

UC Irvine

UC Irvine Previously Published Works

Title

New techniques for jet calibration with the ATLAS detector

Permalink

<https://escholarship.org/uc/item/15n4977w>

Journal

European Physical Journal C, 83(8)

ISSN

1434-6044

Authors

Aad, G

Abbott, B

Abeling, K

et al.

Publication Date

2023

DOI

10.1140/epjc/s10052-023-11837-9

Copyright Information

This work is made available under the terms of a Creative Commons Attribution License, available at <https://creativecommons.org/licenses/by/4.0/>

Peer reviewed



New techniques for jet calibration with the ATLAS detector

ATLAS Collaboration*

CERN, 1211 Geneva 23, Switzerland

Received: 31 March 2023 / Accepted: 13 July 2023 / Published online: 29 August 2023
© CERN for the benefit of the ATLAS Collaboration 2023

Abstract A determination of the jet energy scale is presented using proton–proton collision data with a centre-of-mass energy of $\sqrt{s} = 13$ TeV, corresponding to an integrated luminosity of 140 fb^{-1} collected using the ATLAS detector at the LHC. Jets are reconstructed using the ATLAS particle-flow method that combines charged-particle tracks and topoclusters formed from energy deposits in the calorimeter cells. The anti- k_t jet algorithm with radius parameter $R = 0.4$ is used to define the jet. Novel jet energy scale calibration strategies developed for the LHC Run 2 are reported that lay the foundation for the jet calibration in Run 3. Jets are calibrated with a series of simulation-based corrections, including state-of-the-art techniques in jet calibration such as machine learning methods and novel in situ calibrations to achieve better performance than the baseline calibration derived using up to 81 fb^{-1} of Run 2 data. The performance of these new techniques is then examined in the in situ measurements by exploiting the transverse momentum balance between a jet and a reference object. The b -quark jet energy scale using particle flow jets is measured for the first time with around 1% precision using γ +jet events.

Contents

1	Introduction	1
2	The ATLAS detector	2
3	Data samples and simulated events	2
3.1	Data samples	2
3.2	Monte Carlo simulation	2
3.3	Jet reconstruction	3
4	The simulation-based calibration	4
4.1	Event selection	4
4.2	Estimating the median p_T density and the pile-up density correction	4
4.3	Residual pile-up correction	5
4.3.1	The 1D residual pile-up correction	5
4.3.2	The 3D residual pile-up correction	6

4.3.3	Comparison of the different residual pile-up corrections	7
4.4	The jet energy scale and η calibration	7
4.4.1	Polynomial fits	7
4.4.2	Penalised splines	7
4.4.3	Comparison of calibrations	8
4.4.4	Absolute MC jet η calibration	9
4.5	The global property calibration	10
4.5.1	The global sequential calibration	10
4.5.2	The global neural network calibration	12
4.5.3	Comparison of the methods	12
4.5.4	Flavour uncertainties	14
5	In situ analysis	14
5.1	η -intercalibration	16
5.2	Z/γ +jet balance	17
5.3	b -quark jet energy scale in γ +jet balance	22
6	Conclusion	22
	References	26

1 Introduction

The energetic proton–proton (pp) collisions produced by the Large Hadron Collider (LHC) yield final states that are predominantly characterised by jets, which are collimated sprays of charged and neutral hadrons and their decay products. Jets constitute an essential piece of the physics programme carried out using the ATLAS detector, and a precise understanding of jet reconstruction is critical for a wide variety of processes. Measurements of both the jet energy scale (JES) and resolution (JER) of these complex objects are therefore essential for precision measurements of the Standard Model and for searches for new phenomena beyond it. Several new methods are presented for improving the jet energy scale calibration and evaluating their performance in simulation and data, paving the way to achieving a better precision on the JES for Run 3 and beyond. These techniques were developed for jets reconstructed with the anti- k_t algorithm [1,2] with radius parameter $R = 0.4$ using particle flow inputs [3,4]. Previous

* e-mail: atlas.publications@cern.ch

calibration strategies by the ATLAS Collaboration, that used up to 81 fb^{-1} data [4–9], are extended and improved by taking advantage of the full Run 2 data sample of 140 fb^{-1} .

The jet energy scale calibration consists of a series of calibration steps. The first stage of the calibration uses simulation to derive corrections to the jet energy scale to reduce the impact of pile-up, detector effects, and other parameters. The second stage of the calibration is a residual in situ calibration, correcting for remaining differences between data and Monte Carlo (MC) simulation, derived using well-measured reference objects, including photons and Z bosons.

The structure of the paper is as follows. Section 2 describes the ATLAS detector, and Sect. 3 describes the recorded data and the MC simulation samples, and the inputs and algorithms used to reconstruct the jets. Section 4 describes the methods used and the result of the simulation-based calibration, Sect. 5 describes the in situ calibration, and conclusions are given in Sect. 6.

2 The ATLAS detector

The ATLAS detector [10] at the LHC covers nearly the entire solid angle around the collision point.¹ It consists of an inner tracking detector surrounded by a thin superconducting solenoid, electromagnetic and hadron calorimeters, and a muon spectrometer incorporating three large superconducting air-core toroidal magnets.

The inner-detector system (ID) is immersed in a 2 T axial magnetic field and provides charged-particle tracking in the range of $|\eta| < 2.5$. The high-granularity silicon pixel detector covers the vertex region and typically provides four measurements per track, the first hit normally being in the insertable B-layer (IBL) installed before Run 2 [11, 12]. It is followed by the silicon microstrip tracker (SCT), which usually provides eight measurements per track. These silicon detectors are complemented by the transition radiation tracker (TRT), which enables radially extended track reconstruction up to $|\eta| = 2.0$. The TRT also provides electron identification information based on the fraction of hits above a higher energy-deposit threshold corresponding to transition radiation.

The calorimeter system covers the pseudorapidity range of $|\eta| < 4.9$. In the region $|\eta| < 3.2$, electromagnetic calorimetry is provided by barrel and endcap high-

granularity lead/liquid-argon (LAr) calorimeters, with an additional thin LAr presampler covering $|\eta| < 1.8$ to correct for energy loss in material upstream of the calorimeters. Hadron calorimetry is provided by the steel/scintillator-tile calorimeter, segmented into three barrel structures within $|\eta| = 1.7$, and two copper/LAr hadron endcap calorimeters. The solid angle coverage is completed with forward copper/LAr and tungsten/LAr calorimeter modules optimised for electromagnetic and hadronic energy measurements respectively.

The muon spectrometer (MS) comprises separate trigger and high-precision tracking chambers measuring the deflection of muons in a magnetic field generated by the superconducting air-core toroidal magnets. The field integral of the toroids ranges between 2.0 and 6.0 T m across most of the detector. Three layers of precision chambers, each consisting of layers of monitored drift tubes, cover the region $|\eta| < 2.7$, complemented by cathode-strip chambers in the forward region, where the background is highest. The muon trigger system covers the range of $|\eta| < 2.4$ with resistive-plate chambers in the barrel, and thin-gap chambers in the endcap regions.

Interesting events are selected by the first-level trigger system implemented in custom hardware, followed by selections made by algorithms implemented in software in the high-level trigger [13]. The first-level trigger accepts events from the up-to 40 MHz bunch crossings at a rate below 100 kHz, which is further reduced by the high-level trigger to record events to disk at about 1 kHz.

An extensive software suite [14] is used in the reconstruction and analysis of real and simulated data, in detector operations, and in the trigger and data acquisition systems of the experiment.

3 Data samples and simulated events

3.1 Data samples

The data used for the calibration study were collected by ATLAS in the pp collisions at $\sqrt{s} = 13 \text{ TeV}$ from 2015 to 2018 with all subdetectors operational, corresponding to an integrated luminosity of 140 fb^{-1} . The proton bunch crossing interval was 25 ns during the data taking. The average number of interactions per bunch crossing (μ) for the Run 2 data is 34. These conditions, and those of the detector configuration and reconstruction thresholds, are taken into account in producing and reconstructing simulated data [15].

3.2 Monte Carlo simulation

The Monte-Carlo (MC) based calibration uses MC simulated dijet and multijet events. PYTHIA v8.230 [16] is used as

¹ ATLAS uses a right-handed coordinate system with its origin at the nominal interaction point (IP) in the centre of the detector and the z -axis along the beam pipe. The x -axis points from the IP to the centre of the LHC ring, and the y -axis points upwards. Cylindrical coordinates (r, ϕ) are used in the transverse plane, ϕ being the azimuthal angle around the z -axis. The pseudorapidity is defined in terms of the polar angle θ as $\eta = -\ln \tan(\theta/2)$. Angular distance is measured in units of $\Delta R \equiv \sqrt{(\Delta\eta)^2 + (\Delta\phi)^2}$.

the nominal MC generator for simulating dijets. Samples of $2 \rightarrow 2$ dijet events are simulated using the A14 tune [17] and the NNPDF 2.3 [18] parton distribution function (PDF) set. Decays of heavy-flavour hadrons are modeled using EvtGen [19].

Several alternative samples are used to study the impact of the MC simulation on the calibration, and to determine uncertainties based on this. Two different dijet samples are simulated using the SHERPA 2.2.5 [20] generator. The matrix element calculation was included for the $2 \rightarrow 2$ process at leading-order, and the default SHERPA parton shower [21] based on Catani–Seymour dipole factorisation was used for the showering with p_T ordering, using the CT 14NNLO PDF set [22]. The first of these samples made use of the dedicated SHERPA AHADIC model for hadronisation [23], based on cluster fragmentation. A second sample was generated with the same configuration but using the SHERPA interface to the Lund string fragmentation model of PYTHIA 6 [24] and its decay tables. These two sets of samples were used to evaluate uncertainties stemming from the hadronisation modelling.

Two sets of samples are simulated using HERWIG 7.1.6 [25–27] with the NNPDF 2.3LO PDF set. These samples include the $2 \rightarrow 2$ process at the matrix element level, and either the default angular-ordered parton shower [28] or a dipole parton shower [29,30] using splitting kernels based on the Catani–Seymour subtraction scheme [31,32], and in both the cases, using a cluster hadronisation.

Alternative samples of multijet production at NLO accuracy were produced with POWHEG BOX v2 [33,34] interfaced to PYTHIA 8. These were generated with the dijet process as implemented in POWHEG BOX v2 [35]. The p_T of the underlying Born configuration was taken as the renormalisation and factorisation scales and the NNPDF 3.0NLO [36] PDF was used. PYTHIA with the A14 tune and the NNPDF 2.3LO PDF was used for the shower and multi-parton interactions. These samples included per-event weight variations for different perturbative scales in the matrix element, different parton distribution functions and their uncertainties, and the PYTHIA perturbative shower uncertainties. This multijet sample is used as the nominal MC sample for in situ η -intercalibration, while POWHEG BOX v2 [33,34] interfaced to HERWIG [27] with an angular-ordered parton shower [28] and the NNPDF 2.3LO PDF is used as an alternative sample.

For the in situ Z +jet analysis, samples of Z boson with jets (Z +jet) are produced with MADGRAPH + PYTHIA8 [37] using the NNPDF 3.0NNLO PDF set [20] and the AZNLO set of tuned parameters [38]. SHERPA 2.2.11 is used as the alternative MC sample. The nominal γ +jets samples are generated with PYTHIA v8.230 [16] using the A14 set of tuned parameters [17] and NNPDF 2.3 PDF set. The γ +jet samples are produced for the direct photon component and the fragmen-

tation photon component separately. The alternative sample used in γ +jet events is SHERPA 2.2.2 with NNPDF 3.0NNLO PDF set [20].

All samples are reconstructed using a full detector simulation and superimposed minimum-bias interactions simulated using PYTHIA 8 with the A3 set of tuned parameters [39] and NNPDF 2.3LO PDF set to represent multiple pp interactions during the same or nearby bunch crossings (pile-up). The distribution of the average number of pile-up interactions in simulation is reweighted during data analysis to match that observed in the Run 2 data.

3.3 Jet reconstruction

The jets in these studies are reconstructed with the anti- k_r algorithm with a radius parameter $R = 0.4$ as implemented in the FASTJET software package [40]. Four-momentum objects are used as inputs to the algorithm, and may be particles at the generator level of the MC, charged-particle tracks, calorimeter energy deposits, or algorithmic combinations of the latter two, as in the case of the particle-flow (PFlow) reconstruction technique. Particles at the MC generator level are referred to as truth particles. Reconstructed jets use PFlow objects (PFOs) as inputs to jet reconstruction, which combine measurements from the tracker and the calorimeter to form the input signals for jet reconstruction. Specifically, energy deposited in the calorimeter by charged particles is subtracted from the observed topo-clusters and replaced by the momenta of tracks that are matched to those topo-clusters, as described in Ref. [3], and with the updates described in Ref. [4]. These resulting PFlow jets show improved energy and angular resolution, reconstruction efficiency, and pile-up stability compared with jets reconstructed using only calorimeter information.

Charged particle tracks are used for both the PFlow reconstruction and for deriving calibrations. These tracks are reconstructed in the full acceptance of the inner detector $|\eta| < 2.5$, and are required to have a $p_T > 500$ MeV unless otherwise stated, and must satisfy criteria based on the number of hits in the ID subdetectors. In addition, tracks must satisfy $|z_0 \sin \theta| < 2$ mm, where z_0 is the distance of closest approach of the track to the hard-scatter primary vertex along the z -axis. Tracks used in the calibration are matched to jets using ghost association, a procedure that treats them as four-vectors of infinitesimal magnitude during the jet reconstruction and assigns them to the jet with which they are clustered [41].

Generator-level jets, referred to as truth jets, are reconstructed using stable final-state particles, defined as those with $c\tau > 10$ mm, excluding muons, neutrinos, and particles from pile-up interactions. Generator-level jets are selected with $p_T > 7$ GeV and $|\eta| < 4.5$, while reconstructed jets used for the MC calibration are selected with $|\eta| < 4.5$.

4 The simulation-based calibration

This section details the simulation-based jet energy scale calibration, which restores the average jet energy to that of truth jets. The event selection for these steps is described in Sect. 4.1, and the calibration is done in four steps. The first two steps apply pile-up corrections to remove the excess p_T due to additional pp interactions in the same (in-time) or nearby (out-of-time) bunch crossings. The first pile-up correction, described in Sect. 4.2, applies a subtraction based on the median p_T density measured in the event and the jet area (the ‘pile-up density correction’) [41,42], minimizing the sensitivity to the model of the pile-up used in the simulation. Next, a correction for residual dependence on the number of reconstructed primary vertices in the event (N_{PV}) and μ is applied (the ‘residual pile-up correction’), based on corrections derived using simulated samples, as described in Sect. 4.3. The third step, the absolute JES calibration detailed in Sect. 4.4, corrects jets so that they agree, on average, in energy and direction with truth jets from dijet MC events. Finally, the global calibration improves the jet p_T resolution and related uncertainties by reducing the dependence of the reconstructed jet response on observables constructed using information from the tracking, calorimeter, and muon chamber detector systems, as introduced in Sect. 4.5.

4.1 Event selection

All stages of the simulation-based jet energy scale calibration use the same event selection. The MC simulation is used to determine the energy scale and resolution of jets by comparing PFlow jets with truth jets. Truth and reconstructed jets are required to satisfy $|\eta| < 4.5$ to be fully contained in the detector acceptance, and truth jets are additionally required to have $p_T > 7$ GeV. Uncalibrated jets have a positive energy, but can become negative in energy after applying the corrections described in Sects. 4.2–4.3. Biases are reduced in the determination of the average jet energy response ($E_{\text{reco}}/E_{\text{true}}$) at low energies, by requiring reconstructed jets to have $p_T > 0$ after the pile-up density correction described in Sect. 4.2, but making no requirement on the p_T after the correction described in Sect. 4.3.

Events are required to have at least one reconstructed primary vertex with at least two matched tracks with $p_T > 500$ MeV. In simulated reconstructed events, the primary vertex is the reconstructed primary vertex with the largest sum of squared track momentum, while at the truth level, the primary vertex corresponds to that of the simulated hard-scatter process, and not the collision with the highest momentum transfer. This results in some events where the pile-up collision has a larger momentum transfer than the hard-scatter collision. For MC samples, the reconstructed and truth primary vertices are required to have z -positions within 0.2 mm

of each other. Events are required to have at least two reconstructed jets, and at least one truth jet. Truth jets are geometrically matched to PFlow jets using the angular distance ΔR with the requirement $\Delta R < 0.3$. In addition, truth jets are required to be isolated from all other truth jets by $\Delta R > 1.0$, while reconstructed jets are required to be isolated from all reconstructed jets by $\Delta R > 0.6$. To reduce the contribution of events where the pile-up collision has a larger momentum transfer than the hard-scatter collision, the average p_T of the two leading reconstructed jets is required to be no larger than $1.4 \times p_T^{\text{truth, leading}}$.

4.2 Estimating the median p_T density and the pile-up density correction

Pile-up interactions change the jet energy scale, and jet reconstruction is affected by additional pp interactions in the same or nearby bunch crossings. The first stage of the jet calibration, referred to as the ‘pile-up density correction’, subtracts the expected contribution from pile-up based on the area of the jet and the median p_T density in the event [41]. To compute the jet area A , a dense, uniformly distributed in $\eta \times \phi$ population of infinitesimally soft ghost particles is overlaid on top of the event. Then, A is defined as the transverse momentum of the sum of the four-momenta of all ghost constituents matched with a given jet after clustering, normalised by the ghost constituent transverse momentum density.

The median pile-up p_T density ρ is estimated for each event by the median p_T density (p_T/A) of all jets clustered with the k_t algorithm [43,44] with a radius parameter of 0.4:

$$\rho = \text{median} \left\{ \frac{p_{T,i}^{\text{jet}}}{A_i^{\text{jet}}} \right\},$$

where the index i enumerates over the jets. For this calculation, only jets with $|\eta| < 2$ are used, since ρ falls off steeply beyond this region, due to a combination of physics and detector effects.

Assuming the pile-up is a uniform, diffuse background, the pile-up contribution to the energy of the jet can then be approximated by the product of the jet area times the median p_T density. The pile-up density-corrected jet p_T , p_T^{area} , is then defined as

$$p_T^{\text{area}} = p_T - \rho \times A.$$

The ratio of the p_T^{area} to the uncorrected jet p_T is applied as a scale factor to the jet four-momentum and does hence not affect its direction.

Previously, the inputs to the ρ calculation were the same as those used to build jets: the neutral PFlow objects and the charged PFlow objects that satisfy $|z_0 \sin \theta| < 2$ mm [4]. However, this results in a bias from the inclusion of hard-scatter tracks, shifting the median p_T density to higher

values, particularly when the hard-scatter process has a large jet multiplicity. To prevent such biases, the ‘pile-up sideband’ (PUSB) ρ definition is studied, which uses neutral PFlow objects, and charged PFlow objects that satisfy $2 \text{ mm} < |z_0 \sin \theta| < 4 \text{ mm}$ as inputs. The total amount of pile-up using the sideband cuts is expected to be similar to the nominal criteria, since a similar amount of pile-up will meet these criteria. This ensures a minimal loss of the event-by-event correlation of the charged pile-up component that is not removed by charged hadron subtraction cuts.

The JES is measured in specific event selections as described in Sect. 5, but these calibrations are applied to many final states. An uncertainty is required to cover potential inadequacy of the model used in the simulation of the difference between the bias in ρ in different event topologies, specifically the difference between the bias in ρ between dijet and $Z(\rightarrow \mu\mu)$ +jets events. The $Z(\rightarrow \mu\mu)$ +jets events are distinctive from dijet events in several ways, including the quark-gluon composition, colour flow, and momentum transfer of the process, making it a good topology to use to estimate the magnitude of a potential bias. This bias at a given μ is estimated by comparing the difference between the value of $\langle \rho \rangle$ for the data sample in two different event topologies to measure potential inadequacy of the model used in the MC simulation of the data:

$$\Delta \rho(\mu) = \left(\rho(\mu)_{\text{MC}}^{\text{dijet}} - \rho(\mu)_{\text{MC}}^{Z+\text{jet}} \right) - \left(\rho(\mu)_{\text{data}}^{\text{dijet}} - \rho(\mu)_{\text{data}}^{Z+\text{jet}} \right),$$

and the bias that is propagated to the jet energy scale uncertainties is the bias determined at the average value of μ for the data sample:

$$\text{bias}[\text{GeV}] = \Delta \rho(\langle \mu \rangle).$$

The $Z(\rightarrow \mu\mu)$ +jets selection uses the lowest unscaled single muon trigger, and requires two muons with $p_T^{\mu 1} > 30 \text{ GeV}$, $p_T^{\mu 2} > 25 \text{ GeV}$, $80 < m^{\mu\mu} < 100 \text{ GeV}$, and $p_T^{\mu\mu} > 25 \text{ GeV}$, and the dijet selection uses the lowest unscaled single jet trigger, and requires a leading jet with $p_T > 500 \text{ GeV}$, $|\eta| < 2.4$, and greater than 5% of the momentum carried by charged particle flow objects.

Figure 1 shows the dependence of $\langle \rho \rangle$ in the two processes as a function of μ for data and simulation, comparing both the ρ definitions described above. The lower panels compare the values of $\langle \rho \rangle$ in the two processes. Two different SHERPA dijet samples are shown: a 2.1.1 sample [45] that was used in the previous calibration [4], and the 2.2.5 sample used now, while for $Z(\rightarrow \mu\mu)$ +jets, only SHERPA 2.2.1 is used. The SHERPA 2.2.X samples include an improvement to the multi-parton interaction (MPI) model, which directly affects the bias in ρ . Significantly larger differences are seen between the dijet SHERPA 2.1.1 sample and the $Z(\rightarrow \mu\mu)$ +jets SHERPA 2.2.1 sample than between the dijet SHERPA 2.2.5 sample and the $Z(\rightarrow \mu\mu)$ +jets SHERPA 2.2.1

sample. Previously, the bias was determined using the dijet SHERPA 2.1.1 sample and the $Z(\rightarrow \mu\mu)$ +jets SHERPA 2.2.1 sample, which have different MPI models. Using the updated dijet SHERPA sample that uses a consistent MPI model with $Z(\rightarrow \mu\mu)$ +jets results in a factor of four reduction in the bias, showing the importance of MPI modelling in MC simulations. The new ρ definition, ρ^{PUSB} , results in significantly smaller differences between the different topologies, and a better description of the data by the simulation. Similarly, the improvements to the ρ definition result in almost a factor of three improvement to the uncertainty, as seen by the difference between data and SHERPA for the two different ρ definitions. Together, these improvements reduce the JES uncertainty from the ρ modelling by a factor of nearly seven.

4.3 Residual pile-up correction

To further reduce the impact of pile-up, a residual pile-up correction is applied, based on N_{PV} , μ , the reconstructed jet p_T (p_T^{reco}), and the reconstructed jet η (η^{reco}). Due to the fast response of the silicon tracking detectors used to reconstruct the tracks used to find the primary vertices, N_{PV} is sensitive to the in-time pile-up, while μ is sensitive to the out-of-time pile-up, since it accounts for the average amount of pile-up around a given bunch crossing. Typically, in-time pile-up increases the energy of the jet, and out-of-time pile-up decreases it. The negative dependence of the jet energy scale on μ for out-of-time pile-up is a result of the liquid-argon calorimeter’s pulse shape, which is negative during the period soon after registering a signal [46]. Two options for the residual pile-up correction are compared.

4.3.1 The 1D residual pile-up correction

The first strategy, referred to as the ‘1D residual pile-up calibration’, follows the method outlined in Ref. [4], where additional corrections are applied based on the μ and N_{PV} of the event, with

$$p_T^{\text{1D residual}} = p_T^{\text{area}} - (\partial p_T / \partial N_{\text{PV}}) \times (N_{\text{PV}} - 1) - (\partial p_T / \partial \mu) \times \mu,$$

where $\partial p_T / \partial N_{\text{PV}}$ and $\partial p_T / \partial \mu$ are determined as follows. To determine $\partial p_T / \partial N_{\text{PV}}$, first, the dependence of $\Delta p_T^{\text{area-truth}} = p_T^{\text{area}} - p_T^{\text{truth}}$ on N_{PV} is fit with a line in bins of μ , p_T^{truth} , and η . The slope of this function is taken as the dependence $\partial p_T / \partial N_{\text{PV}}$ per μ bin. The average of these slopes across μ is taken to be the p_T dependence on N_{PV} for a given p_T^{truth} and η bin. For each η bin, the average p_T dependence is fit as a function of p_T^{truth} with a logarithmic function for $20 \text{ GeV} < p_T^{\text{truth}} < 200 \text{ GeV}$. The value of the logarithmic fit at 25 GeV is taken as the nominal correction, since pile-up effects are most relevant for low- p_T jets. Finally, a piecewise

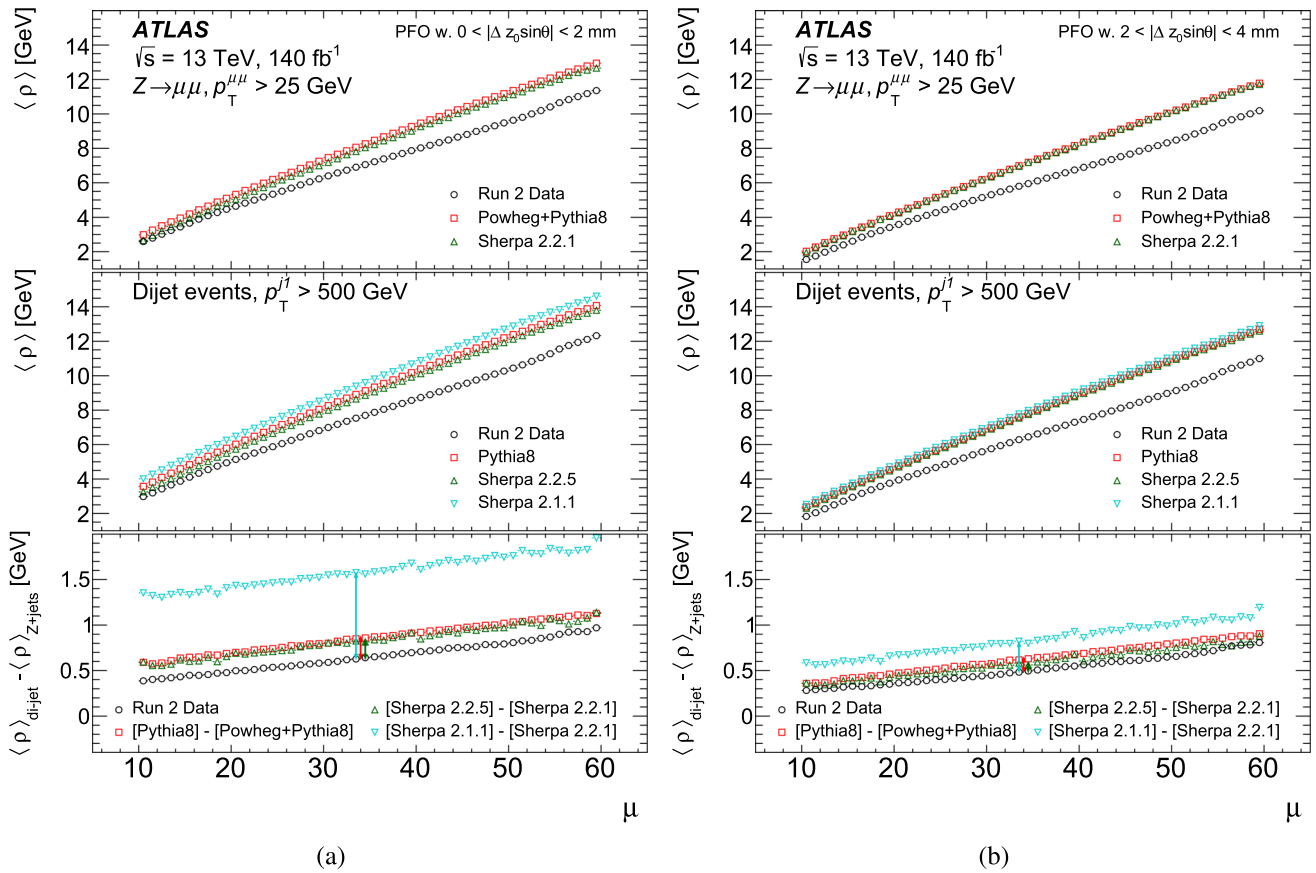


Fig. 1 The distribution of ρ as a function of μ for the (top) $Z(\rightarrow \mu\mu)+$ jets and (middle) dijet selections for data, PYTHIA 8, SHERPA 2.2.5, and SHERPA 2.1.1. The lower panel shows the difference between the two topologies which is used to determine the uncertainty from the extrapolation across topologies, indicated by the vertical arrows. The

left plot shows ρ built from the jet constituents: neutral PFOs and charged PFOs with $|z_0 \sin \theta| < 2$ mm, and the right plot shows ρ built using neutral PFOs and charged PFOs satisfying the new sideband selection

linear function is fit over the per $|\eta|$ bin values of $\partial p_T / \partial N_{PV}$, reducing statistical fluctuations and providing a continuous correction over the full η range. To determine $\partial p_T / \partial \mu$, the same process is repeated with N_{PV} and μ switched, to get the dependence of $\Delta p_T^{\text{area-truth}}$ on μ .

closest non-empty (μ, N_{PV}) bin (with the same p_T), and the result is smoothed. This average difference, $\Delta p_T^{\text{area-truth}}$, is fit as a function of p_T^{area} using a linear plus logarithmic function, in bins of N_{PV} , μ , and p_T^{reco} , determined using jets with $10 \text{ GeV} < p_T^{\text{truth}} < 200 \text{ GeV}$. The corrected value is given by

4.3.2 The 3D residual pile-up correction

The 1D residual pile-up correction does not account for correlations between μ and N_{PV} , and does not account for changes in the pile-up contribution as a function of jet p_T . The 3D residual pile-up correction is designed to include these correlations. In this calibration, derived in bins of η^{reco} , the jet p_T scale is shifted to match the truth jet scale as a function of $(N_{PV}, \mu, p_T^{\text{area}})$, simultaneously correcting for pile-up and detector effects. The p_T^{truth} is used as a reference to compute a correction given by $\Delta p_T^{\text{area-truth}}$. For extreme values of μ and N_{PV} , where there are insufficient events to determine an accurate correction, the correction is determined using the

$$p_T^{\text{3D residual}} = p_T^{\text{area}} - \Delta p_T^{\text{area-truth}}(N_{PV}, \mu, p_T^{\text{area}}).$$

By construction, this residual pile-up calibration corrects the jet energy scale to the truth jet scale, combining corrections due to pile-up with corrections due to detector effects. This is contrasted with the 1D residual pile-up correction, which is designed to exclusively remove the impact of pile-up on the jet p_T scale. Several options to only correct for the pile-up p_T were studied, but these were found to either increase the pile-up dependence or result in problematic effects such as a large fraction of jets with negative p_T .

4.3.3 Comparison of the different residual pile-up corrections

A comparison of the different options for the residual pile-up corrections is shown in Fig. 2. As seen in this figure, the residual pile-up calibration is especially useful for improving the pile-up dependence for jets with $|\eta^{\text{reco}}| > 2.5$. Overall, for the 1D residual pile-up correction, the absolute pile-up dependence increases for higher p_T jets, but the relative impact on the p_T response is smaller. While the 1D residual pile-up correction performs best for the p_T range which it is optimised for (20–30 GeV), it has a sizeable pile-up dependence at other jet p_T . In addition, since the 1D residual pile-up correction is optimised for this same bin, its performance appears enhanced by construction, while a more differential binning would show a worse performance. The 3D residual pile-up correction significantly reduces the pile-up dependence of the calibration, particularly at high p_T . Based on these results, the 3D residual pile-up calibration is used for the remainder of the reported studies.

4.4 The jet energy scale and η calibration

The absolute jet energy scale (MCJES) and η corrections provide calibration functions for the energy and η as a function of η^{det} and E^{reco} such that jets agree on average with the truth jet energy and η . Since the calorimeters measure the energy of particles, and not the transverse momenta, this correction is determined as a function of the jet energy. The jet energy response \mathcal{R} , defined as the mean of a fit with a Gaussian function to the core of the $E^{\text{reco}}/E^{\text{true}}$ distribution, is measured in E_{true} and η^{det} bins, where η^{det} is the jet η pointing from the geometric centre of the detector, which is used to remove any ambiguity about which region of the detector is measuring the jet. The difference in \mathcal{R} from the expected value of one is referred to as the non-closure, and regions where the \mathcal{R} has the expected value within the uncertainties are said to demonstrate closure.

The jet energy response after the application of the residual calibration is shown as a function of E_{true} and η^{det} in Fig. 3. This differs from previous JES calibrations by the ATLAS experiment, in that the jet energy response is already close to unity, meaning that the correction is relatively small. This is a feature of the 3D residual pile-up correction, which shifts the energy scale of the jets close to the truth scale, although there is some significant difference from one at high η , where the residual calibration insufficiently captures the behaviour of the energy response. Since the Δp_T term in the 3D residual pile-up correction is determined using jets with $p_T^{\text{truth}} < 200$ GeV, the jet energy response shifts away from one at energies corresponding to $p_T > 200$ GeV.

Directly predicting the jet energy response from E_{reco} depends on the distribution of E_{true} used to derive the cali-

bration. Overall, the distribution of the response is approximately Gaussian for a given E_{true} , but not for a given E_{reco} [47]. Therefore, the calibration uses a numerical inversion technique [5], where, for each η bin, the jet energy response is fit as a function of E_{true} , and the jet calibration factor as a function of E_{reco} is determined using the inverse of this function. The two methods of determining the fit function, polynomial fits of order N , and penalised splines are compared below.

4.4.1 Polynomial fits

Following the procedure outlined in Ref. [4], polynomial fits are defined as a function of $\log(E)$, where $N_{\text{max}} = 8$ is the maximal order of the fitted polynomials. Out of the given N_{max} fit functions, the best fit function is identified using Pearson’s χ^2 test [48]. The calibration factors are usually frozen at an η -dependent energy between 3 and 4 TeV to reduce statistical fluctuations, while at for $p_T < 8$ GeV, a linear extrapolation of the calibration factor is performed.

4.4.2 Penalised splines

In addition to the polynomial fit functions, a new method using penalised splines is studied. A spline $S(x)$ of degree N is a piecewise polynomial function of degree N , where pieces of the spline meet at points called knots, and the first $N - 1$ derivatives are continuous across the knots. Splines may be defined from b-spline basis functions $B_i(x, t)$ [49] via

$$S(x) = \sum_i^{n-1} a_i B_i(x, t),$$

where n is the number of data fit points, a_i are control points weighting the individual basis functions B_i , and t are the knots. Since $S(x) = 0$ for x outside of the range defined by the knots, an extrapolation to lower (higher) energy values is added using a linear extrapolation based on the first (last) five points for the low (high) end of the spline.

