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Authors

Kluesner, Jared W
Johnson, Samuel Y
Nishenko, Stuart P
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6 **High-Resolution Geophysical and Geochronological Analysis of**
7 **a Relict Shoreface Deposit Offshore Central California:**
8 **Implications for Slip Rate along the Hosgri Fault**

9

10

11 **Jared W. Kluesner¹, Samuel Y. Johnson¹, Stuart P. Nishenko², Elisa Medri³, Alexander R.**
12 **Simms³, H. Gary Greene⁴, Harrison J. Grey⁵, Shannon A. Mahan⁵, Emma T. Krolczyk⁵,**
13 **Daniel S. Brothers¹, James E. Conrad¹**

14

15 *¹U.S. Geological Survey, Pacific Coastal and Marine Science Center, Santa Cruz, CA*

16 *²Pacific Gas and Electric (retired)*

17 *³Department of Earth Sciences, University of California Santa Barbara, Santa Barbara, CA*

18 *⁴Moss Landing Marine Laboratories, Moss Landing, CA*

19 *⁵U.S. Geological Survey, Geosciences and Environmental Change Science Center, Denver, CO,*

20 *80225*

21

22 **ABSTRACT**

23 The Cross-Hosgri Slope (CHS) is a bathymetric lineament that crosses the eastern strand
24 of the Hosgri Fault (EHF) offshore Point Estero, central California. Recently collected CHIRP
25 seismic reflection profiles and sediment cores provide the basis for a reassessment of CHS origin
26 and the proposed lateral slip rate of the EHF based on offset of the CHS lower slope break. The
27 CHS is comprised of two distinct stratigraphic units. The lower unit (Unit 1) overlies the post-
28 Last Glacial Maximum transgressive surface of erosion and is interpreted as a Younger Dryas
29 (~12.8 to 11.5 ka) shoreface deposit based on radiocarbon and OSL ages, seismic facies,
30 sediment texture, sediment infauna, and heavy mineral component. The shoreface was
31 abandoned and partly eroded during rapid sea-level rise from about 11.5 to 7 ka. Unit 2 consists
32 of fine sand and silt deposited in a mid-shelf environment when the rate of sea-level rise slowed
33 between ~7 ka and the present. Although Unit 2 provides a thin, relatively uniform cover over
34 the lower slope break of the older shoreface, this feature still provides a valuable piercing point,
35 and the previously estimated EHF slip rate of 2.6 ± 0.9 mm/yr remains valid. Full-waveform
36 processing of CHIRP data resulted in significantly higher resolution in coarser grained strata that
37 are typically difficult to interpret with more traditional envelope processing. Our novel
38 combination of offshore radiocarbon and OSL dating is the first application for offshore
39 paleoseismic studies, and our results indicate the utility of this approach for future marine
40 neotectonic investigations.

41

42 **INTRODUCTION**

43 Assessment of earthquake hazards requires comprehensive characterization of active
44 faults, including documentation of fault location, length, connectivity, slip rate, and rupture

45 history. Acquiring such documentation for faults in the near offshore can be especially important
46 given the possible impact on commonly dense coastal population and infrastructure. The target
47 of this paper is the slip rate of one such structure, the Hosgri Fault of central California (Fig. 1).
48 We focus on a distinct, linear, southwest-facing bathymetric slope referred to as the cross-Hosgri
49 slope (CHS; Johnson et al., 2014; Fig. 2). The CHS crosses and is offset by the main strand of
50 the Hosgri Fault about 30 to 35 km north of Point Buchon and the Pacific Gas and Electric
51 (PG&E)-operated Diablo Canyon Nuclear Power Plant.

52 Using global sea-level curves (Stanford et al., 2011), Johnson et al. (2014) interpreted the
53 CHS as the relict shoreface of a now-eroded latest Pleistocene sand spit that formed during sea-
54 level rise following the Last Glacial Maximum (LGM). The current depth of the CHS shoreface
55 approximately coincides with sea level during a period of relatively slower eustatic rise, the
56 Younger Dryas stadial ($\sim 12,800\text{--}11,500$ yr B.P.), and the CHS was inferred to have been
57 rapidly submerged and preserved during the following period of rapid eustatic rise, from about
58 11,500 to 7,000 yr B.P. Using high-resolution bathymetry (Fig. 2 inset), Johnson et al. (2014)
59 plotted and analyzed closely spaced (12.5 m) normal-to-slope profiles bracketing the trace of the
60 Hosgri fault where it crosses the lower slope break of the CHS. They documented 30.3 ± 9.4 m
61 (mean and 2 standard deviations) of dextral fault offset and about 70 to 110 cm of vertical offset.
62 They then used the estimated Younger Dryas age of the CHS and the amount of dextral offset to
63 infer a slip rate of 2.6 ± 0.9 mm/yr on the main strand of the Hosgri fault.

64 Both the depositional model of the CHS and the Hosgri slip rate in Johnson et al. (2014)
65 were derived by correlation with a global sea-level curve. Highest-resolution Compressed High
66 Intensity Radar Pulse (CHIRP) imagery and ground truthing constraints from CHS sediments,

67 including physical properties, geochronology, and petrology, were lacking. Given the limitations
68 and uncertainty associated with the Johnson et al. (2014) approach, USGS, PG&E, and the
69 University of California at Santa Barbara collaborated through Cooperative Research and
70 Development Agreements on a 2019 geophysical and coring program focused primarily on the
71 CHS. The goals of the program were to better understand CHS stratigraphy and sedimentology,
72 and to provide samples for geochronologic studies. Initial results of this recent work are
73 presented in Medri et al. (2023), who show that the CHS is underlain by distinct, post-LGM,
74 stratigraphic units deposited in environments ranging from shoreface to mid-shelf and ranging in
75 age from about 12,400 yr B.P. to Modern based on radiocarbon dating. Medri et al. (2023) focus
76 on the sedimentology of a middle unit of mid-shelf sediments deposited between about 7,000 and
77 1,000 yr. B.P., which was interpreted as a subaqueous clinoform built by wave-supported gravity
78 flows. Our paper primarily focuses on the underlying lowest post-LGM stratigraphic unit, which
79 is used for estimating fault slip rate. We supplement the radiocarbon dating of Medri et al. (2023)
80 with new Optically Stimulated Luminescence (OSL) dating of quartz sediment grains to provide
81 further constraints on the geochronology of the CHS. In addition, we use advanced processing
82 techniques on full-waveform CHIRP data to greatly improve the vertical resolution and clarity of
83 the seismic stratigraphy and fault offsets within the coarse grain and high-energy environment.
84 Results document the CHS and Hosgri Fault in far greater detail and reveal that the CHS is more
85 complex than previously envisioned; however, the additional chronology supports the fault slip
86 rate of 2.6 ± 0.9 mm/yr proposed in Johnson et al. (2014).

87

88 **THE HOSGRI FAULT AND CROSS-HOSGRI SLOPE**

89 The Hosgri fault is the westernmost active fault within the broader San Andreas Fault
90 System and is the southern portion of the San Gregorio-Hosgri fault system (SGHF, Fig. 1). The
91 SGHF extends along the central California coast for about 360 km and is one of the more active
92 structures within the distributed plate boundary along the west coast of North America
93 (Dickinson et al., 2005). Johnson et al. (2018) recognized three distinct sections of the fault
94 based on different fault azimuths and geomorphology: (1) ~110-km-long southern section,
95 including the CHS, that extends from offshore Point Arguello to San Simeon, has a mean
96 transtensional azimuth of $336^\circ \pm 8^\circ$, and is characterized by moderate coastal relief, large coastal
97 embayments, the mouths of the Santa Maria and Santa Ynez rivers, and a wider continental
98 shelf, (2) a ~80-km-long middle section (the “Big Sur Bend”) that extends from San Simeon to
99 Point Sur, has a mean transpressional azimuth of $321^\circ \pm 9^\circ$, and is marked by steep coastal relief,
100 small coastal watersheds, and a narrow and deeply incised continental shelf, and (3) a ~170-km-
101 long northern section extends from Point Sur to Bolinas, where the fault converges with the San
102 Andreas fault. The northern section has a mean transtensional azimuth of $337^\circ \pm 6^\circ$ and is
103 generally characterized by lower coastal relief (south of San Andreas fault influence), the large
104 coastal embayment of Monterey Bay, the mouths of the Salinas and Pajaro rivers, and a wider
105 continental shelf.

106 Based on offset piercing points, SGHF slip rate over the last 10 Ma averages about 14
107 mm/yr (e.g., Clark, 1997; Dickinson et al., 2005), but this rate must have decreased significantly
108 in the late Neogene. SGHF lateral slip-rate estimates based on late Quaternary features vary and
109 increase from south to north, with a slip rate of about 2.6 mm/yr at the CHS location (Johnson et
110 al., 2014). To the north at San Simeon (Fig. 1), Hall et al. (1994) and Hanson and Lettis (1994)
111 estimated “best constrained” rates of 0.9 to 3.4 mm/yr and 1 to 3 mm/yr , respectively. Johnson

112 et al. (2018) derived a slip rate estimate of about 3.35 mm/yr based on offset canyons heads
113 offshore Lopez Point in the Big Sur Bend. PG&E (2014) estimated slip rates of 1.6 mm/yr and
114 1.8 mm/yr in Estero Bay and offshore Point Sal, respectively. These rates are notably lower than
115 the ~11 mm/yr late Quaternary slip rate estimated by Weber (1990) for the SGHF at Point Año
116 Nuevo. Increases in fault slip rate from south to north have been attributed to “adding slip” from
117 northwest-converging faults, including the Lions Head, Casmalia, Shoreline, Oceano, Los Osos,
118 Oceanic, Nacimiento, and Rinconada-Monterey Bay faults (Johnson and Watt, 2012;
119 Langenheim et al., 2013; Watt et al., 2015; Colgan and Stanley, 2016; Johnson et al., 2015; 2018,
120 2019; Nishenko et al., 2018).

121 The CHS occurs at ~68 to 76 m water depth in the northern part of the southern SGHF
122 section, about 20 km south of San Simeon and 5 km northwest of Estero Point (Johnson et al.,
123 2014; Figs 1 and 2). In this region the Hosgri fault splits into western and eastern strands, the
124 latter which crosses the CHS (Fig. 2). The CHS has a linear, southwest-facing trend (Fig. 2), a
125 height of 7-9 m, a length of 1700 m, and a width of 250-280 m. The feature is characterized by a
126 slope dipping 1.6° - 2.0° to the SW, a considerably steeper angle than the surrounding seafloor to
127 the northeast and southwest, which dips more gently at 0.4° - 0.6° . Seismic-reflection CHIRP
128 and coring data (see below) indicate the CHS is underlain by post-LGM sediment and is crossed
129 by the main (eastern) strand of the Hosgri fault.

130

131 **METHODS**

132 **High-Resolution CHIRP Data**

133 Approximately 450 km of high-resolution sub-bottom profiles were collected along the
134 central California shelf onboard the M/V *Bold Horizon* in 2019, including 7 profiles across the

135 CHS (Fig. 2; Snyder et al., 2022). Sub-bottom profiles were collected using an EdgeTech 2300
136 SB-516 CHIRP system. Initially, deterministic deconvolution processing (match filtering) was
137 conducted onboard the CHIRP to remove the sweep signature from the data. Edgetech output
138 JSF formatted files were then converted to SGY to access the raw correlated signal. Processing
139 steps in Shearwater Reveal were applied to the real (non-Hilbert transformed) portion of the data
140 (Henkart, 2006) and included towed depth correction, trace balancing, multi-step static (swell)
141 removal, predictive deconvolution, automatic gain adjustment, and water column mute. As the
142 CHIRP was towed off the starboard side of the ship (vs the stern), predictive deconvolution was
143 applied to remove a short-path multiple likely caused by a reflection off the side of the hull.
144 Horizon mapping and interpretation was done using HIS Kingdom Suite seismic interpretation
145 software.

146

147 **Bathymetry Data**

148 Bathymetric data used in this study includes high-resolution sonar data collected by the
149 California State University of Monterey Bay Seafloor Mapping Lab (CSUMB) between 2009
150 and 2010 as part of the California Seafloor Mapping Program (CSUMB, 2012; Johnson et al.,
151 2017). These data consist of a combination of sonar systems that collected data out to 5.6 km
152 offshore (the 3 nautical mile limit California's State Waters). Additional bathymetric data were
153 collected by the USGS in 2012 using a Reson 7111 multibeam sonar that extended coverage
154 beyond the 3 nautical mile limit and included a patch along the CHS (Figure 2; Hartwell et al.,
155 2013).

156

157 **Sediment Coring and Imaging**

158 A total of 23 sediment cores were collected across the shelf using a Rossfelder P-5
159 vibracorer, including 6 cores on the CHS, 1 core at about 80 m water depth located below the
160 CHS lower slope break, and 1 core located to the south at 113 m near the shelf break (Fig. 2;
161 Snyder et al., 2022). Cores longer than 1.5 m were split into multiple sections to permit scanning
162 in a Geotek RXCT X-ray tomography system. Full 3D CT scans of all cores were generated,
163 along with XZ and YZ 2D plane slices for each (Fig. 3). Cores were split and then photographed
164 using the Geotek multi-sensor core logger (MSCL) system. P-wave velocity was measured using
165 the MSCL on un-split cores HF-1, HF-3, HF-5, and HF-7, providing a mean velocity of 1617
166 m/s.

167

168 **Sand Petrology**

169 Samples from sandy intervals in cores were collected and processed to generate fractions
170 of uniform fine grain size (125-250 microns). Steps included wet sieving, air and oven drying,
171 followed by dry sieving. Standard thin sections with clear epoxy impregnation and potassium
172 feldspar stain were examined petrographically and point counted, 323 to 408 points per sample
173 (Table 1; Snyder et al., 2022).

