1	0.85	1.7007	5.1	0.5304	3.2
ECMB3_24	3.4	5.5	63.3132	7.9	0.7430	7.3	0.92	1.3459	7.3	0.6183	3.2
ECMB3_26	20.3	51.3	16.7809	4.9	0.4150	3.7	0.77	2.4096	3.7	0.2934	3.1
ECMB3_27	8.1	18.1	30.5038	8.2	0.5290	7.6	0.92	1.8904	7.6	0.4184	3.3
ECMB3_28	6.0	10.9	39.1760	6.5	0.5610	5.6	0.86	1.7825	5.6	0.5067	3.3
ECMB3_29	60.4	174.2	9.5511	4.5	0.3602	3.2	0.72	2.7762	3.2	0.1924	3.1
ECMB3_30	28.5	54.8	12.2504	4.5	0.3667	3.1	0.69	2.7270	3.1	0.2424	3.3
ECMB3_31	112.7	387.0	7.7152	5.0	0.3445	3.9	0.77	2.9028	3.9	0.1625	3.2
ECMB3_32	10.4	28.3	25.2184	5.2	0.4710	4.1	0.78	2.1231	4.1	0.3885	3.2
ECMB3_33	8.6	20.1	28.1872	6.2	0.4850	5.2	0.84	2.0619	5.2	0.4217	3.4
ECMB3_34	10.0	31.6	22.8309	9.2	0.4400	8.5	0.93	2.2727	8.5	0.3765	3.4
ECMB3_35	22.9	62.7	15.1554	4.4	0.4009	3.2	0.71	2.4944	3.2	0.2743	3.1
ECMB3_36	40.7	128.6	11.3333	4.5	0.3622	3.3	0.74	2.7609	3.3	0.2270	3.0
ECMB3_38	45.8	70.3	8.9418	4.4	0.3530	3.1	0.71	2.8329	3.1	0.1838	3.1
ECMB3_39	25.8	85.4	13.5569	4.4	0.3798	3.2	0.72	2.6330	3.2	0.2590	3.1
ECMB3_40	36.3	128.5	12.1928	4.5	0.3797	3.3	0.73	2.6337	3.3	0.2330	3.0
ECMB3_42	4.2	6.3	58.7790	6.7	0.7150	5.9	0.88	1.3986	5.9	0.5965	3.1
ECMB3_43	33.8	90.3	11.9596	4.4	0.3699	3.1	0.71	2.7034	3.1	0.2346	3.1
ECMB3_44	23.1	72.5	14.4674	4.7	0.3827	3.5	0.75	2.6130	3.5	0.2743	3.1
ECMB3_45	8.6	11.8	28.0617	5.5	0.5030	4.6	0.83	1.9881	4.6	0.4048	3.1
ECMB3_46	4.3	6.7	50.1885	6.5	0.6330	5.7	0.88	1.5798	5.7	0.5753	3.1
ECMB3_47	7.6	9.7	30.6232	5.0	0.5000	4.0	0.79	2.0000	4.0	0.4444	3.1
ECMB3_48	16.9	32.0	17.7004	5.0	0.4131	3.8	0.75	2.4207	3.8	0.3109	3.3
ECMB3_49	5.2	9.3	44.3465	8.6	0.5950	8.0	0.93	1.6807	8.0	0.5408	3.1
ECMB3_50	9.4	17.7	29.5612	6.0	0.5490	5.0	0.83	1.8215	5.0	0.3907	3.4
ECMB3_51	18.1	33.5	19.5908	4.7	0.4972	3.6	0.76	2.0113	3.6	0.2859	3.1
ECMB3_52	42.7	97.2	11.2835	4.5	0.3959	3.3	0.73	2.5259	3.3	0.2068	3.1

ECMB4											
ECMB4_1	42.2	132.0	7.9502	4.4	0.3331	3.1	0.71	3.0021	3.1	0.1732	3.1
ECMB4_2	20.5	32.1	9.9057	4.6	0.3583	3.2	0.71	2.7910	3.2	0.2006	3.2
ECMB4_3	86.8	161.3	6.4967	4.3	0.3282	3.1	0.71	3.0469	3.1	0.1436	3.0
ECMB4_4	23.4	75.2	10.1149	4.8	0.3677	3.6	0.74	2.7196	3.6	0.1996	3.3
ECMB4_5	6.8	16.9	23.2267	6.0	0.4630	4.3	0.71	2.1598	4.3	0.3640	4.2
ECMB4_6	54.4	168.0	7.0172	4.5	0.3241	3.2	0.71	3.0855	3.2	0.1571	3.1
ECMB4_7	42.2	92.3	8.1260	4.4	0.3428	3.2	0.71	2.9172	3.2	0.1720	3.1
ECMB4_8	43.1	47.8	7.8004	4.5	0.3365	3.3	0.73	2.9718	3.3	0.1682	3.1
ECMB4_9	77.4	218.3	6.9050	4.3	0.3200	3.1	0.71	3.1250	3.1	0.1566	3.1
ECMB4_10	243.4	362.4	5.3710	4.3	0.3120	3.1	0.71	3.2051	3.1	0.1249	3.0