A spline will overfit the data, since the basis function is required to exactly pass through the knots, where in this case, the knots correspond to the energies where the response is determined. For a set of points x_i and their corresponding values y_i , this can be mitigated by using penalised b-splines (p-splines), which include an additional smoothness penalisation term P , minimizing

$$L = \underbrace{\chi^2}_{\text{least squares}} + \alpha \underbrace{P}_{\text{penalisation}} = \sum_{i=0}^n (y_i - S(x_i))^2 + \alpha \int_a^b (S''(x))^2 dx$$

with a and b corresponding to the range over which the penalisation term is included, with $a < x_i < b$, and the penalisa-

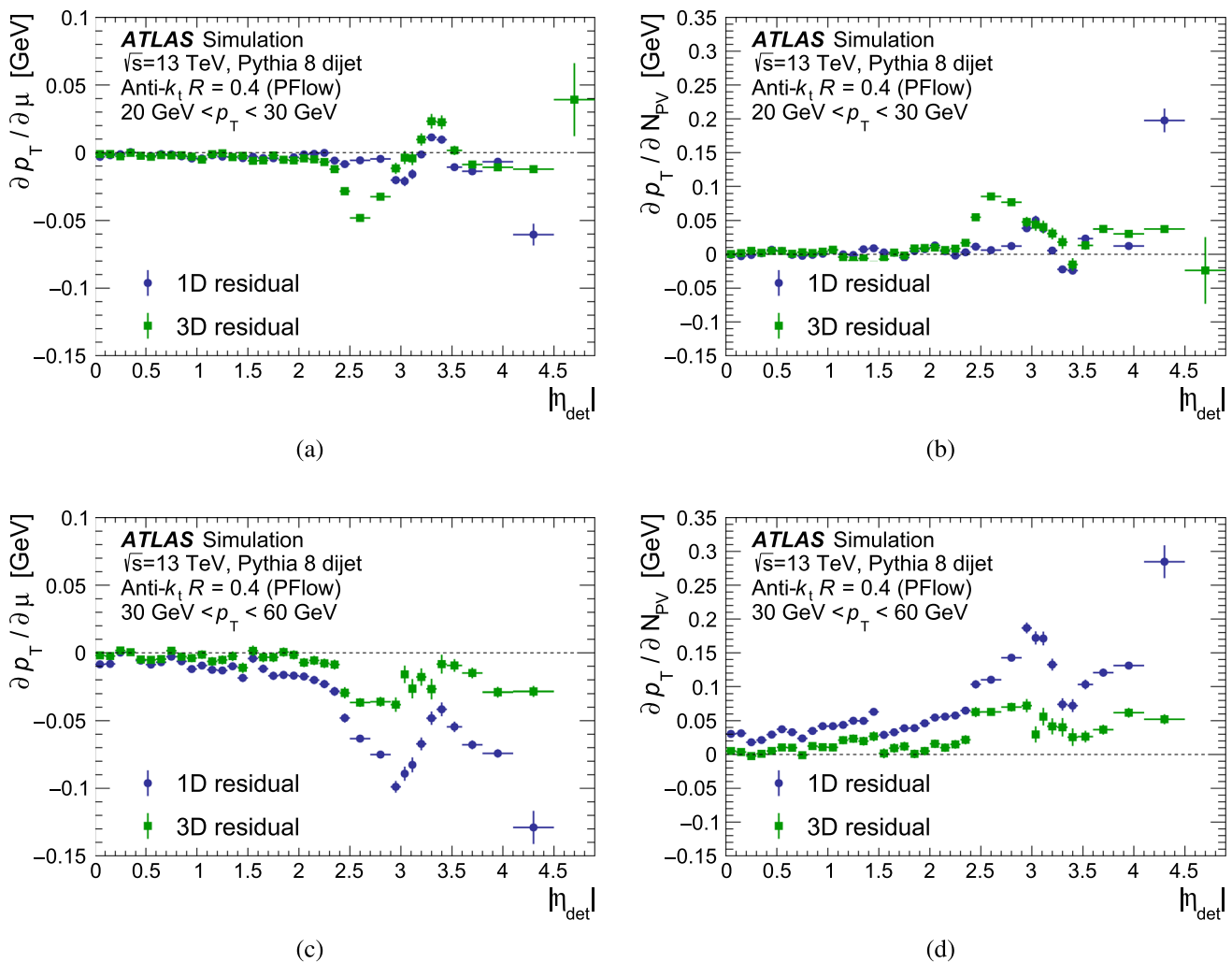


Fig. 2 Dependence of p_T on (left) μ and (right) N_{pV} after the different residual pile-up corrections. The circles indicate the 1D residual pile-up correction, and the squares indicate the 3D residual pile-up correction. This is shown for jets with (top) $20 \text{ GeV} < p_T < 30 \text{ GeV}$, (bottom) $30 \text{ GeV} < p_T < 60 \text{ GeV}$

tion parameter $\alpha \geq 0$ is chosen and fixed. For these studies, the x values correspond to E_{true} , and the y values correspond to the jet energy response. As α increases from zero to ∞ , the result moves from a spline to a linear regression, and this parameter enables a compromise between the curvature penalisation and a close fit to the data. The penalisation parameter α is defined dynamically for each η bin as

$$\alpha = \frac{\lambda}{n} \cdot \sum_{i=0}^n w_i,$$

where i runs over the n data fit points, λ is a regulative parameter, and w_i are the point weights defined as $w = \sigma_y^{-1/2}$, where σ_y is the response fit uncertainty from the iterative fit to a Gaussian function.

For these studies, the splines are implemented using the Splinter framework [50], and a spline of degree three is used, with λ empirically set to 0.1. To check for overfitting, the

calibration and the closure test are performed on statistically independent events.

4.4.3 Comparison of calibrations

A comparison of the MCJES closure for the fitting techniques at different energy values is presented in Fig. 4. Both the strategies can provide closure of 1% for high energies, while at low energies, the p-spline approach provides better closure than the polynomial fit. Overall, the p-spline correction provides closure within 1% across the p_T and η range considered, except for a small number of bins where the calibration becomes difficult due to quickly changing response, and non-Gaussian terms in the energy response. The rest of the studies use the correction determined from the p-spline fit, since it provides the best overall closure.

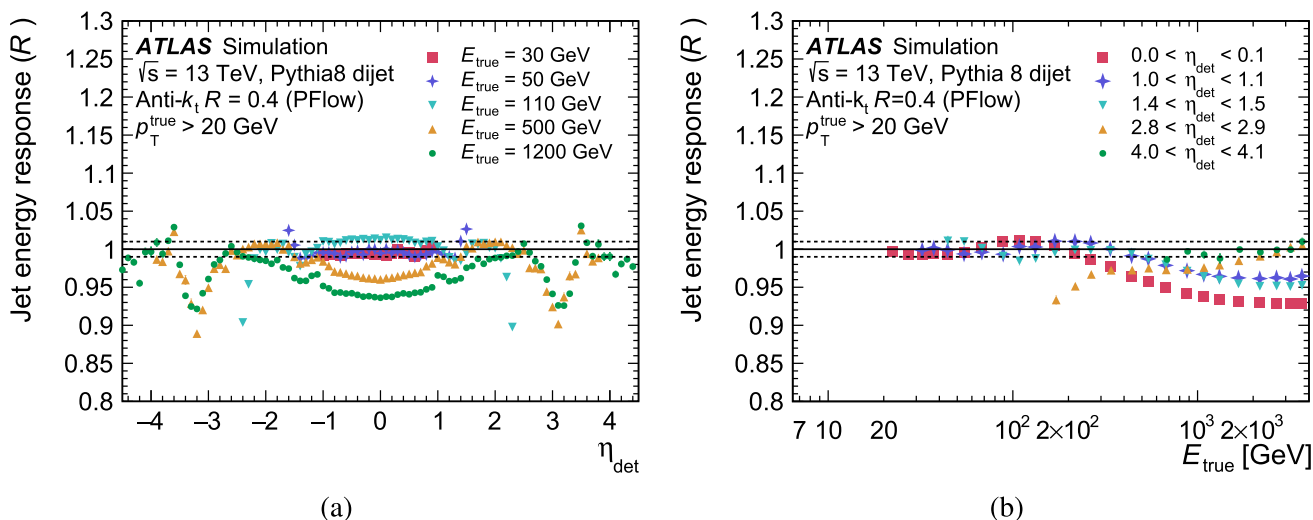


Fig. 3 The jet energy response before the MCJES calibration **a** at fixed energies as a function of η_{det} , and **b** at fixed η_{det} as a function of truth jet energy. **a** The square shows the response for $E_{\text{true}} = 30$ GeV, the plus-sign shows the response for $E_{\text{true}} = 50$ GeV, the down-triangle shows the response for $E_{\text{true}} = 110$ GeV, the up-triangle shows the response for $E_{\text{true}} = 500$ GeV, and the circle shows the response for $E_{\text{true}} =$

1200 GeV. **b** The square shows the response for $0.0 < \eta_{\text{det}} < 0.1$, the plus-sign shows the response for $1.0 < \eta_{\text{det}} < 1.1$, the down-triangle shows the response for $1.4 < \eta_{\text{det}} < 1.5$, the up-triangle shows the response for $2.8 < \eta_{\text{det}} < 2.9$, and the circle shows the response for $4.0 < \eta_{\text{det}} < 4.1$

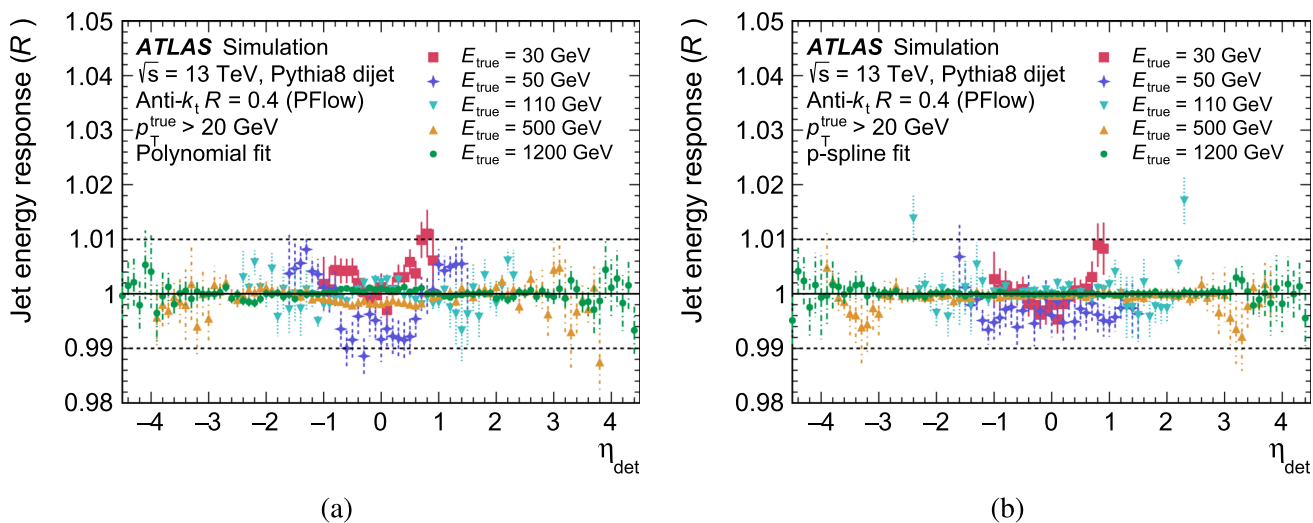


Fig. 4 Jet response at fixed energies as a function of η **a** after the polynomial MCJES calibration step, and **b** after the p-spline MCJES calibration. The square sign shows the response for $E_{\text{true}} = 30$ GeV, the plus sign shows the response for $E_{\text{true}} = 50$ GeV, the down-pointing

triangle shows the response for $E_{\text{true}} = 110$ GeV, the up-pointing triangle shows the response for $E_{\text{true}} = 500$ GeV, and the circle shows the response for $E_{\text{true}} = 1200$ GeV

4.4.4 Absolute MC jet η calibration

In addition to the jet energy, the jet pseudorapidity η is calibrated with a similar approach as the JES calibration to correct for biases in the η reconstruction, following the strategy in Ref. [4]. This bias is most pronounced in the transition region between different parts of the calorimeter, where the discrepant response of the different detectors arti-

ficially shifts the reconstructed energy on one side of the jet, changing the reconstructed η . These η corrections are particularly needed in the barrel-endcap ($|\eta| \sim 1.4$) and forward-endcap ($|\eta| \sim 3.1$) transitions. The bias in η is defined as $\eta_{\text{bias}} = \langle \eta^{\text{reco}} - \eta^{\text{true}} \rangle$, as determined by an iterative fit to a Gaussian function, and the correction is performed on a jet-by-jet basis via $\eta^{\text{calib}} = \eta^{\text{reco}} - \mathcal{R}_{\eta}$. This correction is only applied to the p_{T} and η of the jet, and is parameterised as a

function of E^{true} and η^{det} . For this correction, only polynomial fits are studied, using up to an order three polynomial. There are small correlations between the corrections in η and E , so this correction is derived simultaneously with the JES.

4.5 The global property calibration

The absolute MCJES calibration corrects the jet energy response based on the E and η^{det} of the jet. However, there are many other factors that contribute to the jet response, including the distribution of energy in the jet, the distribution of energy deposits across different calorimeter layers, and the types of hadrons produced in the jet. Many of these characteristics depend on whether the jet is quark- or gluon-initiated. This can be seen in Fig. 5, which shows an example of the jet response distribution for jets with different initiating partons, and the jet p_T response as a function of p_T^{true} , where the parton label is defined by the highest energy parton ghost-associated with the truth jet. Not only are there differences between different jet flavours (i.e. the flavour of the initiating parton), but the behaviours change with the p_T of the jet. Quark-initiated jets tend to have fewer hadrons, each with a higher fraction of the jet p_T , which typically results in contributions further into the calorimeter. In contrast, gluon-initiated jets typically have more, lower- p_T hadrons, leading to a lower calorimeter response and a wider transverse profile. These behaviours are further complicated by the use of particle flow reconstruction, which adds further dependence based on the charged particles in the jet.

The jet p_T response is also impacted by the MC model, as seen by the differences between the jet p_T response shown in Fig. 6. Overall, most MC predictions have similar behaviour for quark-initiated jets, while the differences between gluon-initiated jets can be sizeable. This is due to differences between predictions between MC generators for the amount of soft radiation and its topological distribution in the jet. There is some separation in the behaviour of models with the Lund string model for hadronisation compared with the other models, where the Lund string model tends to predict a higher gluon p_T response, with larger differences for jets with $p_T < 100$ GeV. This can primarily be attributed to the fraction of jet energy carried by baryons and kaons [51].

The global jet property calibration applies further corrections to jets based on their individual characteristics. While these corrections only have a small effect on the overall closure of the calibration, the closure is significantly improved for different classes of jets, improving the JER. In addition, this calibration reduces differences between MC predictions for the JES, resulting in smaller modelling uncertainties. Two methods for deriving the global calibration are outlined below: the global sequential calibration (GSC), which was described in previous work [4], and a new method, the global neural network calibration (GNNC). Both the corrections

are derived in $|\eta^{\text{det}}|$ bins corresponding to different detector regions, creating a balance between the statistical uncertainty and the generality of the results.

4.5.1 The global sequential calibration

The GSC is a series of multiplicative corrections to account for the differences between the calorimeter response to different types of jets, which improves the jet resolution without changing the jet energy response. The GSC is based on global jet observables such as the longitudinal profile of the energy deposits in the calorimeters, tracking information matched to the jet, and information related to the activity in the muon chambers behind a jet. Six observables that improve the JER and reduce modelling uncertainties are used as inputs to the GSC. Each GSC correction to the jet four-momentum is derived and applied independently and sequentially, using the following procedure. First, for a given GSC observable, the jet p_T response distribution is fit for each bin of $|\eta|$, p_T^{true} , and the GSC observable, using the same procedure as in the MCJES calibration. Next, these fitted values are divided by the inclusive value of the response in a given $|\eta|$ and p_T^{true} bin to avoid changing the p_T scale of the jet calibration. Then, for each bin of the GSC observable, the numerical inversion of the jet p_T response is performed after a linear smoothing in a given bin of $(p_T^{\text{true}}, |\eta^{\text{det}}|)$. The resulting responses for a given $|\eta^{\text{det}}|$ bin are then smoothed simultaneously in p_T and the GSC observable using a Gaussian kernel. Because the GSC is applied sequentially, it is possible to validate each GSC correction in a systematic way, testing the impact of any mismodelling of the input variables using data. Such studies were performed to validate the sequential correction procedure.

The six stages of the GSC, in the order of application, are

- f^{charged} : the fraction of the jet p_T carried by charged particles, as measured using ghost-associated tracks with $p_T > 500$ MeV, $|\eta^{\text{det}}| < 2.5$,
- f_{Tile0} : the fraction of jet energy (E_{frac}) measured in the first layer of the hadronic tile calorimeter, $|\eta^{\text{det}}| < 1.8$,
- f_{LAR3} : the E_{frac} measured in the third layer of the electromagnetic LAr calorimeter, $|\eta^{\text{det}}| < 3.5$,
- N_{track} : the number of tracks with $p_T > 1$ GeV ghost-associated with the jet, $|\eta^{\text{det}}| < 2.5$,
- w_{track} : also known as track width, the average p_T -weighted transverse distance in the η - ϕ plane, between the jet axis and all tracks of $p_T > 1$ GeV ghost-associated with the jet, $|\eta^{\text{det}}| < 2.5$,
- N_{segments} : the number of muon track segments ghost-associated with the jet, $|\eta^{\text{det}}| < 2.8$.

The N_{segments} correction, also known as the punch-through correction, reduces the tails of the response distribution

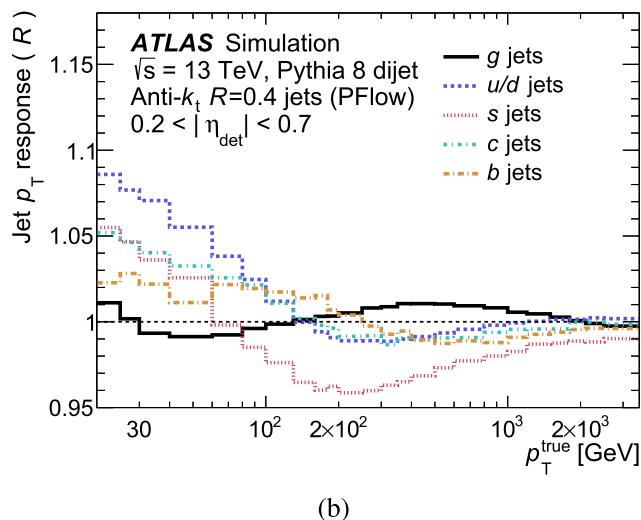
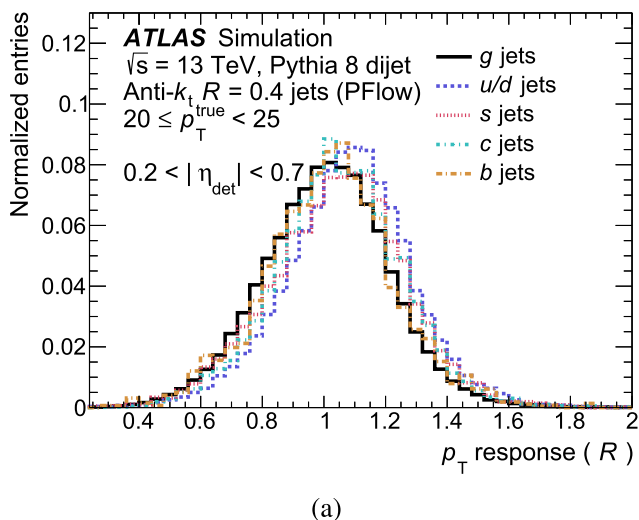


Fig. 5 **a** The jet p_T response distribution for different jet flavours for jets with $20 \text{ GeV} < p_T^{\text{true}} < 25 \text{ GeV}$, and **b** the jet p_T response for several different flavours of jets as a function of p_T^{true} . The solid line shows the response for gluon jets, the long dashed line shows the response for

light quark jets, the short dashed line shows strange jets, the alternating medium and short dashed line shows charm jets, and the alternating long and short dashed line shows bottom jets

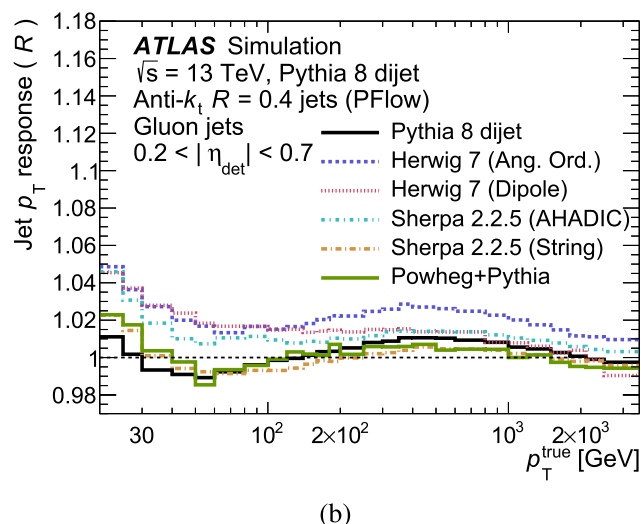
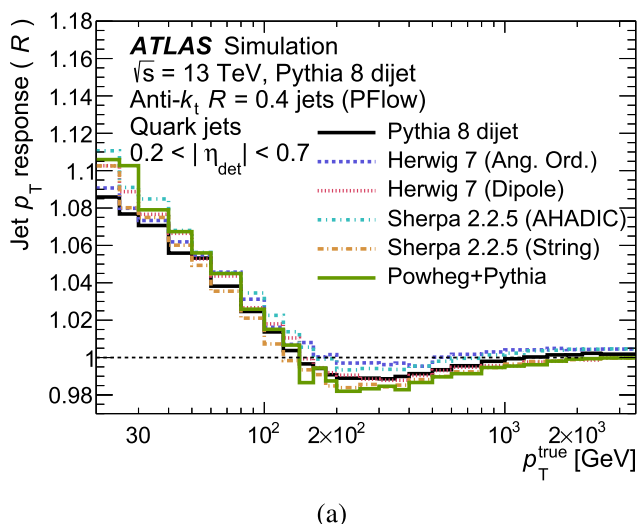


Fig. 6 The jet p_T response as a function of p_T^{true} for several different MC predictions for **a** quark jets, and **b** gluon jets. The solid line shows PYTHIA 8, the long dashed line shows HERWIG with an angular ordered parton shower, the short dashed line shows HERWIG with

a dipole shower, the alternating medium and short dashed line shows Sherpa with the AHADIC hadronisation model, the alternating long and short dashed line shows Sherpa with the string hadronisation model, and the long dash with three short dashes shows Powheg+PYTHIA

caused by high- p_T jets that are not fully contained in the calorimeter. Unlike the other corrections, the N_{segments} correction is applied as a function of the jet energy instead of the jet p_T , since this effect is more strongly correlated with energy escaping the calorimeters.

The jet p_T response for PFlow jets in MC simulation after each of the GSC corrections is shown in Fig. 7 for one $|\eta|$ bin. While the jet energy scale is within 1% at low energies, a small amount of non-closure is introduced when determin-

ing the response using p_T instead of energy. The fractional jet resolution, denoted by σ , is used to determine the magnitude of the fluctuations in the jet energy reconstruction, where σ is the width of the fit to a Gaussian function for the jet p_T response distribution divided by the mean of the fit. This is shown for PFlow jets with $0.2 < |\eta_{\text{det}}| < 0.7$ in MC simulation in Fig. 7. As more corrections are applied, the fractional jet resolution improves and the jet response dependence on the jet flavour is reduced as the calibration is improved for

jets with varying features. The impact of f_{charged} and f_{Tile0} are most apparent in Fig. 7, but the relative impact of the different corrections varies as a function of $|\eta^{\text{det}}|$. In addition, these corrections reduces effects that are less evident in the inclusive case. For instance, the punch-through correction scales with energy, and so it primarily impacts analyses that are sensitive to high-energy jets, but its impact is not obvious in the inclusive distribution.

4.5.2 The global neural network calibration

The GSC is limited to using relatively uncorrelated variables for the correction, since otherwise, each sequential step would potentially interfere with previous corrections due to correlations between observables. This constraint is fundamental to the method, limiting the set of corrections that may be applied. However when adding additional observables, and to account for their correlations, a simultaneous calibration is more appropriate [52]. As an alternative to the sequential calibration, a deep neural network (DNN) is trained to determine a simultaneous correction based on a wide variety of jet properties, enabling the use of correlated variables for determining the global jet property correction. Since analyses make selections based on the jet p_T , the DNN is designed to correct the jet p_T response, in contrast to the GSC, which leaves the energy response unchanged.

To improve the performance based on the detector geometry, a DNN is trained for each $|\eta^{\text{det}}|$ region used to derive the GSC to provide a correction to the jet p_T based on various jet- and event-level features. The DNNs are trained with Keras [53], using the Adam [54] optimisation algorithm. The network has three hidden layers with swish activation functions [55] and a single-node output layer with linear activation. The number of nodes is optimised for each $|\eta^{\text{det}}|$ bin, and ranges between 100 and 300. The network uses the leaky Gaussian kernel (LGK) loss function [56]

$$\text{Loss}(x^{\text{target}}, x^{\text{pred}}) = -\frac{1}{\sqrt{2\pi}} \exp\left(-\frac{(x^{\text{target}} - x^{\text{pred}})^2}{2\alpha^2}\right) + \beta|x^{\text{target}} - x^{\text{pred}}|,$$

where x^{target} is the jet p_T response, x^{pred} is the corresponding NN prediction, and α and β are tunable parameters. As $\alpha \rightarrow 0$, the LGK loss learns the mode, and the second term ensures that the gradient of the error function relative to the current weight does not vanish for large $x^{\text{target}} - x^{\text{pred}}$. Learning the mode is less biased by cases where the response is not a perfect Gaussian distribution, resulting in better closure than a loss function that learns the mean of the distribution.

The architecture of the network was chosen as the result of a hyperparameter optimisation based on the closure of the result, where hyperparameters are parameters involving the

network structure. The training is done with a batch size of 10^4 jets, and a learning rate of 10^{-4} . For the LGK loss, the parameters are chosen to be $\alpha = 10^{-1}$, and $\beta = 10^{-6}$ based on the hyperparameter scan. The training is done to minimise the LGK loss function, and training continues until there are no improvements to the loss for five epochs. Increasing the patience did not have a noticeable effect on the quality of the results. Unweighted events are used because this avoids issues in the training due to large differences between the event weights. Since the target is the p_T response, not the jet p_T itself, the uniform weights do not have a large impact on the final result. Only the two leading jets in the event are used in the training, since the events were simulated using a dijet process, and so this avoids potential biases from using jets that originate purely from the parton shower. For each $|\eta^{\text{det}}|$ bin, several networks were trained, and the one with the best closure was chosen for the final result.

Several sets of variables were considered as inputs to the NN, and the final list of variables used in the training is given in Table 1. This list includes all of the variables used in the GSC calibration, with the addition of more information about the jet kinematics, more granular information about the energy deposits in different calorimeter layers, and measures of pile-up. While the residual pile-up correction removes most of the pile-up dependence, some dependence is reintroduced by the absolute MCJES calibration, and so N_{PV} and μ are included in the training. Some calorimeter layers are not present for certain $|\eta^{\text{det}}|$ regions, in which case their E_{frac} is set to zero. Explicitly removing these observables from the list of input variables used in the NN training had a negligible impact on the results, and so the set of training variables is kept the same for all $|\eta^{\text{det}}|$ regions.

The jet p_T closure from this calibration is typically better than 1%, but it also has some fluctuations, which can sometimes slightly exceed this. The magnitude of these fluctuations varies with each DNN training but were persistent across different DNN hyperparameters, loss functions, and training targets. To mitigate this, an additional p_T calibration is derived after the GNNC, using the p-spline method outlined in Sect. 4.4, but using the truth jet p_T as the target instead of the energy. This is derived in $|\eta^{\text{det}}|$ bins with width of 0.1, which provides better performance than using the same $|\eta^{\text{det}}|$ bins as the GNNC correction. This has a negligible effect on the jet p_T resolution, and only serves to improve the closure and smoothness of the calibration.

4.5.3 Comparison of the methods

Figure 8 shows a comparison of the jet p_T response after the MCJES, GSC and GNNC for one representative $|\eta^{\text{det}}|$ bins. As designed, the GSC does not change the energy response of the jets. Since the JES calibration moves the reconstructed energy scale to match that of the truth scale, this can result in

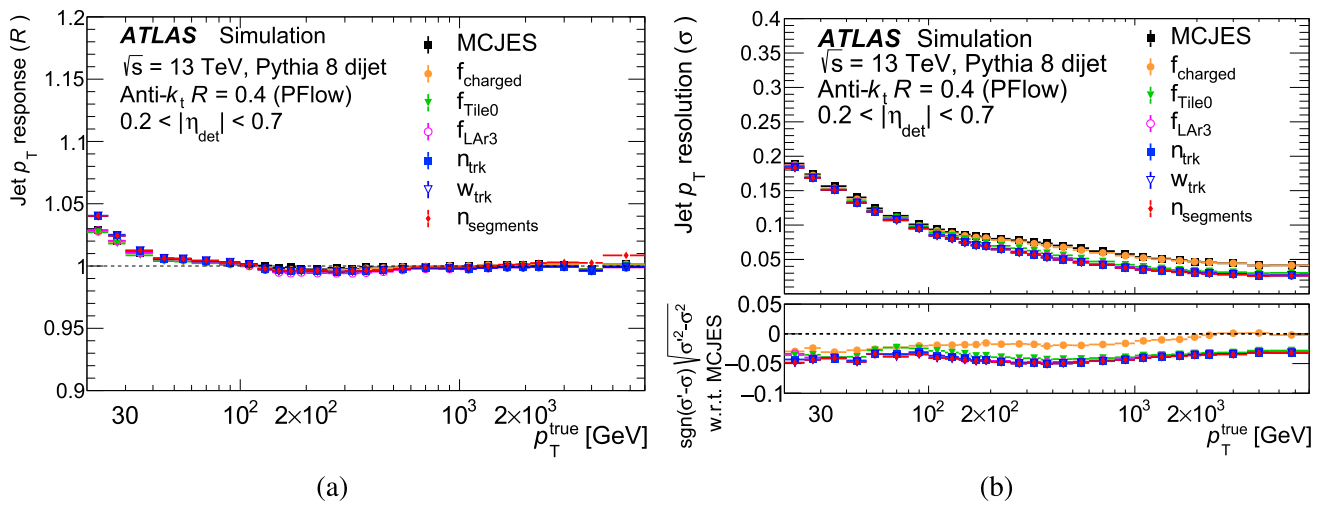


Fig. 7 **a** The jet p_T response after each stage of the GSC calibration, and **b** the jet p_T resolution after the MCJES, after the f_{charged} correction, after the f_{Tile0} correction, after the f_{LAr3} correction, after the N_{track} correction, after the w_{track} correction, and after the N_{segments} correction

Table 1 List of variables used as input to the GNNC. Variables with a * correspond to those that are also used by the GSC

Calorimeter	$f_{\text{LAr0-3}*}$	The E_{frac} measured in the 0th-3rd layer of the EM LAr calorimeter
	$f_{\text{Tile0*-2}}$	The E_{frac} measured in the 0th-2nd layer of the hadronic tile calorimeter
	$f_{\text{HEC,0-3}}$	The E_{frac} measured in the 0th-3rd layer of the hadronic end cap calorimeter
	$f_{\text{FCAL,0-2}}$	The E_{frac} measured in the 0th-2nd layer of the forward calorimeter
	$N_{90\%}$	The minimum number of clusters containing 90% of the jet energy
Jet kinematics	$p_T^{\text{JES}*}$	The jet p_T after the MCJES calibration
	η^{det}	The detector η
Tracking	w_{track}^*	The average p_T -weighted transverse distance in the η - ϕ plane between the jet axis and all tracks of $p_T > 1$ GeV ghost-associated with the jet
	N_{track}^*	The number of tracks with $p_T > 1$ GeV ghost-associated with the jet
	f_{charged}^*	The fraction of the jet p_T measured from ghost-associated tracks
Muon segments	N_{segments}^*	The number of muon track segments ghost-associated with the jet
Pile-up	μ	The average number of interactions per bunch crossing
	N_{PV}	The number of reconstructed primary vertices

some nonclosure in the jet p_T , which is particularly evident at low p_T . The GNNC is designed to change the p_T scale of the jets to match the truth jets, and so the closure in p_T is better than that of the GSC closure. It is worth noting that while the GSC can instead be applied in a way that corrects the jet p_T scale, this does not impact the resolution. Other $|\eta^{\text{det}}|$ bins show similar qualitative features, though the exact nonclosure seen in the p_T response after the MCJES and GSC varies slightly.

Figure 9 show a comparison of the jet p_T resolution after the MCJES, GSC and GNNC for several representative $|\eta^{\text{det}}|$ bins. In a few cases, the jet p_T resolution becomes worse in the lowest p_T bins, but this is also where the p_T nonclosure is most significant, making it difficult to have an accurate estimate of the resolution, particularly since the p_T scale of the GNNC is different than that of the MCJES and GSC. Since the p_T scale of the MCJES and GSC is above one and has a negative slope, the measured resolution is slightly underestimated [47] in these bins, while the GNNC resolution is correctly estimated, since the response closes. In the $0.2 < |\eta| < 0.7$ bin, the GNNC has an average improvement in the jet p_T resolution of over 15%, and maximum improvements of over 25%, when compared with the GSC. Other $|\eta^{\text{det}}|$ bins show similar average improvements of around 15–25%, with maximum improvements often over 30%, and the improvement generally becomes more pronounced at higher $|\eta^{\text{det}}|$, where the resolution improvements are significant, mostly due to the improvements from the additional detector information. Studies comparing the GNNC performance with only the GSC observables as inputs find a similar performance to the GSC, indicating that the improvement in the resolution of GNNC compared with GSC is due to the inclusion of additional observables. This is made possible by a simultaneous correction that accounts for correlations

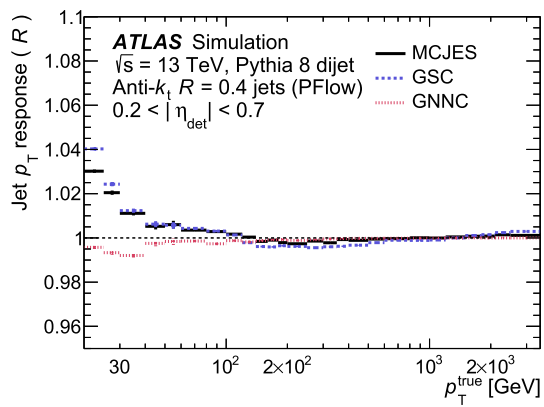


Fig. 8 The jet p_T closure for $0.2 < |\eta^{\text{det}}| < 0.7$. The solid line shows the MCJES, the long dashed line shows the GSC, and the short dashed line shows the GNNC

between observables. The GNNC provides a larger improvement to the jet energy resolution than the GSC, and so it is used for the remainder of the paper.