174

175 **Radiocarbon Dating**

176 A total of 30 radiocarbon samples and ages were collected from the 7 cores collected in
177 the CHS region (Snyder et al., 2022; Medri et al., 2023). This included 23 gastropod shells, one
178 bivalve shell, and six wood fragments. Gastropods that showed no evidence of reworking were
179 used and species were evaluated to assess habitat (Medri et al., 2023). Radiocarbon ages were
180 determined using atomic mass spectrometry at the University of California Irvin carbon cycle

181 accelerator mass spectrometer (KCCAMS) facility. ^{14}C ages calculated from shells used the
182 Marine 20 calibration curve of (Heaton et al., 2020) for calibration, whereas wood fragments
183 used the Intcal calibration curve of (Reimer et al., 2020) within the Calib 8.2 program (Stuiver et
184 al., 2022).

185

186 **Optically Stimulated Luminescence (OSL) Dating**

187 To supplement the radiocarbon dating, we applied OSL dating to determine when quartz
188 grains were last exposed to daylight (Murray et al., 2021). Following the methods of Nelson et
189 al. (2019), we split cores and extracted samples for equivalent dose and dose rate measurements
190 under dark-room conditions. Elemental concentrations were determined by inductively-coupled
191 plasma mass spectrometry on samples immediately above, below, and surrounding the
192 equivalent dose sample and dose rates calculated using the Dose Rate and Age Calculator
193 (Durcan et al., 2015). We applied standard methods to determine the equivalent dose (i.e.,
194 Murray et al., 2021) using the single aliquot regeneration protocol (Murray and Wintle, 2000) on
195 ≤ 2 mm sized masked aliquots of quartz sand. This means that for each aliquot less than 300
196 grains were analyzed in an averaged luminescence signal (Duller, 2008). OSL characteristics of
197 the samples were favorable with a dominant fast component (Durcan and Duller, 2011) and high
198 percentages of aliquots passed our rejection criteria (Wintle and Murray, 2006). We did not
199 observe substantial evidence of partial bleaching, which would be manifested by high
200 overdispersion and skewed dose distributions (e.g., Fig. 4). We calculated ages using the central
201 age model (CAM) and minimum age model (MAM) from functions in the "luminescence"
202 package for the programming language R (Kreutzer et al., 2013). We found that the CAM
203 produced the best agreement with the radiocarbon ages and thus argue that this model produces

204 the most accurate, best-bleached exposure age estimates for the sampled core sediment. For
205 additional details and figures related to the methodology please see the supplemental report and
206 Snyder et al. (2022).

207

208

209 **RESULTS**

210 Below we present results that establish the stratigraphic and chronologic development of
211 the CHS and characterize fault activity. Within the CHS, Medri et al. (2023) defined three
212 seismic stratigraphic units (S1, S2, S3) above a post-LGM transgressive surface of erosion based
213 on seismic-reflection data, coring, and geochronology. Their S1 unit is here labeled “Unit 1,” and
214 their S2 and S3 units are here grouped into our “Unit 2,” as described below. The measured
215 mean P-wave velocity of 1,617 m/s is used for time-to-depth conversion for thickness estimates
216 from the CHIRP imaging.

217

218 **Transgressive Surface of Erosion**

219 Global sea level was about 120 to 130 m lower than present about 21 ka BP during the
220 LGM (Stanford et al., 2011), at which time most of the central California shelf was emergent and
221 the shoreline was at or near the current shelfbreak (Johnson et al., 2019). Landward migration of
222 the shoreline across the ~ 6-km wide emergent shelf in this area is marked by a prominent, shelf-
223 wide, unconformity (Johnson and Watt, 2012; Johnson et al., 2019). This unconformity, a
224 transgressive surface of erosion in seismic-stratigraphic nomenclature (e.g., Catuneanu, 2006), is
225 imaged on CHIRP seismic-reflection profiles as a relatively continuous moderate- to high-
226 amplitude reflection bounding and truncating reflection-free and weakly reflective material below

227 and an overlying zone of moderate- to high-amplitude discontinuous reflections (Figs. 5, 6, 7, 8).
228 The transgressive surface can be traced across the region from the mid to the outer shelf on
229 CHIRP profiles and was penetrated in four cores collected at water depths of 101 to 113 m. In
230 core HF12, the surface is represented as a shell lag that yielded both pre-LGM and post-LGM
231 radiocarbon dates (e.g., Fig. 8). Such shell lags commonly occur along and above transgressive
232 erosion surfaces as discontinuous scour fills of variable thickness (Catuneanu, 2006), and we
233 infer that they are the source of the high-amplitude reflections common in the lower part of Unit
234 1 (see below).

235

236 **Unit 1 Shoreface Deposits**

237 Unit 1 (Unit S1 of Medri et al., 2023) overlies the transgressive surface of erosion,
238 forming the lower stratigraphic unit within the CHS. Unit 1 is characterized by variable seismic
239 facies including parallel to locally divergent, low- to high-amplitude reflections and
240 homogeneous “reflection-free” zones (Figs. 5, 6, 7). Reflections commonly dip offshore about
241 1.8°, similar to the slope of the CHS seafloor. Laterally, the ~1300-m-long CHIRP profile that
242 extends across the mid-CHS (Fig. 7) shows that the lower boundary of the unit has an erosional,
243 channeled surface, whereas the upper boundary is mostly smooth. The thickness of Unit 1 ranges
244 laterally from about 100 to 310 cm, with much of the variation is due to the channeled relief on
245 the lower erosional boundary. Unit 1’s lower and upper boundaries also converge in both the
246 onshore (upslope of the CHS) and offshore (downslope of the CHS) directions, so that the
247 stratigraphic horizon forms a distinct lens with maximum thickness located beneath the CHS
248 (Figs. 5, 6)

249 Within the CHS, Unit 1 was penetrated the lower 12 cm of HF-1, the lower 24 cm of HF-
250 2, the lower 29 cm of core HF-4, the basal core cutter of HF-3, (Figs. 3, 5A), and in the lower 50
251 cm in core HF-7, located about 80 m below the CHS lower slope break (Fig. 6B). Medri et al.
252 (2023) describe the cored sediment as the “black sand facies,” composed of well-sorted fine-
253 grained sand (mean grain size 130 μm) with less than 5% silt, no clay, and sparse shell
254 fragments. They report that shells within Unit 1 include *Clathurella canfieldi* and *Truncatella*
255 *californica*, which prefer a sandy habitat, among surf grass roots in the upper intertidal zone
256 (Guz, 2007). Also found was *Turitella Cooperi*, which can occur at intertidal to mid-shelf water
257 depths.

258 Seismic, sedimentologic, and faunal data indicate Unit 1 was deposited in a shallow
259 water shoreface environment like modern shorefaces along the high-energy California coast
260 (e.g., Barnard et al. 2007, 2009). Seaward-dipping reflections (Figs. 5, 6) indicate primary
261 depositional slope and progradation. Sediment texture (grain size and sorting), the relative lack
262 of infauna, and the significant proportion of heavy minerals (~8.4%, see below) in the sand are
263 also consistent with a high-energy depositional environment.

264 We obtained nine radiocarbon dates and five OSL dates from unit 1 to constrain its age
265 (Tables 1, 2). Eleven samples are from the upper 28 cm of Unit 1 beneath the CHS in cores HF-
266 1, 2, 3, and 4. Three samples are from the upper 50 cm of unit 1 in core HF-7, located about 80
267 m below the CHS lower slope break. The mean age of the nine calibrated radiocarbon dates is
268 10,830 yr B.P with a variation in mean ages ranging from 9,690 to 11,968 yr B.P. The mean age
269 of the five OSL dates is 11.8 ka., with a range of 10.6 ± 0.5 to 12.4 ± 0.4 ka, and four of the dates
270 fall between 11.8 and 12.4 ka. As discussed above, the OSL dates indicate the time that sand
271 grains were last exposed to sunlight on the beach or in very shallow water, whereas the

272 radiocarbon dates mainly indicate the time that marine organisms were living on or within the
273 sediment.

274 Geochronologic data provide a good fit for deposition during or immediately after the
275 Younger Dryas stadial, about 12,800 to 11,500 yr B.P.. The rate of sea-level rise significantly
276 declined during this stadial (Stanford et al., 2011), providing an opportunity for shoreface
277 progradation. All the dates are from the upper 50 cm of a unit that is as thick as 310 cm. Four of
278 the five OSL dates and two of the nine radiocarbon dates fall within the Younger Dryas time
279 interval, five of the seven remaining radiocarbon dates are slightly younger, between ~10,600
280 and 11,000 yr B.P., and two dates are about 9,700 to 9,800 yr B.P. The younger seven
281 radiocarbon dates are from the uppermost part of the unit and likely reflect late- and post-
282 Younger Dryas physical and biological reworking (e.g., burrowing, grazing) of the upper part of
283 the unit.

284

285 **Unit 2 – Mid-Shelf Deposits**

286 Unit 2 overlies Unit 1 beneath the CHS (Medri et al., 2023). The lower contact is marked
287 by a discontinuous reflection that coincides with an upward transition from moderate- to high-
288 amplitude reflections at the top of Unit 1 to the zone of diffuse, seafloor-parallel, low- to
289 medium-amplitude reflections that characterize Unit 2 (Figs. 5, 6, 7). The lower contact of Unit 2
290 is parallel to low-angle discordant, locally truncating reflections in the upper part of Unit 1. The
291 top of Unit 2 is the seafloor.

292 Unit 2 is up to 300 cm thick, reaching a maximum thickness on the upper CHS (Fig. 5A).
293 Both CHIRP and coring data (Figs. 5, 6) indicate that Unit 2 thins markedly downslope, nearly
294 pinching out at the lower slope break. Paired cores HF-5 and HF-6 on the upper CHS penetrated

295 a complete 300-cm-thick section of Unit 2. Paired cores HF-3 and HF-4 on the mid slope, and
296 paired cores HF-1 and HF-2 on the lower slope penetrated complete 140-cm- and 80-cm-thick
297 sections, respectively. Core HF-7, located about 80 m below the CHS lower slope break,
298 penetrated a complete 30-cm-thick section of Unit 2. CHIRP profiles (Fig. 5, 6, 7) confirm this
299 downslope thinning, and show variable Unit 2 thinning across the CHS.

300 Medri et al. (2023) subdivided Unit 2 into 3 distinct subunits based on their sediment
301 characteristics. The lowest subunit consists of 50% sand, 40% silt, and 10% clay, and contains
302 abundant shell fragments throughout. Strata are organized in 10– to 20-cm-thick fining-upward
303 sequences bounded by erosion surfaces. Internally, sediments are parallel- and ripple-laminated.
304 Shell fragments are abundant as are whole shells of gastropods (*Amphissa versicolor*, *Callianax*
305 *Baetica*) that commonly occur in mid-shelf water depths (~40 to 70 m). This lower subunit thins
306 from 200 cm in HF-5 on the upper slope, to 70 cm in HF-3 on the mid slope, to 60 cm in HF-1
307 on the lower slope (Figs 3, 5A). This subunit is not present in core HF-7, located about 80 m
308 below the lower slope break (Figs 2, 6B).

309 The middle subunit of Unit 2 consists of thin (~2-cm-thick) beds of sandy shell hash
310 comprised of granule to cobble-sized shell fragments in a sandy matrix, bounded by thin beds of
311 parallel-laminated sands. This subunit forms one distinctive, low-angle, ~250-m-wide channel,
312 filled with relatively higher-amplitude reflections, imaged only on the eastern part of the CHS
313 (Figs. 5A, 7A). Where it does occur, the thickness of this subunit decreases from 85 cm in core
314 HF-5 on the upper CHS to 30 cm in HF-3 on the mid CHS, and it is not present in core HF-1 on
315 the lower CHS or in core HF-7, about 80 m below the lower slope break of the CHS. Impedance
316 contrasts associated with the contrasting “hard” shell fragments and “softer” interbeds of fine-

317 grained sediment are the likely source of the discontinuous, moderate- to high-amplitude
318 reflections that characterize this subunit.

319 The upper subunit of Unit 2 consists of massive sandy silt with local burrows and
320 scattered shell fragments. This subunit occurs in all the cores, ranging in thickness from 15 cm
321 on the upper CHS in core HF-5, to 40 cm in core HF-3 on the mid CHS, to 20 cm in core HF-1
322 on the lower slope, to 30 cm in core HF-7 about 80 m below the CHS lower slope break.

323 We obtained twenty radiocarbon dates and three OSL dates from Unit 2 to constrain its
324 age (Tables 1, 2). The calibrated radiocarbon dates range from about 7,400 yr B.P to Modern,
325 and the OSL dating yields ages of 8.55 to 2.47 ka. We exclude consideration of one out-of-
326 sequence radiocarbon date from HF-5 (~9,400 yr B.P.) because seven deeper samples from this
327 core yielded younger, in-sequence ages, from ~7,400 to 3,400 yr B.P. As noted above, the timing
328 of the onset of Unit 2 deposition (~7,400 yr. B.P.) coincides approximately with the substantial
329 decrease in the rate of sea-level rise, from more than 10 m/1000 years to about 1 m/1,000 years
330 (Peltier and Fairbanks, 2006; Stanford et al., 2011; Reynolds and Simms, 2015). Correlation with
331 global and California sea-level curves indicate unit 2 was deposited at mid-shelf water depths.

332 Medri et al. (2023) propose that Holocene unit 2 nucleated on the pre-existing relict of a
333 latest Pleistocene (Younger Dryas) shoreface. They infer that sediment transport of the lowest
334 subunit was initiated on the mid shelf by winter-storm waves, with deposition occurring on the
335 CHS as wave-supported gravity flows (WGSF). The overlying shell-hash subunit was interpreted
336 as wave-winnowed shelf deposits. The lowest Unit 2 subunit thins markedly downward on the
337 CHS, and the middle shell-hash subunit pinches out both downward and laterally on the CHS.
338 Medri et al. (2023) interpret the upper Unit 2 sandy silt subunit as a young drape that was
339 deposited by suspension on the CHS and adjacent shelf.

