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TUNABLE FAE-INFRARED RADIATION GENERATED FROM THE DIFFERENCE FREQUENCY BETWEEN TWO RUBY LASERS

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TUNABLE FAR-INFRARED RADIATION GENERATED FROM THE  
DIFFERENCE FREQUENCY BETWEEN TWO RUBY LASERS

D. W. Farries\*, P. L. Richards, Y. R. Shen, and K. H. Yang

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Tunable Far-Infrared Radiation Generated from the  
Difference Frequency Between Two Ruby Lasers

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ABSTRACT

Far-infrared radiation generated from the difference frequency between two temperature-tuned ruby lasers operated on the  $R_1$  and  $R_2$  transitions has been observed. This radiation is continuously tunable over the frequency range from 20 to 38  $\text{cm}^{-1}$ . Lithium niobate was used as a mixing crystal. The expected frequencies of the far-infrared radiation were measured using a Fabry-Perot interferometer. The phase matching conditions were also verified.

In a previous paper we described the generation of tunable far-infrared radiation over the frequency range from 1.5 to  $8.1 \text{ cm}^{-1}$  by beating two temperature-tuned ruby laser beams in a nonlinear crystal.<sup>1</sup> By operating one laser on the  $R_1$  transition and the other on the  $R_2$  transition, we have now obtained tunable radiation from 20 to  $38 \text{ cm}^{-1}$ . In order to produce laser action on the  $R_2$  transition,<sup>2-4</sup> we used essentially a Lyot-Ohman<sup>5,6</sup> filter in the laser cavity to discriminate against the  $R_1$  transition.<sup>7</sup> The power output of the  $R_2$  laser was about 1/2 MW compared with the 1 MW output of the  $R_1$  laser. Lithium niobate was used for the non-linear crystal. The additional experimental apparatus and the measurement technique were essentially the same as described earlier.<sup>1</sup>

If both lasers are operated at room temperature, the difference-frequency radiation generated should be at  $29 \text{ cm}^{-1}$ . In order to phase-match this difference-frequency generation process, a 1.5 mm. slice of  $\text{LiNbO}_3$  was cut with the c-axis tilted approximately  $18^\circ$  away from the normal to the surface. The frequency of the far-infrared output was measured using a Fabry-Perot interferometer with electroformed metal mesh mirrors. The measured transmission curve<sup>8</sup> of the Fabry-Perot interferometer is compared in Fig. 1(a) to a theoretical curve calculated for a frequency of  $28.8 \text{ cm}^{-1}$ . Since the interferometer has a finesse of about 4, the spectral purity of the far-infrared radiation could not be measured. In Fig. 1(b) the Fabry-Perot transmission is shown for radiation generated with the  $R_1$  laser at room temperature and the  $R_2$  laser at  $-23^\circ\text{C}$ . The calculated transmission for  $35.8 \text{ cm}^{-1}$ , the

frequency expected from the known temperature dependence of the  $R_2$  transition, is also shown.<sup>9</sup> The same  $\text{LiNbO}_3$  crystal was used, but it was oriented for phase matching at  $35.8 \text{ cm}^{-1}$ .

The phase matching curve for production of the difference frequency at  $29 \text{ cm}^{-1}$  is shown in Fig. 2. The absorption coefficient for  $\text{LiNbO}_3$  at  $29 \text{ cm}^{-1}$  is  $18 \text{ cm}^{-1}$ .<sup>10</sup> Normalized theoretical curves obtained by solving Maxwell's equations in the plane wave approximation, with and without absorption,<sup>1,11</sup> are shown for comparison. The effect of absorption changes the width at half-maximum very little, but shows a definite difference at the wings of the curves. The shapes of the curves would not be changed appreciably by including diffraction and boundary conditions. The observed peak power in this case was about 5 mW, compared with a previously observed value of  $\sim 1 \text{ mW}$  at  $8.1 \text{ cm}^{-1}$  from a 0.47 mm. thick crystal.<sup>1</sup>