4.5.4 Flavour uncertainties

The two flavour-dependence uncertainties in the JES are derived from simulation and account for relative flavour fractions and differing responses to quark- and gluon-initiated jets. The flavour response uncertainty accounts for the fact that, unlike the quark-initiated jet response \mathcal{R}_q , the gluon-initiated jet response \mathcal{R}_g is found to differ significantly between generators. This uncertainty is defined as

$$\sigma_{\text{response}} = f_g (\mathcal{R}_{g,\text{PYTHIA8}} - \mathcal{R}_{g,\text{HERWIG}}),$$

where f_g is the fraction of gluon-initiated jets, and $\mathcal{R}_{g,\text{PYTHIA8}}$ and $\mathcal{R}_{g,\text{HERWIG}}$ are the gluon-initiated jet response \mathcal{R}_g in PYTHIA 8 and HERWIG respectively. The flavour composition uncertainty accounts for the fact that the jet response is different for quark- and gluon-initiated jets. This is determined based on the fraction of gluon-initiated jets f_g , where \mathcal{R}_q and \mathcal{R}_g are the quark and gluon jet responses measured in PYTHIA 8, and σ_g^f is the uncertainty in f_g in the sample, with the uncertainty defined as

$$\sigma_{\text{composition}} = \sigma_g^f \frac{\mathcal{R}_q - \mathcal{R}_g}{f_g \mathcal{R}_g + (1 - f_g) \mathcal{R}_q}.$$

Figure 10 shows a comparison of the flavour composition and flavour response uncertainties for the MCJES, GNNC and GSC. After the MCJES calibration, $\mathcal{R}_q - \mathcal{R}_g$ becomes negative for jets above 100 GeV, which appears as a dip in the flavour composition uncertainty. Both the GSC and GNNC can reduce these uncertainties, with the GNNC providing a greater reduction. For each $|\eta^{\text{det}}|$ bin, when compared with the GSC, the GNNC results in an average improvement of around 15–25% in the $40 \leq p_T < 300$ GeV range for the

flavour response uncertainty, and up to 25% improvements for the flavour composition uncertainty.

5 In situ analysis

The final calibration step accounts for differences in the jet response between simulation and data. Such differences arise due to the imperfect simulation of detector response and detector material, and the modelling of physics processes involved: hard scatter, underlying events, pile-up, jet formation and particle interactions with detector material. For the remainder of these studies, a single jet calibration is studied, using the sideband ρ definition in Sect. 4.2 and the 3D residual calibration in Sect. 4.3, the absolute MC calibration implemented with p-splines in Sect. 4.4, and the GNNC for the global calibration in Sect. 4.5. To fully understand the impact of these changes relative to the calibration procedure in Ref. [4], on the calibration and corresponding uncertainties, the in situ calibration is studied. The in situ calibration provides important validation of the new MC calibration of jets by comparing the data-to-MC difference between the p_T balance of a jet against a well-calibrated object or system. In addition, novel studies are done to disentangle the physics effects and detector effects in the η -intercalibration to reduce the systematic uncertainties. Furthermore, the b -jet JES is evaluated in situ using PFlow jets, which is performed using γ + jet events for the first time.

The in situ calibration response $\mathcal{R}_{in\ situ}$ is defined as the average ratio of the jet p_T to the transverse momentum of the reference object p_T^{ref} , derived as a function of p_T^{ref} . The $\mathcal{R}_{in\ situ}$ response is susceptible to effects such as the radiation of additional partons or the loss of energy outside the reconstructed jet cone. Dedicated event selections are applied to mitigate these effects. A double ratio, insensitive to these secondary effects provided they are well-modelled in simulations, is defined

$$\mathcal{C} = \frac{\mathcal{R}_{in\ situ}^{\text{data}}}{\mathcal{R}_{in\ situ}^{\text{MC}}}.$$

The calibration factor to the jet four-momentum can be obtained by a numerical inversion of this double ratio as a function of jet p_T , and as a function of η_{det} in η -intercalibration.

Two stages of in situ analyses are done sequentially to assess the performance of MC calibrations. First, a relative in situ calibration referred to as the η -intercalibration is done, which corrects the energy scale of forward jets ($0.8 < |\eta_{\text{det}}| < 4.5$) to match that of the central jets ($|\eta_{\text{det}}| < 0.8$) using the p_T balance in a dijet system. Second, an absolute calibration is done by measuring the p_T balance of a central jet against a well-calibrated Z boson or a photon. The missing- E_T projection fraction (MPF) method

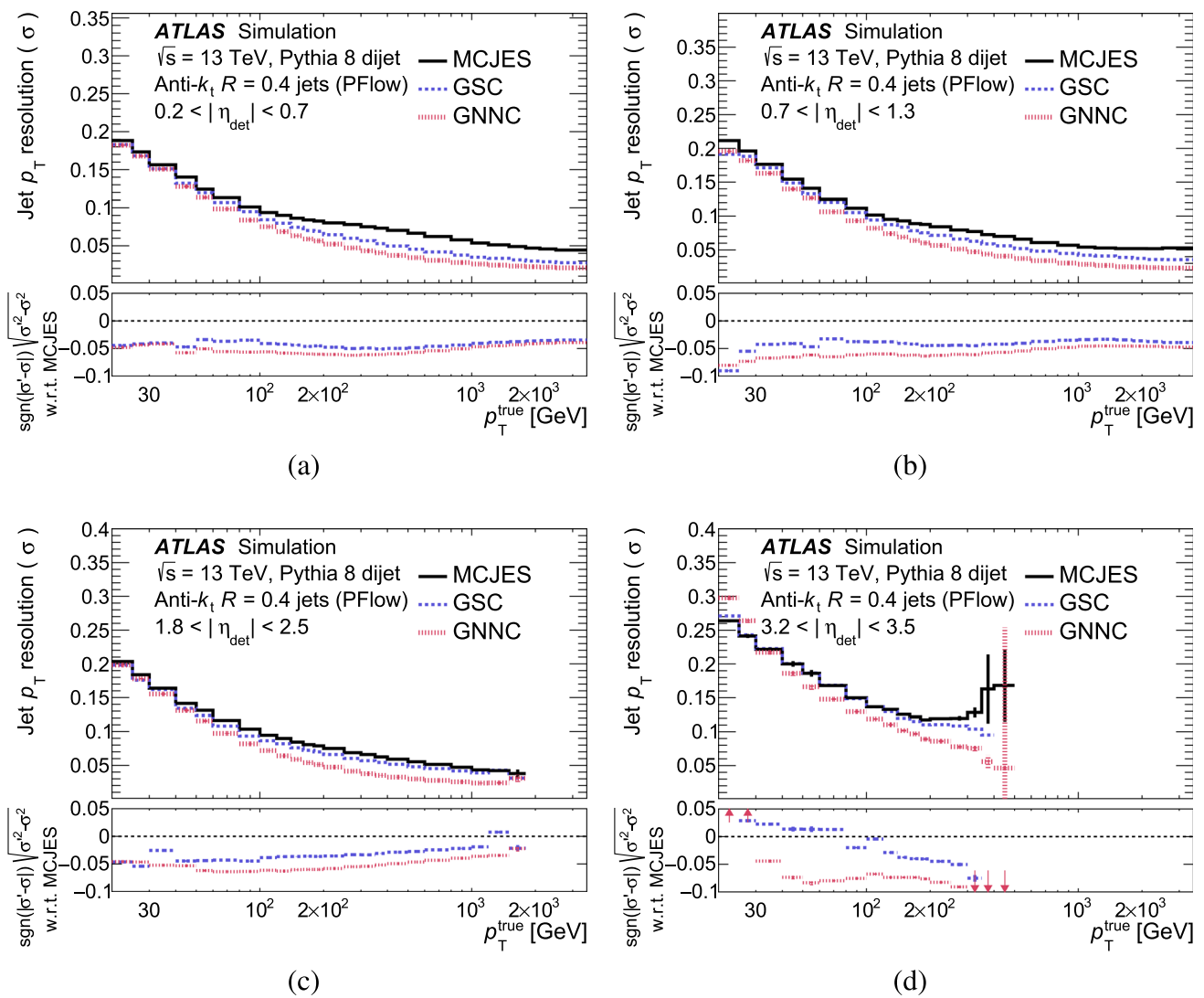


Fig. 9 The jet p_T resolution for **a** $0.2 < |\eta^{\text{det}}| < 0.7$, **b** $0.7 < |\eta^{\text{det}}| < 1.3$, **c** $1.8 < |\eta^{\text{det}}| < 2.5$, and **d** $3.2 < |\eta^{\text{det}}| < 3.5$. The solid line shows the MCJES, the long dashed line shows the GSC, and the short dashed line shows the GNNC

[57] is used in Z/γ +jet events to calculate the p_T balance between the full hadronic recoil and a Z boson or a photon. The method is less susceptible to effects of pile-up and the threshold of the jet reconstruction than the direct balance method, allowing a reliable measurement of the low- p_T jet response below 100 GeV. The direct balance (DB) method measures the balance between a (b -)jet recoiled against a photon in γ +jet events. By using the DB instead of MPF, the response of the b -jet itself is studied without including the effects of the hadronic recoil.

For each in situ analysis, main sources of systematic uncertainties arise from the MC model of physics processes, the measurement of the reference object and the p_T balance due to the selected event topology. Uncertainties related to MC model of physics effects are addressed by taking the difference between the predictions between two distinct MC event

generators. The difference between jet response in simulations depends on hadronisation models that cause different jet contents [51]. Uncertainties in the reference object are estimated by propagating its own $\pm 1\sigma$ calibration uncertainties through the analysis. Uncertainties due to the selected event topology are evaluated by varying the event selection criteria and comparing the impact on the response ratios between data and MC simulation. To reduce the statistical fluctuations when applying the systematic variations, a rebinning procedure similar to that used in previous publications [4] is employed to obtain statistically significant results using pseudo-experiments. This rebinning procedure is only performed in regions where no sharp variations in p_T response are observed to ensure no real physics effects are removed.

Events must satisfy the common selection requirements in the in situ analysis. Each event is required to have at least

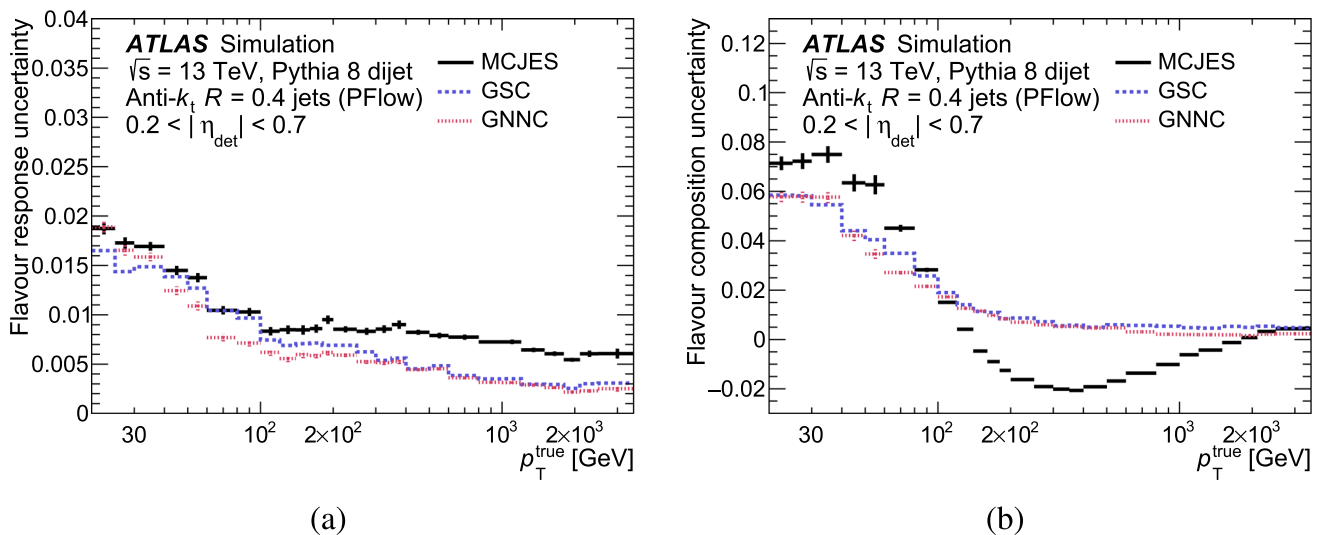


Fig. 10 **a** The flavour response uncertainty, and **b** the flavour composition uncertainty for central jets. The solid line shows the MCJES, the long dashed line shows the GSC, and the short dashed line shows the GNCC

one reconstructed primary vertex with at least two matched tracks of $p_T > 500$ MeV. Jets arising from cosmic rays, non-collision background and calorimeter noise are vetoed by applying data-quality requirements [58]. In addition, jets with $20 < p_T < 60$ GeV and $|\eta_{\text{det}}| < 2.4$ are required to satisfy the criteria of jet vertex tagging (JVT) [59,60]. The JVT criteria rejects jets from pile-up interactions by matching jets with the primary vertex; it has a selection efficiency of 97% for hard scatter jets at the nominal operating point.

5.1 η -intercalibration

The jet response in the forward region ($0.8 < |\eta_{\text{det}}| < 4.5$) is typically less understood due to the more complicated detector structure. The η -intercalibration provides a correction for forward jets ($0.8 < |\eta_{\text{det}}| < 4.5$) to bring them to the same energy scale as central jets ($|\eta_{\text{det}}| < 0.8$). This calibration uses events with a dijet topology, requiring two back-to-back jets in the transverse plane in different η_{det} regions. In order to increase the statistical precision, there is no requirement on whether or not one of the two jets is in the central reference region: instead, all regions will be calibrated relative to one another by solving a set of linear equations. This is referred to as the matrix method [4]. The momentum asymmetry is defined to measure the jet p_T balance between the two jets in two distinct detector regions (symbolically labelled *left* and *right* for simplicity)

$$\mathcal{A} = \frac{p_T^{\text{left}} - p_T^{\text{right}}}{p_T^{\text{avg}}},$$

where p_T^{avg} is the average of the transverse momentum of the *left* and *right* jets ($p_T^{\text{avg}} = (p_T^{\text{left}} + p_T^{\text{right}})/2$). For a narrow

bin approximation in p_T^{avg} , the relative response \mathcal{R} of the *left* and *right* jets in terms of the calibration factor for each jet and the mean of \mathcal{A} can be defined as

$$\mathcal{R} = \frac{c^{\text{right}}}{c^{\text{left}}} = \frac{2 + \langle \mathcal{A} \rangle}{2 - \langle \mathcal{A} \rangle} \approx \frac{\langle p_T^{\text{left}} \rangle}{\langle p_T^{\text{right}} \rangle}$$

where \mathcal{R} is measured in terms of η_{det} for *left* and *right* jets and p_T^{avg} . The intercalibration factor c is defined as $c = \frac{c^{\text{right}}}{c^{\text{left}}}$ and hence the relative response \mathcal{R} satisfies $\mathcal{R} = 1/c$.

Dijet events are selected using a combination of forward and central single-jet triggers, where each trigger is considered in the range of p_T^{avg} that has an efficiency of at least 99%. Prescaled jet triggers are used to accommodate bandwidth limits, and each selected event is weighted accordingly. The trigger combination method [4,61] is used to maximise the statistical precision. Each event must have at least two leading jets with $p_T^{\text{avg}} > 25$ GeV and $|\eta_{\text{det}}| < 4.5$. Events containing a third jet with $p_T^{\text{jet } 3} / p_T^{\text{avg}} > 0.25$ are excluded. The two leading jets must be back-to-back in the transverse plane satisfying a requirement on their azimuthal angle difference $\Delta\phi^{1,2} > 2.5$.

The nominal calibration is estimated by taking the ratio of the simulated response in POWHEG+PYTHIA8 to the measured response in data. The binning in η_{det} and p_T^{avg} is chosen to ensure enough of a sample size in scarce reference regions and to capture granular variations in detector response. A two-dimensional Gaussian kernel is optimised to smooth statistical fluctuations while also capturing notable detector features.

The 2017 data sample is representative of the high pile-up conditions and thus discussed here. The relative response, parameterised by η_{det} in two p_T^{avg} regions and by p_T^{avg} in two

η_{det} regions between the 2017 data sample and MC simulations from POWHEG+PYTHIA8 and POWHEG+HERWIG 7 with an angular ordered shower, is shown in Fig. 11. The predicted response in the two MC simulations is found to capture the overall shape of the η_{det} dependence. However, the response predicted from simulations is consistently lower than that measured in data for the forward detector regions across all $p_{\text{T}}^{\text{avg}}$ bins.

Uncertainties can arise due to the inaccurate description of physics, detector response and the dijet topology on the momentum balance. They are evaluated in terms of $p_{\text{T}}^{\text{avg}}$ and η_{det} . Uncertainties arising from the MC mismodelling are estimated by taking the difference between the smoothed residual correction between POWHEG+PYTHIA8 and POWHEG+HERWIG 7 with angular ordered shower. Other uncertainties due to the mismodelling in physics and event topology are estimated by modifying the requirement on the third jet veto, the $\Delta\phi^{1,2}$ separation, and JVT.

Further studies are performed at particle and reconstruction level separately to disentangle physics and detector effects. The particle level can be used to study physics effects affecting the dijet balance due to additional parton radiations or out-of-cone corrections. It is performed using the same procedure as the reconstruction level except that no JVT requirements are applied. The matrix method is used as the nominal method while the central reference method [4] is used as a cross-check. These physics effects induce a smooth and non trivial structure of the relative response $1/c$ with a slight asymmetry between positive and negative η_{det} in the forward region as shown in Fig. 12a. A similar structure with sharper variations due to convolution with detector effects is also present at the reconstruction level shown in Fig. 11a. The systematic uncertainty Δc on the intercalibration factor at particle level is $\Delta c = \left| \frac{c^{\text{syst}}}{c^{\text{nominal}}} - 1 \right|$, where c^{syst} is the intercalibration coefficient obtained with a different selection of the events, either a different selection on $p_{\text{T}}^{\text{jet } 3}/p_{\text{T}}^{\text{avg}}$ or $\Delta\phi^{1,2}$. By comparing Fig. 12a, b, the magnitude of these physics effects at particle level is similar to the magnitude of the systematic uncertainties designed to cover them, which are evaluated by varying the selection criteria for $p_{\text{T}}^{\text{jet } 3}/p_{\text{T}}^{\text{avg}}$ and $\Delta\phi^{1,2}$ at reconstruction level in data and MC simultaneously. Therefore, these uncertainties are not underestimated.

Variations in parton showering and hadronisation models can affect dijet balance that convolves both the physics and detector effects. The MC modelling uncertainty derived at particle level as a function of η_{det} only considers the physics effects on the dijet balance and excludes impacts on the detector response which were evaluated in the jet flavour response uncertainty using various MC simulations discussed in Sect. 4.5. Such a procedure will significantly reduce the MC modelling uncertainty shown in Fig. 12c and avoid possible double counting of uncertainties.

Figure 13 shows the fractional uncertainties derived as a function of η_{det} for two representative p_{T} values. The systematic uncertainty for $|\eta_{\text{det}}| < 0.8$ is set to zero as they are determined from the absolute in situ JES measurements such as Z/γ +jet analysis. The fractional uncertainties increase with η_{det} for $|\eta_{\text{det}}| > 0.8$ and illustrate a significant decrease with increasing p_{T} . Dominant uncertainties arise from the choice of event generators and variations in the selection criteria on $p_{\text{T}}^{\text{jet } 3}/p_{\text{T}}^{\text{avg}}$. The total systematic uncertainty is significantly reduced by using MC modelling uncertainty estimated at particle level instead of reconstruction level. It is worth noting that systematic variations in the selection criterion such as $p_{\text{T}}^{\text{jet } 3}/p_{\text{T}}^{\text{avg}}$ are performed simultaneously in data and simulation at reconstruction level while Fig. 12b shows only the relative impact at particle level. If there is a difference between up and down variations, then the systematic uncertainty is taken to be the larger absolute value. Systematic uncertainties are symmetrised around $\eta_{\text{det}} = 0$ between the positive and negative η_{det} values using the most conservative approach, as whether the asymmetry of the systematic uncertainty in η_{det} arises from statistical fluctuations or detector effects is unknown.

5.2 Z/γ +jet balance

The next step in the jet calibration brings the absolute jet energy scale in data to the same scale in simulation by exploiting the p_{T} balance between the hadronic recoil and a well-calibrated object such as a Z boson or a photon. The jet used in the in situ analysis is required to be from the central detector region ($|\eta| < 0.8$), in which the derived correction can be applied to jets in the forward region via the η intercalibration. The Z/γ +jet balance measurement is built upon the precise determination of the energy of the photon or e/μ pair from a Z boson decay. These measurements benefit from the accurate knowledge of the energy scale and resolution of the leptons. The calibration of electrons and photons is accurately known through measurements using $Z \rightarrow ee$ data and other final states [62] while the muon calibration is determined to high precision through studies of $J/\Psi \rightarrow \mu\mu$, $Z \rightarrow \mu\mu$ and $\Upsilon \rightarrow \mu\mu$ [63].

Three independent measurements consisting of Z +jet for the $Z \rightarrow ee$ and $Z \rightarrow \mu\mu$ decay channels, and γ +jet are used to do the absolute in situ calibration. The Z +jet measurement provides enough of a sample size at low and medium jet p_{T} covering $17 < p_{\text{T}} < 800$ GeV with limited precision above 800 GeV. The γ +jet analysis provides a complimentary measurement at medium and high p_{T} covering $30 < p_{\text{T}} < 2000$ GeV, with limited precision below 100 GeV due to prescaled low- p_{T} triggers, jets misidentified as photons, and MC event generator choices.

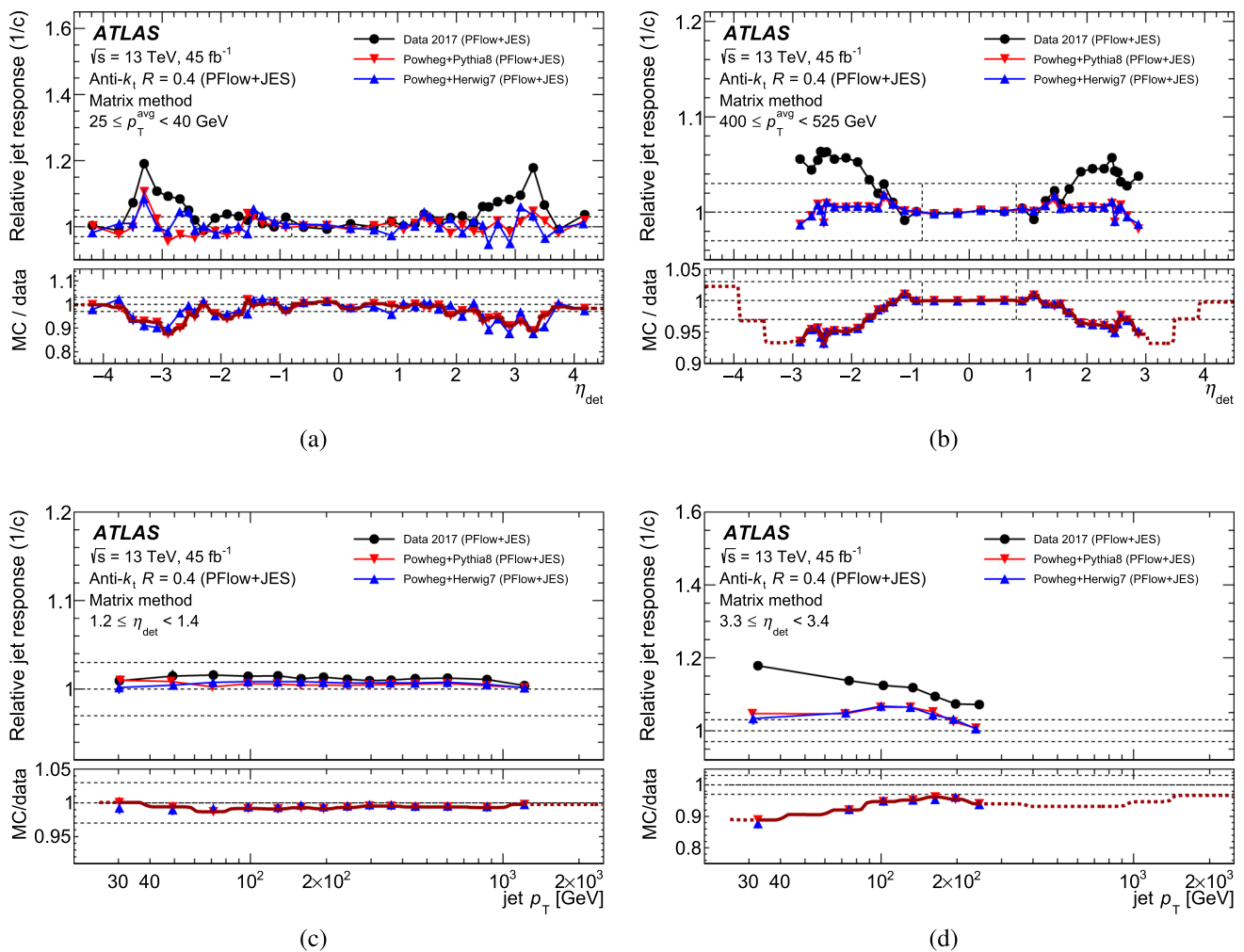


Fig. 11 Relative jet response, $1/c$, calibrated with PFlow+JES as a function of η_{det} in the ranges of **a** $25 \text{ GeV} < p_{\text{T}}^{\text{jet}} < 40 \text{ GeV}$ and **b** $400 \text{ GeV} < p_{\text{T}}^{\text{jet}} < 525 \text{ GeV}$, and as a function of p_{T} in ranges of **c** $1.2 < \eta_{\text{det}} < 1.4$ and **d** $3.3 < \eta_{\text{det}} < 3.4$. The top panel presents the measured relative response for data (dots), POWHEG+PYTHIA8 (triangles) and POWHEG+HERWIG 7 (triangles). The bottom panel presents the MC-to-data response ratios represented by triangles and the smoothed

in situ corrections are represented by overlaid curves, in which the solid line shows the derived calibration and the dashed line shows the extrapolated calibration to sparse detector regions using the two dimensional Gaussian Kernels. Two perpendicular lines are drawn at $\eta_{\text{det}} = \pm 0.8$ to indicate the central ($|\eta_{\text{det}}| < 0.8$) and the forward ($0.8 < |\eta_{\text{det}}| < 4.5$) detector region. Three horizontal dashed lines are drawn at 0.97, 1, and 1.03 to provide reference points for the viewer

The MPF method measures the p_{T} balance between the reference objects and the full hadronic recoil in Z/γ +jet events. This technique allows the calorimeter response to the hadronic showers to be computed directly. It has low susceptibility to pile-up and underlying event which is uniform across the detector and thus cancelled out in the MPF method. According to conservation of transverse momentum, the transverse momentum of all of the hadronic activity in a Z/γ +jet event, $p_{\text{T}}^{\text{recoil}}$, should be equal and opposite to the transverse momentum of the reference boson, $p_{\text{T}}^{\text{ref}}$, at particle level, such that

$$\vec{p}_{\text{T,truth}}^{\text{ref}} + \vec{p}_{\text{T,truth}}^{\text{recoil}} = 0. \quad (1)$$

At the detector level, the well-calibrated objects have a response of one while the calorimeter response to the hadronic recoil r_{MPF} is lower than unity, resulting in possible missing transverse momentum $\vec{E}_{\text{T}}^{\text{miss}}$ in the event. Therefore, equation 1 can be written as:

$$\vec{p}_{\text{T}}^{\text{ref}} + r_{\text{MPF}} \vec{p}_{\text{T}}^{\text{recoil}} = -\vec{E}_{\text{T}}^{\text{miss}}.$$

Projecting the vector terms along the direction of the reference boson using a unit vector \hat{n}_{ref} in the transverse plane, then r_{MPF} is only dependent on the missing transverse momentum and the transverse momentum of the reference boson. The average of r_{MPF} , \mathcal{R}_{MPF} , is measured as a function

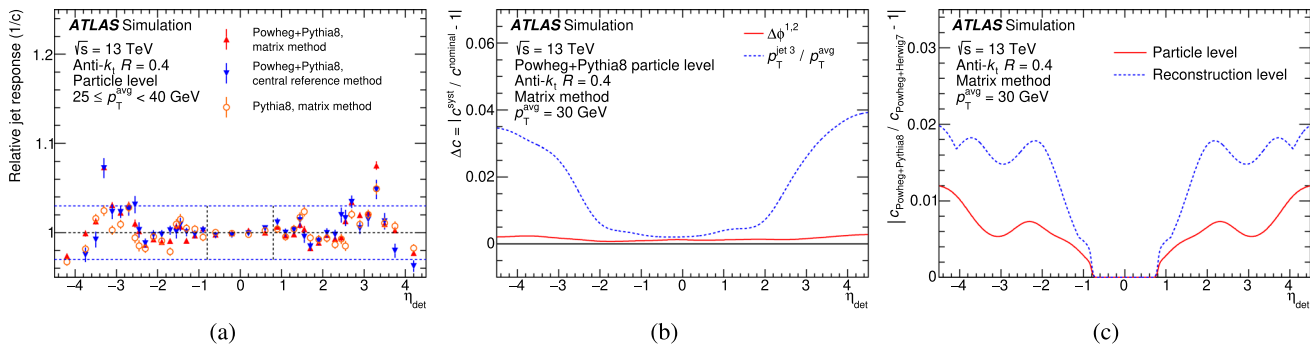


Fig. 12 Intercalibration coefficients and uncertainties derived at both the particle and reconstruction level, using MC simulated events reconstructed with the conditions of 2017 data-taking period. **a** Intercalibration coefficients obtained for different generators and different methods. POWHEG+PYTHIA8 is used as the nominal MC generator. Three horizontal dashed lines are drawn at 0.97, 1, and 1.03 to provide reference points for the viewer. **b** Systematic uncertainties obtained with

the matrix method and with POWHEG+PYTHIA8 at particle level. The up and down variations are symmetrised by taking the maximum of either of them. A 2D smoothing is applied (in p_T^{avg} and in η) with a Gaussian kernel. **c** MC modelling uncertainty, evaluated either at reconstruction level in the dashed line or at particle level in the solid line. A smoothing is applied after the computation of $\left| \frac{C_{POWHEG+PYTHIA8}}{C_{POWHEG+HERWIG7}} - 1 \right|$

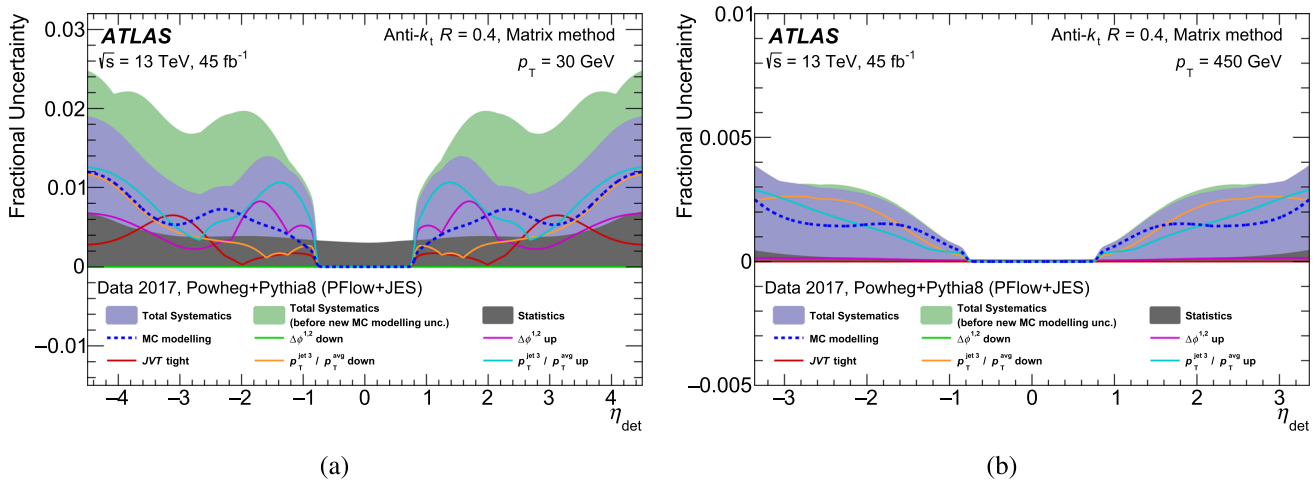


Fig. 13 Systematic uncertainties in the η intercalibration as a function of η_{det} for PFlow+JES jets of **a** $p_T = 35$ GeV, **b** $p_T = 450$ GeV. The total systematic uncertainty is represented by the middle shaded band, which is a quadrature sum of different components of systematic uncertainties marked by coloured lines. The statistical uncertainty is indicated by the lowest shaded band. The highest shaded band shows the quadrature sum of different components of systematic uncertainties using MC modelling uncertainty estimated at reconstruction level. A smoothing procedure is applied to the systematic uncertainty to suppress statistical fluctuations. The precision is limited by the MC modelling uncertainty estimated at particle level and variations in the selection criteria for p_T^{jet3} / p_T^{avg}

of p_T^{ref} , the p_T of the reference Z/γ boson:

$$\mathcal{R}_{MPF} = \left\langle 1 + \frac{\hat{n}_{ref} \cdot \vec{E}_T^{miss}}{p_T^{ref}} \right\rangle$$

where \vec{E}_T^{miss} is computed using particle-flow objects calibrated at the EM scale.