340

341 **Sand Petrology**

342 Petrographic analysis of fourteen samples of fine-grained sand from Units 1 and 2 is
343 summarized in Table 3 and Figure 9. The siliciclastic framework grain proportions of Unit 1 and
344 Unit 2 are nearly identical (mean QFL = 30,18,52), with somewhat more variation in the upper
345 unit. The sand is derived from the varied bedrock in the Franciscan Complex in the adjacent,
346 southern Santa Lucia Range (Graymer et al., 2014). Prominent Franciscan lithologies in this
347 source area include melange, sandstone, graywacke, conglomerate, greenstone, diabase, chert,
348 serpentinite, and glaucophane schist.

349 Although the framework of the siliciclastic component of Units 1 and 2 are similar, there
350 are significant differences in the proportions of bioclasts and heavy minerals between the two
351 units. Mean bioclast proportion in Unit 1 and Unit 2 are 5.3% and 31.9%, respectively. Bioclasts
352 consists entirely of shell fragments in Unit 1, but Unit 2 includes both shell fragments and
353 common whole shells (including foraminifera tests). Heavy minerals identified based on color
354 (plain light), relief, pleochroism, and birefringence, include a mix of amphibole (e.g.,
355 hornblende) and pyroxene (e.g., augite), and less common mineral types including epidote,
356 sphene, and magnetite. The mean proportion of heavy minerals in Unit 1 and Unit 2 are 8.4%
357 and 2.2%.

358 The proportions of heavy minerals and bioclasts are consistent with the contrasting
359 depositional environments for Unit 1 and Unit 2. Because of their specific gravity, heavy
360 minerals should have increased presence in a high-energy shoreface environment (Unit 1) that is
361 also not hospitable for significant shell-bearing fauna. In contrast, the inferred mid-shelf
362 environment of Unit 2 provides a less energetic location for shell-bearing fauna (the source of

363 bioclasts) and is a less likely location for heavy minerals given the need for significant offshore
364 sediment transport.

365

366 **Hosgri Fault Characterization**

367 The CHIRP profiles that cross the CHS provide ultra high-resolution imagery of the
368 Hosgri fault in the upper 1 to 6 m of the subsurface. Thus, the seismic profiles are representative
369 in scale to roadcut or trench exposures than to seismic-reflection profiles generated by other
370 high-resolution seismic-reflection systems such as mini-sparkers (e.g., Johnson and Watt, 2012;
371 Kluesner et al., 2019). The appearance of the fault varies significantly in the CHIRP imagery
372 within the small CHS area (Fig. 2) as it cuts through different lithologies and stratigraphy at
373 different orientations. Although most motion on the Hosgri Fault is strike slip, the transgressive
374 surface of erosion and the contact between Units 1 and 2 provide markers for documenting
375 apparent vertical fault slip; these markers do not provide information on the more significant
376 strike-slip offset along the fault.

377 The Hosgri Fault cuts across or is located just below the CHS as exhibited on 3 of the 4
378 slope-normal CHIRP profiles. Just below the lower slope break on profile HFC-3 (Figs. 2, 5B),
379 the Hosgri Fault is imaged as a sharp contact that warps reflectors in Unit 1, and vertically offsets
380 (up to the west) the transgressive surface of erosion about 110 cm and the shallow (~30 cm deep)
381 Unit 1-Unit 2 contact about 60-70 cm. Based on bathymetric slope profiles, Johnson et al. (2014,
382 their Fig. 10) previously estimated ≥ 70 cm (as much as 130 cm) of up-to-the-west vertical offset
383 on the lower slope break along the Hosgri.

384 Profile HFC-4 images the Hosgri Fault where it crosses the CHS about 20 m below the
385 upper slope break (Figs. 2, 6A). The fault is again imaged as a sharp near-vertical plane,

386 offsetting the transgressive surface of erosion and the Unit 1-Unit 2 contact about 120 cm and 60
387 cm, respectively. Profile HFC-2 (Figs. 2, 6B) crosses the Hosgri Fault about 60 m above the CHS
388 upper slope break, imaging the fault as a ~120-m-wide zone of chaotic reflections across which
389 reliable vertical offsets cannot be determined.

390 Vertical offsets of the Unit 1-Unit 2 contact, along with sediment ages, indicate the
391 Hosgri Fault has been active in the last ~7,000 years. The lack of continuous, traceable internal
392 reflections within Unit 2 limits our ability to discern and measure vertical offsets from any single
393 post-7 ka earthquakes. That the Hosgri Fault does not offset the seafloor on any of the profiles
394 likely indicates that shelf processes (deposition and erosion) have been sufficient to smooth over
395 any vertical relief that might have been generated during the most recent earthquake(s).

396

397

398 **DISCUSSION**

399 **Depositional Model**

400 The comprehensive geophysical, sedimentologic, and geochronologic data reported
401 above, and the results of Medri et al., 2023, provide the basis for a substantial revision of the
402 Johnson et al. (2014) CHS depositional model. These data suggest the following CHS
403 depositional history, with key stages shown in Figure 10.

404 A. Rapid sea-level rise from the LGM (~21 ka) to the beginning of the Younger Dryas
405 stadial (~12.8 ka) provided significant accommodation for available coastal sediment supply. A
406 transgressive surface of erosion was generated at the high-energy shoreline, preserved at least
407 locally as a thin shell lag (Fig. 8), the likely source of a nearly continuous, moderate amplitude
408 reflection on CHIRP data (Fig. 4). Clastic sediment derived from the coastal landscape and the

409 emergent part of the shelf was eroded and redeposited offshore as the shoreline migrated
410 landward.

411 B. Sea-level rise slowed considerably during the Younger Dryas stadial, which extended
412 from about 12.8 to 11.5 ka (Stanford et al., 2011). As a result, landward migration of the
413 transgressive surface and shoreline slowed and a prograding beach and shoreface developed (Fig.
414 9B). The CHS lower slope break in Unit 1 occurs at depths of about 73 m below modern sea level
415 and is inferred to represent the base of the shoreface at water depths of about 10 m, based on the
416 depths of modern shorefaces (5 to 15 m) along the California coast (Barnard et al., 2007, 2009).
417 Thus, sea level should have been about 58 to 68 m below present at 11.5 ka when Unit 1
418 deposition ended. Given the presence of an emergent rocky point at this sea level at the west end
419 of the CHS (Hosgri Ridge on Fig. 2), the prograding beach may have formed a sand spit bounded
420 landward by a lagoon, as suggested by Johnson et al. (2014, their Fig. 7). A preserved portion of
421 this shoreface forms Unit 1 of the CHS, consistent with the sedimentologic, petrologic, and
422 geochronologic data outlined above.

423 C. Sea level rose rapidly between 11.5 to 7.0 ka (Stanford et al., 2011; Reynolds and
424 Simms, 2015). This change in base level provided abundant accommodation space for deposition
425 of coastal sediment on what is now the inner continental shelf, effectively shutting down
426 sediment supply to the progressively deeper CHS. The upper part of the Younger Dryas beach-
427 shoreface complex was eroded and reworked by wave energy in the earliest part of this time-
428 period, forming the erosional surface on seismic reflection profiles that represents the contact
429 between Units 1 and 2 (Figs. 5, 6, 7). The sediment-starved, wave-erosion surface and
430 underlying uppermost part of Unit 1 were then colonized by shallow-water fauna, providing the
431 material that yielded radiocarbon dates slightly younger than the Younger Dryas depositional age

432 of unit 1 (Table 1). The wave-erosion surface converges with the underlying transgressive
433 surface of erosion both seaward and landward of the CHS, so that the Unit 1 shoreface was
434 buried by younger sediment and preserved as an elongate lens.

435 D. From about 7.0 ka to the present, the rate of sea-level rise dramatically decreased to
436 about 1 m/1,000 years (Reynolds and Simms, 2015). The effects of this decrease included: (1) a
437 substantial decrease in the rate of creation of new sediment accommodation space, and (2) an
438 associated increase in the time available for offshore shelf-sediment transport. The net result was
439 movement of coastal sediment to mid-shelf environments and deposition of Unit 2 mid-shelf
440 deposits as a low-angle clinoform above the relict CHS shoreface (Medri et al., 2023).

441

442

443 **Hosgri Fault Slip Rate and Earthquake History**

444 Determining the slip-rate of a strike-slip fault requires documentation of the amount of
445 lateral slip of a piercing point, and the time-period over which that slip occurred. Johnson et al.
446 (2014) showed that the lower slope break of the CHS was a mappable, linear geomorphic feature
447 that crossed the Hosgri Fault, forming a unique piercing point (Fig. 2 inset). Understanding the
448 development and history of the CHS lower slope break (Fig. 10) is thus essential for developing
449 and (or) revising a Hosgri Fault slip rate.

450 Previously, Johnson et al. (2014) inferred that the CHS was underlain by a relict Younger
451 Dryas (latest Pleistocene) shoreface deposit (i.e., Unit 1 of this study), and noted lateral offset of
452 both the lower and upper slope breaks. The offset was determined by plotting data points and
453 regression lines obtained from normal-to-slope profiles located 12.5 m apart (Fig. 2; Johnson et
454 al., 2014, their Figs. 8, 9, 10, 11). They restricted the slip-rate analysis to a set of 34 (of 93)

455 points located on slope profiles from 225 m west of the fault trace to 200 m to the east of the
456 fault, to limit the influence of slope curvature and (or) the depositional or erosional irregularities
457 that occur along the lower slope break farther from the fault. They considered the lower slope
458 break (the inferred base of the paleoshoreface) as a far more reliable piercing point because, as
459 sea level rose in the latest Pleistocene, it would sink earlier below wave base and thereby would
460 have experienced less reworking by storm waves and subsequent sedimentation. They measured
461 30.3 ± 9.4 m of lateral offset of the lower slope break based on the slope profiles, an amount
462 similar to their estimate (34 m) based on a simpler analysis of digital elevation models and slope
463 maps derived from high-resolution bathymetric surveys (Johnson et al., 2014, their Fig. 8).

464 The new data presented in this study show that the Johnson et al. (2014) analysis
465 significantly underestimated the amount of erosion of the latest Pleistocene Unit 1 shoreface
466 deposits and were not aware of the thickness and sedimentology of the overlying Unit 2 middle
467 to late Holocene mid-shelf deposits (Medri et al., 2023). This new information is incorporated in
468 the depositional model of Figure 10, providing essential context for understanding the origin and
469 history of the lower slope break and its viability as a piercing point.

470 Together, the CHIRP (Figs. 5,6,7) and core data (Fig. 3) indicate that the lower slope
471 break represents the base of the Unit 1 shoreface (Fig. 10B, C), covered by downslope-thinning
472 Unit 2 shelf deposits (Fig. 10D). In the 425-m-wide local area Johnson et al. (2014) used to
473 measure fault offset, Unit 2 thickness decreases from about 108 to 124 cm thick along the HF-
474 25b profile (Fig. 2), to about 81 to 88 cm in cores HF-1 and HF-2 about 20 m farther downslope
475 (Figs. 2, 3). Unit 2 is about 50 to 60 cm thick at the lower slope break, and finally to about 30
476 cm, 80 m seaward of the CHS lower slope break in core HF-7 (Medri et al. 2023). Core and
477 CHIRP data indicate that at the lower slope break, the 50-60 cm of Unit 2 consists of about half

478 lower subunit (wave supported gravity flow deposits) and about half suspension deposits of the
479 upper subunit. In addition, no significant change in Unit 2 thickness is observed across the fault
480 zone on CHIRP profile HFC-25b (Fig. 7B), which is the closest and approximately parallel
481 profile to the lower slope break points used by Johnson et al., 2014 (Fig. 2).

482 Given the density of CHIRP profile lines and the number and location of cores collected,
483 we cannot quantify the amount (if any; see Fig. 7B) of variation in Unit 2 thickness across the
484 local CHS area that was used to determine lateral fault slip (Fig. 2). For our slip-rate analysis, we
485 therefore assume the Unit 2 thicknesses outlined above is uniform in this local, fault-adjacent
486 area. Thus, simple uniform removal of the thin Unit 2 cover does not change the locations of the
487 lower slope break relative to one another on bathymetric slope profiles, so the amount of
488 estimated lateral offset of this linear feature determined by Johnson et al. (2014) remains valid.
489 The thin Unit 2 drape (Fig. 10D) may slightly smooth the lower slope break (Fig. 10D), but it
490 does not compromise its value as a piercing point.

491 Given issues with small ($\sim 1^\circ$) slope changes, projections, and rounding errors, Johnson
492 et al. (2014) estimated that uncertainties in locating the lower slope break based on slope profiles
493 could be as much as 10 m for some data points, but that the effects of such errors were
494 minimized by analyzing large numbers of slope profiles at tight profile spacing. For our
495 assessment, it seems possible that undetected variations in Unit 2 thickness could lead to greater
496 uncertainty in locating the minimally buried base of the latest Pleistocene shoreface, but that
497 increase cannot be quantified with the current data.

498 The OSL and radiocarbon dates reported herein strongly support a Younger Dryas age,
499 $\sim 12,800$ to $11,500$ yr. B.P., for the latest Pleistocene shoreface (Unit 1 above), with the end of

500 shoreface deposition resulting from rapid sea-level rise at the onset of Meltwater Pulse 1B (e.g.,
501 Alley, 2000; Stanford et al., 2013; Liu and Milliman, 2004). There are variable estimates of sea-
502 level rise following the Younger Dryas. Stanford and others (2011) described Meltwater Pulse
503 1B as "robustly expressed," a multi-millennial interval of enhanced rates between 11,500 and
504 8,800 calendar years ago with peak rates of rise of up to 25 m/1,000 yrs. Using 11,500 yr B.P. as
505 the age of the CHS lower slope break and 30.3 ± 9.4 (two standard deviations, 95% confidence
506 limit) as the amount of offset yields a slip rate of 2.6 ± 0.8 mm/yr, essentially the same as the
507 previously suggested rate of 2.6 ± 0.9 mm/yr (Johnson et al., 2014). The previous rate was
508 determined using a Monte Carlo simulation in which the age of the lower slope break was
509 assigned a range of $12,000 \pm 500$ years, so that possible slip ranged from 20.9 m in 12,500 years
510 to 39.7 m in 11,500 years. The wealth of new geochronologic data obtained for this study
511 (Tables 2, 3) reduces age uncertainty and makes the simpler and more direct statistical
512 methodology approach we use in this study feasible.