ECMB4_11	128.9	247.7	6.2358	4.3	0.3301	3.1	0.71	3.0294	3.1	0.1371	3.1
ECMB4_12	7.9	25.3	19.9491	6.3	0.4280	5.2	0.83	2.3364	5.2	0.3382	3.5
ECMB4_13	2.6	4.4	41.3289	8.9	0.5880	8.3	0.93	1.7007	8.3	0.5100	3.3
ECMB4_14	7.8	16.4	22.8506	5.5	0.4660	4.3	0.78	2.1459	4.3	0.3558	3.4
ECMB4_15	42.9	96.1	7.9985	4.5	0.3554	3.2	0.72	2.8137	3.2	0.1633	3.1
ECMB4_16	71.2	137.0	6.5874	4.3	0.3256	3.1	0.71	3.0713	3.1	0.1468	3.0
ECMB4_17	67.9	117.3	6.4340	4.3	0.3129	3.1	0.72	3.1959	3.1	0.1492	3.0
ECMB4_18	43.6	44.7	7.5746	4.6	0.3448	3.2	0.69	2.9002	3.2	0.1594	3.3
ECMB4_19	49.5	147.3	7.8430	4.9	0.3605	3.8	0.78	2.7739	3.8	0.1579	3.1
ECMB4_20	12.3	20.5	11.8828	5.4	0.3752	4.2	0.78	2.6652	4.2	0.2298	3.3
ECMB4_21	15.2	37.2	12.2956	4.8	0.3612	3.6	0.75	2.7685	3.6	0.2470	3.2
Standards											
MAD_1	28.4	713.3	1.5713	3.3	0.0860	2.3	0.69	11.6306	2.3	0.1326	2.4
MAD_2	28.0	709.2	1.5796	3.2	0.0855	2.2	0.68	11.7000	2.2	0.1341	2.4
MAD_3	27.8	700.3	1.5342	3.4	0.0846	2.2	0.65	11.8217	2.2	0.1316	2.6
MAD_4	27.8	690.6	1.5090	3.3	0.0831	2.2	0.66	12.0279	2.2	0.1317	2.5
MAD_5	27.9	696.4	1.5248	3.3	0.0863	2.4	0.71	11.5875	2.4	0.1282	2.4
MAD_6	28.1	693.6	1.5438	3.3	0.0855	2.3	0.70	11.6945	2.3	0.1310	2.3
MAD_7	27.1	673.7	1.4123	3.2	0.0839	2.2	0.67	11.9246	2.2	0.1222	2.4
MAD_8	27.6	692.1	1.5598	3.3	0.0860	2.3	0.67	11.6279	2.3	0.1316	2.5
MAD_9	28.2	704.0	1.5330	3.5	0.0846	2.3	0.66	11.8217	2.3	0.1315	2.6
MAD_10	28.1	697.2	1.5045	3.3	0.0847	2.1	0.65	11.8078	2.1	0.1289	2.5
MAD_11	28.0	691.9	1.5490	3.3	0.0852	2.2	0.65	11.7440	2.2	0.1320	2.5
MAD_12	28.0	702.9	1.5404	3.2	0.0836	2.2	0.70	11.9617	2.2	0.1337	2.3
MAD_13	28.3	712.3	1.4711	3.4	0.0834	2.2	0.66	11.9919	2.2	0.1280	2.5
MAD_14	28.0	701.2	1.5407	3.4	0.0843	2.3	0.67	11.8610	2.3	0.1326	2.6
MAD_15	28.1	704.5	1.5307	3.4	0.0847	2.2	0.66	11.8036	2.2	0.1311	2.5
MAD_16	28.1	700.1	1.5184	3.5	0.0851	2.1	0.61	11.7454	2.1	0.1294	2.8
MAD_17	27.2	680.8	1.4555	3.3	0.0847	2.2	0.66	11.8078	2.2	0.1247	2.5
MAD_18	28.7	725.3	1.6133	3.2	0.0858	2.2	0.68	11.6605	2.2	0.1365	2.3
McClure_1	31.5	65.0	2.5769	3.4	0.1009	2.2	0.64	9.9157	2.2	0.1854	2.6
McClure_2	16.5	34.0	3.5750	3.5	0.1089	2.4	0.69	9.1827	2.4	0.2382	2.5
McClure_3	15.2	33.9	3.7798	3.7	0.1101	2.6	0.71	9.0827	2.6	0.2491	2.6
McClure_4	13.6	29.4	3.8994	3.5	0.1125	2.6	0.72	8.8889	2.6	0.2515	2.4
McClure_5	14.6	31.3	3.7320	3.4	0.1099	2.4	0.71	9.0992	2.4	0.2464	2.4
McClure_6	15.0	31.0	3.7084	4.2	0.1085	2.4	0.57	9.2166	2.4	0.2480	3.4
McClure_7	14.5	30.3	3.5285	3.7	0.1078	2.7	0.72	9.2764	2.7	0.2375	2.5
McClure_8	14.7	32.0	4.0311	3.6	0.1105	2.5	0.68	9.0498	2.5	0.2647	2.7