Z +jet events are selected using either the lowest- p_T unrescaled dielectron or dimuon trigger. The lowest- p_T threshold in the dielectron trigger corresponds to 15 GeV for each electron while the lowest- p_T threshold in the dimuon trigger corresponds to 14 GeV [64,65] for each muon. Both

the leptons are required to have $p_T > 20$ GeV to have fully efficient triggers. Electrons or muons must satisfy loose identification and isolation criteria [62,63]. Electrons are required to fall within $|\eta| = 2.47$ and are rejected if they fall in the calorimeter crack region $1.37 < |\eta| < 1.52$. Muons must fall within $|\eta| = 2.4$. The oppositely charged electron and muon pair is required to have an invariant mass around the Z boson mass, $66 < m_{ee/\mu\mu} < 116$ GeV. γ +jet events are selected using a combination of prescaled and unrescaled single photon triggers in which the lowest prescaled trigger E_T threshold is 10 GeV. Photon candidates entering the analysis are required to have $E_T^{jet} > 25$ GeV and $|\eta^\gamma| < 1.37$ and

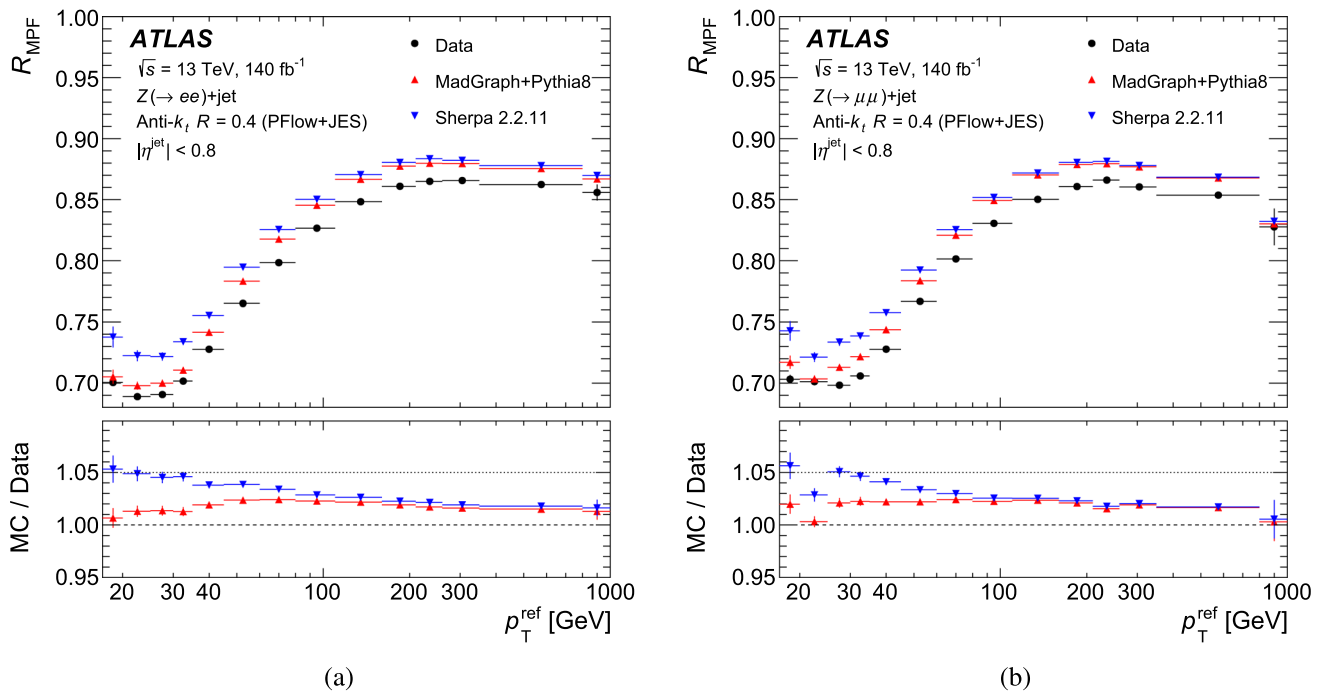


Fig. 14 The MPF response as a function of p_T^{ref} measured in data and simulations in $Z+\text{jet}$ events for **a** $Z \rightarrow ee$ and **b** $Z \rightarrow \mu\mu$. The data are represented by the black dots. The MADGRAPH+PYTHIA8 predictions are represented by the triangles while the SHERPA predictions are rep-

resented by the inverted triangles. The MC-to-data response ratios are shown in the bottom panel. The error bars correspond to the statistical uncertainties

to satisfy the tight identification and isolation selection criteria [62]. The jet is removed if it falls within $\Delta R = 0.4$ (0.35) of a photon (lepton).

Further selection criteria are imposed in the $Z/\gamma+\text{jet}$ measurements to reduce the impact from pile-up and additional parton radiations. To suppress contamination from pile-up, jets are required to satisfy the cleaning criteria and to satisfy the JVT requirement. Events must contain a jet with p_T greater than 10 GeV that falls within $|\eta| = 0.8$. To suppress effects from additional parton radiations, further requirements are imposed on the azimuthal angle between the reference boson and the leading jet $\Delta\phi^{\text{ref, jet}} > 2.9$ and p_T of the subleading jet $p_T < \max(0.3 \times p_T^{\text{ref}}, 12)$ GeV, where the subleading jet falls within $|\eta| = 4.5$.

The MPF response as a function of reference boson p_T is shown in Figs. 14 and 15 using $Z+\text{jet}$ and $\gamma+\text{jet}$ events for data and two distinct MC samples. The MC sample used to derive the nominal calibration for $Z+\text{jet}$ ($\gamma+\text{jet}$) corresponds to MADGRAPH+PYTHIA8 (PYTHIA8). The alternative MC sample corresponds to SHERPA to determine the uncertainty from the MC event modelling. The dip in the MPF response at low p_T^{ref} arises due to two opposing effects: the jet reconstruction threshold which tends to increase the response at the lowest jet p_T values between 17 GeV and 20 GeV and the apparent rise in MPF response as a function of p_T . The MC-

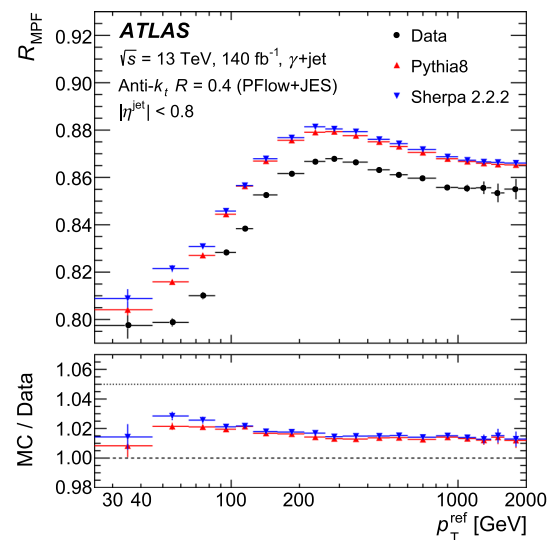
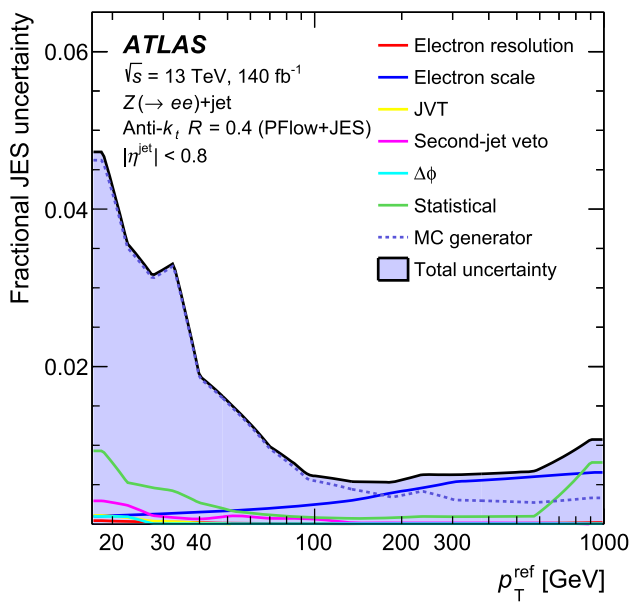
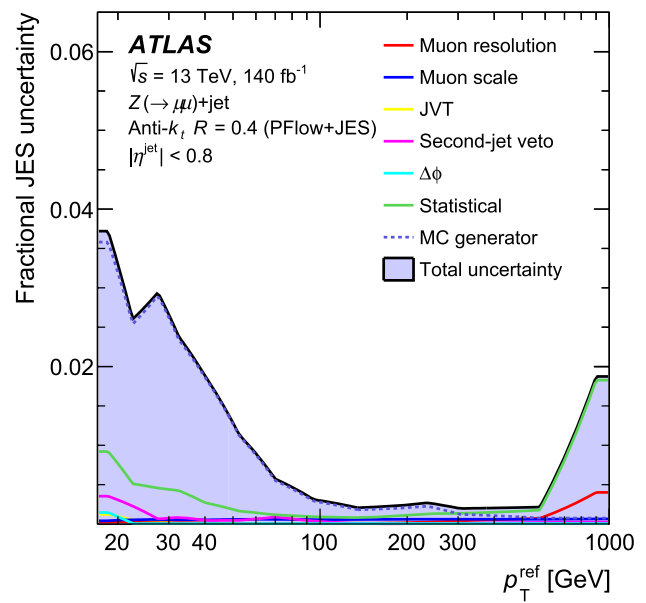


Fig. 15 The MPF response as a function of p_T^{ref} measured in data and simulations for $\gamma+\text{jet}$. The data are represented by the black dots. The PYTHIA8 predictions are represented by the triangles while the SHERPA predictions are represented by the inverted triangles. The MC-to-data response ratios are shown in the bottom panel. The error bars correspond to the statistical uncertainties

to-data response ratio are rather consistent between $Z+\text{jet}$ and $\gamma+\text{jet}$.



(a)



(b)

Fig. 16 Systematic uncertainty in the MPF response ratios as a function of p_T for jets calibrated up to, and including, the η intercalibration for **a** $Z(\rightarrow ee)+jet$ events and **b** $Z(\rightarrow \mu\mu)+jet$ events. Uncertainties arise from JVT, the subleading jet veto and $\Delta\phi^{ref, jet}$ requirement in the analysis selection. Uncertainties due to electron and muon energy

scale and resolution are propagated through the analysis. The statistical uncertainty of the MC-to-data response ratios and the uncertainties due to choice of event generators are shown. Each uncertainty is smoothed to suppress statistical fluctuations

Several sources of systematic uncertainties are considered. Uncertainties due to the energy scale and resolution of the reference objects $e/\mu/\gamma$ are derived from existing calibrations for each object and propagated through the corresponding analysis. The impact of additional parton radiation on the response measurement is evaluated by varying the selection criteria for the subleading jet veto and $\Delta\phi^{ref, jet}$. Uncertainties arising from pile-up suppression are estimated by comparing the response measurement between tighter and looser JVT working points. Uncertainties arising from photon purity in $\gamma+jet$ events are assessed using the same methodology documented in [66], in which one of the jets is misreconstructed as a photon. The pseudo-experiments are implemented in the estimate of uncertainties to reduce statistical fluctuations.

The uncertainties for the calibration are presented for the $Z \rightarrow ee$ and $Z \rightarrow \mu\mu$ measurements in Fig. 16 and for the $\gamma+jet$ measurement in Fig. 17. Uncertainties are dominated by the modelling in MC simulations in the low- and medium- p_T regions, and by the energy scale of the photon/electron for $p_T > 100$ GeV.

The derived calibrations are stable over a range of pile-up conditions in Run 2. Figure 18 shows the MC-to-data response ratios as a function of μ and N_{PV} in $\gamma+jet$ event for

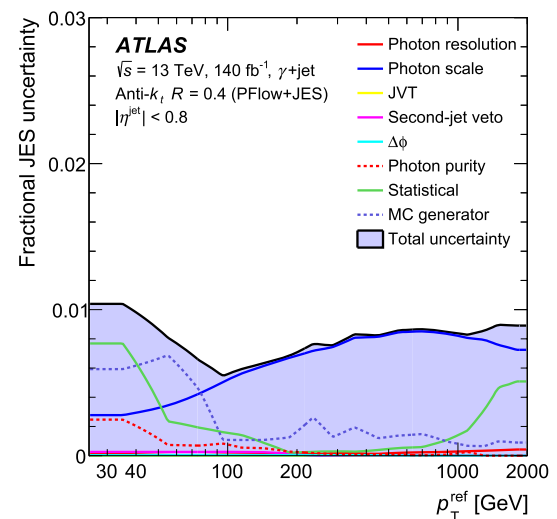


Fig. 17 Systematic uncertainty in the MPF response ratios as a function of p_T for jets calibrated up to, and including, the η intercalibration in $\gamma+jet$ events. Uncertainties arise from JVT, the subleading jet veto and $\Delta\phi^{ref, jet}$ requirement in the analysis selection. Uncertainties due to photon energy scale and resolution are propagated through the analysis. The statistical uncertainty of the MC-to-data response ratios and the uncertainties due to choice of event generators and photon purities are shown. Each uncertainty is smoothed to suppress statistical fluctuation

$45 < p_T^{\text{ref}} < 65$ GeV. The in situ calibration is consistent as a function of μ or N_{PV} , demonstrating the expected stability.

5.3 b -quark jet energy scale in γ +jet balance

The measurement of the top-quark mass is limited by the b -quark jet energy scale (b JES) and a measurement of the b JES can potentially improve the precision. The direct balance (DB) technique is used in γ +jet events to measure the balance of a (b -tagged) jet against a well-calibrated photon. It represents the first measurement determining the b -tagged jet energy scale using the PFlow jets in this event topology. p_T^{ref} is defined in terms of the reference object p_T , $p_T^{\text{ref}} = p_T \times \cos \Delta\phi$, where p_T is the transverse momentum of the photon, and $\Delta\phi$ is the azimuthal angle difference between the photon and the leading jet.

The selections are similar to the γ +jet selection in the MPF method unless stated otherwise. Events must have a jet with $p_T > 20$ GeV instead of $p_T > 10$ GeV in the central detector region ($|\eta_{\text{det}}| < 0.8$). The higher jet p_T threshold arises due to the tighter requirement on the jet transverse momentum in the b -tagging algorithm used. To suppress additional radiation, the DB technique requires the subleading jet $p_T^{\text{j2}} < \max(0.1 \times p_T^{\text{ref}}, 15)$ GeV and $\Delta\phi^{\gamma,\text{jet}} > 2.8$. Jets in the inner tracker coverage ($|\eta| < 2.5$) containing b -hadrons are identified (b -tagged) by a multivariate algorithm (DL1r) using information of impact parameters of tracks and displaced vertices [67]. The b -tagging working points with an average efficiency of 60%, 70%, 77% and 85% are used. The events are classified into inclusive and b -tagged categories. The b -tagged category is predominantly composed of the b - and c - quark jets while the light quark and gluon jets dominates the inclusive categories. A jet is labelled as b - (c -) quark jets if any $b(c)$ parton or hadron at particle level is found to be within a cone of $\Delta R < 0.3$ around a reconstructed jet, otherwise it is labelled as light quark or gluons. A summary of the jet flavour composition for inclusive and b -tagged jets is documented in Table 2. The 85% b -tagging working points are dominated by the presence of c -quark jets and the measurement can be used to constrain the c JES in $H \rightarrow c\bar{c}$ analysis [68] for instance.

Figure 19 shows the DB response as a function of the reference photon p_T for the inclusive and b -tagged jets using b -tagging working points with an average efficiency of 77%. The MC simulations are in reasonable agreement with data. The MC-to-data response ratios are found to be slightly below one for b -tagged jets and above one for inclusive jets in almost all bins. The difference between DB response between PYTHIA8 and SHERPA arises due to different b -quark fragmentation and decay models. Checks on the apparent rise of the DB response around 150 GeV for b -tagged jets are done such as the quality of the DB response

fit, b -tagging scale factors applied in simulations, the jet flavour composition between neighbouring p_T bins and a looser second jet veto with $p_T^{\text{j2}} < 0.2 \times p_T^{\text{ref}}$. None of the checks mentioned above is responsible for the DB response rise around 150 GeV. Hence these checks suggest that the feature is due to statistical fluctuations. Figure 20 shows the uncertainties for the b -tagged case with a precision between 1% and 5% and inclusive jets with a precision up to 1% for the chosen p_T range. For b -tagged jets, uncertainties are dominated by the event generator modelling everywhere, while for inclusive jets the precision is limited by the event generator modelling, photon purity and the subleading jet veto at lower p_T and photon energy scale for $p_T > 70$ GeV.

A new observable $\tilde{R}_{b\text{JES}}$ is defined as a double ratio of b -tagged response to the inclusive jet response to further measure the energy scale differences between the b -tagged and inclusive jets,

$$\tilde{R}_{b\text{JES}} = \frac{\mathcal{R}_{b\text{-tagged}}^{\text{MC}}/\mathcal{R}_{b\text{-tagged}}^{\text{data}}}{\mathcal{R}_{\text{inclusive}}^{\text{MC}}/\mathcal{R}_{\text{inclusive}}^{\text{data}}}.$$

As the nominal jet calibration is determined relative to the inclusive jet, such a double ratio can be applied on top of the nominal jet calibration to correct b JES. The value of $\tilde{R}_{b\text{JES}}$ is determined to be below one using both the MC samples with a slightly higher response in PYTHIA8 than SHERPA shown in Fig. 21. The difference between the two event generators arises from different fragmentation and decay models. The ratio, $\tilde{R}_{b\text{JES}}$, is also determined inclusively for photon p_T^{ref} between 85 and 1000 GeV for various b -tagging working points in Table 3 to increase statistical precision for PYTHIA8 and SHERPA, respectively. It is foreseen to provide MC specific calibrations for the b JES to reduce the effects arising from MC modelling. The ratio, $\tilde{R}_{b\text{JES}}$, was measured with unprecedented precision up to 1%. This in turn will improve precision in measurements of top mass.

6 Conclusion

The determination of the jet energy scale (JES) is presented using data recorded by the ATLAS experiment in pp collisions at $\sqrt{s} = 13$ TeV. The calibration scheme used for anti- k_t jets reconstructed using radius parameter $R = 0.4$ consists of two steps: a Monte-Carlo-based calibration that corrects jets to the truth jet scale, and an in situ calibration correcting the scale of jets in data.

The simulation-based calibration implements several new strategies to improve the pile-up stability at higher p_T , closure, energy resolution, and modelling uncertainties of the jets. Biases related to the determination of the pile-up p_T density were a dominant source of uncertainty for jets with p_T below 30 GeV. The new procedure presented, combined

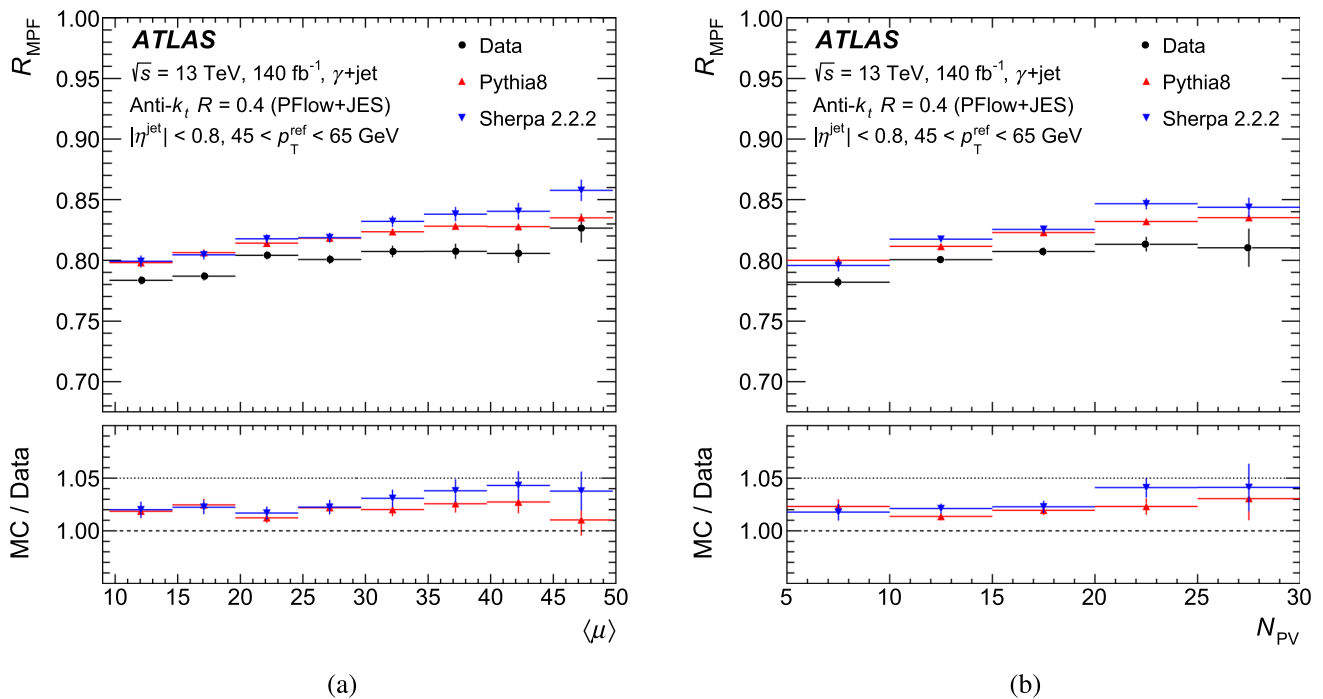


Fig. 18 The MPF response for $45 < p_T^{\text{ref}} < 65$ GeV measured in data and simulations as a function of **a** μ and **b** N_{PV} . The data are represented by the black dots. The PYTHIA8 predictions are represented

by the triangles while the SHERPA predictions are represented by the inverted triangles. The MC-to-data response ratios are shown in the bottom panel. The error bars correspond to the statistical uncertainties

Table 2 The average fractions of jet flavours for various b -tagging working points and inclusive jet

Hadron Label	Inclusive jet		WP 60%		WP 70%		WP 77%		WP 85%	
	PYTHIA8	SHERPA	PYTHIA8	SHERPA	PYTHIA8	SHERPA	PYTHIA8	SHERPA	PYTHIA8	SHERPA
b	1.9%	2.0%	92.1%	85.8%	81.3%	79.2%	59.2%	61.7%	36.3%	39.2%
c	14.2%	10.6%	2.3%	1.6%	13.4%	9.2%	35.4%	27.0%	57.4%	49.7%
Light q or gluon	83.9%	87.4%	5.6%	12.6%	5.3%	11.6%	5.4%	11.3%	6.3%	11.1%

with improvements to the multi-parton interactions model in Monte Carlo simulation, reduces this uncertainty by a factor of seven. Following this, a new residual calibration is applied, which reduces the effects of pile-up by simultaneously correcting for μ , N_{PV} , and p_T . For the absolute MCJES, a new fit method based on splines is used, leading to better closure for jets with p_T below 30 GeV. Finally, for the global calibration, which improves the resolution of jets and reduces the difference between the energy scale for quark- and gluon-initiated jets, a new method using a DNN is used, which allows information from correlated observables to be used for this calibration step. This DNN results in an average improvement of the JER of around 15% improvement compared with previous methods, with maximum improvement of over 40%.

Following these simulation-based calibration steps, the full Run 2 data sample is used to do a residual in situ calibration to correct the data-MC differences and constrain the

uncertainties. Dijet events are used to calibrate jets in the forward region relative to the central region as a function of jet transverse momentum and pseudorapidity. The precision is improved by up to a factor of two in the forward detector region at low p_T by evaluating the MC modelling uncertainty at particle level instead of reconstruction level. Central jets are calibrated by exploiting the balance between jets recoiling against either a photon or a Z boson. Unprecedented precision up to 1% is achieved in the in situ analysis. For the first time, the energy scale of b -tagged jets relative to inclusive jets is determined with precision up to 1% in γ +jet events. This result is important for improving precision in analysis sensitive to b -JES such as the top quark mass and $H \rightarrow b\bar{b}$ measurements.

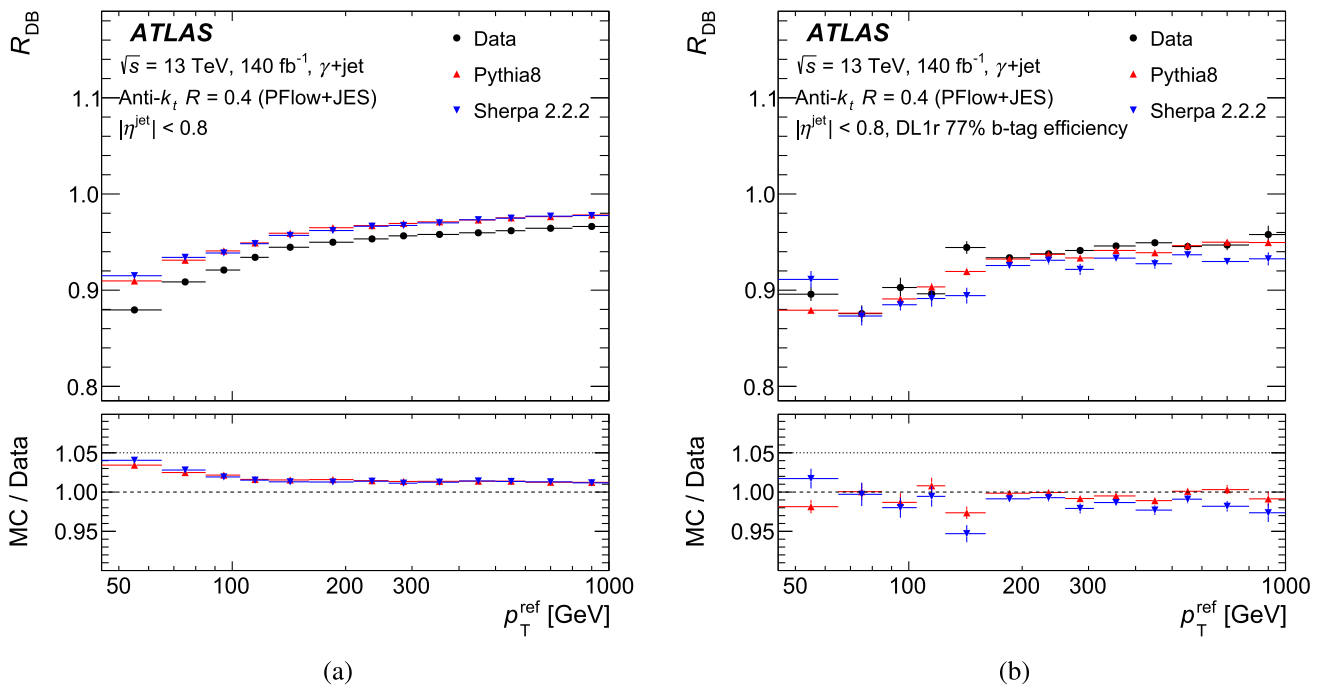


Fig. 19 The DB response as a function of p_T^{ref} measured in data and simulations for **a** γ +jet and **b** γ + b -tagged jet. The data are represented by the black dots and the PYTHIA8 are represented by the triangles, and

the (SHERPA) predictions are represented by the inverted triangles. The MC-to-data response ratios are shown in the bottom panel. The error bars correspond to the statistical uncertainties

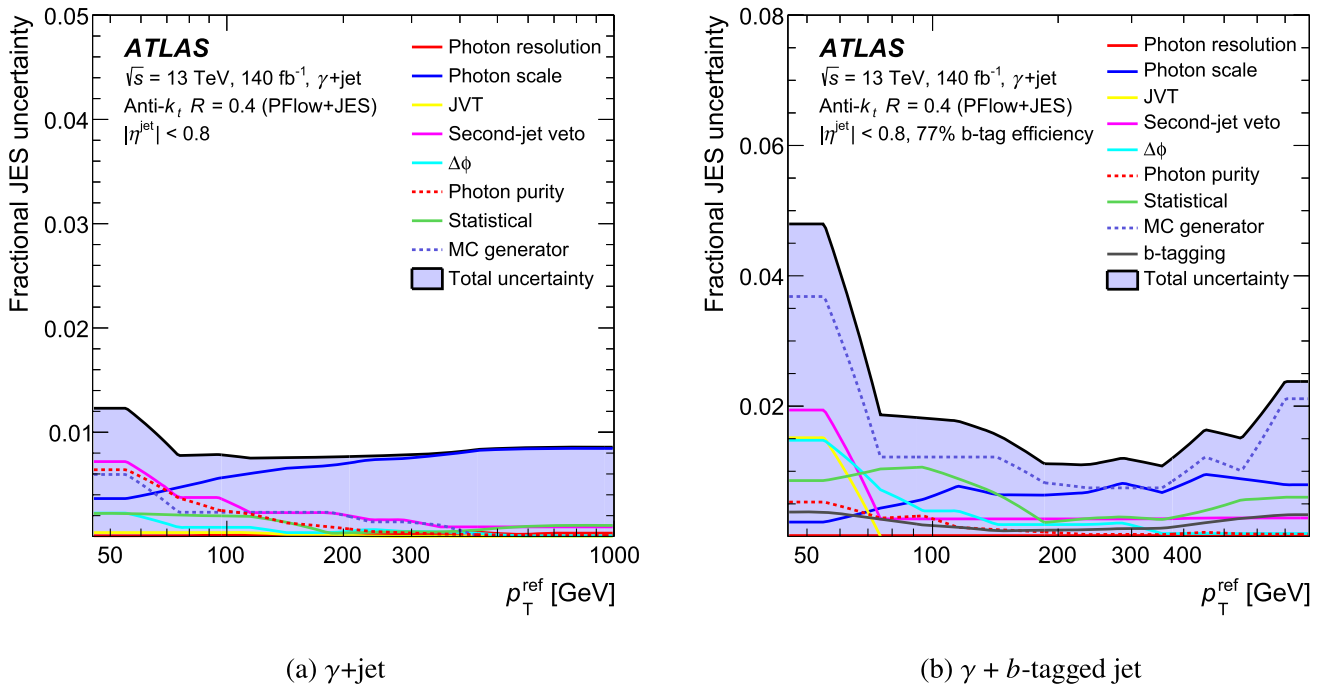


Fig. 20 Summaries of uncertainties in the MC-to-data response ratio as a function of p_T^{ref} for **a** γ +jet and **b** γ + b -tagged jet. Uncertainties can arise from photon purity, b -tagging, JVT, $\Delta\phi$ and subleading jet veto

requirement. Uncertainties due the photon energy scale and resolution are propagated through the analysis

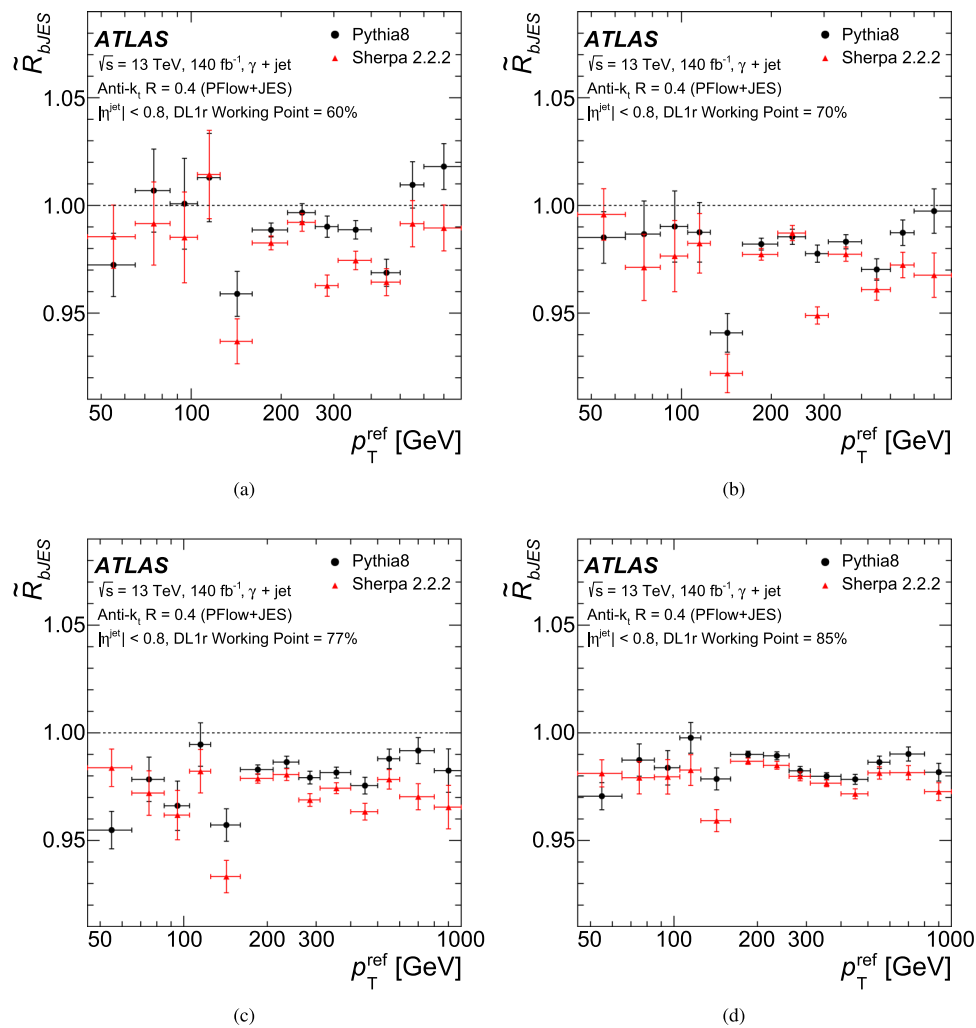


Fig. 21 \tilde{R}_{bJES} as a function of reference photon p_T determined using either PYTHIA8 or SHERPA for b-tagging working point with an efficiency of **a** 60%, **b** 70%, **c** 77% and **d** 85%. The error bars correspond to the statistical uncertainties

Table 3 \tilde{R}_{bJES} obtained for various b-tagging working points using PYTHIA8 and SHERPA separately for $85 < p_T < 1000$ GeV

WP	PYTHIA 8	Sherpa
60%	0.990 ± 0.010 (stat.) ± 0.013 (syst.)	0.984 ± 0.010 (stat.) ± 0.013 (syst.)
70%	0.984 ± 0.010 (stat.) ± 0.011 (syst.)	0.974 ± 0.010 (stat.) ± 0.012 (syst.)
77%	0.978 ± 0.006 (stat.) ± 0.011 (syst.)	0.966 ± 0.006 (stat.) ± 0.011 (syst.)
85%	0.989 ± 0.004 (stat.) ± 0.007 (syst.)	0.979 ± 0.004 (stat.) ± 0.007 (syst.)

Acknowledgements We thank CERN for the very successful operation of the LHC, as well as the support staff from our institutions without whom ATLAS could not be operated efficiently. We acknowledge the support of ANPCyT, Argentina; YerPhI, Armenia; ARC, Australia; BMWFW and FWF, Austria; ANAS, Azerbaijan; CNPq and FAPESP, Brazil; NSERC, NRC and CFI, Canada; CERN; ANID, Chile; CAS, MOST and NSFC, China; Minciencias, Colombia; MEYS CR, Czech Republic; DNRf and DNSRC, Denmark; IN2P3-CNRS and CEA-DRF/IRFU, France; SRNSFG, Georgia; BMBF, HGF and MPG, Germany; GSRI, Greece; RGC and Hong Kong SAR, China; ISF and Benozio Center, Israel; INFN, Italy; MEXT and JSPS, Japan; CNRST, Morocco; NWO, Netherlands; RCN, Norway; MEiN, Poland; FCT, Por-

tugal; MNE/IFA, Romania; MESTD, Serbia; MSSR, Slovakia; ARRS and MIZŠ, Slovenia; DSI/NRF, South Africa; MICINN, Spain; SRC and Wallenberg Foundation, Sweden; SERI, SNSF and Cantons of Bern and Geneva, Switzerland; MOST, Taiwan; TENMAK, Türkiye; STFC, United Kingdom; DOE and NSF, United States of America. In addition, individual groups and members have received support from BCKDF, CANARIE, Compute Canada and CRC, Canada; PRIMUS 21/SCI/017 and UNCE SCI/013, Czech Republic; COST, ERC, ERDF, Horizon 2020 and Marie Skłodowska-Curie Actions, European Union; Investissements d’Avenir Labex, Investissements d’Avenir IDEX and ANR, France; DFG and AvH Foundation, Germany; Herakleitos, Thales and Aristeia programmes co-financed by EU-ESF and the Greek NSRF,

Greece; BSF-NSF and MINERVA, Israel; Norwegian Financial Mechanism 2014–2021, Norway; NCN and NAWA, Poland; La Caixa Banking Foundation, CERCA Programme Generalitat de Catalunya and PROMETEO and GenT Programmes Generalitat Valenciana, Spain; Göran Gustafssons Stiftelse, Sweden; The Royal Society and Leverhulme Trust, United Kingdom. The crucial computing support from all WLCG partners is acknowledged gratefully, in particular from CERN, the ATLAS Tier-1 facilities at TRIUMF (Canada), NDGF (Denmark, Norway, Sweden), CC-IN2P3 (France), KIT/GridKA (Germany), INFN-CNAF (Italy), NL-T1 (Netherlands), PIC (Spain), ASGC (Taiwan), RAL (UK) and BNL (USA), the Tier-2 facilities worldwide and large non-WLCG resource providers. Major contributors of computing resources are listed in Ref. [69].