513 CHIRP profiles clearly indicate Holocene (since at least ~7 ka) activity of the Hosgri
514 fault. Reflection-defined surfaces interpreted as the latest Pleistocene transgressive surface of
515 erosion and an early Holocene wave-erosion surface (Unit 1-Unit 2 contact) are vertically offset
516 on all fault crossings (Figs. 5, 6, 7). The amount of vertical offset is consistently larger for the
517 older surfaces, although the amounts of vertical offset on each surface varies from profile to
518 profile. Maximum offsets for the transgressive surface and wave-erosion surface are 125 cm and
519 73 cm.

520

521 **Advances in Geophysical Imaging and Geochronology**

522 Conducting paleoseismology in the marine environment is difficult, with a significant
523 portion of studies relying on secondary evidence of fault activity (e.g., turbidites; Goldfinger,
524 2011) or the correlation of offset sedimentary features to approximated sea-level curves to
525 provide timing constraints (e.g., Johnson et al., 2014, Nishenko et al., 2018). However, ultra-
526 high-resolution sub-bottom geophysical imaging, combined with precise age control from dated
527 sediment cores, offers the opportunity to provide primary, near-field evidence of time-
528 constrained coseismic fault offset (e.g., Brothers et al., 2011;2015, Watt et al., 2016).

529 Although Johnson et al. (2014) used single-channel sparker data to image and interpret
530 the CHS, CHIRP results from this study show that the sparker data lacks the vertical resolution
531 needed to make accurate interpretations of the post-LGM stratigraphy (Fig. 11A). Furthermore,
532 most offshore CHIRP studies utilize “envelope” imagery output by acquisition systems for the
533 final interpreted image. Such envelope imagery is composed of the original or “real” signal
534 merged with an “imaginary” signal, a 90° phase-shifted version (Hilbert transform) of the real
535 signal (Henkart, 2006). This envelope combination results in only positive amplitudes, loss of
536 polarity information, and a lower frequency/resolution image than is possible with more
537 advanced processing (Baradello, 2014). In this study we observed that when processing the
538 “real” or full-waveform CHIRP data, extra care must be taken in calculating static (swell)
539 corrections and deconvolution, so that the high-frequency polarity-preserved reflections align
540 properly from trace to trace. Doing so preserves the lateral high-resolution content and layering,
541 which can be easily lost with less optimal processing. This is especially important for mapping
542 seismic stratigraphy and fault offsets in coarse-grained, marginally stratified, or extensively
543 bioturbated sediments, where signal penetration can be limited, and strata commonly show
544 minimal impedance contrasts. The shoreface and mid-shelf deposits of our study provide good

545 examples of difficult to image depositional systems where advanced processing can be especially
546 important.

547 When comparing the envelope to the full-waveform processed CHIRP data along profile
548 HFC-4, the difference in geological resolvability is stark (Fig. 11b, 11c). For example, on the
549 envelope processed version of HFC-4 the offset of the shoreface deposit (Unit 1) by the Hosgri
550 fault is nearly impossible to resolve (Fig. 11b), whereas this can be confidently interpreted, and
551 offsets measured on the full waveform data (Fig. 11c). Similarly, clearly resolving the
552 transgressive surface of erosion and internal progradational bedding of the shoreface deposit
553 (Unit 1) is difficult at best on the envelope-processed data (Fig. 11b). No evidence of offset, nor
554 internal shoreface deposit patterns are discernable on the sparker profile, which is also marred by
555 short period multiples (Fig. 11a) that are characteristic of this source (Kluesner et al., 2019).
556 These comparative subbottom profiles show that when attempting to image and study recent or
557 active faulting of Holocene sediments, special care must be taken in selecting the proper sound
558 source, as well as processing the data to the fullest potential.

559 Although it is common practice for terrestrial paleoseismic work to use the combination
560 of radiocarbon and OSL dating techniques (e.g., Gray et al., 2015; Bennett et al., 2018; DuRoss
561 et al., 2022), this approach is currently in the incipient stages for offshore paleoseismic studies,
562 with this study being the first known application. Other offshore sedimentological studies have
563 looked at sediment ages with both radiocarbon and OSL techniques (e.g., Alappat et al., 2010; Yi
564 et al., 2013; Yang et al., 2015), including a study of sediment transport down the Monterey
565 canyon system, offshore central California (Stevens et al., 2013). Stevens et al., 2013 found that
566 differences in OSL and radiocarbon ages indicated sediment transport timing varied from decadal
567 (canyon head)- to millennial (>1.1 km depth)-scale transport down the canyon system.

568 In contrast, our study involves a somewhat different environment that includes on a relict
569 shoreface deposit and issues of mid-shelf sediment transport. We find generally good agreement
570 between radiocarbon and OSL CAM model ages, with OSL generally producing slightly older
571 ages. For example, within Unit 1 the mean OSL age is ~ 1,000 yr older than the mean
572 radiocarbon age. As discussed earlier, this time gap may reflect the time between last quartz
573 grain sun exposure on land or shallow water and when the dated radiocarbon material was
574 deposited in deeper water. Another possibility is that the OSL was incompletely reset prior to
575 deposition with an inherited 1000 years of apparent age, although we did not see strong evidence
576 for partial bleaching. It is also worth noting that we assumed full saturated water content over the
577 OSL burial period. It is possible that the sampled sediment was drier during the subaqueous
578 episode prior to eustatic sea level rise. This unknown drier period would result in a slightly high
579 dose rate for that period and a slight overestimation of the true burial age. However, the
580 generally good correspondence between radiocarbon and OSL suggests that the assumptions
581 behind OSL dating are applicable in this environment. This study shows that combining the two
582 techniques can provide additional information to help resolve the depositional history and further
583 strengthen the geochronological model, both of which are crucial when conducting offshore
584 paleoseismic studies.

585

586

587 **CONCLUSIONS**

588 New data obtained from CHIRP seismic reflection profiles and shallow sediment cores
589 provide the basis for a reassessment of the slip rate of the Hosgri Fault offshore central
590 California. The Hosgri fault cuts the cross-Hosgri slope (CHS), a subtle seafloor lineament

591 previously interpreted as a latest Pleistocene (Younger Dryas, ~12.8 to 11.5 ka) shoreface
592 deposit. Offset of the lower slope break of this inferred shoreface provided the basis for inferring
593 a lateral slip rate of 2.6 ± 0.9 mm/yr. The CHS has a complex depositional history that comprises
594 two distinct stratigraphic units. The lower unit (Unit 1) overlies the post-LGM transgressive
595 surface of erosion and is interpreted as a shoreface deposit based on seismic facies (offshore-
596 dipping reflections), sediment texture (clean fine sand), sediment infauna, and significant
597 component (~8.4%) of heavy minerals. Radiocarbon and OSL dates from the top of Unit 1 are
598 consistent with deposition during the Younger Dryas. The shoreface was abandoned and partly
599 eroded during the subsequent pulse of rapid sea-level rise and transgression that ended about 7
600 ka. Unit 2 consists of fine to very fine sand and silt deposited in a mid-shelf environment
601 between ~ 7 ka and the present, when the rate of sea-level rise slowed dramatically. Deposition
602 primarily occurred as wave-supported gravity flows transported sandy sediment down the pre-
603 existing shoreface slope, and through deposition of muddy sediment from suspension (Medri et
604 al., 2023). Although Unit 2 provides a thin (~50 - 60 cm) cover over the lower slope break of the
605 shoreface, it does not compromise the value of this feature as a piercing point and the previously
606 proposed slip rate remains valid. This work benefitted significantly from full-waveform
607 processing of CHIRP data, resulting in significantly higher resolution in coarser grained strata
608 that are typically difficult to image; application of these techniques should have importance in
609 neotectonic studies of shelf and high-energy environments elsewhere. Our novel combination of
610 radiocarbon and OSL dating also provided important insights into depositional history and
611 strengthened the geochronological model, both of which can have similar value in other offshore
612 paleosiesmic investigations.

613

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UCIAM S #	Core	Depth (cm)	Material/species	D ¹⁴ C ‰	±	¹⁴ C age	±	Calibrated age (Cal years B.P.)
254929	HF-1	46	<i>Amphissa versicolor</i>	-186.7	1.3	1660	15	1180-917
254931	HF-1	61	<i>Neverita lewisi</i>	-519.68	1.0	5890	20	6264-5961*
254930	HF-1	80	<i>Amphissa versicolor</i>	-408.7	1.3	4220	20	4295-3955
254932	HF-1	90	<i>Clathurella canfieldi</i>	-705.8	0.6	9825	20	10774-10442
254936	HF-1	97	<i>Truncatella californica</i>	-734.9	0.6	10665	20	12027-11623
260440	HF-2	103.5	Gastropod	-730.1	0.7	10520	25	11808-11386
260441	HF-2	106	Gastropod	-712.3	0.7	10010	20	11085-10714
246214	HF-3	Cutter nose	<i>Callianax baetica</i>	-710.30	0.7	9950	20	11044-10640
230405	HF-3	VC cutter nose	Gastropod	-714.6	0.9	10075	25	11163-10795
243289	HF-3	137	<i>Amphissa versicolor</i>	-555.3	0.8	6510	15	6940-6634
230406	HF-3	131	Gastropod	-568.3	1.1	6750	25	7224-6901
243298	HF-3	110	Wood	-213.6	1.2	1930	15	1924-1794
243290	HF-3	80	Bivalve	-196.2	1.4	1755	15	1279-1025
243291	HF-3	40	<i>Amphissa versicolor</i>	-140.9	1.5	1220	15	721-515
260442	HF-4	154	<i>Turritella cooperi</i>	-683.2	0.7	9235	20	10035-9620
260443	HF-4	171.5	Gastropod	-679.2	0.7	9135	20	9857-9524
243299	HF-5	115	Wood	-265.7	1.2	2480	15	2710-2473*
243292	HF-5	122	<i>Amphissa versicolor</i>	-183.2	1.3	1625	15	1157-898
254938	HF-5	155	<i>Amphissa versicolor</i>	-267.4	1.2	2500	15	2125-1835
254933	HF-5	166	<i>Amphissa versicolor</i>	-243.7	1.2	2245	15	1808-1534
254935	HF-5	175	<i>Callianax baetica</i>	-670.7	0.6	8920	20	9532-9301*
254945	HF-5	180	Wood	-323.6	1.1	3140	15	3441-3272
254944	HF-5	199	Wood	-309.4	1.1	2975	15	3211-3076
254937	HF-5	222	<i>Amphissa versicolor</i>	-434.4	0.9	4575	15	4776-4434
246230	HF-5	233	Wood	-451.8	0.9	4830	15	5594-5484
243293	HF-5	255	<i>Callianax baetica</i>	-561.5	0.7	6625	15	7088-6749
243300	HF-5	287	Wood	-492.6	0.9	5450	15	6295-6207*
260444	HF-5	296	<i>Turritella cooperi</i>	-586.1	0.9	7085	20	7516-7257
254934	HF-7	20	<i>Amphissa versicolor</i>	8.2	1.6	Modern		
246207	HF-7	70	<i>Callianax baetica</i>	-737.6	0.7	10750	25	12195-11741
260448	HF-12	59	Bivalve	-798.1	0.6	12855	25	14787-14217
260447	HF-12	63	Bivalve	-852.7	0.6	15385	35	17967-17551
260446	HF-12	67	Shell fragment	-997.9	0.4	49700	1600	54911-48737
260445	HF-12	71	Gastropod	-993.8	0.4	40770	550	44014-42414

Lab Code	n ^a	CAM Over	CAM De ^c	CAM Age ^c
		Dispersion ^b	(Gray)	(ka)
HF-1(92-96cm)	15(29)	19%	11.8 ± 0.6	12.1 ± 0.7
HF-2(59-64cm)	14(28)	25%	2.67 ± 0.18	2.47 ± 0.18
HF-2(96-101cm)	14(35)	27%	13.1 ± 1.0	12.3 ± 1.0
HF-4(119-127cm)	10(15)	20%	7.75 ± 0.51	6.95 ± 0.48
HF-4(159-166cm)	10(15)	23%	14.7 ± 1.1	11.8 ± 0.9
HF-6(239-248cm)	6(10)	12%	8.88 ± 0.49	8.55 ± 0.50
HF-7(22-33cm)	18(25)	16%	11.7 ± 0.5	10.6 ± 0.5
HF-7(68-76cm)	19(30)	12%	14.9 ± 0.5	12.4 ± 0.4
HF-15(95-103cm)	26(32)	14%	21.9 ± 0.6	18.5 ± 0.6
HF-15(120-130cm)	5(20)	29%	119 ± 16	78.8 ± 10.9
HF-15(134-144cm)	11(16)	19%	111 ± 7	67.7 ± 4.2

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894 TABLE 2. Table of Optical Stimulated Luminescence ages derived from core samples. First
895 column denotes core and sample interval. The central age model (CAM) is used as this provides
896 the closest agreement with radiocarbon ages (Table 1). See supplementary materials for
897 additional details. ^aNumber of aliquots meeting acceptance criteria, parentheses indicate total
898 number of aliquots measured. ^bDefined as the statistical dispersion beyond what would be
899 expected for a perfectly bleached sample (Galbraith and Roberts, 2012). ^cDetermined using the
900 function calc CentralDose from the R-Luminescence package. Uncertainty is 2 σ .