McClure_9	15.3	31.7	3.3715	3.8	0.1099	2.5	0.65	9.0992	2.5	0.2226	2.9
OD306_1	24.3	69.8	5.1614	6.1	0.2861	2.4	0.39	3.4953	2.4	0.1309	5.6
OD306_2	24.8	67.1	4.5124	3.4	0.2825	2.3	0.66	3.5398	2.3	0.1159	2.6
OD306_3	14.0	56.0	4.9799	6.3	0.2893	2.6	0.42	3.4566	2.6	0.1249	5.7
OD306_4	22.9	67.7	4.9913	4.1	0.2845	2.4	0.59	3.5149	2.4	0.1273	3.3
OD306_5	23.8	68.6	4.1846	3.1	0.2813	2.1	0.70	3.5549	2.1	0.1079	2.2
OD306_6	18.3	36.9	4.4793	3.3	0.2889	2.3	0.71	3.4614	2.3	0.1125	2.3
OD306_7	20.3	48.6	4.3296	3.3	0.2843	2.3	0.69	3.5174	2.3	0.1105	2.4
OD306_8	19.8	40.4	4.4902	3.4	0.2909	2.3	0.68	3.4376	2.3	0.1120	2.5
OD306_9	23.9	62.4	5.2166	3.8	0.2861	2.2	0.58	3.4953	2.2	0.1323	3.1
OD306_10	19.7	39.8	4.2460	3.7	0.2811	2.4	0.64	3.5575	2.4	0.1096	2.8
OD306_11	22.9	45.4	4.3426	3.2	0.2849	2.2	0.69	3.5100	2.2	0.1106	2.3
OD306_12	24.0	73.8	4.6749	4.0	0.2815	2.2	0.57	3.5524	2.2	0.1205	3.3
OD306_13	11.3	25.1	4.2567	3.4	0.2831	2.5	0.72	3.5323	2.5	0.1091	2.4
OD306_14	24.0	65.6	4.8882	3.6	0.2784	2.4	0.67	3.5920	2.4	0.1274	2.7
OD306_15	26.1	80.7	4.4955	3.1	0.2812	2.1	0.69	3.5562	2.1	0.1160	2.3
OD306_16	19.0	38.6	5.6322	5.0	0.2880	2.4	0.48	3.4722	2.4	0.1419	4.4
OD306_17	9.5	20.9	9.2240	6.4	0.3210	4.0	0.62	3.1153	4.0	0.2085	5.1
OD306_18	24.1	67.8	5.1256	3.5	0.2901	2.2	0.62	3.4471	2.2	0.1282	2.7
401_1	19.3	139.5	1.0279	4.2	0.0890	2.5	0.59	11.2360	2.5	0.0838	3.4
401_2	19.0	144.6	1.2461	4.4	0.0917	2.4	0.55	10.9051	2.4	0.0986	3.6
401_3	19.1	140.0	0.9271	4.1	0.0884	2.4	0.58	11.3122	2.4	0.0761	3.3
401_4	18.3	137.1	0.9784	4.8	0.0883	2.4	0.49	11.3250	2.4	0.0804	4.2
401_5	19.0	132.4	0.8781	3.6	0.0861	2.3	0.62	11.6144	2.3	0.0740	2.8
401_6	19.7	138.0	0.8524	3.8	0.0859	2.6	0.68	11.6414	2.6	0.0720	2.8
401_7	19.6	137.3	0.8867	4.5	0.0871	2.2	0.49	11.4863	2.2	0.0739	3.9
401_8	19.2	140.3	1.0044	4.6	0.0892	2.8	0.60	11.2108	2.8	0.0817	3.7
401_9	19.1	139.2	0.8821	3.6	0.0872	2.2	0.61	11.4679	2.2	0.0734	2.9
401_10	19.8	143.1	0.8847	3.8	0.0855	2.2	0.58	11.6986	2.2	0.0751	3.1
401_11	19.6	141.1	0.8927	3.6	0.0873	2.4	0.67	11.4548	2.4	0.0742	2.7
401_12	20.1	149.2	0.8859	3.7	0.0852	2.2	0.59	11.7302	2.2	0.0754	3.0
401_13	19.1	139.5	0.9152	4.3	0.0877	2.2	0.52	11.3999	2.2	0.0757	3.6
401_14	19.4	140.3	0.8451	4.1	0.0853	2.3	0.55	11.7247	2.3	0.0719	3.4
401_15	19.6	140.1	0.8696	3.8	0.0855	2.3	0.60	11.6959	2.3	0.0738	3.1
401_16	19.6	140.9	0.9218	4.3	0.0851	2.5	0.58	11.7509	2.5	0.0786	3.5

Strikethrough indicates data discarded because final integrations were not flat

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