Open Access This article is licensed under a Creative Commons Attribution 4.0 International License, which permits use, sharing, adaptation, distribution and reproduction in any medium or format, as long as you give appropriate credit to the original author(s) and the source, provide a link to the Creative Commons licence, and indicate if changes were made. The images or other third party material in this article are included in the article's Creative Commons licence, unless indicated otherwise in a credit line to the material. If material is not included in the article's Creative Commons licence and your intended use is not permitted by statutory regulation or exceeds the permitted use, you will need to obtain permission directly from the copyright holder. To view a copy of this licence, visit <http://creativecommons.org/licenses/by/4.0/>.

Funded by SCOAP³. SCOAP³ supports the goals of the International Year of Basic Sciences for Sustainable Development.

References

1. M. Cacciari, G.P. Salam, G. Soyez, The anti- k_t jet clustering algorithm. *JHEP* **04**, 063 (2008). [arXiv:0802.1189](https://arxiv.org/abs/0802.1189) [hep-ph]
2. M. Cacciari, G.P. Salam, Dispelling the N^3 myth for the k_t jet-finder. *Phys. Lett. B* **641**, 57 (2006). [arXiv:hep-ph/0512210](https://arxiv.org/abs/hep-ph/0512210)
3. ATLAS Collaboration, Jet reconstruction and performance using particle flow with the ATLAS detector. *Eur. Phys. J. C* **77**, 466 (2017). [arXiv:1703.10485](https://arxiv.org/abs/1703.10485) [hep-ex]
4. ATLAS Collaboration, Jet energy scale and resolution measured in proton–proton collisions at $\sqrt{s} = 13$ TeV with the ATLAS detector. *Eur. Phys. J. C* **81**, 689 (2020). [arXiv:2007.02645](https://arxiv.org/abs/2007.02645) [hep-ex]
5. ATLAS Collaboration, Jet energy measurement with the ATLAS detector in proton–proton collisions at $\sqrt{s} = 7$ TeV. *Eur. Phys. J. C* **73**, 2304 (2013). [arXiv:1112.6426](https://arxiv.org/abs/1112.6426) [hep-ex]
6. ATLAS Collaboration, Jet energy measurement and its systematic uncertainty in proton–proton collisions at $\sqrt{s} = 7$ TeV with the ATLAS detector. *Eur. Phys. J. C* **75**, 17 (2015). [arXiv:1406.0076](https://arxiv.org/abs/1406.0076) [hep-ex]
7. ATLAS Collaboration, Determination of jet calibration and energy resolution in proton–proton collisions at $\sqrt{s} = 8$ TeV using the ATLAS detector. *Eur. Phys. J. C* **80**, 1104 (2020). [arXiv:1910.04482](https://arxiv.org/abs/1910.04482) [hep-ex]
8. ATLAS Collaboration, Jet energy scale measurements and their systematic uncertainties in proton–proton collisions at $\sqrt{s} = 13$ TeV with the ATLAS detector. *Phys. Rev. D* **96**, 072002 (2017). [arXiv:1703.09665](https://arxiv.org/abs/1703.09665) [hep-ex]
9. ATLAS Collaboration, In situ calibration of large-radius jet energy and mass in 13 TeV proton–proton collisions with the ATLAS detector. *Eur. Phys. J. C* **79**, 135 (2019). [arXiv:1807.09477](https://arxiv.org/abs/1807.09477) [hep-ex]
10. ATLAS Collaboration, The ATLAS Experiment at the CERN Large Hadron Collider. *JINST* **3**, S08003 (2008)
11. ATLAS Collaboration, ATLAS Insertable B-Layer: Technical Design Report, ATLAS-TDR-19; CERN-LHCC-2010-013, (2010). <https://cds.cern.ch/record/1291633> [Addendum: ATLAS-TDR-19-ADD-1, CERN-LHCC-2012-009, 2012, <https://cds.cern.ch/record/1451888>]
12. B. Abbott et al., Production and integration of the ATLAS Insertable B-Layer. *JINST* **13**, T05008 (2018). [arXiv:1803.00844](https://arxiv.org/abs/1803.00844) [physics.ins-det]
13. ATLAS Collaboration, Performance of the ATLAS trigger system in 2015. *Eur. Phys. J. C* **77**, 317 (2017). [arXiv:1611.09661](https://arxiv.org/abs/1611.09661) [hep-ex]
14. ATLAS Collaboration, The ATLAS Collaboration Software and Firmware, ATL-SOFT-PUB-2021-001 (2021). <https://cds.cern.ch/record/2767187>
15. ATLAS Collaboration, ATLAS data quality operations and performance for 2015.2018 data-taking. *JINST* **15**, P04003 (2020). [arXiv:1911.04632](https://arxiv.org/abs/1911.04632) [physics.ins-det]
16. T. Sjostrand et al., An introduction to PYTHIA 8.2. *Comput. Phys. Commun.* **191**, 159 (2015). [arXiv:1410.3012](https://arxiv.org/abs/1410.3012) [hep-ph]
17. ATLAS Collaboration, ATLAS Pythia 8 tunes to 7 TeV data, ATL-PHYS-PUB-2014-021 (2014). <https://cds.cern.ch/record/1966419>
18. R.D. Ball et al., Parton distributions with LHC data. *Nucl. Phys. B* **867**, 244 (2013). [arXiv:1207.1303](https://arxiv.org/abs/1207.1303) [hep-ph]
19. D.J. Lange, The EvtGen particle decay simulation package. *Nucl. Instrum. Meth. A* **462**, 152 (2001)
20. E. Bothmann et al., Event generation with Sherpa 2.2. *SciPost Phys.* **7**, 034 (2019). [arXiv:1905.09127](https://arxiv.org/abs/1905.09127) [hep-ph]
21. S. Schumann, F. Krauss, A parton shower algorithm based on Catani-Seymour dipole factorisation. *JHEP* **03**, 038 (2008). [arXiv:hep-ph/0709.1027](https://arxiv.org/abs/hep-ph/0709.1027) [hep-ph]
22. S. Dulat et al., New parton distribution functions from a global analysis of quantum chromodynamics. *Phys. Rev. D* **93**, 033006 (2016). [arXiv:1506.07443](https://arxiv.org/abs/1506.07443) [hep-ph]
23. J.-C. Winter, F. Krauss, G. Soff, A modified cluster-hadronization model. *Eur. Phys. J. C* **36**, 381 (2004). [arXiv:hep-ph/0311085](https://arxiv.org/abs/hep-ph/0311085)
24. T. Sjöstrand, S. Mrenna, P.Z. Skands, PYTHIA 6.4 physics and manual. *JHEP* **05**, 026 (2006). [arXiv:hep-ph/0603175](https://arxiv.org/abs/hep-ph/0603175)
25. M. Bähr et al., Herwig++ physics and manual. *Eur. Phys. J. C* **58**, 639 (2008). [arXiv:0803.0883](https://arxiv.org/abs/0803.0883) [hep-ph]
26. J. Bellm et al., Herwig 7.0/Herwig++ 3.0 release note. *Eur. Phys. J. C* **76**, 196 (2016). [arXiv:1512.01178](https://arxiv.org/abs/1512.01178) [hep-ph]
27. J. Bellm et al., Herwig 7.1 Release Note (2017). [arXiv:1705.06919](https://arxiv.org/abs/1705.06919) [hep-ph]
28. S. Gieseke, P. Stephens, B. Webber, New formalism for QCD parton showers. *JHEP* **12**, 045 (2003). [arXiv:hep-ph/0310083](https://arxiv.org/abs/hep-ph/0310083)
29. S. Plätzer, S. Gieseke, Coherent parton showers with local recoils. *JHEP* **01**, 024 (2011). [arXiv:0909.5593](https://arxiv.org/abs/0909.5593) [hep-ph]
30. S. Plätzer, S. Gieseke, Dipole showers and automated NLO matching in Herwig++. *Eur. Phys. J. C* **72**, 2187 (2012). [arXiv:1109.6256](https://arxiv.org/abs/1109.6256) [hep-ph]
31. S. Catani, M.H. Seymour, A General algorithm for calculating jet cross-sections in NLO QCD. *Nucl. Phys. B* **485**, 291 (1997) [Erratum: *Nucl. Phys. B* **510**, 503–504 (1998)]. [arXiv:hep-ph/9605323](https://arxiv.org/abs/hep-ph/9605323)
32. S. Catani, S. Dittmaier, M.H. Seymour, Z. Trocsanyi, The dipole formalism for next-to-leading order QCD calculations with massive partons. *Nucl. Phys. B* **627**, 189 (2002). [arXiv:hep-ph/0201036](https://arxiv.org/abs/hep-ph/0201036)
33. P. Nason, A new method for combining NLO QCD with shower Monte Carlo algorithms. *JHEP* **11**, 040 (2004). [arXiv:hep-ph/0409146](https://arxiv.org/abs/hep-ph/0409146)
34. S. Frixione, P. Nason, C. Oleari, Matching NLO QCD computations with parton shower simulations: the POWHEG method. *JHEP* **11**, 070 (2007). [arXiv:0709.2092](https://arxiv.org/abs/0709.2092) [hep-ph]
35. S. Alioli, P. Nason, C. Oleari, E. Re, A general framework for implementing NLO calculations in shower Monte Carlo programs: the POWHEG BOX. *JHEP* **06**, 043 (2010). [arXiv:1002.2581](https://arxiv.org/abs/1002.2581) [hep-ph]
36. R.D. Ball et al., Parton distributions for the LHC run II. *JHEP* **04**, 040 (2015). [arXiv:1410.8849](https://arxiv.org/abs/1410.8849) [hep-ph]

37. J. Alwall et al., The automated computation of tree-level and next-to-leading order differential cross sections, and their matching to parton shower simulations. *JHEP* **07**, 079 (2014). [arXiv:1405.0301](https://arxiv.org/abs/1405.0301) [hep-ph]
38. ATLAS Collaboration, Example ATLAS tunes of PYTHIA8, PYTHIA6 and POWHEG to an observable sensitive to Z boson transverse momentum, ATL-PHYS-PUB-2013-017 (2013). <https://cds.cern.ch/record/1629317>
39. ATLAS Collaboration, The Pythia 8 A3 tune description of ATLAS minimum bias and inelastic measurements incorporating the Donnachie–Landshoff diffractive model, ATL-PHYS-PUB-2016-017 (2016). <https://cds.cern.ch/record/2206965>
40. M. Cacciari, G.P. Salam, G. Soyez, FastJet user manual. *Eur. Phys. J. C* **72**, 1896 (2012). [arXiv:1111.6097](https://arxiv.org/abs/1111.6097) [hep-ph]
41. M. Cacciari, G.P. Salam, Pileup subtraction using jet areas. *Phys. Lett. B* **659**, 119 (2008). [arXiv:0707.1378](https://arxiv.org/abs/0707.1378) [hep-ph]
42. M. Cacciari, G.P. Salam, G. Soyez, The catchment area of jets. *JHEP* **04**, 005 (2008). [arXiv:0802.1188](https://arxiv.org/abs/0802.1188) [hep-ph]
43. S.D. Ellis, D.E. Soper, Successive combination jet algorithm for hadron collisions. *Phys. Rev. D* **48**, 3160 (1993). [arXiv:hep-ph/9305266](https://arxiv.org/abs/hep-ph/9305266)
44. S. Catani, Y.L. Dokshitzer, M.H. Seymour, B.R. Webber, Longitudinally-invariant k_{\perp} -clustering algorithms for hadron-hadron collisions. *Nucl. Phys. B* **406**, 187 (1993)
45. T. Gleisberg et al., Event generation with SHERPA 1.1. *JHEP* **02**, 007 (2009). [arXiv:0811.4622](https://arxiv.org/abs/0811.4622) [hep-ph]
46. H. Abreu et al., Performance of the electronic readout of the ATLAS liquid argon calorimeters. *J. Instrum.* **5**, P09003 (2010)
47. A. Cukierman, B. Nachman, Mathematical properties of numerical inversion for jet calibrations. *Nucl. Instrum. Meth. A* **858**, 1 (2017). [arXiv:1609.05195](https://arxiv.org/abs/1609.05195) [physics.data-an]
48. K. Pearson, X. On the criterion that a given system of deviations from the probable in the case of a correlated system of variables is such that it can be reasonably supposed to have arisen from random sampling. *Lond. Edinb. Dublin Philos. Mag. J. Sci.* **50**, 157 (1900)
49. H.B. Curry, I.J. Schoenberg, On Polya frequency functions IV: the fundamental spline functions and their limits. *J. Anal. Math.* **17**, 71 (1966)
50. B. Grimstad et al., SPLINTER: a library for multivariate function approximation with splines (2015). <http://github.com/bgrimstad/splinter>. Accessed 16 May 2015
51. ATLAS Collaboration, Dependence of the jet energy scale on the particle content of hadronic jets in the ATLAS detector simulation, ATL-PHYS-PUB-2022-021 (2022). <https://cds.cern.ch/record/2808016>
52. ATLAS Collaboration, Generalized numerical inversion: a neural network approach to jet calibration, ATL-PHYS-PUB-2018-013 (2018). <https://cds.cern.ch/record/2630972>
53. F. Chollet et al., Keras (2015). <https://keras.io>
54. D.P. Kingma, J. Ba, Adam: a method for stochastic optimization (2017). [arXiv:1412.6980](https://arxiv.org/abs/1412.6980) [cs.LG]
55. P. Ramachandran, B. Zoph, Q.V. Le, Searching for activation functions (2017). [arXiv:1710.05941](https://arxiv.org/abs/1710.05941)
56. S. Cheong, A. Cukierman, B. Nachman, M. Safdari, A. Schwartzman, Parametrizing the detector response with neural networks. *JINST* **15**, P01030 (2020). [arXiv:1910.03773](https://arxiv.org/abs/1910.03773) [physics.data-an]
57. ATLAS Collaboration, Determination of the jet energy scale and resolution at ATLAS using Z/ γ -jet events in data at $\sqrt{s} = 8$ TeV, ATLAS-CONF-2015-057 (2015). <https://cds.cern.ch/record/2059846>
58. ATLAS Collaboration, Selection of jets produced in 13 TeV proton–proton collisions with the ATLAS detector, ATLAS-CONF-2015-029 (2015). <https://cds.cern.ch/record/2037702>
59. ATLAS Collaboration, Tagging and suppression of pileup jets with the ATLAS detector, ATLAS-CONF-2014-018 (2014). <https://cds.cern.ch/record/1700870>
60. ATLAS Collaboration, Performance of pile-up mitigation techniques for jets in pp collisions at $\sqrt{s} = 8$ TeV using the ATLAS detector. *Eur. Phys. J. C* **76**, 581 (2016). [arXiv:1510.03823](https://arxiv.org/abs/1510.03823) [hep-ex]
61. V. Lendermann et al., Combining triggers in HEP data analysis. *Nucl. Instrum. Meth. A* **604**, 707 (2009). [arXiv:0901.4118](https://arxiv.org/abs/0901.4118) [hep-ex]
62. ATLAS Collaboration, Electron and photon performance measurements with the ATLAS detector using the 2015–2017 LHC proton–proton collision data. *JINST* **14**, P12006 (2019). [arXiv:1908.00005](https://arxiv.org/abs/1908.00005) [hep-ex]
63. ATLAS Collaboration, Muon reconstruction and identification efficiency in ATLAS using the full Run 2 pp collision data set at $\sqrt{s} = 13$ TeV. *Eur. Phys. J. C* **81**, 578 (2021). [arXiv:2012.00578](https://arxiv.org/abs/2012.00578) [hep-ex]
64. ATLAS Collaboration, Performance of the ATLAS muon triggers in Run 2. *JINST* **15**, P09015 (2020). [arXiv:2004.13447](https://arxiv.org/abs/2004.13447) [hep-ex]
65. ATLAS Collaboration, Performance of electron and photon triggers in ATLAS during LHC Run 2. *Eur. Phys. J. C* **80**, 47 (2020). [arXiv:1909.00761](https://arxiv.org/abs/1909.00761) [hep-ex]
66. ATLAS Collaboration, Measurement of the ratio of cross sections for inclusive isolated-photon production in pp collisions at $\sqrt{s} = 13$ and 8 TeV with the ATLAS detector. *JHEP* **04**, 093 (2019). [arXiv:1901.10075](https://arxiv.org/abs/1901.10075) [hep-ex]
67. ATLAS flavour-tagging algorithms for the LHC Run 2 pp collision dataset (2022). [arXiv:2211.16345](https://arxiv.org/abs/2211.16345) [physics.data-an]
68. ATLAS Collaboration, Direct constraint on the Higgs-charm coupling from a search for Higgs boson decays into charm quarks with the ATLAS detector. *Eur. Phys. J. C* **82**, 717 (2022). [arXiv:2201.11428](https://arxiv.org/abs/2201.11428) [hep-ex]
69. ATLAS Collaboration, ATLAS Computing Acknowledgements, ATL-SOFT-PUB-2021-003 (2021). <https://cds.cern.ch/record/2776662>

ATLAS Collaboration*

G. Aad¹⁰², B. Abbott¹²⁰, K. Abeling⁵⁵, N. J. Abicht⁴⁹, S. H. Abidi²⁹, A. Aboulhorma^{35e}, H. Abramowicz¹⁵¹, H. Abreu¹⁵⁰, Y. Abulaiti¹¹⁷, A. C. Abusleme Hoffman^{137a}, B. S. Acharya^{69a,69b,n}, C. Adam Bourdarios⁴, L. Adamczyk^{85a}, L. Adamek¹⁵⁵, S. V. Addepalli²⁶, M. J. Addison¹⁰¹, J. Adelman¹¹⁵, A. Adiguzel^{21c}, T. Adye¹³⁴, A. A. Affolder¹³⁶, Y. Afik³⁶, M. N. Agaras¹³, J. Agarwala^{73a,73b}, A. Aggarwal¹⁰⁰, C. Agheorghiesei^{27c}, A. Ahmad³⁶, F. Ahmadov^{38,y}, W. S. Ahmed¹⁰⁴, S. Ahuja⁹⁵, X. Ai^{62a}, G. Aielli^{76a,76b}, M. Ait Tamlhat^{35c}, B. Aitbenkhik^{35a}, I. Aizenberg¹⁶⁹, M. Akbiyik¹⁰⁰, T. P. A. Åkesson⁹⁸, A. V. Akimov³⁷, D. Akiyama¹⁶⁸, N. N. Akolkar²⁴, K. Al Khoury⁴¹, G. L. Alberghi^{23b}, J. Albert¹⁶⁵, P. Albicocco⁵³, G. L. Albouy⁶⁰, S. Alderweireldt⁵², M. Aleksa³⁶, I. N. Aleksandrov³⁸, C. Alexa^{27b}, T. Alexopoulos¹⁰, A. Alfonsi¹¹⁴, F. Alfonsi^{23b}, M. Alhroob¹²⁰, B. Ali¹³², S. Ali¹⁴⁸, M. Aliev³⁷, G. Alimonti^{71a}, W. Alkakh⁵⁵, C. Allaire⁶⁶, B. M. M. Allbrooke¹⁴⁶, C. A. Allendes Flores^{137f}, P. P. Allport²⁰, A. Aloisio^{72a,72b}, F. Alonso⁹⁰, C. Alpigiani¹³⁸, M. Alvarez Estevez⁹⁹, A. Alvarez Fernandez¹⁰⁰, M. G. Alviggi^{72a,72b}, M. Aly¹⁰¹, Y. Amaral Coutinho^{82b}, A. Ambler¹⁰⁴, C. Amelung³⁶, M. Amerl¹⁰¹, C. G. Ames¹⁰⁹, D. Amidei¹⁰⁶, S. P. Amor Dos Santos^{130a}, K. R. Amos¹⁶³, V. Ananiev¹²⁵, C. Anastopoulos¹³⁹, T. Andeen¹¹, J. K. Anders³⁶, S. Y. Andreev^{47a,47b}, A. Andreatta^{71a,71b}, S. Angelidakis⁹, A. Angerami^{41,aa}, A. V. Anisenkov³⁷, A. Annovi^{74a}, C. Antel⁵⁶, M. T. Anthony¹³⁹, E. Antipov¹⁴⁵, M. Antonelli⁵³, D. J. A. Antrim^{17a}, F. Anulli^{75a}, M. Aoki⁸³, T. Aoki¹⁵³, J. A. Aparisi Pozo¹⁶³, M. A. Aparo¹⁴⁶, L. Aperio Bella⁴⁸, C. Appelt¹⁸, N. Aranzabal³⁶, C. Arcangeletti⁵³, A. T. H. Arce⁵¹, E. Arena⁹², J.-F. Arguin¹⁰⁸, S. Argyropoulos⁵⁴, J.-H. Arling⁴⁸, A. J. Armbruster³⁶, O. Arnaez⁴, H. Arnold¹¹⁴, Z. P. Arrubarrena Tame¹⁰⁹, G. Artoni^{75a,75b}, H. Asada¹¹¹, K. Asai¹¹⁸, S. Asai¹⁵³, N. A. Asbah⁶¹, J. Assahsah^{35d}, K. Assamagan²⁹, R. Astalos^{28a}, S. Atashi¹⁶⁰, R. J. Atkin^{33a}, M. Atkinson¹⁶², N. B. Atlay¹⁸, H. Atmani^{62b}, P. A. Atlasiddha¹⁰⁶, K. Augsten¹³², S. Auricchio^{72a,72b}, A. D. Auriol²⁰, V. A. Austrup¹⁷¹, G. Avolio³⁶, K. Axiotis⁵⁶, G. Azuelos^{108,ac}, D. Babal^{28b}, H. Bachacou¹³⁵, K. Bachas^{152,q}, A. Bachiu³⁴, F. Backman^{47a,47b}, A. Badea⁶¹, P. Bagnaia^{75a,75b}, M. Bahmani¹⁸, A. J. Bailey¹⁶³, V. R. Bailey¹⁶², J. T. Baines¹³⁴, C. Bakalis¹⁰, O. K. Baker¹⁷², E. Bakos¹⁵, D. Bakshi Gupta⁸, R. Balasubramanian¹¹⁴, E. M. Baldin³⁷, P. Balek^{85a}, E. Ballabene^{23a,23b}, F. Balli¹³⁵, L. M. Baltes^{63a}, W. K. Balunas³², J. Balz¹⁰⁰, E. Banas⁸⁶, M. Bandieramonte¹²⁹, A. Bandyopadhyay²⁴, S. Bansal²⁴, L. Barak¹⁵¹, M. Barakat⁴⁸, E. L. Barberio¹⁰⁵, D. Barberis^{57a,57b}, M. Barbero¹⁰², G. Barbour⁹⁶, K. N. Barends^{33a}, T. Barillari¹¹⁰, M.-S. Barisits³⁶, T. Barklow¹⁴³, P. Baron¹²², D. A. Baron Moreno¹⁰¹, A. Baroncelli^{62a}, G. Barone²⁹, A. J. Barr¹²⁶, J. D. Barr⁹⁶, L. Barranco Navarro^{47a,47b}, F. Barreiro⁹⁹, J. Barreiro Guimarães da Costa^{14a}, U. Barron¹⁵¹, M. G. Barros Teixeira^{130a}, S. Barsov³⁷, F. Bartels^{63a}, R. Bartoldus¹⁴³, A. E. Barton⁹¹, P. Bartos^{28a}, A. Basan¹⁰⁰, M. Baselga⁴⁹, A. Bassalat^{66,ag}, M. J. Basso^{156a}, C. R. Basson¹⁰¹, R. L. Bates⁵⁹, S. Batlamous^{35e}, J. R. Batley³², B. Batool¹⁴¹, M. Battaglia¹³⁶, D. Battulga¹⁸, M. Bauce^{75a,75b}, M. Bauer³⁶, P. Bauer²⁴, L. T. Bazzano Hurrell³⁰, J. B. Beacham⁵¹, T. Beau¹²⁷, P. H. Beauchemin¹⁵⁸, F. Becherer⁵⁴, P. Bechtle²⁴, H. P. Beck^{19,p}, K. Becker¹⁶⁷, A. J. Beddall^{21d}, V. A. Bednyakov³⁸, C. P. Bee¹⁴⁵, L. J. Beamster¹⁵, T. A. Beermann³⁶, M. Begalli^{82d}, M. Begel²⁹, A. Behera¹⁴⁵, J. K. Behr⁴⁸, J. F. Beirer⁵⁵, F. Beisiegel²⁴, M. Belfkir¹⁵⁹, G. Bella¹⁵¹, L. Bellagamba^{23b}, A. Bellerive³⁴, P. Bellos²⁰, K. Beloborodov³⁷, N. L. Belyaev³⁷, D. Bencheikroun^{35a}, F. Bendebba^{35a}, Y. Benhammou¹⁵¹, M. Benoit²⁹, J. R. Bensinger²⁶, S. Bentvelsen¹¹⁴, L. Beresford⁴⁸, M. Beretta⁵³, E. Bergeas Kuutmann¹⁶¹, N. Berger⁴, B. Bergmann¹³², J. Beringer^{17a}, G. Bernardi⁵, C. Bernius¹⁴³, F. U. Bernlochner²⁴, F. Bernon^{36,102}, T. Berry⁹⁵, P. Berta¹³³, A. Berthold⁵⁰, I. A. Bertram⁹¹, S. Bethke¹¹⁰, A. Betti^{75a,75b}, A. J. Bevan⁹⁴, M. Bhamjee^{33c}, S. Bhatta¹⁴⁵, D. S. Bhattacharya¹⁶⁶, P. Bhattarai²⁶, V. S. Bhopatkar¹²¹, R. Bi^{29,ae}, R. M. Bianchi¹²⁹, G. Bianco^{23a,23b}, O. Biebel¹⁰⁹, R. Bielski¹²³, M. Biglietti^{77a}, T. R. V. Billoud¹³², M. Bindi⁵⁵, A. Bingul^{21b}, C. Bini^{75a,75b}, A. Biondini⁹², C. J. Birch-sykes¹⁰¹, G. A. Bird^{20,134}, M. Birman¹⁶⁹, M. Biros¹³³, T. Bisanz⁴⁹, E. Bisceglie^{43a,43b}, D. Biswas¹⁴¹, A. Bitadze¹⁰¹, K. Björke¹²⁵, I. Bloch⁴⁸, C. Blocker²⁶, A. Blue⁵⁹, U. Blumenschein⁹⁴, J. Blumenthal¹⁰⁰, G. J. Bobbink¹¹⁴, V. S. Bobrovnikov³⁷, M. Boehler⁵⁴, B. Boehm¹⁶⁶, D. Bogavac³⁶, A. G. Bogdanchikov³⁷, C. Bohm^{47a}, V. Boisvert⁹⁵, P. Bokan⁴⁸, T. Bold^{85a}, M. Bomben⁵, M. Bona⁹⁴, M. Boonekamp¹³⁵, C. D. Booth⁹⁵, A. G. Borbély⁵⁹, I. S. Bordulev³⁷, H. M. Borecka-Bielska¹⁰⁸, L. S. Borgna⁹⁶, G. Borissov⁹¹, D. Bortoletto¹²⁶, D. Boscherini^{23b}, M. Bosman¹³, J. D. Bossio Sola³⁶, K. Bouaouda^{35a}, N. Bouchhar¹⁶³, J. Boudreau¹²⁹, E. V. Bouhova-Thacker⁹¹, D. Boumediene⁴⁰, R. Bouquet⁵, A. Boveia¹¹⁹, J. Boyd³⁶, D. Boye²⁹, I. R. Boyko³⁸, J. Bracinik²⁰, N. Brahimi^{62d}, G. Brandt¹⁷¹, O. Brandt³², F. Braren⁴⁸, B. Brau¹⁰³, J. E. Brau¹²³, R. Brenner¹⁶⁹, L. Brenner¹¹⁴, R. Brenner¹⁶¹, S. Bressler¹⁶⁹, D. Britton⁵⁹

D. Britzger¹¹⁰, I. Brock²⁴, G. Brooijmans⁴¹, W. K. Brooks^{137f}, E. Brost²⁹, L. M. Brown^{165.1}, L. E. Bruce⁶¹, T. L. Bruckler¹²⁶, P. A. Bruckman de Renstrom⁸⁶, B. Brüers⁴⁸, D. Bruncko^{28b,*}, A. Bruni^{23b}, G. Bruni^{23b}, M. Bruschi^{23b}, N. Bruscolo^{75a,75b}, T. Buanes¹⁶, Q. Buat¹³⁸, D. Buchin¹¹⁰, A. G. Buckley⁵⁹, M. K. Bugge¹²⁵, O. Bulekov³⁷, B. A. Bullard¹⁴³, A. T. Bunka¹⁴², S. Burdin⁹², C. D. Burgard⁴⁹, A. M. Burger⁴⁰, B. Burghgrave⁸, O. Burlayenko⁵⁴, J. T. P. Burr³², C. D. Burton¹¹, J. C. Burzynski¹⁴², E. L. Busch⁴¹, V. Büscher¹⁰⁰, P. J. Bussey⁵⁹, J. M. Butler²⁵, C. M. Buttar⁵⁹, J. M. Butterworth⁹⁶, W. Buttinger¹³⁴, C. J. Buxo Vazquez¹⁰⁷, A. R. Buzykaev³⁷, G. Cabras^{23b}, S. Cabrera Urbán¹⁶³, D. Caforio⁵⁸, H. Cai¹²⁹, Y. Cai^{14a,14e}, V. M. M. Cairo³⁶, O. Cakir^{3a}, N. Calace³⁶, P. Calafiura^{17a}, G. Calderini¹²⁷, P. Calfayan⁶⁸, G. Callea⁵⁹, L. P. Caloba^{82b}, D. Calvet⁴⁰, S. Calvet⁴⁰, T. P. Calvet¹⁰², M. Calvetti^{74a,74b}, R. Camacho Toro¹²⁷, S. Camarda³⁶, D. Camarero Munoz²⁶, P. Camarri^{76a,76b}, M. T. Camerlingo^{72a,72b}, D. Cameron¹²⁵, C. Camincher¹⁶⁵, M. Campanelli⁹⁶, A. Camplani⁴², V. Canale^{72a,72b}, A. Canesse¹⁰⁴, M. Cano Bret⁸⁰, J. Cantero¹⁶³, Y. Cao¹⁶², F. Capocasa²⁶, M. Capua^{43a,43b}, A. Carbone^{71a,71b}, R. Cardarelli^{76a}, J. C. J. Cardenas⁸, F. Cardillo¹⁶³, T. Carli³⁶, G. Carlino^{72a}, J. I. Carlotto¹³, B. T. Carlson^{129,r}, E. M. Carlson^{156a,165}, L. Carminati^{71a,71b}, M. Carnesale^{75a,75b}, S. Caron¹¹³, E. Carquin^{137f}, S. Carrá^{71a,71b}, G. Carratta^{23a,23b}, F. Carrio Argos^{33g}, J. W. S. Carter¹⁵⁵, T. M. Carter⁵², M. P. Casado^{13.i}, M. Caspar⁴⁸, E. G. Castiglia¹⁷², F. L. Castillo⁴, L. Castillo Garcia¹³, V. Castillo Gimenez¹⁶³, N. F. Castro^{130a,130e}, A. Catinaccio³⁶, J. R. Catmore¹²⁵, V. Cavaliere²⁹, N. Cavalli^{23a,23b}, V. Cavasinni^{74a,74b}, Y. C. Cekmecelioglu⁴⁸, E. Celebi^{21a}, F. Celli¹²⁶, M. S. Centonze^{70a,70b}, K. Cerny¹²², A. S. Cerqueira^{82a}, A. Cerri¹⁴⁶, L. Cerrito^{76a,76b}, F. Cerutti^{17a}, B. Cervato¹⁴¹, A. Cervelli^{23b}, G. Cesarini⁵³, S. A. Cetin^{21d}, Z. Chadi^{35a}, D. Chakraborty¹¹⁵, M. Chala^{130f}, J. Chan¹⁷⁰, W. Y. Chan¹⁵³, J. D. Chapman³², E. Chapon¹³⁵, B. Chargeishvili^{149b}, D. G. Charlton²⁰, T. P. Charman⁹⁴, M. Chatterjee¹⁹, C. Chauhan¹³³, S. Chekanov⁶, S. V. Chekulaev^{156a}, G. A. Chelkov^{38,a}, A. Chen¹⁰⁶, B. Chen¹⁵¹, B. Chen¹⁶⁵, H. Chen^{14c}, H. Chen²⁹, J. Chen^{62c}, J. Chen¹⁴², M. Chen¹²⁶, S. Chen¹⁵³, S. J. Chen^{14c}, X. Chen^{62c}, X. Chen^{14b,ab}, Y. Chen^{62a}, C. L. Cheng¹⁷⁰, H. C. Cheng^{64a}, S. Cheong¹⁴³, A. Cheplakov³⁸, E. Cheremushkina⁴⁸, E. Cherepanova¹¹⁴, R. Cherkaoui El Moursli^{35e}, E. Cheu⁷, K. Cheung⁶⁵, L. Chevalier¹³⁵, V. Chiarella⁵³, G. Chiarelli^{74a}, N. Chiedde¹⁰², G. Chiodini^{70a}, A. S. Chisholm²⁰, A. Chitan^{27b}, M. Chitishvili¹⁶³, M. V. Chizhov³⁸, K. Choi¹¹, A. R. Chomont^{75a,75b}, Y. Chou¹⁰³, E. Y. S. Chow¹¹⁴, T. Chowdhury^{33g}, L. D. Christopher^{33g}, K. L. Chu¹⁶⁹, M. C. Chu^{64a}, X. Chu^{14a,14e}, J. Chudoba¹³¹, J. J. Chwastowski⁸⁶, D. Cieri¹¹⁰, K. M. Ciesla^{85a}, V. Cindro⁹³, A. Ciocio^{17a}, F. Ciroto^{72a,72b}, Z. H. Citron¹⁶⁹, M. Citterio^{71a}, D. A. Ciubotaru^{27b}, B. M. Ciungu¹⁵⁵, A. Clark⁵⁶, P. J. Clark⁵², J. M. Clavijo Columbie⁴⁸, S. E. Clawson⁴⁸, C. Clement^{47a,47b}, J. Clercx⁴⁸, L. Clissa^{23a,23b}, Y. Coadou¹⁰², M. Cobal^{69a,69c}, A. Coccaro^{57b}, R. F. Coelho Barrue^{130a}, R. Coelho Lopes De Sa¹⁰³, S. Coelli^{71a}, H. Cohen¹⁵¹, A. E. C. Coimbra^{71a,71b}, B. Cole⁴¹, J. Collot⁶⁰, P. Conde Muñio^{130a,130g}, M. P. Connell^{33c}, S. H. Connell^{33c}, I. A. Connelly⁵⁹, E. I. Conroy¹²⁶, F. Conventi^{72a,ad}, H. G. Cooke²⁰, A. M. Cooper-Sarkar¹²⁶, A. Cordeiro Oudot Choi¹²⁷, F. Cormier¹⁶⁴, L. D. Corpe⁴⁰, M. Corradi^{75a,75b}, F. Corriveau^{104,w}, A. Cortes-Gonzalez¹⁸, M. J. Costa¹⁶³, F. Costanza⁴, D. Costanzo¹³⁹, B. M. Cote¹¹⁹, G. Cowan⁹⁵, K. Cranmer¹¹⁷, D. Cremonini^{23a,23b}, S. Crépe-Renaudin⁶⁰, F. Crescioli¹²⁷, M. Cristinziani¹⁴¹, M. Cristoforetti^{78a,78b}, V. Croft¹¹⁴, J. E. Crosby¹²¹, G. Crosetti^{43a,43b}, A. Cueto⁹⁹, T. Cuhadar Donszelmann¹⁶⁰, H. Cui^{14a,14e}, Z. Cui⁷, W. R. Cunningham⁵⁹, F. Curcio^{43a,43b}, P. Czodrowski³⁶, M. M. Czurylo^{63b}, M. J. Da Cunha Sargedas De Sousa^{62a}, J. V. Da Fonseca Pinto^{82b}, C. Da Via¹⁰¹, W. Dabrowski^{85a}, T. Dado⁴⁹, S. Dabhi^{33g}, T. Dai¹⁰⁶, C. Dallapiccola¹⁰³, M. Dam⁴², G. D'amen²⁹, V. D'Amico¹⁰⁹, J. Damp¹⁰⁰, J. R. Dandoy¹²⁸, M. F. Daneri³⁰, M. Danninger¹⁴², V. Dao³⁶, G. Darbo^{57b}, S. Darmora⁶, S. J. Das²⁹, S. D'Auria^{71a,71b}, C. David^{156b}, T. Davidek¹³³, B. Davis-Purcell³⁴, I. Dawson⁹⁴, H. A. Day-hall¹³², K. De⁸, R. De Asmundis^{72a}, N. De Biase⁴⁸, S. De Castro^{23a,23b}, N. De Groot¹¹³, P. de Jong¹¹⁴, H. De la Torre¹⁰⁷, A. De Maria^{14c}, A. De Salvo^{75a}, U. De Sanctis^{76a,76b}, A. De Santo¹⁴⁶, J. B. De Vivie De Regie⁶⁰, D. V. Dedovich³⁸, J. Degens¹¹⁴, A. M. Deiana⁴⁴, F. Del Corso^{23a,23b}, J. Del Peso⁹⁹, F. Del Rio^{63a}, F. Deliot¹³⁵, C. M. Delitzsch⁴⁹, M. Della Pietra^{72a,72b}, D. Della Volpe⁵⁶, A. Dell'Acqua³⁶, L. Dell'Asta^{71a,71b}, M. Delmastro⁴, P. A. Delsart⁶⁰, S. Demers¹⁷², M. Demichev³⁸, S. P. Denisov³⁷, L. D'Eramo⁴⁰, D. Derendarz⁸⁶, F. Derue¹²⁷, P. Dervan⁹², K. Desch²⁴, C. Deutsch²⁴, F. A. Di Bello^{57a,57b}, A. Di Ciaccio^{76a,76b}, L. Di Ciaccio⁴, A. Di Domenico^{75a,75b}, C. Di Donato^{72a,72b}, A. Di Girolamo³⁶, G. Di Gregorio⁵, A. Di Luca^{78a,78b}, B. Di Micco^{77a,77b}, R. Di Nardo^{77a,77b}, C. Diaconu¹⁰², F. A. Dias¹¹⁴, T. Dias Do Vale¹⁴², M. A. Diaz^{137a,137b}, F. G. Diaz Capriles²⁴, M. Didenko¹⁶³, E. B. Diehl¹⁰⁶, L. Diehl⁵⁴, S. Díez Cornell⁴⁸, C. Díez Pardos¹⁴¹, C. Dimitriadis^{24,161}, A. Dimitrievska^{17a}, J. Dingfelder²⁴, I-M. Dinu^{27b}, S. J. Dittmeier^{63b}, F. Dittus³⁶, F. Djama¹⁰², T. Djobava^{149b}