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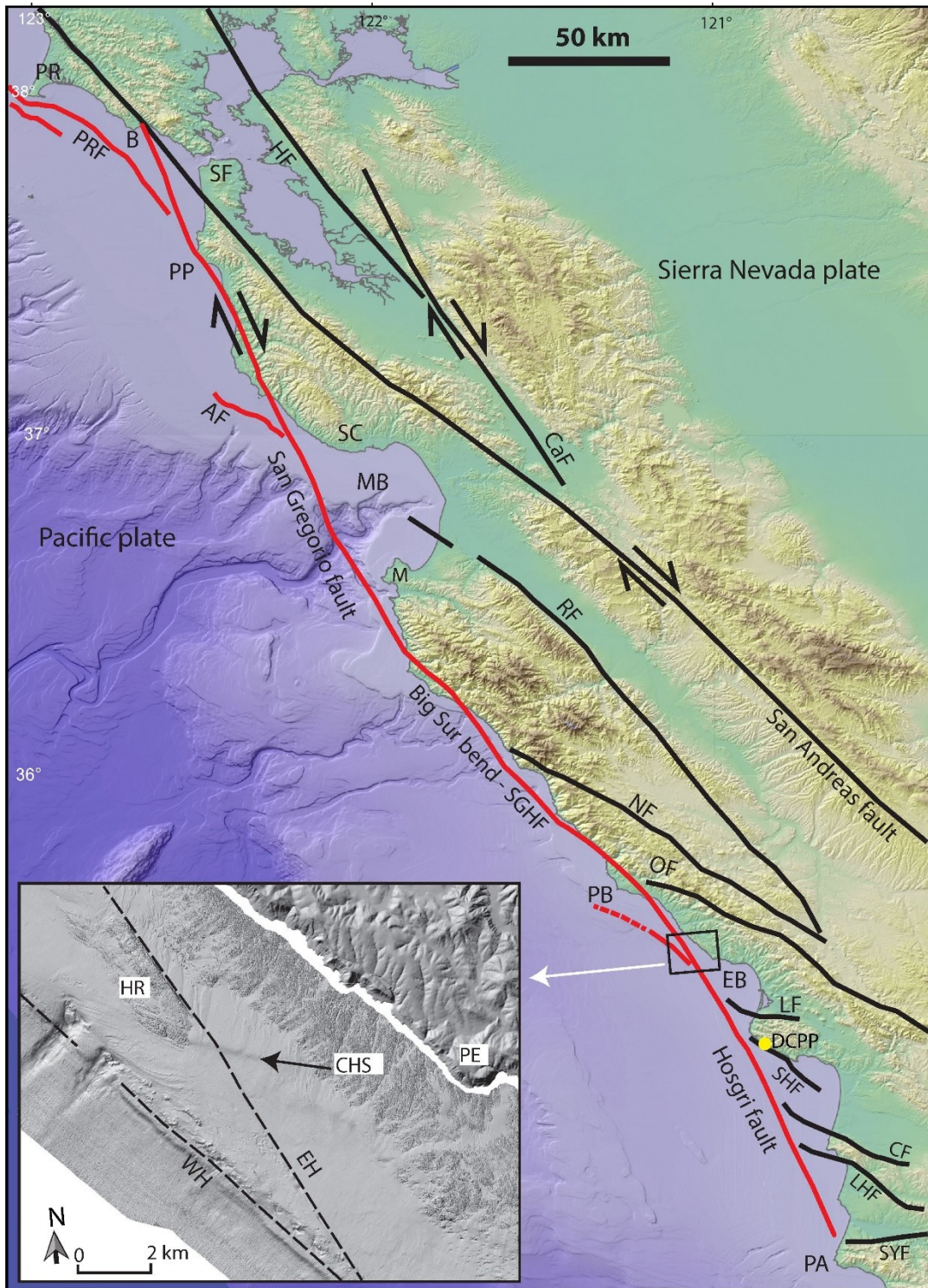
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	Unit 1 (n = 7)		Unit 2 (n = 7)	
	Mean	S.D.	Mean	S.D.
Quartz	20.0	2.0	15.2	4.2
Polycrystalline Quartz	5.5	1.1	4.3	1.9
Plagioclase Feldspar	15.1	1.9	11.8	2.6
Sedimentary Lithic & Chert	39.0	5.2	30.1	2.9
Volcanic Lithic	6.4	1.1	3.9	1.1
Carbonate bioclast	5.3	2.5	31.9	10.0
Heavy Mineral	8.4	1.8	2.2	1.3
Mica	0.2	0.5	0.4	0.5
QmpFLt	30,18,53		30,18,52	

905 TABLE 3. Composition of fine-grained sand sampled from CHS cores. S.D. is standard
906 deviation.

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910 FIGURE 1. Regional map showing the main fault structures that are part of the Pacific-North

911 American Plate Boundary along central California. The San Gregorio-Hosgri Fault (SGHF)

912 system is predominantly located near the coastline and highlighted in red. Inset map shows the
913 focus region of this study where the Cross Hosgri Slope (CHS) is located. AF = Ascension Fault;
914 B = Bolinas; CaF= Calaveras Fault; CF = Casmalia Fault; DCP = Diablo Canyon Power Plant;
915 EB = Estero Bay; EH = Eastern strand of Hosgri Fault; HF = Hayward Fault; HR = Hosgri ridge;
916 LF = Los Osos Fault; LHF = Lions Head Fault; M = Monterey; MB = Monterey Bay; NF =
917 Nacimiento Fault; OF = Oceanic Fault; PA = Point Arguello; PB = Piedras Blancas; PE = Point
918 Estero; PP = Pillar Point; PR = Point Reyes; PRF = Point Reyes Fault; RF = Rinconada Fault;
919 SC = Santa Cruz; SF = San Francisco; SHF = Shoreline Fault; SYF = Santa Ynez Fault; WH =
920 Western strand of Hosgri Fault.

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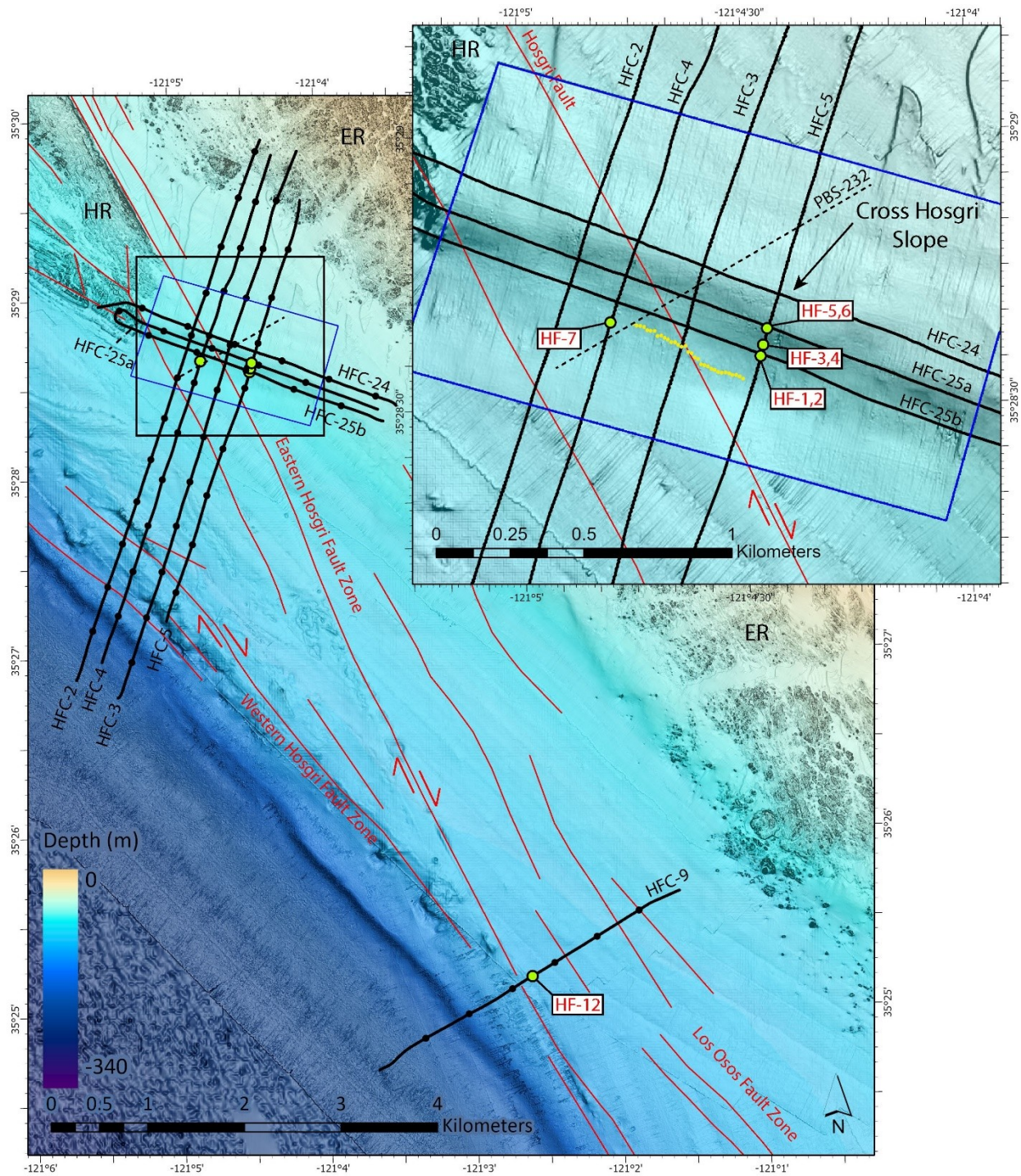
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937 FIGURE 2. Map showing high resolution bathymetry and locations of CHIRP tracklines (black)

938 and sediment cores (green) in the study region. Inset shows details of CHIRP tracklines and core

939 locations along the Cross Hosgri Slope (CHS). Inset location outlined with black rectangle. Red
940 lines denote fault locations from the USGS Quaternary Fault and Fold database (Walton et al.,
941 2020). Black circles along the CHIRP tracklines denote every 500 shots. Blue polygon outlines
942 USGS collected Reson 7111 multibeam bathymetry (Hartwell et al., 2013) and yellow points
943 denote lower slope break points used for slip rate analysis in Johnson et al. (2014). Additional
944 bathymetry source includes data from the California Seafloor Mapping Program (Johnson et al.,
945 2017). Dashed black line shows location of sparker profile used in Johnson et al., 2014 and
946 shown in Figure 11. ER = Estero Rocks, HR = Hosgri Ridge.