From the simple theory with plane-wave approximation, we expect an  $\omega^2$  dependence of the far-infrared output power on frequency. Since the extinction length at  $29 \text{ cm}^{-1}$  is 0.055 cm, as compared to crystal length of 0.47 cm in the  $8.1 \text{ cm}^{-1}$  case, we estimate that the power at  $29 \text{ cm}^{-1}$  should be about 15 times the power at  $8.1 \text{ cm}^{-1}$  for the same laser power. Computer calculations<sup>11</sup> including the effects of absorption, diffraction, boundary conditions, the spatial distribution and mode structure of the laser light, the birefringent property of the  $\text{LiNbO}_3$  crystal, and the collection angle of the detector gives a factor of 12. Since the power of the  $R_2$  transition laser is a factor of about two less than that of

the  $R_1$  transition, the complete theory predicts a factor of 6 compared to the factor 5 observed. This result seems satisfactory considering the uncertainties in our experiment.<sup>11</sup> The absolute value of the power obtained at  $8.1 \text{ cm}^{-1}$  was shown to agree with theory in the previous paper.

The tunable range of the far infrared radiation produced by the beating between  $R_1$  and  $R_1$  lasers and between  $R_1$  and  $R_2$  lasers could be easily extended from 0 to  $50 \text{ cm}^{-1}$  by using liquid nitrogen as a coolant. Other systems might enlarge the range of tunability. Many tunable laser sources exist which can be used to produce difference frequency radiation. Among the more promising are the dye lasers,<sup>12</sup> the spin-flip Raman lasers,<sup>13</sup> stimulated polariton scattering,<sup>14</sup> and parametric oscillators.<sup>15</sup> A system using two dye lasers could be tuned over the entire far infrared region. The power, linewidth, and divergence will perhaps never be as good as the ruby laser, but they may be adequate for some applications. Improvements in the ruby laser sources, particularly the divergence, should produce a significant increase in far infrared power. We can also increase the power by choosing the optimum focusing of the lasers into the crystal, provided that damage can be avoided. It seems clear that this tunable far infrared radiation source could be used for spectroscopy in the 1 to  $50 \text{ cm}^{-1}$  region, especially for saturation and other non-linear phenomena which require large peak power.

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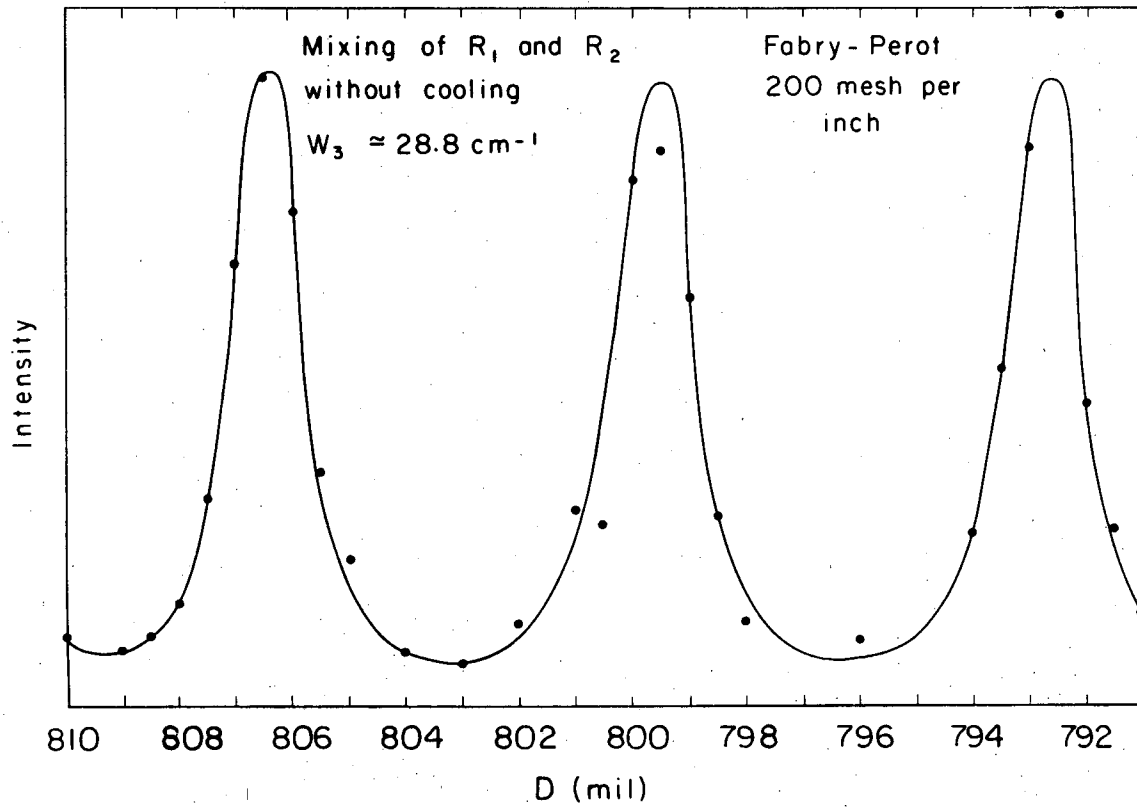
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FIGURE CAPTIONS