J. I. Djuvsland¹⁶, C. Doglioni^{101,98}, J. Dolejsi¹³³, Z. Dolezal¹³³, M. Donadelli^{82c}, B. Dong¹⁰⁷, J. Domini⁴⁰, A. D'Onofrio^{77a,77b}, M. D'Onofrio⁹², J. Dopke¹³⁴, A. Doria^{72a}, N. Dos Santos Fernandes^{130a}, M. T. Dova⁹⁰, A. T. Doyle⁵⁹, M. A. Draguet¹²⁶, E. Dreyer¹⁶⁹, I. Drivas-koulouris¹⁰, A. S. Drobac¹⁵⁸, M. Drozdova⁵⁶, D. Du^{62a}, T. A. du Pree¹¹⁴, F. Dubinin³⁷, M. Dubovsky^{28a}, E. Duchovni¹⁶⁹, G. Duckeck¹⁰⁹, O. A. Ducu^{27b}, D. Duda⁵², A. Dudarev³⁶, E. R. Duden²⁶, M. D'uffizi¹⁰¹, L. Duflot⁶⁶, M. Dührssen³⁶, C. Dülsen¹⁷¹, A. E. Dumitriu^{27b}, M. Dunford^{63a}, S. Dungs⁴⁹, K. Dunne^{47a,47b}, A. Duperrin¹⁰², H. Duran Yildiz^{3a}, M. Düren⁵⁸, A. Durglishvili^{149b}, B. L. Dwyer¹¹⁵, G. I. Dyckes^{17a}, M. Dyndal^{85a}, S. Dysch¹⁰¹, B. S. Dziedzic⁸⁶, Z. O. Earnshaw¹⁴⁶, G. H. Eberwein¹²⁶, B. Eckerova^{28a}, S. Eggebrecht⁵⁵, M. G. Eggleston⁵¹, E. Egidio Purcino De Souza¹²⁷, L. F. Ehrke⁵⁶, G. Eigen¹⁶, K. Einsweiler^{17a}, T. Ekelof¹⁶¹, P. A. Ekman⁹⁸, Y. El Ghazali^{35b}, H. El Jarrari^{35e,148}, A. El Moussaouy^{35a}, V. Ellajosyula¹⁶¹, M. Ellert¹⁶¹, F. Ellinghaus¹⁷¹, A. A. Elliot⁹⁴, N. Ellis³⁶, J. Elmsheuser²⁹, M. Elsing³⁶, D. Emelianov¹³⁴, Y. Enari¹⁵³, I. Ene^{17a}, S. Epari¹³, J. Erdmann⁴⁹, P. A. Erland⁸⁶, M. Errenst¹⁷¹, M. Escalier⁶⁶, C. Escobar¹⁶³, E. Etzion¹⁵¹, G. Evans^{130a}, H. Evans⁶⁸, L. S. Evans⁹⁵, M. O. Evans¹⁴⁶, A. Ezhilov³⁷, S. Ezzarqtouni^{35a}, F. Fabbri⁵⁹, L. Fabbri^{23a,23b}, G. Facini⁹⁶, V. Fadeyev¹³⁶, R. M. Fakhruddinov³⁷, S. Falciano^{75a}, L. F. Falda Ulhoa Coelho³⁶, P. J. Falke²⁴, J. Faltova¹³³, C. Fan¹⁶², Y. Fan^{14a}, Y. Fang^{14a,14e}, M. Fanti^{71a,71b}, M. Faraj^{69a,69b}, Z. Farazpay⁹⁷, A. Farbin⁸, A. Farilla^{77a}, T. Farooque¹⁰⁷, S. M. Farrington⁵², F. Fassi^{35e}, D. Fassouliotis⁹, M. Fauci Giannelli^{76a,76b}, W. J. Fawcett³², L. Fayard⁶⁶, P. Federic¹³³, P. Federicova¹³¹, O. L. Fedin^{37.a}, G. Fedotov³⁷, M. Feickert¹⁷⁰, L. Feligioni¹⁰², A. Fell¹³⁹, D. E. Fellers¹²³, C. Feng^{62b}, M. Feng^{14b}, Z. Feng¹¹⁴, M. J. Fenton¹⁶⁰, A. B. Fenyuk³⁷, L. Ferencz⁴⁸, R. A. M. Ferguson⁹¹, S. I. Fernandez Luengo^{137f}, M. J. V. Fernoux¹⁰², J. Ferrando⁴⁸, A. Ferrari¹⁶¹, P. Ferrari^{113,114}, R. Ferrari^{73a}, D. Ferrere⁵⁶, C. Ferretti¹⁰⁶, F. Fiedler¹⁰⁰, A. Filipčić⁹³, E. K. Filmer¹, F. Filthaut¹¹³, M. C. N. Fiolhais^{130a,130c}, L. Fiorini¹⁶³, W. C. Fisher¹⁰⁷, T. Fitschen¹⁰¹, P. M. Fitzhugh¹³⁵, I. Fleck¹⁴¹, P. Fleischmann¹⁰⁶, T. Flick¹⁷¹, L. Flores¹²⁸, M. Flores^{33d}, L. R. Flores Castillo^{64a}, F. M. Follega^{78a,78b}, N. Fomin¹⁶, J. H. Foo¹⁵⁵, B. C. Forland⁶⁸, A. Formica¹³⁵, A. C. Forti¹⁰¹, E. Fortin³⁶, A. W. Fortman⁶¹, M. G. Foti^{17a}, L. Fountas⁹, D. Fournier⁶⁶, H. Fox⁹¹, P. Francavilla^{74a,74b}, S. Francescato⁶¹, S. Franchellucci⁵⁶, M. Franchini^{23a,23b}, S. Franchino^{63a}, D. Francis³⁶, L. Franco¹¹³, L. Franconi⁴⁸, M. Franklin⁶¹, G. Frattari²⁶, A. C. Freegard⁹⁴, W. S. Freund^{82b}, Y. Y. Frid¹⁵¹, N. Fritzsche⁵⁰, A. Froch⁵⁴, D. Froidevaux³⁶, J. A. Frost¹²⁶, Y. Fu^{62a}, M. Fujimoto¹¹⁸, E. Fullana Torregrosa^{163,*}, E. Furtado De Simas Filho^{82b}, M. Furukawa¹⁵³, J. Fuster¹⁶³, A. Gabrielli^{23a,23b}, A. Gabrielli¹⁵⁵, P. Gadow⁴⁸, G. Gagliardi^{57a,57b}, L. G. Gagnon^{17a}, E. J. Gallas¹²⁶, B. J. Gallop¹³⁴, K. K. Gan¹¹⁹, S. Ganguly¹⁵³, J. Gao^{62a}, Y. Gao⁵², F. M. Garay Walls^{137a,137b}, B. Garcia^{29,ae}, C. García¹⁶³, A. Garcia Alonso¹¹⁴, A. G. Garcia Caffaro¹⁷², J. E. García Navarro¹⁶³, M. Garcia-Sciveres^{17a}, G. L. Gardner¹²⁸, R. W. Gardner³⁹, N. Garelli¹⁵⁸, D. Garg⁸⁰, R. B. Garg¹⁴³, J. M. Gargan⁵², C. A. Garner¹⁵⁵, S. J. Gasiorowski¹³⁸, P. Gaspar^{82b}, G. Gaudio^{73a}, V. Gautam¹³, P. Gauzzi^{75a,75b}, I. L. Gavrilenko³⁷, A. Gavrilyuk³⁷, C. Gay¹⁶⁴, G. Gaycken⁴⁸, E. N. Gazis¹⁰, A. A. Geanta^{27b}, C. M. Gee¹³⁶, C. Gemme^{57b}, M. H. Genest⁶⁰, S. Gentile^{75a,75b}, S. George⁹⁵, W. F. George²⁰, T. Gerialis⁴⁶, P. Gessinger-Befurt³⁶, M. E. Geyik¹⁷¹, M. Ghneimat¹⁴¹, K. Ghorbanian⁹⁴, A. Ghosal¹⁴¹, A. Ghosh¹⁶⁰, A. Ghosh⁷, B. Giacobbe^{23b}, S. Giagu^{75a,75b}, P. Giannetti^{74a}, A. Giannini^{62a}, S. M. Gibson⁹⁵, M. Gignac¹³⁶, D. T. Gil^{85b}, A. K. Gilbert^{85a}, B. J. Gilbert⁴¹, D. Gillberg³⁴, G. Gilles¹¹⁴, N. E. K. Gillwald⁴⁸, L. Ginabat¹²⁷, D. M. Gingrich^{2,ac}, M. P. Giordani^{69a,69c}, P. F. Giraud¹³⁵, G. Giugliarelli^{69a,69c}, D. Giugni^{71a}, F. Giuli³⁶, I. Gkialas^{9,j}, L. K. Gladilin³⁷, C. Glasman⁹⁹, G. R. Gledhill¹²³, M. Glisic¹²³, I. Gnesi^{43b,f}, Y. Go^{29,ae}, M. Goblirsch-Kolb³⁶, B. Gocke⁴⁹, D. Godin¹⁰⁸, B. Gokturk^{21a}, S. Goldfarb¹⁰⁵, T. Golling⁵⁶, M. G. D. Gololo^{33g}, D. Golubkov³⁷, J. P. Gombas¹⁰⁷, A. Gomes^{130a,130b}, G. Gomes Da Silva¹⁴¹, A. J. Gomez Delegido¹⁶³, R. Gonçalo^{130a,130c}, G. Gonella¹²³, L. Gonella²⁰, A. Gongadze³⁸, F. Gonnella²⁰, J. L. Gonski⁴¹, R. Y. González Andana⁵², S. González de la Hoz¹⁶³, S. Gonzalez Fernandez¹³, R. Gonzalez Lopez⁹², C. Gonzalez Renteria^{17a}, R. Gonzalez Suarez¹⁶¹, S. Gonzalez-Sevilla⁵⁶, G. R. Gonzalvo Rodriguez¹⁶³, L. Goossens³⁶, P. A. Gorbounov³⁷, B. Gorini³⁶, E. Gorini^{70a,70b}, A. Gorišek⁹³, T. C. Gosart¹²⁸, A. T. Goshaw⁵¹, M. I. Gostkin³⁸, S. Goswami¹²¹, C. A. Gottardo³⁶, M. Gouighri^{35b}, V. Goumarre⁴⁸, A. G. Goussiou¹³⁸, N. Govender^{33c}, I. Grabowska-Bold^{85a}, K. Graham³⁴, E. Gramstad¹²⁵, S. Grancagnolo^{70a,70b}, M. Grandi¹⁴⁶, V. Gratchev^{37,*}, P. M. Gravila^{27f}, F. G. Gravili^{70a,70b}, H. M. Gray^{17a}, M. Greco^{70a,70b}, C. Grefe²⁴, I. M. Gregor⁴⁸, P. Grenier¹⁴³, C. Grieco¹³, A. A. Grillo¹³⁶, K. Grimm³¹, S. Grinstein^{13,t}, J.-F. Grivaz⁶⁶, E. Gross¹⁶⁹, J. Grosse-Knetter⁵⁵, C. Grud¹⁰⁶, J. C. Grundy¹²⁶, L. Guan¹⁰⁶, W. Guan¹⁷⁰, C. Gubbels¹⁶⁴, J. G. R. Guerrero Rojas¹⁶³, G. Guerrieri^{69a,69b}, F. Guescini¹¹⁰, R. Gugel¹⁰⁰, J. A. M. Guhit¹⁰⁶, A. Guida¹⁸, T. Guillemin⁴, E. Guillon^{167,134}, S. Guindon³⁶, F. Guo^{14a,14e}, J. Guo^{62c}, L. Guo⁴⁸

Y. Guo¹⁰⁶, R. Gupta⁴⁸, S. Gurbuz²⁴, S. S. Gurdasani⁵⁴, G. Gustavino³⁶, M. Guth⁵⁶, P. Gutierrez¹²⁰, L. F. Gutierrez Zagazeta¹²⁸, C. Gutschow⁹⁶, C. Gwenlan¹²⁶, C. B. Gwilliam⁹², E. S. Haaland¹²⁵, A. Haas¹¹⁷, M. Habedank⁴⁸, C. Haber^{17a}, H. K. Hadavand⁸, A. Hadei¹⁰⁰, S. Hadzic¹¹⁰, J. J. Hahn¹⁴¹, E. H. Haines⁹⁶, M. Haleem¹⁶⁶, J. Haley¹²¹, J. J. Hall¹³⁹, G. D. Hallowell¹⁰², L. Halser¹⁹, K. Hamano¹⁶⁵, H. Hamdaoui^{35c}, M. Hamer²⁴, G. N. Hamity⁵², E. J. Hampshire⁹⁵, J. Han^{62b}, K. Han^{62a}, L. Han^{14c}, L. Han^{62a}, S. Han^{17a}, Y. F. Han¹⁵⁵, K. Hanagaki⁸³, M. Hance¹³⁶, D. A. Hangal^{41,aa}, H. Hanif¹⁴², M. D. Hank¹²⁸, R. Hankache¹⁰¹, J. B. Hansen⁴², J. D. Hansen⁴², P. H. Hansen⁴², K. Hara¹⁵⁷, D. Harada⁵⁶, T. Harenberg¹⁷¹, S. Harkusha³⁷, Y. T. Harris¹²⁶, N. M. Harrison¹¹⁹, P. F. Harrison¹⁶⁷, N. M. Hartman¹⁴³, N. M. Hartmann¹⁰⁹, Y. Hasegawa¹⁴⁰, A. Hasib⁵², S. Haug¹⁹, R. Hauser¹⁰⁷, C. M. Hawkes²⁰, R. J. Hawkins³⁶, Y. Hayashi¹⁵³, S. Hayashida¹¹¹, D. Hayden¹⁰⁷, C. Hayes¹⁰⁶, R. L. Hayes¹¹⁴, C. P. Hays¹²⁶, J. M. Hays⁹⁴, H. S. Hayward⁹², F. He^{62a}, Y. He¹⁵⁴, Y. He¹²⁷, N. B. Heatley⁹⁴, V. Hedberg⁹⁸, A. L. Heggelund¹²⁵, N. D. Hehir⁹⁴, C. Heidegger⁵⁴, K. K. Heidegger⁵⁴, W. D. Heidorn⁸¹, J. Heilman³⁴, S. Heim⁴⁸, T. Heim^{17a}, J. G. Heinlein¹²⁸, J. J. Heinrich¹²³, L. Heinrich¹¹⁰, J. Hejbal¹³¹, L. Helary⁴⁸, A. Held¹⁷⁰, S. Hellesund¹⁶, C. M. Helling¹⁶⁴, S. Hellman^{47a,47b}, C. Helsens³⁶, R. C. W. Henderson⁹¹, L. Henkelmann³², A. M. Henriques Correia³⁶, H. Herde⁹⁸, Y. Hernández Jiménez¹⁴⁵, L. M. Herrmann²⁴, T. Herrmann⁵⁰, G. Herten⁵⁴, R. Hertenberger¹⁰⁹, L. Hervas³⁶, M. E. Hesping¹⁰⁰, N. P. Hessey^{156a}, H. Hibi⁸⁴, S. J. Hillier²⁰, J. R. Hinds¹⁰⁷, F. Hinterkeuser²⁴, M. Hirose¹²⁴, S. Hirose¹⁵⁷, D. Hirschbuehl¹⁷¹, T. G. Hitchings¹⁰¹, B. Hiti⁹³, J. Hobbs¹⁴⁵, R. Hobincu^{27c}, N. Hod¹⁶⁹, M. C. Hodgkinson¹³⁹, B. H. Hodgkinson³², A. Hoecker³⁶, J. Hofer⁴⁸, T. Holm²⁴, M. Holzbock¹¹⁰, L. B. A. H. Hommels³², B. P. Honan¹⁰¹, J. Hong^{62c}, T. M. Hong¹²⁹, B. H. Hooberman¹⁶², W. H. Hopkins⁶, Y. Horii¹¹¹, S. Hou¹⁴⁸, A. S. Howard⁹³, J. Howarth⁵⁹, J. Hoya⁶, M. Hrabovsky¹²², A. Hrynevich⁴⁸, T. Hryn'ova⁴, P. J. Hsu⁶⁵, S.-C. Hsu¹³⁸, Q. Hu⁴¹, Y. F. Hu^{14a,14e}, D. P. Huang⁹⁶, S. Huang^{64b}, X. Huang^{14c}, Y. Huang^{62a}, Y. Huang^{14a}, Z. Huang¹⁰¹, Z. Hubacek¹³², M. Huebner²⁴, F. Huegging²⁴, T. B. Huffman¹²⁶, C. A. Hugli⁴⁸, M. Huhtinen³⁶, S. K. Huiberts¹⁶, R. Hulsken¹⁰⁴, N. Huseynov^{12,a}, J. Huston¹⁰⁷, J. Huth⁶¹, R. Hyneman¹⁴³, G. Iacobucci⁵⁶, G. Iakovidis²⁹, I. Ibragimov¹⁴¹, L. Iconomidou-Fayard⁶⁶, P. Iengo^{72a,72b}, R. Iguchi¹⁵³, T. Iizawa⁵⁶, Y. Ikegami⁸³, N. Ilic¹⁵⁵, H. Imam^{35a}, T. Ingebretsen Carlson^{47a,47b}, G. Introzzi^{73a,73b}, M. Iodice^{77a}, V. Ippolito^{75a,75b}, R. K. Irwin⁹², M. Ishino¹⁵³, W. Islam¹⁷⁰, C. Issever^{18,48}, S. Istin^{21a}, H. Ito¹⁶⁸, J. M. Iturbe Ponce^{64a}, R. Iuppa^{78a,78b}, A. Ivina¹⁶⁹, J. M. Izen⁴⁵, V. Izzo^{72a}, P. Jacka^{131,132}, P. Jackson¹, R. M. Jacobs⁴⁸, B. P. Jaeger¹⁴², C. S. Jagfeld¹⁰⁹, P. Jain⁵⁴, G. Jäkel¹⁷¹, K. Jakobs⁵⁴, T. Jakoubek¹⁶⁹, J. Jamieson⁵⁹, K. W. Janas^{85a}, A. E. Jaspan⁹², M. Javurkova¹⁰³, F. Jeanneau¹³⁵, L. Jeanty¹²³, J. Jejelava^{149a,z}, P. Jenni^{54,g}, C. E. Jessiman³⁴, S. Jézéquel⁴, C. Jia^{62b}, J. Jia¹⁴⁵, X. Jia⁶¹, X. Jia^{14a,14e}, Z. Jia^{14c}, Y. Jiang^{62a}, S. Jiggins⁴⁸, J. Jimenez Pena¹¹⁰, S. Jin^{14c}, A. Jinaru^{27b}, O. Jinnouchi¹⁵⁴, P. Johansson¹³⁹, K. A. Johns⁷, J. W. Johnson¹³⁶, D. M. Jones³², E. Jones⁴⁸, P. Jones³², R. W. L. Jones⁹¹, T. J. Jones⁹², R. Joshi¹¹⁹, J. Jovicevic¹⁵, X. Ju^{17a}, J. J. Junggeburth³⁶, T. Junkermann^{63a}, A. Juste Rozas^{13,t}, M. K. Jurek⁸⁶, S. Kabana^{137e}, A. Kaczmarska⁸⁶, M. Kado¹¹⁰, H. Kagan¹¹⁹, M. Kagan¹⁴³, A. Kahn⁴¹, A. Kahn¹²⁸, C. Kahra¹⁰⁰, T. Kaji¹⁶⁸, E. Kajomovitz¹⁵⁰, N. Kakati¹⁶⁹, I. Kalaitzidou⁵⁴, C. W. Kalderon²⁹, A. Kamenshchikov¹⁵⁵, S. Kanayama¹⁵⁴, N. J. Kang¹³⁶, D. Kar^{33g}, K. Karava¹²⁶, M. J. Kareem^{156b}, E. Karentzos⁵⁴, I. Karkanas¹⁵², O. Karkout¹¹⁴, S. N. Karpov³⁸, Z. M. Karpova³⁸, V. Kartvelishvili⁹¹, A. N. Karyukhin³⁷, E. Kasimi¹⁵², J. Katzy⁴⁸, S. Kaur³⁴, K. Kawade¹⁴⁰, T. Kawamoto¹³⁵, E. F. Kay³⁶, F. I. Kaya¹⁵⁸, S. Kazakov¹³, V. F. Kazanin³⁷, Y. Ke¹⁴⁵, J. M. Keaveney^{33a}, R. Keeler¹⁶⁵, G. V. Kehris⁶¹, J. S. Keller³⁴, A. S. Kelly⁹⁶, J. J. Kempster¹⁴⁶, K. E. Kennedy⁴¹, P. D. Kennedy¹⁰⁰, O. Kepka¹³¹, B. P. Kerridge¹⁶⁷, S. Kersten¹⁷¹, B. P. Kerševan⁹³, S. Keshri⁶⁶, L. Keszeghova^{28a}, S. Ketabchi Haghighat¹⁵⁵, M. Khandoga¹²⁷, A. Khanov¹²¹, A. G. Kharlamov³⁷, T. Kharlamova³⁷, E. E. Khoda¹³⁸, T. J. Khoo¹⁸, G. Khorauli¹⁶⁶, J. Khubua^{149b}, Y. A. R. Khwaira⁶⁶, M. Kiehn³⁶, A. Kilgallon¹²³, D. W. Kim^{47a,47b}, Y. K. Kim³⁹, N. Kimura⁹⁶, A. Kirchhoff⁵⁵, C. Kirfel²⁴, F. Kirfel²⁴, J. Kirk¹³⁴, A. E. Kiryunin¹¹⁰, C. Kitsaki¹⁰, O. Kivernyk²⁴, M. Klassen^{63a}, C. Klein³⁴, L. Klein¹⁶⁶, M. H. Klein¹⁰⁶, M. Klein⁹², S. B. Klein⁵⁶, U. Klein⁹², P. Klimek³⁶, A. Klimentov²⁹, T. Klioutchnikova³⁶, P. Kluit¹¹⁴, S. Kluth¹¹⁰, E. Kneringer⁷⁹, T. M. Knight¹⁵⁵, A. Knue⁵⁴, R. Kobayashi⁸⁷, S. F. Koch¹²⁶, M. Kocian¹⁴³, P. Kodyš¹³³, D. M. Koeck¹²³, P. T. Koenig²⁴, T. Koffas³⁴, M. Kolb¹³⁵, I. Koletsou⁴, T. Komarek¹²², K. Köneke⁵⁴, A. X. Y. Kong¹, T. Kono¹¹⁸, N. Konstantinidis⁹⁶, B. Konya⁹⁸, R. Kopeliansky⁶⁸, S. Koperny^{85a}, K. Korcyl⁸⁶, K. Kordas^{152,e}, G. Koren¹⁵¹, A. Korn⁹⁶, S. Korn⁵⁵, I. Korolkov¹³, N. Korotkova³⁷, B. Kortman¹¹⁴, O. Kortner¹¹⁰, S. Kortner¹¹⁰, W. H. Kostecka¹¹⁵, V. V. Kostyukhin¹⁴¹, A. Kotskechagia¹³⁵, A. Kotwal⁵¹, A. Koulouris³⁶, A. Kourkouveli-Charalampidi^{73a,73b}, C. Kourkouvelis⁹, E. Kourlitis⁶, O. Kovanda¹⁴⁶, R. Kowalewski¹⁶⁵, W. Kozanecki¹³⁵, A. S. Kozhin³⁷

V. A. Kramarenko³⁷, G. Kramberger⁹³, P. Kramer¹⁰⁰, M. W. Krasny¹²⁷, A. Krasznahorkay³⁶, J. A. Kremer¹⁰⁰, T. Kresse⁵⁰, J. Kretzschmar⁹², K. Kreul¹⁸, P. Krieger¹⁵⁵, S. Krishnamurthy¹⁰³, M. Krivos¹³³, K. Krizka²⁰, K. Kroeninger⁴⁹, H. Kroha¹¹⁰, J. Kroll¹³¹, J. Kroll¹²⁸, K. S. Krowpman¹⁰⁷, U. Kruchonak³⁸, H. Krüger²⁴, N. Krumnack⁸¹, M. C. Kruse⁵¹, J. A. Krzysiak⁸⁶, O. Kuchinskaia³⁷, S. Kuday^{3a}, S. Kuehn³⁶, R. Kuesters⁵⁴, T. Kuhl⁴⁸, V. Kukhtin³⁸, Y. Kulchitsky^{37,a}, S. Kuleshov^{137d,137b}, M. Kumar^{33g}, N. Kumari¹⁰², A. Kupco¹³¹, T. Kupfer⁴⁹, A. Kupich³⁷, O. Kuprash⁵⁴, H. Kurashige⁸⁴, L. L. Kurchaninov^{156a}, O. Kurdysh⁶⁶, Y. A. Kurochkin³⁷, A. Kurova³⁷, M. Kuze¹⁵⁴, A. K. Kvam¹⁰³, J. Kvita¹²², T. Kwan¹⁰⁴, N. G. Kyriacou¹⁰⁶, L. A. O. Laatu¹⁰², C. Lacasta¹⁶³, F. Lacava^{75a,75b}, H. Lacker¹⁸, D. Lacour¹²⁷, N. N. Lad⁹⁶, E. Ladygin³⁸, B. Laforge¹²⁷, T. Lagouri^{137e}, S. Lai⁵⁵, I. K. Lakomic^{85a}, N. Lalloue⁶⁰, J. E. Lambert^{165,1}, S. Lammers⁶⁸, W. Lampl⁷, C. Lampoudis^{152,e}, A. N. Lancaster¹¹⁵, E. Lançon²⁹, U. Landgraf⁵⁴, M. P. J. Landon⁹⁴, V. S. Lang⁵⁴, R. J. Langenberg¹⁰³, O. K. B. Langrekken¹²⁵, A. J. Lankford¹⁶⁰, F. Lanni³⁶, K. Lantzsch²⁴, A. Lanza^{73a}, A. Lapertosa^{57a,57b}, J. F. Laporte¹³⁵, T. Lari^{71a}, F. Lasagni Manghi^{23b}, M. Lassnig³⁶, V. Latonova¹³¹, A. Laudrain¹⁰⁰, A. Laurier¹⁵⁰, S. D. Lawlor⁹⁵, Z. Lawrence¹⁰¹, M. Lazzaroni^{71a,71b}, B. Le¹⁰¹, E. M. Le Boulicaut⁵¹, B. Leban⁹³, A. Lebedev⁸¹, M. LeBlanc³⁶, F. Ledroit-Guillon⁶⁰, A. C. A. Lee⁹⁶, S. C. Lee¹⁴⁸, S. Lee^{47a,47b}, T. F. Lee⁹², L. L. Leeuw^{33c}, H. P. Lefebvre⁹⁵, M. Lefebvre¹⁶⁵, C. Leggett^{17a}, K. Lehmann¹⁴², G. Lehmann Miotto³⁶, M. Leigh⁵⁶, W. A. Leight¹⁰³, W. Leinonen¹¹³, A. Leisos^{152,s}, M. A. L. Leite^{82c}, C. E. Leitgeb⁴⁸, R. Leitner¹³³, K. J. C. Leney⁴⁴, T. Lenz²⁴, S. Leone^{74a}, C. Leonidopoulos⁵², A. Leopold¹⁴⁴, C. Leroy¹⁰⁸, R. Les¹⁰⁷, C. G. Lester³², M. Levchenko³⁷, J. Levêque⁴, D. Levin¹⁰⁶, L. J. Levinson¹⁶⁹, M. P. Lewicki⁸⁶, D. J. Lewis⁴, A. Li⁵, B. Li^{62b}, C. Li^{62a}, C.-Q. Li^{62c}, H. Li^{62a}, H. Li^{62b}, H. Li^{14c}, H. Li^{62b}, K. Li¹³⁸, L. Li^{62c}, M. Li^{14a,14e}, Q. Y. Li^{62a}, S. Li^{14a,14e}, S. Li^{62c,62d,d}, T. Li^{5,b}, X. Li¹⁰⁴, Z. Li¹²⁶, Z. Li¹⁰⁴, Z. Li⁹², Z. Li^{14a,14e}, Z. Liang^{14a}, M. Liberatore⁴⁸, B. Liberti^{76a}, K. Lie^{64c}, J. Lieber Marin^{82b}, H. Lien⁶⁸, K. Lin¹⁰⁷, R. E. Lindley⁷, J. H. Lindon², A. Linss⁴⁸, E. Lipeles¹²⁸, A. Lipniacka¹⁶, A. Lister¹⁶⁴, J. D. Little⁴, B. Liu^{14a}, B. X. Liu¹⁴², D. Liu^{62c,62d}, J. B. Liu^{62a}, J. K. K. Liu³², K. Liu^{62c,62d}, M. Liu^{62a}, M. Y. Liu^{62a}, P. Liu^{14a}, Q. Liu^{62c,62d,138}, X. Liu^{62a}, Y. Liu^{14d,14e}, Y. L. Liu¹⁰⁶, Y. W. Liu^{62a}, J. Llorente Merino¹⁴², S. L. Lloyd⁹⁴, E. M. Lobodzinska⁴⁸, P. Loch⁷, S. Loffredo^{76a,76b}, T. Lohse¹⁸, K. Lohwasser¹³⁹, E. Loiacono⁴⁸, M. Lokajicek¹³¹, J. D. Lomas²⁰, J. D. Long¹⁶², I. Longarini¹⁶⁰, L. Longo^{70a,70b}, R. Longo¹⁶², I. Lopez Paz⁶⁷, A. Lopez Solis⁴⁸, J. Lorenz¹⁰⁹, N. Lorenzo Martinez⁴, A. M. Lory¹⁰⁹, O. Loseva³⁷, X. Lou^{47a,47b}, X. Lou^{14a,14e}, A. Lounis⁶⁶, J. Love⁶, P. A. Love⁹¹, G. Lu^{14a,14e}, M. Lu⁸⁰, S. Lu¹²⁸, Y. J. Lu⁶⁵, H. J. Lubatti¹³⁸, C. Luci^{75a,75b}, F. L. Lucio Alves^{14c}, A. Lucotte⁶⁰, F. Luehring⁶⁸, I. Luise¹⁴⁵, O. Lukianchuk⁶⁶, O. Lundberg¹⁴⁴, B. Lund-Jensen¹⁴⁴, N. A. Luongo¹²³, M. S. Lutz¹⁵¹, D. Lynn²⁹, H. Lyons⁹², R. Lysak¹³¹, E. Lytken⁹⁸, V. Lyubushkin³⁸, T. Lyubushkina³⁸, M. M. Lyukova¹⁴⁵, H. Ma²⁹, L. L. Ma^{62b}, Y. Ma¹²¹, D. M. Mac Donell¹⁶⁵, G. Maccarrone⁵³, J. C. MacDonald¹⁰⁰, R. Madar⁴⁰, W. F. Mader⁵⁰, J. Maeda⁸⁴, T. Maeno²⁹, M. Maerker⁵⁰, H. Maguire¹³⁹, A. Maio^{130a,130b,130d}, K. Maj^{85a}, O. Majersky⁴⁸, S. Majewski¹²³, N. Makovec⁶⁶, V. Maksimovic¹⁵, B. Malaescu¹²⁷, Pa. Malecki⁸⁶, V. P. Maleev³⁷, F. Malek⁶⁰, M. Mali⁹³, D. Malito^{95,o}, U. Mallik⁸⁰, S. Maltezos¹⁰, S. Malyukov³⁸, J. Mamuzic¹³, G. Mancini⁵³, G. Manco^{73a,73b}, J. P. Mandalia⁹⁴, I. Mandić⁹³, L. Manhaes de Andrade Filho^{82a}, I. M. Maniatis¹⁶⁹, J. Manjarres Ramos¹⁰², D. C. Mankad¹⁶⁹, A. Mann¹⁰⁹, B. Mansoulie¹³⁵, S. Manzoni³⁶, A. Marantis¹⁵², G. Marchiori⁵, M. Marcisovsky¹³¹, C. Marcon^{71a,71b}, M. Marinescu²⁰, M. Marjanovic¹²⁰, E. J. Marshall⁹¹, Z. Marshall^{17a}, S. Marti-Garcia¹⁶³, T. A. Martin¹⁶⁷, V. J. Martin⁵², B. Martin dit Latour¹⁶, L. Martinelli^{75a,75b}, M. Martinez^{13,t}, P. Martinez Agullo¹⁶³, V. I. Martinez Outschoorn¹⁰³, P. Martinez Suarez¹³, S. Martin-Haugh¹³⁴, V. S. Martoiu^{27b}, A. C. Martyniuk⁹⁶, A. Marzin³⁶, D. Mascione^{78a,78b}, L. Masetti¹⁰⁰, T. Mashimo¹⁵³, J. Masik¹⁰¹, A. L. Maslennikov³⁷, L. Massa^{23b}, P. Massarotti^{72a,72b}, P. Mastrandrea^{74a,74b}, A. Mastroberardino^{43a,43b}, T. Masubuchi¹⁵³, T. Mathisen¹⁶¹, J. Matousek¹³³, N. Matsuzawa¹⁵³, J. Maurer^{27b}, B. Maček⁹³, D. A. Maximov³⁷, R. Mazini¹⁴⁸, I. Maznas¹⁵², M. Mazza¹⁰⁷, S. M. Mazza¹³⁶, E. Mazzeo^{71a,71b}, C. Mc Ginn²⁹, J. P. Mc Gowan¹⁰⁴, S. P. Mc Kee¹⁰⁶, E. F. McDonald¹⁰⁵, A. E. McDougall¹¹⁴, J. A. Mcfayden¹⁴⁶, R. P. McGovern¹²⁸, G. Mchedlidze^{149b}, R. P. McKenzie^{33g}, T. C. McLachlan⁴⁸, D. J. McLaughlin⁹⁶, K. D. McLean¹⁶⁵, S. J. McMahon¹³⁴, P. C. McNamara¹⁰⁵, C. M. Mcpartland⁹², R. A. McPherson^{165,w}, S. Mehlhase¹⁰⁹, A. Mehta⁹², D. Melini¹⁵⁰, B. R. Mellado Garcia^{33g}, A. H. Melo⁵⁵, F. Meloni⁴⁸, A. M. Mendes Jacques Da Costa¹⁰¹, H. Y. Meng¹⁵⁵, L. Meng⁹¹, S. Menke¹¹⁰, M. Mentink³⁶, E. Meoni^{43a,43b}, C. Merlassino¹²⁶, L. Merola^{72a,72b}, C. Meroni^{71a}, G. Merz¹⁰⁶, O. Meshkov³⁷, J. Metcalfe⁶, A. S. Mete⁶, C. Meyer⁶⁸, J.-P. Meyer¹³⁵, R. P. Middleton¹³⁴, L. Mijović⁵², G. Mikenberg¹⁶⁹, M. Mikesikova¹³¹, M. Mikuz⁹³, H. Mildner¹⁰⁰, A. Milic³⁶, C. D. Milke⁴⁴, D. W. Miller³⁹, L. S. Miller³⁴