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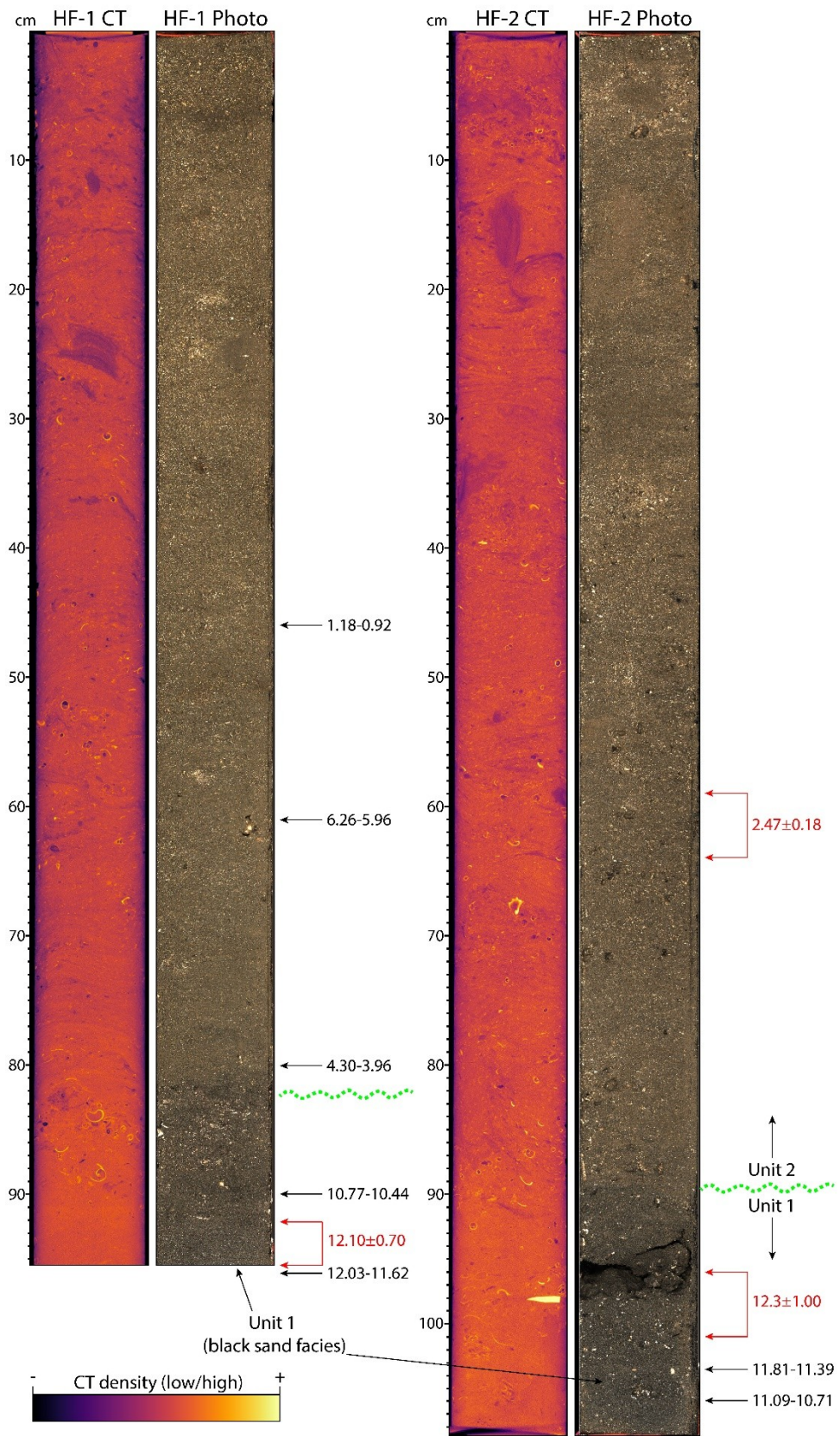
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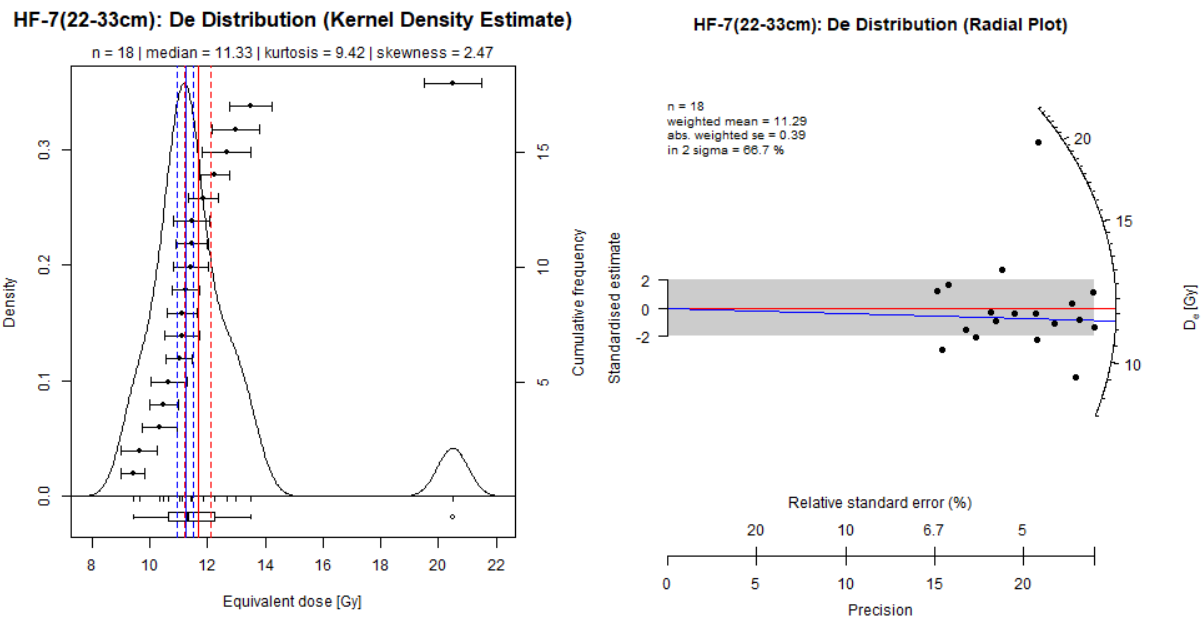
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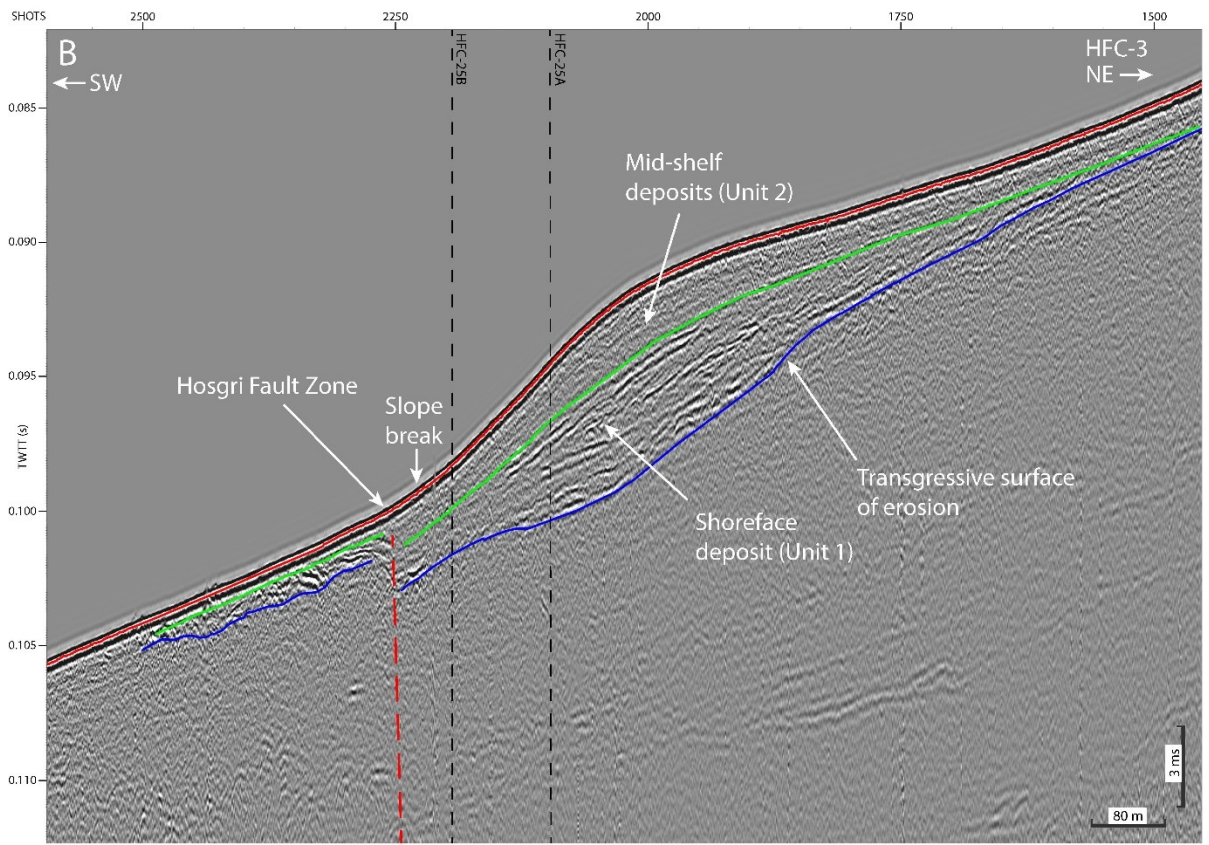
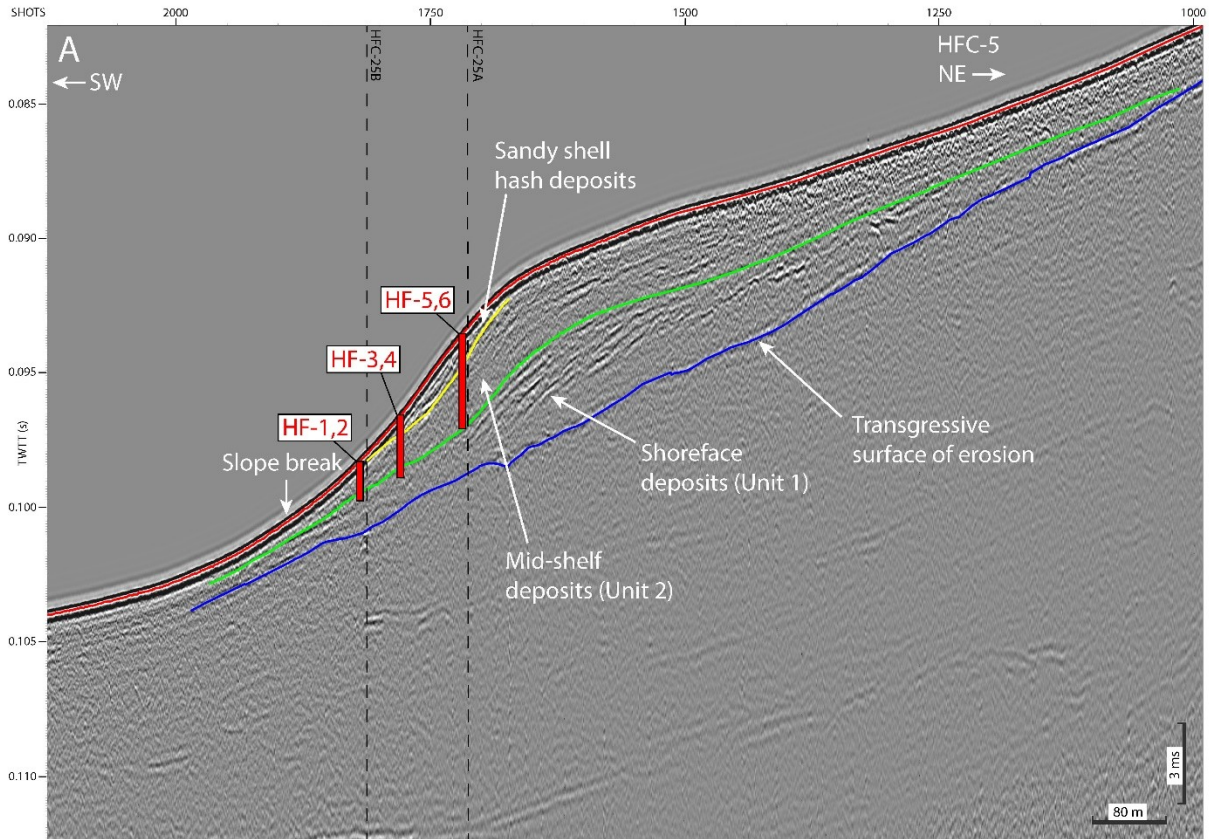
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963 FIGURE 3. CT scan and photos of sediment cores HF-1 and HF-2. Radiocarbon ages are noted
964 in black, whereas OSL ages are provided in red. Wavy dashed green line denotes the transition
965 zone from Unit 1 below to Unit 2 above. Ages are shown in ka.
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969 FIGURE 4. Plots of Equivalent Dose for HF-7(22-33cm). Red lines show CAM (Central Age
970 Model) equivalent dose, blue lines show MAM (Minimum Age Model) equivalent dose, solid
971 line shows the age, and dashed lines show uncertainty. We analyzed HF-7 (22-33cm) using
972 Quartz OSL on 125-150 μm size grains. Out of 25 aliquots, 18 produced acceptable
973 luminescence characteristics and passed acceptance criteria (Supplementary Material). Aliquots
974 that passed acceptance criteria are plotted as a kernel density estimate (left) and as a radial plot
975 (right).
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980 FIGURE 5. CHIRP profiles across the CHS. (A) CHIRP profile HFC-5 that crosses core sites HF-
981 1 through HF-6. (B) CHIRP profile HFC-3 showing offset transgressive surface of erosion on
982 Unit 1 near the toe of the CHS. Blue horizon denotes the transgressive surface of erosion, the
983 green horizon traces the top of paleo shoreface deposits (Unit 1), the yellow horizon traces the
984 bottom of the sandy shell hash deposits, and the seafloor is delineated in red. Core locations are
985 shown in red, and the Hosgri Fault Zone is marked with a dashed red line on panel B. Vertical
986 dashed black lines show location of crossing CHIRP profiles HFC-25a and HFC-25b