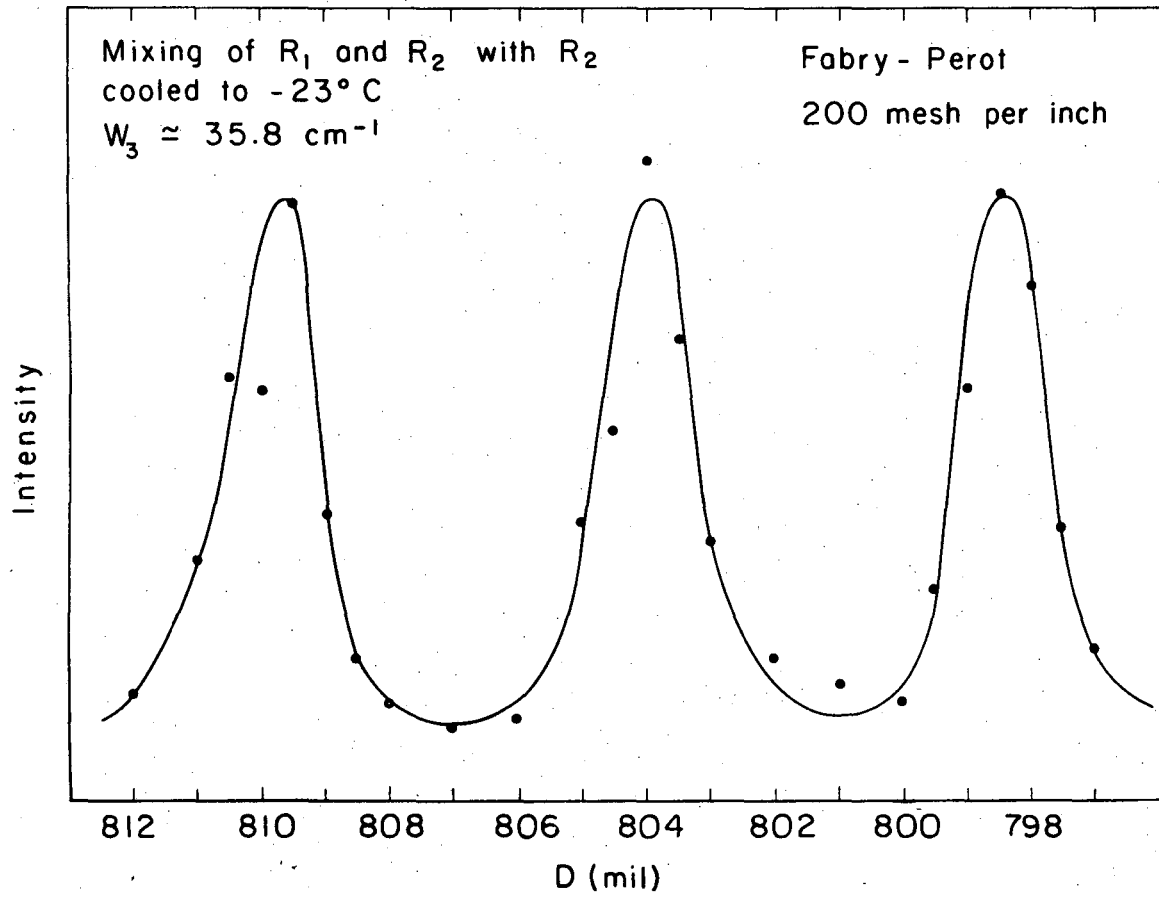
Fig. 1(a) Fabry-Perot scan of difference-frequency output using  $R_1$  and  $R_2$  lines. The first scan (a) shows a difference frequency of  $28.8 \text{ cm}^{-1}$ ; the second scan (b) shows  $35.8 \text{ cm}^{-1}$ . The solid curve is calculated from the dimensions of the wire grid interferometer and from the difference frequency expected from the laser temperature.