A. Milov¹⁶⁹, D. A. Milstead^{47a,47b}, T. Min^{14c}, A. A. Minaenko³⁷, I. A. Minashvili^{149b}, L. Mince⁵⁹, A. I. Mincer¹¹⁷, B. Mindur^{85a}, M. Mineev³⁸, Y. Mino⁸⁷, L. M. Mir¹³, M. Miralles Lopez¹⁶³, M. Mironova^{17a}, A. Mishima¹⁵³, M. C. Missio¹¹³, T. Mitani¹⁶⁸, A. Mitra¹⁶⁷, V. A. Mitsou¹⁶³, O. Miú¹⁵⁵, P. S. Miyagawa⁹⁴, Y. Miyazaki⁸⁹, A. Mizukami⁸³, T. Mkrtchyan^{63a}, M. Mlinarevic⁹⁶, T. Mlinarevic⁹⁶, M. Mlynarikova³⁶, S. Mobius⁵⁵, K. Mochizuki¹⁰⁸, P. Moder⁴⁸, P. Mogg¹⁰⁹, A. F. Mohammed^{14a,14e}, S. Mohapatra⁴¹, G. Mokgatitwane^{33g}, L. Moleri¹⁶⁹, B. Mondal¹⁴¹, S. Mondal¹³², G. Monig¹⁴⁶, K. Möni⁴⁸, E. Monnier¹⁰², L. Monsonis Romero¹⁶³, J. Montejo Berlingen^{13,83}, M. Montella¹¹⁹, F. Monticelli⁹⁰, S. Monzani^{69a,69c}, N. Morange⁶⁶, A. L. Moreira De Carvalho^{130a}, M. Moreno Llácer¹⁶³, C. Moreno Martinez⁵⁶, P. Morettini^{57b}, S. Morgenstern³⁶, M. Morii⁶¹, M. Morinaga¹⁵³, A. K. Morley³⁶, F. Morodei^{75a,75b}, L. Morvaj³⁶, P. Moschovakos³⁶, B. Moser³⁶, M. Mosidze^{149b}, T. Moskalets⁵⁴, P. Moskvitina¹¹³, J. Moss^{31,m}, E. J. W. Moyse¹⁰³, O. Mtintsilana^{33g}, S. Muanza¹⁰², J. Mueller¹²⁹, D. Muenstermann⁹¹, R. Müller¹⁹, G. A. Mullier¹⁶¹, A. J. Mullin³², J. J. Mullin¹²⁸, D. P. Mungo¹⁵⁵, D. Munoz Perez¹⁶³, F. J. Munoz Sanchez¹⁰¹, M. Murin¹⁰¹, W. J. Murray^{167,134}, A. Murrone^{71a,71b}, J. M. Muse¹²⁰, M. Muškinja^{17a}, C. Mwewa²⁹, A. G. Myagkov^{37,a}, A. J. Myers⁸, A. A. Myers¹²⁹, G. Myers⁶⁸, M. Myska¹³², B. P. Nachman^{17a}, O. Nackenhorst⁴⁹, A. Nag⁵⁰, K. Nagai¹²⁶, K. Nagano⁸³, J. L. Nagle^{29,ae}, E. Nagy¹⁰², A. M. Nairz³⁶, Y. Nakahama⁸³, K. Nakamura⁸³, K. Nakkalil⁵, H. Nanjo¹²⁴, R. Narayan⁴⁴, E. A. Narayanan¹¹², I. Naryshkin³⁷, M. Naseri³⁴, S. Nasri¹⁵⁹, C. Nass²⁴, G. Navarro^{22a}, J. Navarro-Gonzalez¹⁶³, R. Nayak¹⁵¹, A. Nayaz¹⁸, P. Y. Nechaeva³⁷, F. Nechansky⁴⁸, L. Nedic¹²⁶, T. J. Neep²⁰, A. Negri^{73a,73b}, M. Negrini^{23b}, C. Nellist¹¹⁴, C. Nelson¹⁰⁴, K. Nelson¹⁰⁶, S. Nemecek¹³¹, M. Nessi^{36,h}, M. S. Neubauer¹⁶², F. Neuhaus¹⁰⁰, J. Neundorfer⁴⁸, R. Newhouse¹⁶⁴, P. R. Newman²⁰, C. W. Ng¹²⁹, Y. W. Y. Ng⁴⁸, B. Ngair^{35e}, H. D. N. Nguyen¹⁰⁸, R. B. Nickerson¹²⁶, R. Nicolaidou¹³⁵, J. Nielsen¹³⁶, M. Niemeyer⁵⁵, J. Niemann^{55,36}, N. Nikiforou³⁶, V. Nikolaenko^{37,a}, I. Nikolic-Audit¹²⁷, K. Nikolopoulos²⁰, P. Nilsson²⁹, I. Ninca⁴⁸, H. R. Nindhito⁵⁶, G. Ninio¹⁵¹, A. Nisati^{75a}, N. Nishu², R. Nisius¹¹⁰, J.-E. Nitschke⁵⁰, E. K. Nkadimeng^{33g}, S. J. Noacco Rosende⁹⁰, T. Nobe¹⁵³, D. L. Noel³², T. Nommensen¹⁴⁷, M. B. Norfolk¹³⁹, R. R. B. Norisam⁹⁶, B. J. Norman³⁴, J. Novak⁹³, T. Novak⁴⁸, L. Novotny¹³², R. Novotny¹¹², L. Nozka¹²², K. Ntekas¹⁶⁰, N. M. J. Nunes De Moura Junior^{82b}, E. Nurse⁹⁶, J. Ocariz¹²⁷, A. Ochi⁸⁴, I. Ochoa^{130a}, S. Oerdek¹⁶¹, J. T. Offermann³⁹, A. Ogrodnik¹³³, A. Oh¹⁰¹, C. C. Ohm¹⁴⁴, H. Oide⁸³, R. Oishi¹⁵³, M. L. Ojeda⁴⁸, Y. Okazaki⁸⁷, M. W. O'Keefe⁹², Y. Okumura¹⁵³, L. F. Oleiro Seabra^{130a}, S. A. Olivares Pino^{137d}, D. Oliveira Damazio²⁹, D. Oliveira Goncalves^{82a}, J. L. Oliver¹⁶⁰, M. J. R. Olsson¹⁶⁰, A. Olszewski⁸⁶, Ö. O. Öncel⁵⁴, D. C. O'Neil¹⁴², A. P. O'Neill¹⁹, A. Onofre^{130a,130e}, P. U. E. Onyisi¹¹, M. J. Oreglia³⁹, G. E. Orellana⁹⁰, D. Orestano^{77a,77b}, N. Orlando¹³, R. S. Orr¹⁵⁵, V. O'Shea⁵⁹, R. Ospanov^{62a}, G. Otero y Garzon³⁰, H. Otono⁸⁹, P. S. Ott^{63a}, G. J. Ottino^{17a}, M. Ouchrif^{35d}, J. Ouellette²⁹, F. Ould-Saada¹²⁵, M. Owen⁵⁹, R. E. Owen¹³⁴, K. Y. Oyulmaz^{21a}, V. E. Ozcan^{21a}, N. Ozturk⁸, S. Ozturk^{21d}, H. A. Pacey³², A. Pacheco Pages¹³, C. Padilla Aranda¹³, G. Padovano^{75a,75b}, S. Pagan Griso^{17a}, G. Palacino⁶⁸, A. Palazzo^{70a,70b}, S. Palestini³⁶, J. Pan¹⁷², T. Pan^{64a}, D. K. Panchal¹¹, C. E. Pandini¹¹⁴, J. G. Panduro Vazquez⁹⁵, H. Pang^{14b}, P. Pani⁴⁸, G. Panizzo^{69a,69c}, L. Paolozzi⁵⁶, C. Papadatos¹⁰⁸, S. Parajuli⁴⁴, A. Paramonov⁶, C. Paraskevopoulos¹⁰, D. Paredes Hernandez^{64b}, T. H. Park¹⁵⁵, M. A. Parker³², F. Parodi^{57a,57b}, E. W. Parrish¹¹⁵, V. A. Parrish⁵², J. A. Parsons⁴¹, U. Parzefall⁵⁴, B. Pascual Dias¹⁰⁸, L. Pascual Dominguez¹⁵¹, F. Pasquali¹¹⁴, E. Pasqualucci^{75a}, S. Passaggio^{57b}, F. Pastore⁹⁵, P. Pasuwan^{47a,47b}, P. Patel⁸⁶, U. M. Patel⁵¹, J. R. Pater¹⁰¹, T. Pauly³⁶, J. Parkes¹⁴³, M. Pedersen¹²⁵, R. Pedro^{130a}, S. V. Peleganchuk³⁷, O. Penc³⁶, E. A. Pender⁵², H. Peng^{62a}, K. E. Pensi¹⁰⁹, M. Penzin³⁷, B. S. Peralva^{82d,82d}, A. P. Pereira Peixoto⁶⁰, L. Pereira Sanchez^{47a,47b}, D. V. Perepelitsa^{29,ae}, E. Perez Codina^{156a}, M. Perganti¹⁰, L. Perini^{71a,71b,*}, H. Pernegger³⁶, S. Perrella³⁶, A. Perrevoort¹¹³, O. Perrin⁴⁰, K. Peters⁴⁸, R. F. Y. Peters¹⁰¹, B. A. Petersen³⁶, T. C. Petersen⁴², E. Petit¹⁰², V. Petousis¹³², C. Petridou^{152,e}, A. Petrukhin¹⁴¹, M. Pettee^{17a}, N. E. Pettersson³⁶, A. Petukhov³⁷, K. Petukhova¹³³, A. Peyaud¹³⁵, R. Pezoa^{137f}, L. Pezzotti³⁶, G. Pezzullo¹⁷², T. M. Pham¹⁷⁰, T. Pham¹⁰⁵, P. W. Phillips¹³⁴, G. Piacquadio¹⁴⁵, E. Pianori^{17a}, F. Piazza^{71a,71b}, R. Piegaia³⁰, D. Pietreanu^{27b}, A. D. Pilkington¹⁰¹, M. Pinamonti^{69a,69c}, J. L. Pinfold², B. C. Pinheiro Pereira^{130a}, A. E. Pinto Pinoargote¹³⁵, C. Pitman Donaldson⁹⁶, D. A. Pizzi³⁴, L. Pizzimento^{76a,76b}, A. Pizzini¹¹⁴, M.-A. Pleier²⁹, V. Plesanovs⁵⁴, V. Pleskot¹³³, E. Plotnikova³⁸, G. Poddar⁴, R. Poettgen⁹⁸, L. Poggioli¹²⁷, D. Pohl²⁴, I. Pokharel⁵⁵, S. Polacek¹³³, G. Polesello^{73a}, A. Poley^{142,156a}, R. Polifka¹³², A. Polini^{23b}, C. S. Pollard¹⁶⁷, Z. B. Pollock¹¹⁹, V. Polychronakos²⁹, E. Pompa Pacchi^{75a,75b}, D. Ponomarenko¹¹³, L. Pontecorvo³⁶, S. Popa^{27a}, G. A. Popeneciu^{27d}, D. M. Portillo Quintero^{156a}, S. Pospisil¹³², M. A. Postill¹³⁹, P. Postolache^{27c}, K. Potamianos¹⁶⁷, P. P. Potepa^{85a}, I. N. Potrap³⁸

C. J. Potter³², H. Potti¹, T. Poulsen⁴⁸, J. Poveda¹⁶³, M. E. Pozo Astigarraga³⁶, A. Prades Ibanez¹⁶³, M. M. Prapa⁴⁶, J. Pretel⁵⁴, D. Price¹⁰¹, M. Primavera^{70a}, M. A. Principe Martin⁹⁹, R. Privara¹²², T. Procter⁵⁹, M. L. Proffitt¹³⁸, N. Proklova¹²⁸, K. Prokofiev^{64c}, G. Proto^{76a,76b}, S. Protopopescu²⁹, J. Proudfoot⁶, M. Przybycien^{85a}, W. W. Przygoda^{85b}, J. E. Puddefoot¹³⁹, D. Pudzha³⁷, D. Pyatiizbyantseva³⁷, J. Qian¹⁰⁶, D. Qichen¹⁰¹, Y. Qin¹⁰¹, T. Qiu⁵², A. Quadt⁵⁵, M. Queitsch-Maitland¹⁰¹, G. Quetant⁵⁶, G. Rabanal Bolanos⁶¹, D. Rafanoharana⁵⁴, F. Ragusa^{71a,71b}, J. L. Rainbolt³⁹, J. A. Raine⁵⁶, S. Rajagopalan²⁹, E. Ramakoti³⁷, K. Ran^{48,14e}, N. P. Rapheeha^{33g}, H. Rasheed^{27b}, V. Raskina¹²⁷, D. F. Rassloff^{63a}, S. Rave¹⁰⁰, B. Ravina⁵⁵, I. Ravinovich¹⁶⁹, M. Raymond³⁶, A. L. Read¹²⁵, N. P. Readioff¹³⁹, D. M. Rebutti^{73a,73b}, G. Redlinger²⁹, A. S. Reed¹¹⁰, K. Reeves²⁶, J. A. Reidelsturz¹⁷¹, D. Reikher¹⁵¹, A. Rej¹⁴¹, C. Rembser³⁶, A. Renardi⁴⁸, M. Renda^{27b}, M. B. Rendel¹¹⁰, F. Renner⁴⁸, A. G. Rennie⁵⁹, S. Resconi^{71a}, M. Ressegotti^{57a,57b}, E. D. Resseguie^{17a}, S. Rettie³⁶, J. G. Reyes Rivera¹⁰⁷, B. Reynolds¹¹⁹, E. Reynolds^{17a}, M. Rezaei Estabragh¹⁷¹, O. L. Rezanova³⁷, P. Reznicek¹³³, N. Ribaric⁹¹, E. Ricci^{78a,78b}, R. Richter¹¹⁰, S. Richter^{47a,47b}, E. Richter-Was^{85b}, M. Ridel¹²⁷, S. Ridouani^{35d}, P. Rieck¹¹⁷, P. Riedler³⁶, M. Rijssenbeek¹⁴⁵, A. Rimoldi^{73a,73b}, M. Rimoldi⁴⁸, L. Rinaldi^{23a,23b}, T. T. Rinn²⁹, M. P. Rinnagel¹⁰⁹, G. Ripellino¹⁶¹, I. Riu¹³, P. Rivadeneira⁴⁸, J. C. Rivera Vergara¹⁶⁵, F. Rizatdinova¹²¹, E. Rizvi⁹⁴, C. Rizzi⁵⁶, B. A. Roberts¹⁶⁷, B. R. Roberts^{17a}, S. H. Robertson^{104.w}, M. Robin⁴⁸, D. Robinson³², C. M. Robles Gajardo^{137f}, M. Robles Manzano¹⁰⁰, A. Robson⁵⁹, A. Rocchi^{76a,76b}, C. Roda^{74a,74b}, S. Rodriguez Bosca^{63a}, Y. Rodriguez Garcia^{22a}, A. Rodriguez Rodriguez⁵⁴, A. M. Rodríguez Vera^{156b}, S. Roe³⁶, J. T. Roemer¹⁶⁰, A. R. Roepke-Gier¹³⁶, J. Roggel¹⁷¹, O. Röhne¹²⁵, R. A. Rojas¹⁰³, C. P. A. Roland⁶⁸, J. Roloff²⁹, A. Romaniouk³⁷, E. Romano^{73a,73b}, M. Romano^{23b}, A. C. Romero Hernandez¹⁶², N. Rompotis⁹², L. Roos¹²⁷, S. Rosati^{75a}, B. J. Rosser³⁹, E. Rossi¹²⁶, E. Rossi^{72a,72b}, L. P. Rossi^{57b}, L. Rossini⁴⁸, R. Rosten¹¹⁹, M. Rotaru^{27b}, B. Rottler⁵⁴, C. Rougier¹⁰², D. Rousseau⁶⁶, D. Rousso³², A. Roy¹⁶², S. Roy-Garand¹⁵⁵, A. Rozanov¹⁰², Y. Rozen¹⁵⁰, X. Ruan^{33g}, A. Rubio Jimenez¹⁶³, A. J. Ruby⁹², V. H. Ruelas Rivera¹⁸, T. A. Ruggeri¹, A. Ruggiero¹²⁶, A. Ruiz-Martinez¹⁶³, A. Rummler³⁶, Z. Rurikova⁵⁴, N. A. Rusakovich³⁸, H. L. Russell¹⁶⁵, G. Russo^{75a,75b}, J. P. Rutherford⁷, S. Rutherford Colmenares³², K. Rybacki⁹¹, M. Rybar¹³³, E. B. Rye¹²⁵, A. Ryzhov³⁷, J. A. Sabater Iglesias⁵⁶, P. Sabatini¹⁶³, L. Sabetta^{75a,75b}, H.F.-W. Sadrozinski¹³⁶, F. Safai Tehrani^{75a}, B. Safarzadeh Samani¹⁴⁶, M. Safdari¹⁴³, S. Saha¹⁰⁴, M. Sahinsoy¹¹⁰, M. Saimpert¹³⁵, M. Saito¹⁵³, T. Saito¹⁵³, D. Salamani³⁶, A. Salnikov¹⁴³, J. Salt¹⁶³, A. Salvador Salas¹³, D. Salvatore^{43a,43b}, F. Salvatore¹⁴⁶, A. Salzburger³⁶, D. Sammel⁵⁴, D. Sampsonidis^{152.e}, D. Sampsonidou¹²³, J. Sánchez¹⁶³, A. Sanchez Pineda⁴, V. Sanchez Sebastian¹⁶³, H. Sandaker¹²⁵, C. O. Sander⁴⁸, J. A. Sandesara¹⁰³, M. Sandhoff¹⁷¹, C. Sandoval^{22b}, D. P. C. Sankey¹³⁴, T. Sano⁸⁷, A. Sansoni⁵³, L. Santi^{75a,75b}, C. Santoni⁴⁰, H. Santos^{130a,130b}, S. N. Santpur^{17a}, A. Santra¹⁶⁹, K. A. Saoucha¹³⁹, J. G. Saraiva^{130a,130d}, J. Sardain⁷, O. Sasaki⁸³, K. Sato¹⁵⁷, C. Sauer^{63b}, F. Sauerburger⁵⁴, E. Sauvan⁴, P. Savard^{155.ac}, R. Sawada¹⁵³, C. Sawyer¹³⁴, L. Sawyer⁹⁷, I. Sayago Galvan¹⁶³, C. Sbarra^{23b}, A. Sbrizzi^{23a,23b}, T. Scanlon⁹⁶, J. Schaarschmidt¹³⁸, P. Schacht¹¹⁰, D. Schaefer³⁹, U. Schäfer¹⁰⁰, A. C. Schaffer^{66,44}, D. Schaile¹⁰⁹, R. D. Schamberger¹⁴⁵, C. Scharf¹⁸, M. M. Schefer¹⁹, V. A. Schegelsky³⁷, D. Scheirich¹³³, F. Schenck¹⁸, M. Schernau¹⁶⁰, C. Scheulen⁵⁵, C. Schiavi^{57a,57b}, E. J. Schioppa^{70a,70b}, M. Schioppa^{43a,43b}, B. Schlag¹⁴³, K. E. Schleicher⁵⁴, S. Schlenker³⁶, J. Schmeing¹⁷¹, M. A. Schmidt¹⁷¹, K. Schmieden¹⁰⁰, C. Schmitt¹⁰⁰, S. Schmitt⁴⁸, L. Schoeffel¹³⁵, A. Schoening^{63b}, P. G. Scholer⁵⁴, E. Schopf¹²⁶, M. Schott¹⁰⁰, J. Schovancova³⁶, S. Schramm⁵⁶, F. Schroeder¹⁷¹, H.-C. Schultz-Coulon^{63a}, M. Schumacher⁵⁴, B. A. Schumm¹³⁶, Ph. Schune¹³⁵, A. J. Schuy¹³⁸, H. R. Schwartz¹³⁶, A. Schwartzman¹⁴³, T. A. Schwarz¹⁰⁶, Ph. Schwemling¹³⁵, R. Schwienhorst¹⁰⁷, A. Sciandra¹³⁶, G. Sciolla²⁶, F. Scuri^{74a}, C. D. Sebastiani⁹², K. Sedlaczek¹¹⁵, P. Seema¹⁸, S. C. Seidel¹¹², A. Seiden¹³⁶, B. D. Seidlitz⁴¹, C. Seitz⁴⁸, J. M. Seixas^{82b}, G. Sekhniaidze^{72a}, S. J. Sekula⁴⁴, L. Selem⁶⁰, N. Semprini-Cesari^{23a,23b}, D. Sengupta⁵⁶, V. Senthilkumar¹⁶³, L. Serin⁶⁶, L. Serkin^{69a,69b}, M. Sessa^{76a,76b}, H. Severini¹²⁰, F. Sforza^{57a,57b}, A. Sfyrta⁵⁶, E. Shabalina⁵⁵, R. Shaheen¹⁴⁴, J. D. Shahinian¹²⁸, D. Shaked Renous¹⁶⁹, L. Y. Shan^{14a}, M. Shapiro^{17a}, A. Sharma³⁶, A. S. Sharma¹⁶⁴, P. Sharma⁸⁰, S. Sharma⁴⁸, P. B. Shatalov³⁷, K. Shaw¹⁴⁶, S. M. Shaw¹⁰¹, A. Shcherbakova³⁷, Q. Shen^{62c,5}, P. Sherwood⁹⁶, L. Shi⁹⁶, X. Shi^{14a}, C. O. Shimmin¹⁷², Y. Shimogama¹⁶⁸, J. D. Shinner⁹⁵, I. P. J. Shipsey¹²⁶, S. Shirabe^{56.h}, M. Shiyakova³⁸, J. Shlomi¹⁶⁹, M. J. Shochet³⁹, J. Shojaii¹⁰⁵, D. R. Shope¹²⁵, S. Shrestha^{119.af}, E. M. Shrif^{33g}, M. J. Shroff¹⁶⁵, P. Sicho¹³¹, A. M. Sickles¹⁶², E. Sideras Haddad^{33g}, A. Sidoti^{23b}, F. Siegert⁵⁰, Dj. Sijacki¹⁵, R. Sikora^{85a}, F. Sili⁹⁰, J. M. Silva²⁰, M. V. Silva Oliveira²⁹, S. B. Silverstein^{47a}, S. Simion⁶⁶, R. Simoniello³⁶, E. L. Simpson⁵⁹, H. Simpson¹⁴⁶, L. R. Simpson¹⁰⁶, N. D. Simpson⁹⁸, S. Simsek^{21d}, S. Sindhu⁵⁵, P. Sinervo¹⁵⁵, S. Singh¹⁵⁵

S. Sinha⁴⁸, S. Sinha¹⁰¹, M. Sioli^{23a,23b}, I. Siral³⁶, E. Sitnikova⁴⁸, S.Yu. Sivoklov^{37,*}, J. Sjölin^{47a,47b}, A. Skaf⁵⁵, E. Skorda⁹⁸, P. Skubic¹²⁰, M. Slawinska⁸⁶, V. Smakhtin¹⁶⁹, B. H. Smart¹³⁴, J. Smiesko³⁶, S.Yu. Smirnov³⁷, Y. Smirnov³⁷, L. N. Smirnova^{37,a}, O. Smirnova⁹⁸, A. C. Smith⁴¹, E. A. Smith³⁹, H. A. Smith¹²⁶, J. L. Smith⁹², R. Smith¹⁴³, M. Smizanska⁹¹, K. Smolek¹³², A. A. Snesarev³⁷, S. R. Snider¹⁵⁵, H. L. Snoek¹¹⁴, S. Snyder²⁹, R. Sobie^{165,w}, A. Soffer¹⁵¹, C. A. Solans Sanchez³⁶, E.Yu. Soldatov³⁷, U. Soldevila¹⁶³, A. A. Solodkov³⁷, S. Solomon²⁶, A. Soloshenko³⁸, K. Solovieva⁵⁴, O. V. Solovyanov⁴⁰, V. Solovyev³⁷, P. Sommer³⁶, A. Sonay¹³, W. Y. Song^{156b}, J. M. Sonneveld¹¹⁴, A. Sopczak¹³², A. L. Sopio⁹⁶, F. Sopkova^{28b}, V. Sothilingam^{63a}, S. Sottocornola⁶⁸, R. Soualah^{116b}, Z. Soumami^{35e}, D. South⁴⁸, S. Spagnolo^{70a,70b}, M. Spalla¹¹⁰, D. Sperlich⁵⁴, G. Spigo³⁶, M. Spina¹⁴⁶, S. Spinali⁹¹, D. P. Spiteri⁵⁹, M. Spousta¹³³, E. J. Staats³⁴, A. Stabile^{71a,71b}, R. Stamen^{63a}, M. Stamenkovic¹¹⁴, A. Stampekis²⁰, M. Standke²⁴, E. Stanecka⁸⁶, M. V. Stange⁵⁰, B. Stanislaus^{17a}, M. M. Stanitzki⁴⁸, M. Stankaityte¹²⁶, B. Stapf⁴⁸, E. A. Starchenko³⁷, G. H. Stark¹³⁶, J. Stark¹⁰², D. M. Starko^{156b}, P. Staroba¹³¹, P. Starovoitov^{63a}, S. Stärz¹⁰⁴, R. Staszewski⁸⁶, G. Stavropoulos⁴⁶, J. Steentoft¹⁶¹, P. Steinberg²⁹, B. Stelzer^{142,156a}, H. J. Stelzer¹²⁹, O. Stelzer-Chilton^{156a}, H. Stenzel⁵⁸, T. J. Stevenson¹⁴⁶, G. A. Stewart³⁶, J. R. Stewart¹²¹, M. C. Stockton³⁶, G. Stoica^{27b}, M. Stolarski^{130a}, S. Stonjek¹¹⁰, A. Straessner⁵⁰, J. Strandberg¹⁴⁴, S. Strandberg^{47a,47b}, M. Strauss¹²⁰, T. Streblner¹⁰², P. Strizenc^{28b}, R. Ströhmer¹⁶⁶, D. M. Strom¹²³, L. R. Strom⁴⁸, R. Stroynowski⁴⁴, A. Strubig^{47a,47b}, S. A. Stucci²⁹, B. Stugu¹⁶, J. Stupak¹²⁰, N. A. Styles⁴⁸, D. Su¹⁴³, S. Su^{62a}, W. Su^{62d}, X. Su^{62a,66}, K. Sugizaki¹⁵³, V. V. Sulim³⁷, M. J. Sullivan⁹², D. M. S. Sultan^{78a,78b}, L. Sultanaliyeva³⁷, S. Sultansoy^{3b}, T. Sumida⁸⁷, S. Sun¹⁰⁶, S. Sun¹⁷⁰, O. Sunneborn Gudnadottir¹⁶¹, M. R. Sutton¹⁴⁶, M. Svatos¹³¹, M. Swiatlowski^{156a}, T. Swirski¹⁶⁶, I. Sykora^{28a}, M. Sykora¹³³, T. Sykora¹³³, D. Ta¹⁰⁰, K. Tackmann^{48,u}, A. Taffard¹⁶⁰, R. Tafirout^{156a}, J. S. Tafoya Vargas⁶⁶, R. Takashima⁸⁸, E. P. Takeva⁵², Y. Takubo⁸³, M. Talby¹⁰², A. A. Talyshev³⁷, K. C. Tam^{64b}, N. M. Tamir¹⁵¹, A. Tanaka¹⁵³, J. Tanaka¹⁵³, R. Tanaka⁶⁶, M. Tanasini^{57a,57b}, Z. Tao¹⁶⁴, S. Tapia Araya^{137f}, S. Tapprogge¹⁰⁰, A. Tarek Abouelfadl Mohamed¹⁰⁷, S. Tarem¹⁵⁰, K. Tariq^{62b}, G. Tarna^{102,27b}, G. F. Tartarelli^{71a}, P. Tas¹³³, M. Tasevsky¹³¹, E. Tassi^{43a,43b}, A. C. Tate¹⁶², G. Tateno¹⁵³, Y. Tayalati^{35e,v}, G. N. Taylor¹⁰⁵, W. Taylor^{156b}, H. Teagle⁹², A. S. Tee¹⁷⁰, R. Teixeira De Lima¹⁴³, P. Teixeira-Dias⁹⁵, J. J. Teoh¹⁵⁵, K. Terashi¹⁵³, J. Terron⁹⁹, S. Terzo¹³, M. Testa⁵³, R. J. Teuscher^{155,w}, A. Thaler⁷⁹, O. Theiner⁵⁶, N. Themistokleous⁵², T. Thevenaux-Pelzer¹⁰², O. Thielmann¹⁷¹, D. W. Thomas⁹⁵, J. P. Thomas²⁰, E. A. Thompson^{17a}, P. D. Thompson²⁰, E. Thomson¹²⁸, Y. Tian⁵⁵, V. Tikhomirov^{37,a}, Yu.A. Tikhonov³⁷, S. Timoshenko³⁷, D. Timoshyn¹³³, E. X. L. Ting¹, P. Tipton¹⁷², S. H. Tlou^{33g}, A. Tmourji⁴⁰, K. Todome^{23a,23b}, S. Todorova-Nova¹³³, S. Todt⁵⁰, M. Togawa⁸³, J. Tojo⁸⁹, S. Tokár^{28a}, K. Tokushuku⁸³, O. Toldaiev⁶⁸, R. Tombs³², M. Tomoto^{83,111}, L. Tompkins¹⁴³, K. W. Topolnicki^{85b}, E. Torrence¹²³, H. Torres¹⁰², E. Torró Pastor¹⁶³, M. Toscani³⁰, C. Toscirri³⁹, M. Tost¹¹, D. R. Tovey¹³⁹, A. Traeet¹⁶, I. S. Trandafir^{27b}, T. Trefzger¹⁶⁶, A. Tricoli²⁹, I. M. Trigger^{156a}, S. Trincaz-Duvoid¹²⁷, D. A. Trischuk²⁶, B. Trocme⁶⁰, C. Troncon^{71a}, L. Truong^{33c}, M. Trzebinski⁸⁶, A. Trzupek⁸⁶, F. Tsai¹⁴⁵, M. Tsai¹⁰⁶, A. Tsiamis^{152,e}, P. V. Tsiarshka³⁷, S. Tsigaridas^{156a}, A. Tsigiridis^{152,s}, V. Tsiskaridze¹⁵⁵, E. G. Tskhadadze^{149a}, M. Tsopoulou^{152,e}, Y. Tsujikawa⁸⁷, I. I. Tsukerman³⁷, V. Tsulaia^{17a}, S. Tsuno⁸³, O. Tsur¹⁵⁰, K. Tsuru¹¹⁸, D. Tsybychev¹⁴⁵, Y. Tu^{64b}, A. Tudorache^{27b}, V. Tudorache^{27b}, A. N. Tuna³⁶, S. Turchikhin³⁸, I. Turk Cakir^{3a}, R. Turra^{71a}, T. Turtuvshin^{38,x}, P. M. Tuts⁴¹, S. Tzamarias^{152,e}, P. Tzanis¹⁰, E. Tzovara¹⁰⁰, K. Uchida¹⁵³, F. Ukegawa¹⁵⁷, P. A. Ulloa Poblete^{137c,137b}, E. N. Umaka²⁹, G. Unal³⁶, M. Unal¹¹, A. Undrus²⁹, G. Unel¹⁶⁰, J. Urban^{28b}, P. Urquijo¹⁰⁵, G. Usai⁸, R. Ushioda¹⁵⁴, M. Usman¹⁰⁸, Z. Uysal^{21b}, L. Vacavant¹⁰², V. Vacek¹³², B. Vachon¹⁰⁴, K. O. H. Vadla¹²⁵, T. Vafeiadis³⁶, A. Vaitkus⁹⁶, C. Valderanis¹⁰⁹, E. Valdes Santurio^{47a,47b}, M. Valente^{156a}, S. Valentini^{23a,23b}, A. Valero¹⁶³, E. Valiente Moreno¹⁶³, A. Vallier¹⁰², J. A. Valls Ferrer¹⁶³, D. R. Van Arneman¹¹⁴, T. R. Van Daalen¹³⁸, A. Van Der Graaf⁴⁹, P. Van Gemmeren⁶, M. Van Rijnbach^{125,36}, S. Van Stroud⁹⁶, I. Van Vulpen¹¹⁴, M. Vanadia^{76a,76b}, W. Vandelli³⁶, M. Vandenbroucke¹³⁵, E. R. Vandewall¹²¹, D. Vannicola¹⁵¹, L. Vannoli^{57a,57b}, R. Vari^{75a}, E. W. Varnes⁷, C. Varni^{17a}, T. Varol¹⁴⁸, D. Varouchas⁶⁶, L. Varriale¹⁶³, K. E. Varvell¹⁴⁷, M. E. Vasile^{27b}, L. Vaslin⁴⁰, G. A. Vasquez¹⁶⁵, F. Vazeille⁴⁰, T. Vazquez Schroeder³⁶, J. Veatch³¹, V. Vecchio¹⁰¹, M. J. Veen¹⁰³, I. Veliscek¹²⁶, L. M. Veloce¹⁵⁵, F. Veloso^{130a,130c}, S. Veneziano^{75a}, A. Ventura^{70a,70b}, A. Verbytskyi¹¹⁰, M. Verducci^{74a,74b}, C. Vergis²⁴, M. Verissimo De Araujo^{82b}, W. Verkerke¹¹⁴, J. C. Vermeulen¹¹⁴, C. Vernieri¹⁴³, P. J. Verschuur⁹⁵, M. Vessella¹⁰³, M. C. Vetterli^{142,ac}, A. Vgenopoulos^{152,e}, N. Viaux Maira^{137f}, T. Vickey¹³⁹, O. E. Vickey Boeriu¹³⁹, G. H. A. Viehhauser¹²⁶, L. Vigani^{63b}, M. Villa^{23a,23b}, M. Villaplana Perez¹⁶³, E. M. Villhauer⁵², E. Vilucchi⁵³, M. G. Vincker³⁴