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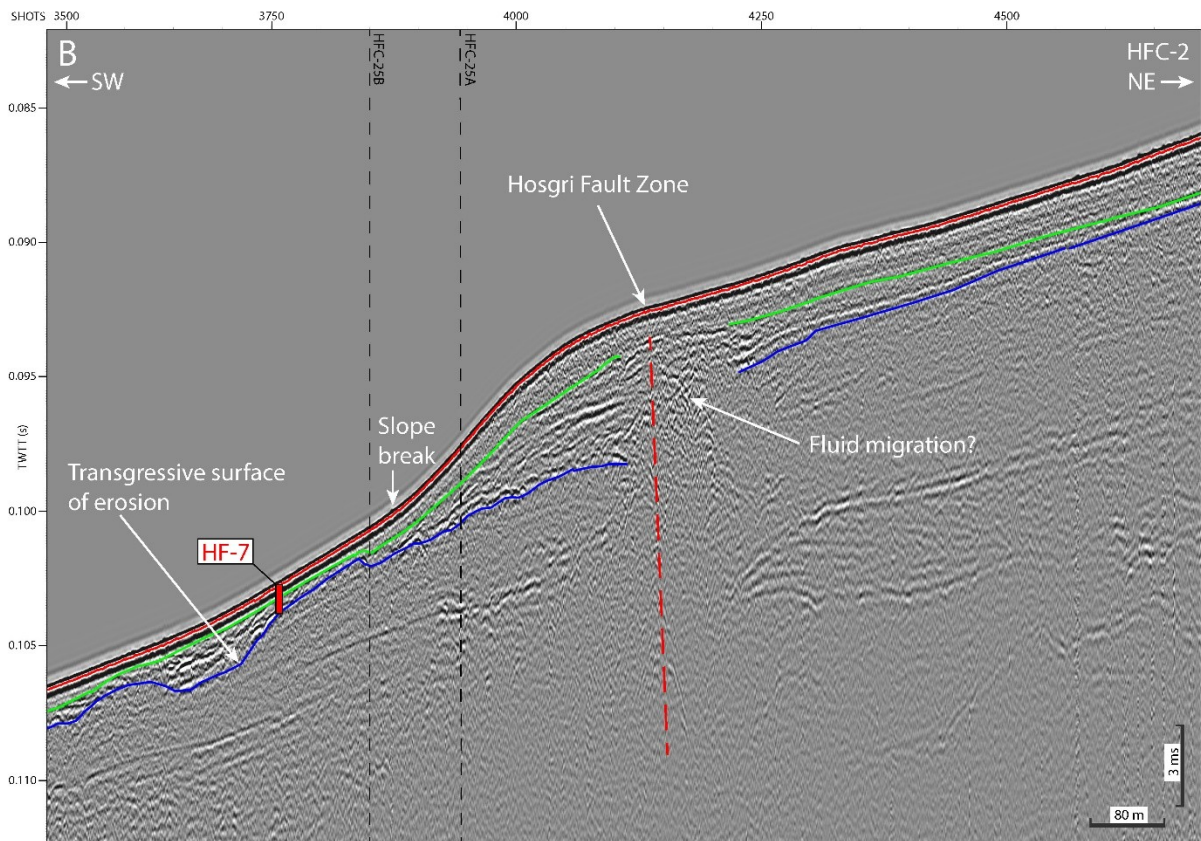
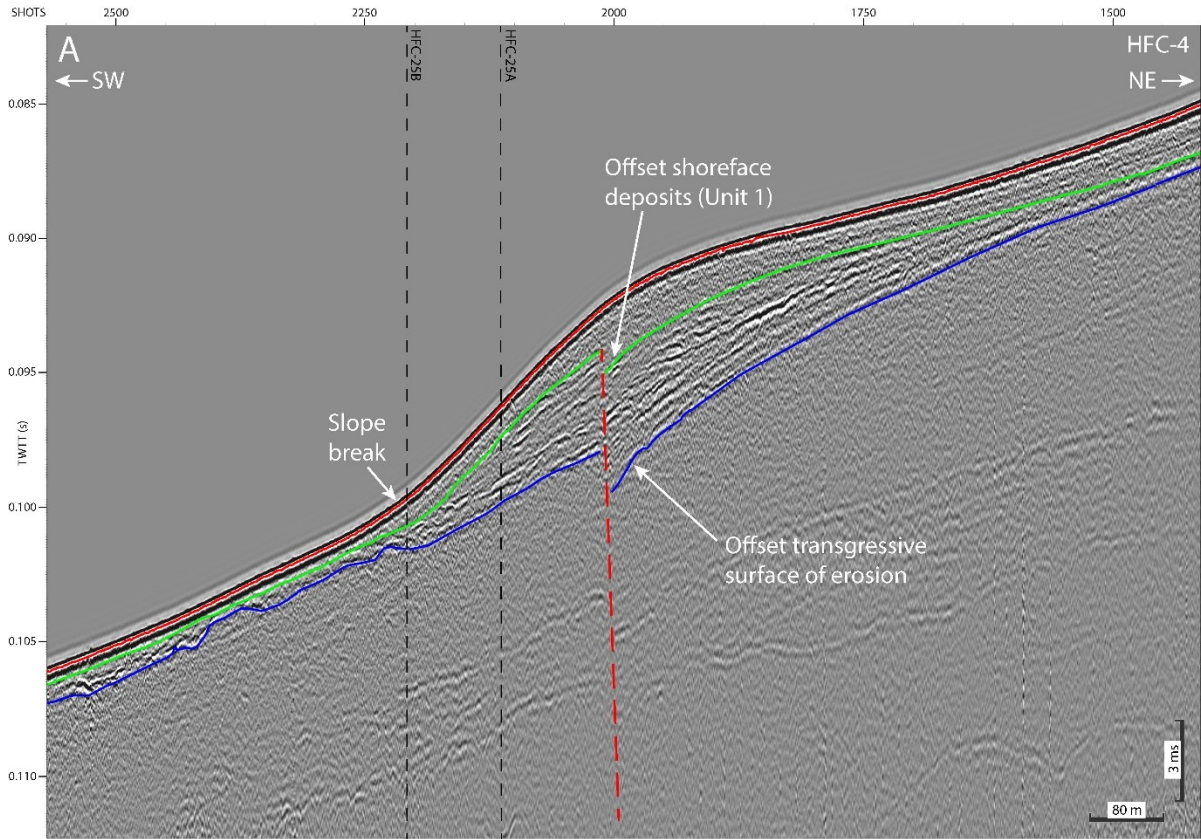
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1004 FIGURE 6. CHIRP profiles across the CHS. (A) CHIRP profile HFC-4 that shows clear offset of
1005 the transgressive surface of erosion and paleo shoreface deposits. (B) CHIRP profile HFC-2 that
1006 crosses the Hosgri Fault along the upper portion of the CHS. Blue horizon denotes the
1007 transgressive surface of erosion, the green horizon traces the top of paleo shoreface deposits
1008 (Unit 1). HF-7 core location is shown in red, and the Hosgri Fault is marked with a dashed red
1009 line. Vertical dashed black lines show location of crossing CHIRP profiles HFC-25a and HFC-
1010 25b.

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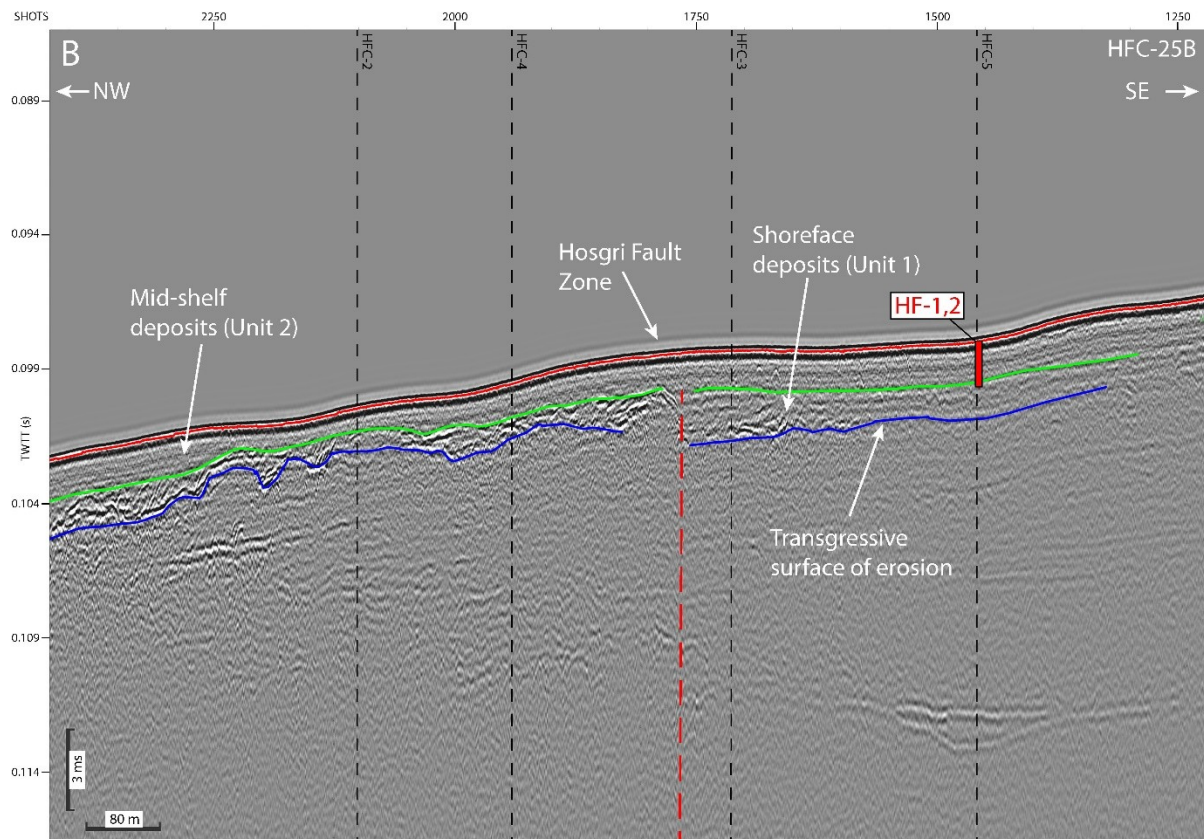
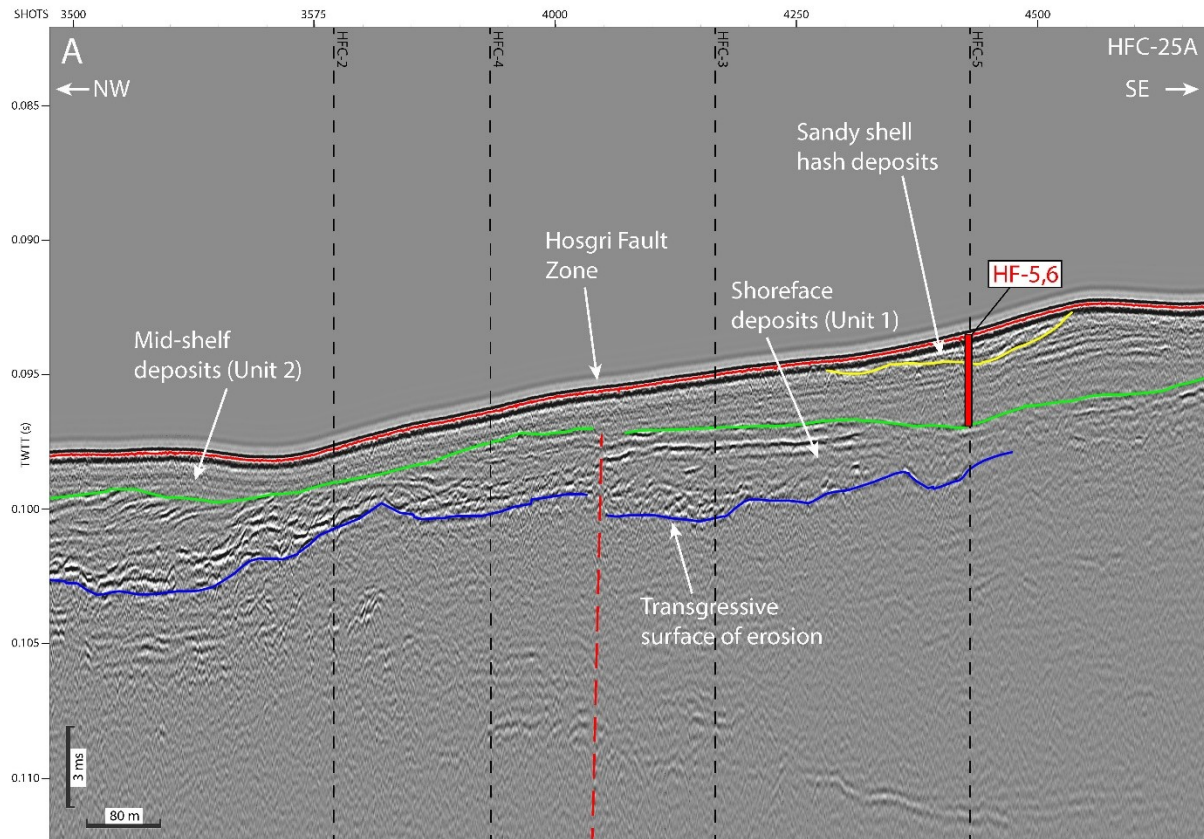
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1028 FIGURE 7. Along slope CHIRP profiles across the CHS. (A) CHIRP profile HFC-25a that
1029 transects the middle portion of the CHS. (B) CHIRP profile HFC-25b that transects the lower
1030 portion of the CHS. Blue horizon denotes the transgressive surface of erosion, the green horizon
1031 traces the top of paleo shoreface deposits (Unit 1), and the yellow horizon traces the base of the
1032 sandy shell hash deposits. Core locations are shown in red, and the Hosgri Fault is marked with a
1033 dashed red line. Vertical dashed black lines show location of crossing CHIRP profiles HFC-2,
1034 HFC-3, HFC-4, and HFC-5.

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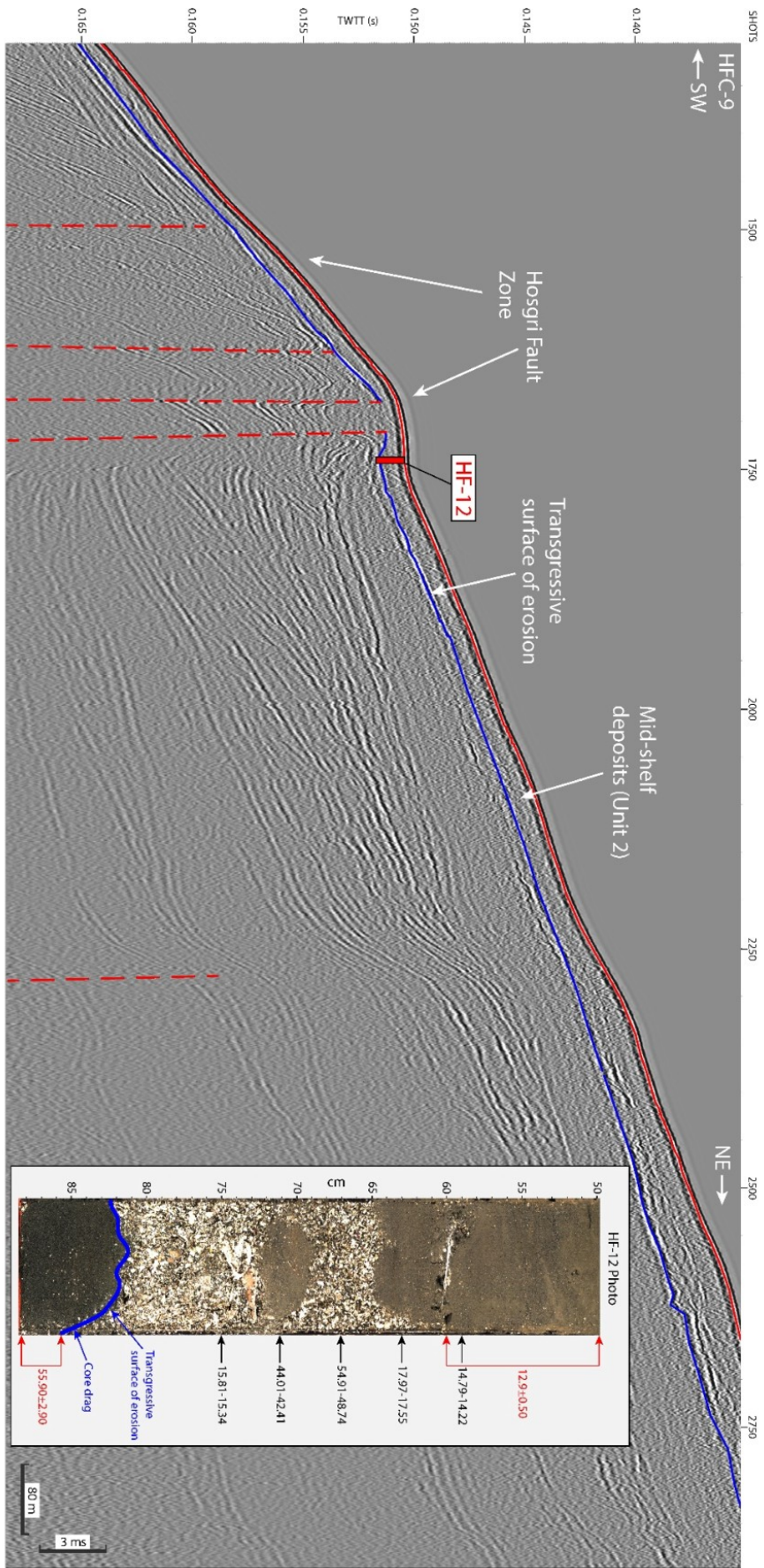
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1052 FIGURE 8. CHIRP profile HFC-9 located south the CHS. Profile crosses the mid-shelf region
1053 and images the extensive transgressive surface of erosion unconformity also seen below the CHS.
1054 Sediment core HF-12 is located along the profile on the flank of the Hosgri Fault Zone. Inset
1055 shows the bottom portion of HF-12 and the associated radiocarbon and OSL ages. Note the jump
1056 in ages above and below the blue horizon. Radiocarbon ages are noted in black, whereas OSL
1057 ages are provided in red.