Fig. 2. Phase-matching curve for  $29 \text{ cm}^{-1}$  in  $0.15 \text{ cm LiNbO}_3$  crystal. The dashed line is  $(\frac{\sin \eta}{\eta})^2$ , the theoretical curve for no absorption. The solid line is the theoretical curve including an absorption coefficient of  $\alpha = 18 \text{ cm}^{-1}$ .



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Fig. 1(a)



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Fig. 1(b)

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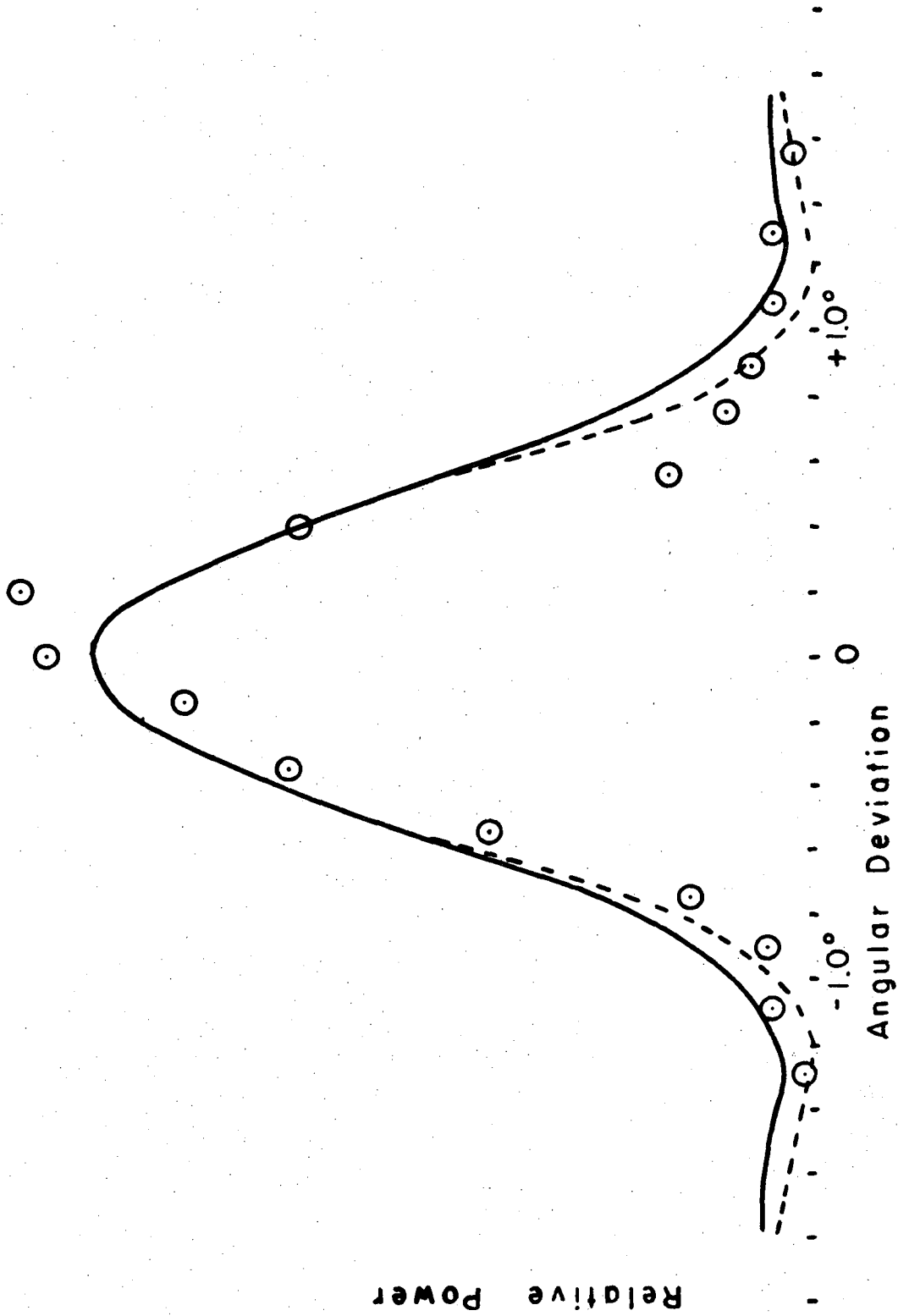


Fig. 2

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