G. S. Virdee²⁰, A. Vishwakarma⁵², C. Vittori³⁶, I. Vivarelli¹⁴⁶, V. Vladimirov¹⁶⁷, E. Voevodina¹¹⁰, F. Vogel¹⁰⁹, P. Vokac¹³², J. Von Ahnen⁴⁸, E. Von Toerne²⁴, B. Vormwald³⁶, V. Vorobel¹³³, K. Vorobev³⁷, M. Vos¹⁶³, K. Voss¹⁴¹, J. H. Vosseveld⁹², M. Vozak¹¹⁴, L. Vozdecky⁹⁴, N. Vranjes¹⁵, M. Vranjes Milosavljevic¹⁵, M. Vreeswijk¹¹⁴, R. Vuillermet³⁶, O. Vujanovic¹⁰⁰, I. Vukotic³⁹, S. Wada¹⁵⁷, C. Wagner¹⁰³, J. M. Wagner^{17a}, W. Wagner¹⁷¹, S. Wahdan¹⁷¹, H. Wahlberg⁹⁰, R. Wakasa¹⁵⁷, M. Wakida¹¹¹, J. Walder¹³⁴, R. Walker¹⁰⁹, W. Walkowiak¹⁴¹, A. Wall¹²⁸, T. Wamorkar⁶, A. Z. Wang¹⁷⁰, C. Wang¹⁰⁰, C. Wang^{62c}, H. Wang^{17a}, J. Wang^{64a}, R.-J. Wang¹⁰⁰, R. Wang⁶¹, R. Wang⁶, S. M. Wang¹⁴⁸, S. Wang^{62b}, T. Wang^{62a}, W. T. Wang⁸⁰, X. Wang^{14c}, X. Wang¹⁶², X. Wang^{62c}, Y. Wang^{62d}, Y. Wang^{14c}, Z. Wang¹⁰⁶, Z. Wang^{51,62c,62d}, Z. Wang¹⁰⁶, A. Warburton¹⁰⁴, R. J. Ward²⁰, N. Warrack⁵⁹, A. T. Watson²⁰, H. Watson⁵⁹, M. F. Watson²⁰, G. Watts¹³⁸, B. M. Waugh⁹⁶, C. Weber²⁹, H. A. Weber¹⁸, M. S. Weber¹⁹, S. M. Weber^{63a}, C. Wei^{62a}, Y. Wei¹²⁶, A. R. Weidberg¹²⁶, E. J. Weik¹¹⁷, J. Weingarten⁴⁹, M. Weirich¹⁰⁰, C. Weiser⁵⁴, C. J. Wells⁴⁸, T. Wenaus²⁹, B. Wendland⁴⁹, T. Wengler³⁶, N. S. Wenke¹¹⁰, N. Wermes²⁴, M. Wessels^{63a}, K. Whalen¹²³, A. M. Wharton⁹¹, A. S. White⁶¹, A. White⁸, M. J. White¹, D. Whiteson¹⁶⁰, L. Wickremasinghe¹²⁴, W. Wiedenmann¹⁷⁰, C. Wiel⁵⁰, M. Wielers¹³⁴, C. Wiglesworth⁴², D. J. Wilbern¹²⁰, H. G. Wilkens³⁶, D. M. Williams⁴¹, H. H. Williams¹²⁸, S. Williams³², S. Willocq¹⁰³, B. J. Wilson¹⁰¹, P. J. Windischhofer³⁹, F. I. Winkel³⁰, F. Winklmeier¹²³, B. T. Winter⁵⁴, J. K. Winter¹⁰¹, M. Wittgen¹⁴³, M. Wobisch⁹⁷, Z. Wolffs¹¹⁴, R. Wölker¹²⁶, J. Wollrath¹⁶⁰, M. W. Wolter⁸⁶, H. Wolters^{130a,130c}, A. F. Wongel⁴⁸, S. D. Worm⁴⁸, B. K. Wosiek⁸⁶, K. W. Woźniak⁸⁶, S. Wozniowski⁵⁵, K. Wraight⁵⁹, C. Wu²⁰, J. Wu^{14a,14e}, M. Wu^{64a}, M. Wu¹¹³, S. L. Wu¹⁷⁰, X. Wu⁵⁶, Y. Wu^{62a}, Z. Wu¹³⁵, J. Wuerzinger¹¹⁰, T. R. Wyatt¹⁰¹, B. M. Wynne⁵², S. Xella⁴², L. Xia^{14c}, M. Xia^{14b}, J. Xiang^{64c}, X. Xiao¹⁰⁶, M. Xie^{62a}, X. Xie^{62a}, S. Xin^{14a,14e}, J. Xiong^{17a}, D. Xu^{14a}, H. Xu^{62a}, L. Xu^{62a}, R. Xu¹²⁸, T. Xu¹⁰⁶, Y. Xu^{14b}, Z. Xu⁵², Z. Xu^{14a}, B. Yabsley¹⁴⁷, S. Yacoub^{33a}, N. Yamaguchi⁸⁹, Y. Yamaguchi¹⁵⁴, E. Yamashita¹⁵³, H. Yamauchi¹⁵⁷, T. Yamazaki^{17a}, Y. Yamazaki⁸⁴, J. Yan^{62c}, S. Yan¹²⁶, Z. Yan²⁵, H. J. Yang^{62c,62d}, H. T. Yang^{62a}, S. Yang^{62a}, T. Yang^{64c}, X. Yang^{62a}, X. Yang^{14a}, Y. Yang⁴⁴, Y. Yang^{62a}, Z. Yang^{62a}, W.-M. Yao^{17a}, Y. C. Yap⁴⁸, H. Ye^{14c}, H. Ye⁵⁵, J. Ye⁴⁴, S. Ye²⁹, X. Ye^{62a}, Y. Yeh⁹⁶, I. Yeletsikh³⁸, B. K. Yeo^{17a}, M. R. Yexley⁹⁶, P. Yin⁴¹, K. Yorita¹⁶⁸, S. Younas^{27b}, C. J. S. Young⁵⁴, C. Young¹⁴³, Y. Yu^{62a}, M. Yuan¹⁰⁶, R. Yuan^{62b,k}, L. Yue⁹⁶, M. Zaazoua^{62a}, B. Zabinski⁸⁶, E. Zaid⁵², T. Zakareishvili^{149b}, N. Zakharchuk³⁴, S. Zambito⁵⁶, J. A. Zamora Saa^{137d,137b}, J. Zang¹⁵³, D. Zanzi⁵⁴, O. Zaplatilek¹³², C. Zeitnitz¹⁷¹, H. Zeng^{14a}, J. C. Zeng¹⁶², D. T. Zenger Jr²⁶, O. Zenin³⁷, T. Ženiš^{28a}, S. Zenz⁹⁴, S. Zerradi^{35a}, D. Zerwas⁶⁶, M. Zhai^{14a,14e}, B. Zhang^{14c}, D. F. Zhang¹³⁹, J. Zhang^{62b}, J. Zhang⁶, K. Zhang^{14a,14e}, L. Zhang^{14c}, P. Zhang^{14a,14e}, R. Zhang¹⁷⁰, S. Zhang¹⁰⁶, T. Zhang¹⁵³, X. Zhang^{62c}, X. Zhang^{62b}, Y. Zhang^{62c,5}, Y. Zhang⁹⁶, Z. Zhang^{17a}, Z. Zhang⁶⁶, H. Zhao¹³⁸, P. Zhao⁵¹, T. Zhao^{62b}, Y. Zhao¹³⁶, Z. Zhao^{62a}, A. Zhemchugov³⁸, K. Zheng¹⁶², X. Zheng^{62a}, Z. Zheng¹⁴³, D. Zhong¹⁶², B. Zhou¹⁰⁶, H. Zhou⁷, N. Zhou^{62c}, Y. Zhou⁷, C. G. Zhu^{62b}, J. Zhu¹⁰⁶, Y. Zhu^{62c}, Y. Zhu^{62a}, X. Zhuang^{14a}, K. Zhukov³⁷, V. Zhulanov³⁷, N. I. Zimine³⁸, J. Zinsser^{63b}, M. Ziolkowski¹⁴¹, L. Živković¹⁵, A. Zoccoli^{23a,23b}, K. Zoch⁵⁶, T. G. Zorbas¹³⁹, O. Zormpa⁴⁶, W. Zou⁴¹, L. Zwalinski³⁶

¹ Department of Physics, University of Adelaide, Adelaide, Australia

² Department of Physics, University of Alberta, Edmonton, AB, Canada

³ (a) Department of Physics, Ankara University, Ankara, Türkiye; (b) Division of Physics, TOBB University of Economics and Technology, Ankara, Türkiye

⁴ LAPP, Univ. Savoie Mont Blanc, CNRS/IN2P3, Annecy, France

⁵ APC, Université Paris Cité, CNRS/IN2P3, Paris, France

⁶ High Energy Physics Division, Argonne National Laboratory, Argonne, IL, USA

⁷ Department of Physics, University of Arizona, Tucson, AZ, USA

⁸ Department of Physics, University of Texas at Arlington, Arlington, TX, USA

⁹ Physics Department, National and Kapodistrian University of Athens, Athens, Greece

¹⁰ Physics Department, National Technical University of Athens, Zografou, Greece

¹¹ Department of Physics, University of Texas at Austin, Austin, TX, USA

¹² Institute of Physics, Azerbaijan Academy of Sciences, Baku, Azerbaijan

¹³ Institut de Física d'Altes Energies (IFAE), Barcelona Institute of Science and Technology, Barcelona, Spain

¹⁴ (a) Institute of High Energy Physics, Chinese Academy of Sciences, Beijing, China; (b) Physics Department, Tsinghua University, Beijing, China; (c) Department of Physics, Nanjing University, Nanjing, China; (d) School of Science,

- Shenzhen Campus of Sun Yat-sen University, Shenzhen, China; ^(e)University of Chinese Academy of Science (UCAS), Beijing, China
- ¹⁵ Institute of Physics, University of Belgrade, Belgrade, Serbia
- ¹⁶ Department for Physics and Technology, University of Bergen, Bergen, Norway
- ¹⁷ ^(a)Physics Division, Lawrence Berkeley National Laboratory, Berkeley, CA, USA; ^(b)University of California, Berkeley, CA, USA
- ¹⁸ Institut für Physik, Humboldt Universität zu Berlin, Berlin, Germany
- ¹⁹ Albert Einstein Center for Fundamental Physics and Laboratory for High Energy Physics, University of Bern, Bern, Switzerland
- ²⁰ School of Physics and Astronomy, University of Birmingham, Birmingham, UK
- ²¹ ^(a)Department of Physics, Bogazici University, Istanbul, Türkiye; ^(b)Department of Physics Engineering, Gaziantep University, Gaziantep, Türkiye; ^(c)Department of Physics, Istanbul University, Istanbul, Türkiye; ^(d)Istinye University, Sariyer, Istanbul, Türkiye
- ²² ^(a)Facultad de Ciencias y Centro de Investigaciones, Universidad Antonio Nariño, Bogotá, Colombia; ^(b)Departamento de Física, Universidad Nacional de Colombia, Bogotá, Colombia; ^(c)Pontificia Universidad Javeriana, Bogotá, Colombia
- ²³ ^(a)Dipartimento di Fisica e Astronomia A. Righi, Università di Bologna, Bologna, Italy; ^(b)INFN Sezione di Bologna, Bologna, Italy
- ²⁴ Physikalisches Institut, Universität Bonn, Bonn, Germany
- ²⁵ Department of Physics, Boston University, Boston, MA, USA
- ²⁶ Department of Physics, Brandeis University, Waltham, MA, USA
- ²⁷ ^(a)Transilvania University of Brasov, Brasov, Romania; ^(b)Horia Hulubei National Institute of Physics and Nuclear Engineering, Bucharest, Romania; ^(c)Department of Physics, Alexandru Ioan Cuza University of Iasi, Iasi, Romania; ^(d)Physics Department, National Institute for Research and Development of Isotopic and Molecular Technologies, Cluj-Napoca, Romania; ^(e)University Politehnica Bucharest, Bucharest, Romania; ^(f)West University in Timisoara, Timisoara, Romania; ^(g)Faculty of Physics, University of Bucharest, Bucharest, Romania
- ²⁸ ^(a)Faculty of Mathematics, Physics and Informatics, Comenius University, Bratislava, Slovak Republic; ^(b)Department of Subnuclear Physics, Institute of Experimental Physics of the Slovak Academy of Sciences, Kosice, Slovak Republic
- ²⁹ Physics Department, Brookhaven National Laboratory, Upton, NY, USA
- ³⁰ Departamento de Física, y CONICET, Facultad de Ciencias Exactas y Naturales, Instituto de Física de Buenos Aires (IFIBA), Universidad de Buenos Aires, Buenos Aires, Argentina
- ³¹ California State University, Northridge, CA, USA
- ³² Cavendish Laboratory, University of Cambridge, Cambridge, UK
- ³³ ^(a)Department of Physics, University of Cape Town, Cape Town, South Africa; ^(b)iThemba Labs, Western Cape, South Africa; ^(c)Department of Mechanical Engineering Science, University of Johannesburg, Johannesburg, South Africa; ^(d)National Institute of Physics, University of the Philippines, Diliman, Philippines; ^(e)Department of Physics, University of South Africa, Pretoria, South Africa; ^(f)University of Zululand, KwaDlangezwa, South Africa; ^(g)School of Physics, University of the Witwatersrand, Johannesburg, South Africa
- ³⁴ Department of Physics, Carleton University, Ottawa, ON, Canada
- ³⁵ ^(a)Faculté des Sciences Ain Chock, Réseau Universitaire de Physique des Hautes Energies, Université Hassan II, Casablanca, Morocco; ^(b)Faculté des Sciences, Université Ibn-Tofail, Kénitra, Morocco; ^(c)Faculté des Sciences Semlalia, Université Cadi Ayyad, LPHEA-Marrakech, Morocco; ^(d)LPMR, Faculté des Sciences, Université Mohamed Premier, Oujda, Morocco; ^(e)Faculté des sciences, Université Mohammed V, Rabat, Morocco; ^(f)Institute of Applied Physics, Mohammed VI Polytechnic University, Ben Guerir, Morocco
- ³⁶ CERN, Geneva, Switzerland
- ³⁷ Affiliated with an institute covered by a cooperation agreement with CERN, Geneva, Switzerland
- ³⁸ Affiliated with an international laboratory covered by a cooperation agreement with CERN, Geneva, Switzerland
- ³⁹ Enrico Fermi Institute, University of Chicago, Chicago, IL, USA
- ⁴⁰ LPC, Université Clermont Auvergne, CNRS/IN2P3, Clermont-Ferrand, France
- ⁴¹ Nevis Laboratory, Columbia University, Irvington, NY, USA
- ⁴² Niels Bohr Institute, University of Copenhagen, Copenhagen, Denmark
- ⁴³ ^(a)Dipartimento di Fisica, Università della Calabria, Rende, Italy; ^(b)INFN Gruppo Collegato di Cosenza, Laboratori Nazionali di Frascati, Frascati, Italy
- ⁴⁴ Physics Department, Southern Methodist University, Dallas, TX, USA

- 45 Physics Department, University of Texas at Dallas, Richardson, TX, USA
- 46 National Centre for Scientific Research “Demokritos”, Agia Paraskevi, Greece
- 47 (a)Department of Physics, Stockholm University, Stockholm, Sweden; (b)Oskar Klein Centre, Stockholm, Sweden
- 48 Deutsches Elektronen-Synchrotron DESY, Hamburg and Zeuthen, Germany
- 49 Fakultät Physik , Technische Universität Dortmund, Dortmund, Germany
- 50 Institut für Kern- und Teilchenphysik, Technische Universität Dresden, Dresden, Germany
- 51 Department of Physics, Duke University, Durham, NC, USA
- 52 SUPA-School of Physics and Astronomy, University of Edinburgh, Edinburgh, UK
- 53 INFN e Laboratori Nazionali di Frascati, Frascati, Italy
- 54 Physikalisches Institut, Albert-Ludwigs-Universität Freiburg, Freiburg, Germany
- 55 II. Physikalisches Institut, Georg-August-Universität Göttingen, Göttingen, Germany
- 56 Département de Physique Nucléaire et Corpusculaire, Université de Genève, Geneva, Switzerland
- 57 (a)Dipartimento di Fisica, Università di Genova, Genoa, Italy; (b)INFN Sezione di Genova, Genoa, Italy
- 58 II. Physikalisches Institut, Justus-Liebig-Universität Giessen, Giessen, Germany
- 59 SUPA - School of Physics and Astronomy, University of Glasgow, Glasgow, UK
- 60 LPSC, Université Grenoble Alpes, CNRS/IN2P3, Grenoble INP, Grenoble, France
- 61 Laboratory for Particle Physics and Cosmology, Harvard University, Cambridge, MA, USA
- 62 (a)Department of Modern Physics and State Key Laboratory of Particle Detection and Electronics, University of Science and Technology of China, Hefei, China; (b)Institute of Frontier and Interdisciplinary Science and Key Laboratory of Particle Physics and Particle Irradiation (MOE), Shandong University, Qingdao, China; (c)School of Physics and Astronomy, Shanghai Jiao Tong University, Key Laboratory for Particle Astrophysics and Cosmology (MOE), SKLPPC, Shanghai, China; (d)Tsung-Dao Lee Institute, Shanghai, China
- 63 (a)Kirchhoff-Institut für Physik, Ruprecht-Karls-Universität Heidelberg, Heidelberg, Germany; (b)Physikalisches Institut, Ruprecht-Karls-Universität Heidelberg, Heidelberg, Germany
- 64 (a)Department of Physics, Chinese University of Hong Kong, Shatin, N.T., Hong Kong, China; (b)Department of Physics, University of Hong Kong, Hong Kong, China; (c)Department of Physics and Institute for Advanced Study, Hong Kong University of Science and Technology, Clear Water Bay, Kowloon, Hong Kong, China
- 65 Department of Physics, National Tsing Hua University, Hsinchu, Taiwan
- 66 IJCLab, Université Paris-Saclay, CNRS/IN2P3, 91405 Orsay, France
- 67 Centro Nacional de Microelectrónica (IMB-CNM-CSIC), Barcelona, Spain
- 68 Department of Physics, Indiana University, Bloomington, IN, USA
- 69 (a)INFN Gruppo Collegato di Udine, Sezione di Trieste, Udine, Italy; (b)ICTP, Trieste, Italy; (c)Dipartimento Politecnico di Ingegneria e Architettura, Università di Udine, Udine, Italy
- 70 (a)INFN Sezione di Lecce, Lecce, Italy; (b)Dipartimento di Matematica e Fisica, Università del Salento, Lecce, Italy
- 71 (a)INFN Sezione di Milano, Milan, Italy; (b)Dipartimento di Fisica, Università di Milano, Milan, Italy
- 72 (a)INFN Sezione di Napoli, Naples, Italy; (b)Dipartimento di Fisica, Università di Napoli, Naples, Italy
- 73 (a)INFN Sezione di Pavia, Pavia, Italy; (b)Dipartimento di Fisica, Università di Pavia, Pavia, Italy
- 74 (a)INFN Sezione di Pisa, Pisa, Italy; (b)Dipartimento di Fisica E. Fermi, Università di Pisa, Pisa, Italy
- 75 (a)INFN Sezione di Roma, Rome, Italy; (b)Dipartimento di Fisica, Sapienza Università di Roma, Rome, Italy
- 76 (a)INFN Sezione di Roma Tor Vergata, Rome, Italy; (b)Dipartimento di Fisica, Università di Roma Tor Vergata, Rome, Italy
- 77 (a)INFN Sezione di Roma Tre, Rome, Italy; (b)Dipartimento di Matematica e Fisica, Università Roma Tre, Rome, Italy
- 78 (a)INFN-TIFPA, Povo, Italy; (b)Università degli Studi di Trento, Trento, Italy
- 79 Department of Astro and Particle Physics, Universität Innsbruck, Innsbruck, Austria
- 80 University of Iowa, Iowa City, IA, USA
- 81 Department of Physics and Astronomy, Iowa State University, Ames, IA, USA
- 82 (a)Departamento de Engenharia Elétrica, Universidade Federal de Juiz de Fora (UFJF), Juiz de Fora, Brazil; (b)Universidade Federal do Rio de Janeiro COPPE/EE/IF, Rio de Janeiro, Brazil; (c)Instituto de Física, Universidade de São Paulo, São Paulo, Brazil; (d)Rio de Janeiro State University, Rio de Janeiro, Brazil
- 83 KEK, High Energy Accelerator Research Organization, Tsukuba, Japan
- 84 Graduate School of Science, Kobe University, Kobe, Japan
- 85 (a)Faculty of Physics and Applied Computer Science, AGH University of Science and Technology, Krakow, Poland; (b)Marian Smoluchowski Institute of Physics, Jagiellonian University, Krakow, Poland

- 86 Institute of Nuclear Physics Polish Academy of Sciences, Krakow, Poland
- 87 Faculty of Science, Kyoto University, Kyoto, Japan
- 88 Kyoto University of Education, Kyoto, Japan
- 89 Research Center for Advanced Particle Physics and Department of Physics, Kyushu University, Fukuoka, Japan
- 90 Instituto de Física La Plata, Universidad Nacional de La Plata and CONICET, La Plata, Argentina
- 91 Physics Department, Lancaster University, Lancaster, UK
- 92 Oliver Lodge Laboratory, University of Liverpool, Liverpool, UK
- 93 Department of Experimental Particle Physics, Jožef Stefan Institute and Department of Physics, University of Ljubljana, Ljubljana, Slovenia
- 94 School of Physics and Astronomy, Queen Mary University of London, London, UK
- 95 Department of Physics, Royal Holloway University of London, Egham, UK
- 96 Department of Physics and Astronomy, University College London, London, UK
- 97 Louisiana Tech University, Ruston, LA, USA
- 98 Fysiska institutionen, Lunds universitet, Lund, Sweden
- 99 Departamento de Física Teórica C-15 and CIAFF, Universidad Autónoma de Madrid, Madrid, Spain
- 100 Institut für Physik, Universität Mainz, Mainz, Germany
- 101 School of Physics and Astronomy, University of Manchester, Manchester, UK
- 102 CPPM, Aix-Marseille Université, CNRS/IN2P3, Marseille, France
- 103 Department of Physics, University of Massachusetts, Amherst, MA, USA
- 104 Department of Physics, McGill University, Montreal, QC, Canada
- 105 School of Physics, University of Melbourne, Victoria, Australia
- 106 Department of Physics, University of Michigan, Ann Arbor, MI, USA
- 107 Department of Physics and Astronomy, Michigan State University, East Lansing, MI, USA
- 108 Group of Particle Physics, University of Montreal, Montreal, QC, Canada
- 109 Fakultät für Physik, Ludwig-Maximilians-Universität München, Munich, Germany
- 110 Max-Planck-Institut für Physik (Werner-Heisenberg-Institut), Munich, Germany
- 111 Graduate School of Science and Kobayashi-Maskawa Institute, Nagoya University, Nagoya, Japan
- 112 Department of Physics and Astronomy, University of New Mexico, Albuquerque, NM, USA
- 113 Institute for Mathematics, Astrophysics and Particle Physics, Radboud University/Nikhef, Nijmegen, The Netherlands
- 114 Nikhef National Institute for Subatomic Physics and University of Amsterdam, Amsterdam, The Netherlands
- 115 Department of Physics, Northern Illinois University, DeKalb, IL, USA
- 116 ^(a)New York University Abu Dhabi, Abu Dhabi, United Arab Emirates; ^(b)University of Sharjah, Sharjah, United Arab Emirates
- 117 Department of Physics, New York University, New York, NY, USA
- 118 Ochanomizu University, Otsuka, Bunkyo-ku, Tokyo, Japan
- 119 Ohio State University, Columbus, OH, USA
- 120 Homer L. Dodge Department of Physics and Astronomy, University of Oklahoma, Norman, OK, USA
- 121 Department of Physics, Oklahoma State University, Stillwater, OK, USA
- 122 Palacký University, Joint Laboratory of Optics, Olomouc, Czech Republic
- 123 Institute for Fundamental Science, University of Oregon, Eugene, OR, USA
- 124 Graduate School of Science, Osaka University, Osaka, Japan
- 125 Department of Physics, University of Oslo, Oslo, Norway
- 126 Department of Physics, Oxford University, Oxford, UK
- 127 LPNHE, Sorbonne Université, Université Paris Cité, CNRS/IN2P3, Paris, France
- 128 Department of Physics, University of Pennsylvania, Philadelphia, PA, USA
- 129 Department of Physics and Astronomy, University of Pittsburgh, Pittsburgh, PA, USA
- 130 ^(a)Laboratório de Instrumentação e Física Experimental de Partículas-LIP, Lisbon, Portugal; ^(b)Departamento de Física, Faculdade de Ciências, Universidade de Lisboa, Lisbon, Portugal; ^(c)Departamento de Física, Universidade de Coimbra, Coimbra, Portugal; ^(d)Centro de Física Nuclear da Universidade de Lisboa, Lisbon, Portugal; ^(e)Departamento de Física, Universidade do Minho, Braga, Portugal; ^(f)Departamento de Física Teórica y del Cosmos, Universidad de Granada, Granada, Spain; ^(g)Departamento de Física, Instituto Superior Técnico, Universidade de Lisboa, Lisbon, Portugal
- 131 Institute of Physics of the Czech Academy of Sciences, Prague, Czech Republic
- 132 Czech Technical University in Prague, Prague, Czech Republic

- 133 Faculty of Mathematics and Physics, Charles University, Prague, Czech Republic
- 134 Particle Physics Department, Rutherford Appleton Laboratory, Didcot, UK
- 135 IRFU, CEA, Université Paris-Saclay, Gif-sur-Yvette, France
- 136 Santa Cruz Institute for Particle Physics, University of California Santa Cruz, Santa Cruz, CA, USA
- 137 ^(a)Departamento de Física, Pontificia Universidad Católica de Chile, Santiago, Chile; ^(b)Millennium Institute for Subatomic Physics at High Energy Frontier (SAPHIR), Santiago, Chile; ^(c)Instituto de Investigación Multidisciplinario en Ciencia y Tecnología y Departamento de Física, Universidad de La Serena, La Serena, Chile; ^(d)Department of Physics, Universidad Andres Bello, Santiago, Chile; ^(e)Instituto de Alta Investigación, Universidad de Tarapacá, Arica, Chile; ^(f)Departamento de Física, Universidad Técnica Federico Santa María, Valparaíso, Chile
- 138 Department of Physics, University of Washington, Seattle, WA, USA
- 139 Department of Physics and Astronomy, University of Sheffield, Sheffield, UK
- 140 Department of Physics, Shinshu University, Nagano, Japan
- 141 Department Physik, Universität Siegen, Siegen, Germany
- 142 Department of Physics, Simon Fraser University, Burnaby, BC, Canada
- 143 SLAC National Accelerator Laboratory, Stanford, CA, USA
- 144 Department of Physics, Royal Institute of Technology, Stockholm, Sweden
- 145 Departments of Physics and Astronomy, Stony Brook University, Stony Brook, NY, USA
- 146 Department of Physics and Astronomy, University of Sussex, Brighton, UK
- 147 School of Physics, University of Sydney, Sydney, Australia
- 148 Institute of Physics, Academia Sinica, Taipei, Taiwan
- 149 ^(a)E. Andronikashvili Institute of Physics, Iv. Javakhishvili Tbilisi State University, Tbilisi, Georgia; ^(b)High Energy Physics Institute, Tbilisi State University, Tbilisi, Georgia; ^(c)University of Georgia, Tbilisi, Georgia
- 150 Department of Physics, Technion, Israel Institute of Technology, Haifa, Israel
- 151 Raymond and Beverly Sackler School of Physics and Astronomy, Tel Aviv University, Tel Aviv, Israel
- 152 Department of Physics, Aristotle University of Thessaloniki, Thessaloniki, Greece
- 153 International Center for Elementary Particle Physics and Department of Physics, University of Tokyo, Tokyo, Japan
- 154 Department of Physics, Tokyo Institute of Technology, Tokyo, Japan
- 155 Department of Physics, University of Toronto, Toronto, ON, Canada
- 156 ^(a)TRIUMF, Vancouver, BC, Canada; ^(b)Department of Physics and Astronomy, York University, Toronto, ON, Canada
- 157 Division of Physics and Tomonaga Center for the History of the Universe, Faculty of Pure and Applied Sciences, University of Tsukuba, Tsukuba, Japan
- 158 Department of Physics and Astronomy, Tufts University, Medford, MA, USA
- 159 United Arab Emirates University, Al Ain, United Arab Emirates
- 160 Department of Physics and Astronomy, University of California Irvine, Irvine, CA, USA
- 161 Department of Physics and Astronomy, University of Uppsala, Uppsala, Sweden
- 162 Department of Physics, University of Illinois, Urbana, IL, USA
- 163 Instituto de Física Corpuscular (IFIC), Centro Mixto Universidad de Valencia-CSIC, Valencia, Spain
- 164 Department of Physics, University of British Columbia, Vancouver, BC, Canada
- 165 Department of Physics and Astronomy, University of Victoria, Victoria, BC, Canada
- 166 Fakultät für Physik und Astronomie, Julius-Maximilians-Universität Würzburg, Würzburg, Germany
- 167 Department of Physics, University of Warwick, Coventry, UK
- 168 Waseda University, Tokyo, Japan
- 169 Department of Particle Physics and Astrophysics, Weizmann Institute of Science, Rehovot, Israel
- 170 Department of Physics, University of Wisconsin, Madison, WI, USA
- 171 Fakultät für Mathematik und Naturwissenschaften, Fachgruppe Physik, Bergische Universität Wuppertal, Wuppertal, Germany
- 172 Department of Physics, Yale University, New Haven, CT, USA

^a Also at Also Affiliated with an institute covered by a cooperation agreement with CERN, Geneva, Switzerland

^b Also at APC, Université Paris Cité, CNRS/IN2P3, Paris, France

^c Also at Borough of Manhattan Community College, City University of New York, New York, NY, USA

^d Also at Center for High Energy Physics, Peking University, Beijing, China

^e Also at Center for Interdisciplinary Research and Innovation (CIRI-AUTH), Thessaloniki, Greece

- ^f Also at Centro Studi e Ricerche Enrico Fermi, Rome, Italy
- ^g Also at CERN, Geneva, Switzerland
- ^h Also at Département de Physique Nucléaire et Corpusculaire, Université de Genève, Geneva, Switzerland
- ⁱ Also at Departament de Física de la Universitat Autònoma de Barcelona, Barcelona, Spain
- ^j Also at Department of Financial and Management Engineering, University of the Aegean, Chios, Greece
- ^k Also at Department of Physics and Astronomy, Michigan State University, East Lansing, MI, USA
- ^l Also at Department of Physics and Astronomy, University of Victoria, Victoria, BC, Canada
- ^m Also at Department of Physics, California State University, Sacramento, USA
- ⁿ Also at Department of Physics, King's College London, London, UK
- ^o Also at Department of Physics, Royal Holloway University of London, Egham, UK
- ^p Also at Department of Physics, University of Fribourg, Fribourg, Switzerland
- ^q Also at Department of Physics, University of Thessaly, Thessaly, Greece
- ^r Also at Department of Physics, Westmont College, Santa Barbara, USA
- ^s Also at Hellenic Open University, Patras, Greece
- ^t Also at Institutio Catalana de Recerca i Estudis Avancats, ICREA, Barcelona, Spain
- ^u Also at Institut für Experimentalphysik, Universität Hamburg, Hamburg, Germany
- ^v Also at Institute of Applied Physics, Mohammed VI Polytechnic University, Ben Guerir, Morocco
- ^w Also at Institute of Particle Physics (IPP), Toronto, Canada
- ^x Also at Institute of Physics and Technology, Ulaanbaatar, Mongolia
- ^y Also at Institute of Physics, Azerbaijan Academy of Sciences, Baku, Azerbaijan
- ^z Also at Institute of Theoretical Physics, Ilia State University, Tbilisi, Georgia
- ^{aa} Also at Lawrence Livermore National Laboratory, Livermore, USA
- ^{ab} Also at The Collaborative Innovation Center of Quantum Matter (CICQM), Beijing, China
- ^{ac} Also at TRIUMF, Vancouver, BC, Canada
- ^{ad} Also at Università di Napoli Parthenope, Naples, Italy
- ^{ae} Also at Department of Physics, University of Colorado Boulder, Boulder, CO, USA
- ^{af} Also at Washington College, Chestertown, MD, USA
- ^{ag} Also at Physics Department, An-Najah National University, Nablus, Palestine
- * Deceased