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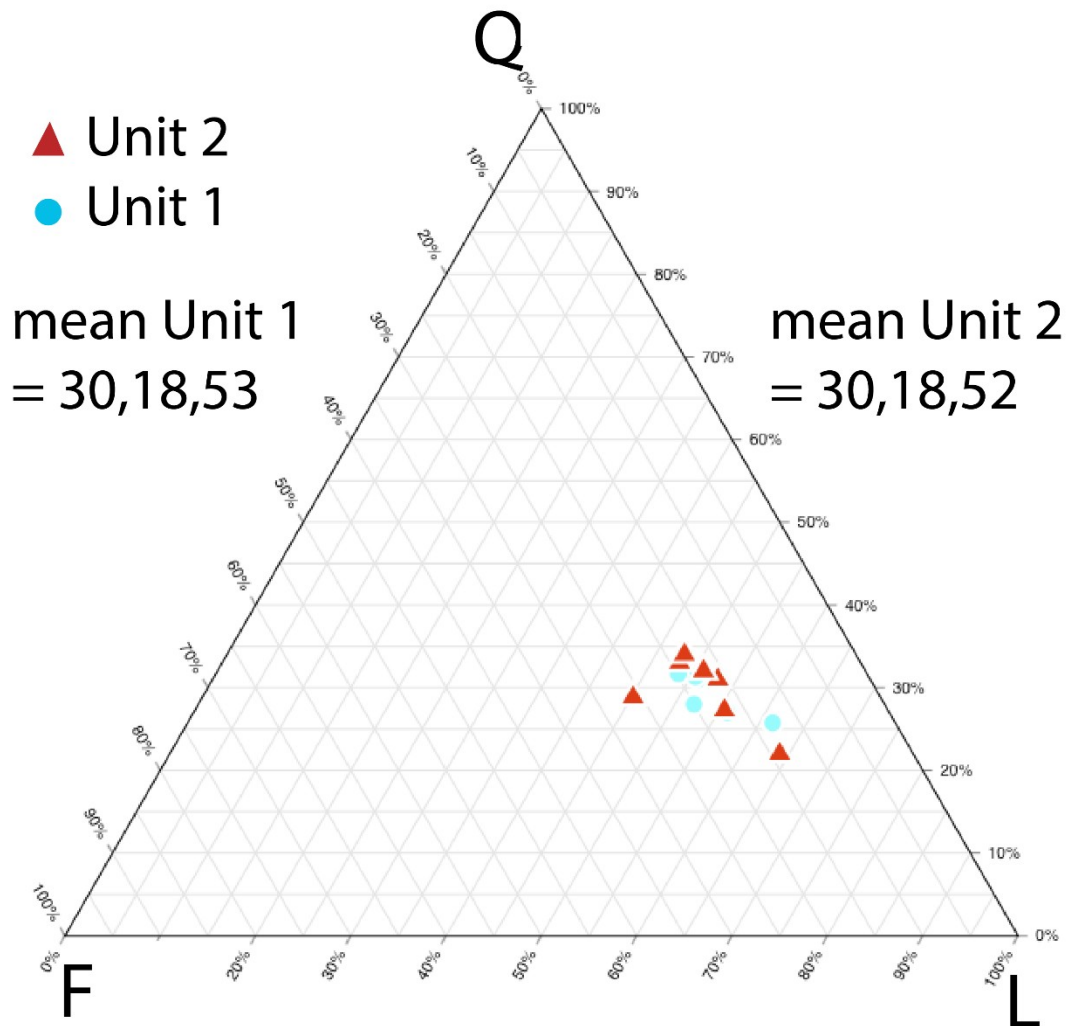
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1077 FIGURE 9. Ternary plot showing CHS sand composition. Q = monocrySTALLine and

1078 polycrySTALLine quartz; L = siliciclastic lithic fragments, chert, and volcanic lithic fragments.

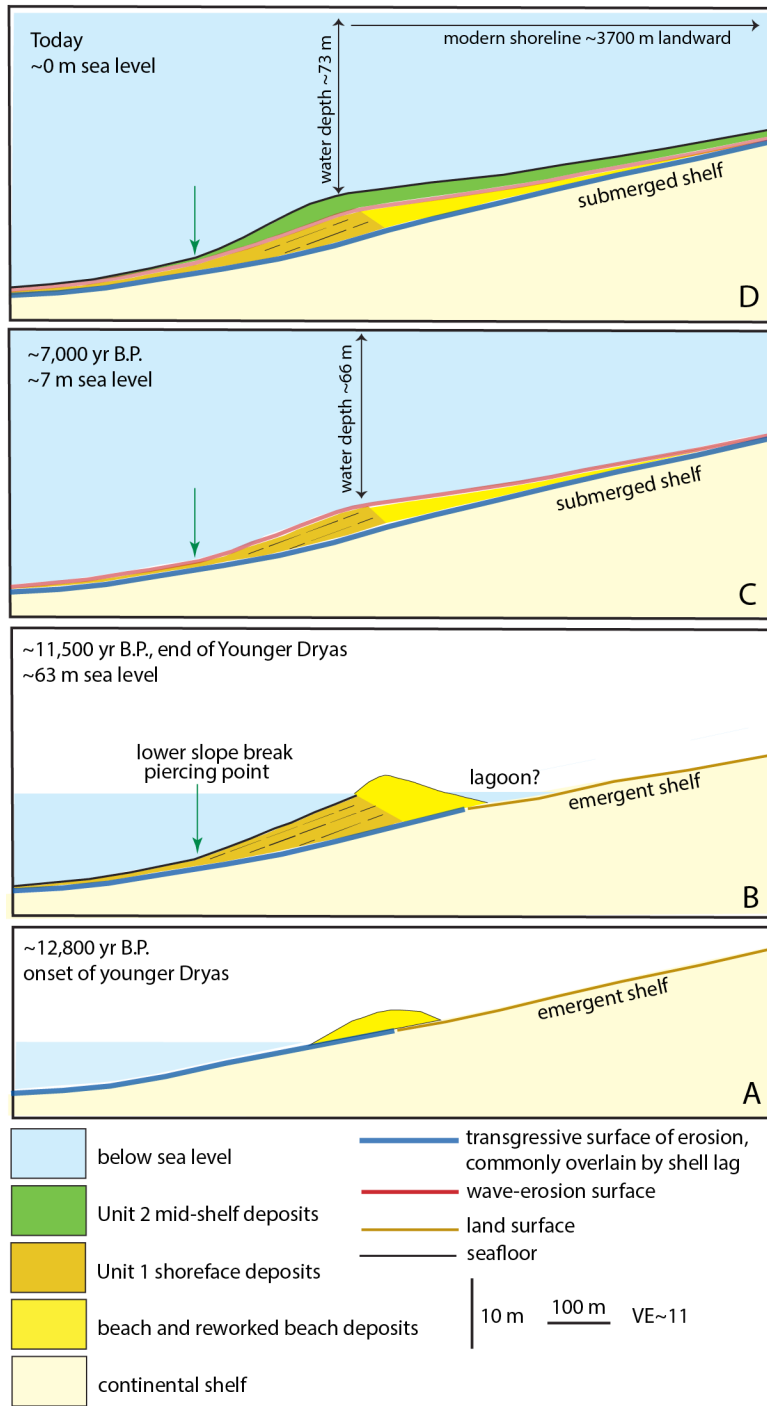
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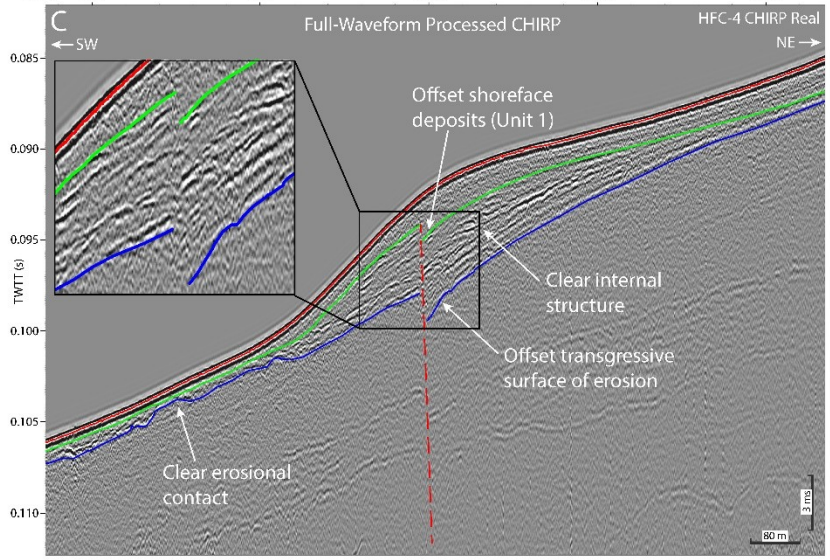
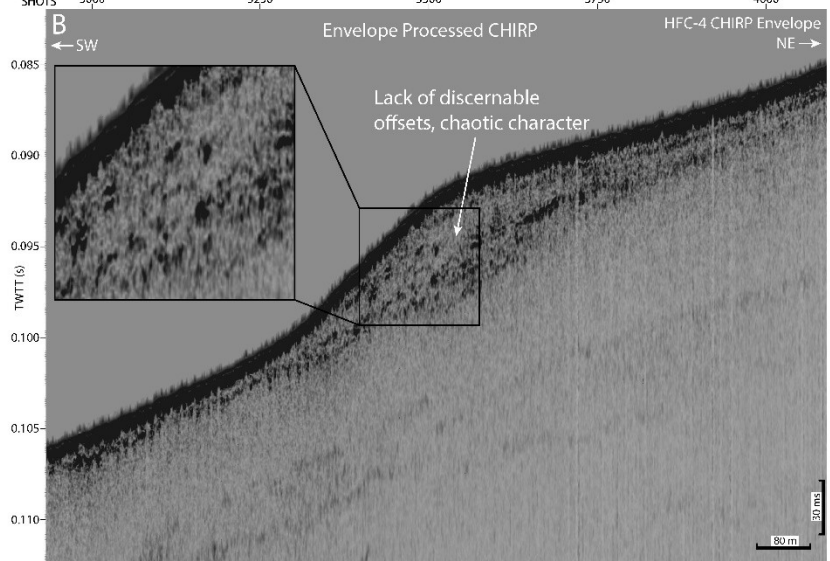
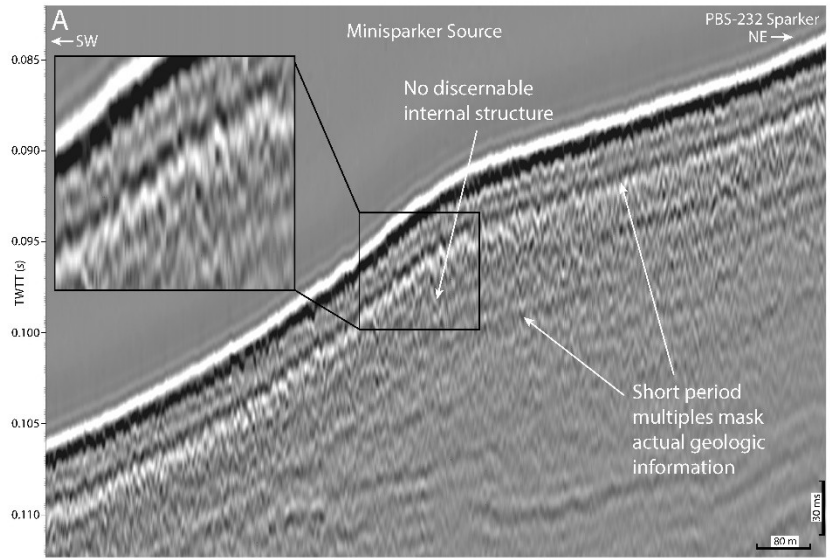


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1085 FIGURE 10. CHS depositional model. See text for discussion.

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1090 FIGURE 11. Three panels showing the differences of resolution and geologic resolvability for
1091 sparker and chirp data collected across the CHS. (A) Sparker profile PBS-232 collected across
1092 the CHS and shown in Johnson et al., 2014. (B) CHIRP profile HFC-4 shown in envelope with 1-
1093 pass swell correction. (C) CHIRP profile HFC-4 processed using the full waveform data and
1094 multiple swell correction steps. Note the jump in resolution from sparker to envelope CHIRP,
1095 and further improvement with full waveform (real) CHIRP data. Insets show 200% zoomed in
1096 portions of the profile with locations delineated by the black rectangles. See Figure 2 for
1097 locations of profiles.

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