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IN ASYMPTOTIC LIGHT FIELDS

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Simple Formulas for the Volume Absorption Coefficient in Asymptotic Light Fields

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INTRODUCTION

It is now a well-established fact that the properties of the light field in optically deep homogeneous stratified media (such as deep cloud layers, oceans, lakes, etc) become extremely regular and predictable at great depths. The conceptual and practical consequences of this fact, however, are only in the first stages of exploration. In this note we derive some further 2,3
3,3
3,3 consequences from this regularity. In particular, we use the established regularity of the light field to derive several simple, exact, formulas relating the volume absorption coefficient to the common limiting value $\oint_{\mathcal{C}_{\infty}}$ of the \mathcal{K} -functions for irradiance. In this way we supplement the exact formula:

$$
a(z) = \int \frac{1}{h(z)} \frac{d\overrightarrow{H}(z, +)}{dz} \qquad (1)
$$

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irradiance. In this way we supplement the exact formula: $\mathcal{L}_{\mathcal{A}}$ for $\mathcal{L}_{\mathcal{A}}$ for $\mathcal{L}_{\mathcal{A}}$ for $\mathcal{L}_{\mathcal{A}}$

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for the volume absorption function with several alternate formulas which are especially suitable for engineering calculations and as handy rules of thumb relating α and k_{∞} . These formulas are as follows:

$$
\begin{array}{lll}\n\text{(I)} & \alpha = \mathcal{A}_{\infty} & \frac{\mathsf{H}(\vec{z},-)}{h(\vec{z})} \quad , \quad \vec{z} \geq \vec{z}_{\text{o}} \; , \\
\end{array}
$$

(II)
$$
\alpha = k_{\infty} \frac{(1 - R_{\infty})}{D(-) + R_{\infty} D(4)},
$$

$$
\text{(III)} \qquad \alpha = \frac{3}{4} k_{\infty} ,
$$

where:

 $\oint_{\bm{a}}$

is the common limit, as
$$
z \rightarrow \infty
$$
, of the \mathcal{K} -functions $\mathcal{K}(\mathcal{Z}, \pm)$, $\mathcal{R}(\mathcal{Z})$ for irradiance and scalar irradiance, respectively.

 $-2-$

 $D(\pm)$ are the limits, as $\lambda \rightarrow \infty$, of the distribution functions $D(\mathcal{Z}, \perp)$.

$$
R_{\infty} \quad \text{is the limit, as } \mathbb{Z} \to \infty \quad \text{, of the reflex-}
$$

tance function
$$
R(\mathbb{Z}, -) = H(\mathbb{Z}, +) / H(\mathbb{Z}, -).
$$

$$
H(\mathcal{Z}, -) = H(\mathcal{Z}, -) - H(\mathcal{Z}, +)
$$
, the net downward
irradiance at any depth $\mathcal{Z} \approx \mathcal{Z}_o$, where
 \mathcal{Z}_o is the depth below which the light
field has essentially attained its
asymptotic structure.

The details of the derivation of (1) , (II) , and (III) will now be given.

DERIVATION OF (I)

Short Derivation

The short derivation starts with (1) , and the fact that there exists a depth Z_0 below which the logarithmic derivatives of $H(z,-)$, $H(z,+)$, and $h(z)$ are constant and equal to a common va'ue \mathcal{L}_{∞} (reference 3, equations (22), (24), and (35)). Therefore

 $\ddot{}$

$$
\frac{dH(Z, +)}{dZ} = \frac{dH(Z, +)}{dZ} - \frac{dH(Z, -)}{dZ}
$$
\n
$$
= -k_{\omega}H(Z, +) + k_{\omega}H(Z, -) = k_{\omega}H(Z, -),
$$
\n(2)

for all $\bar{z} = \bar{z}_0$. Hence:

$$
\alpha = \mathcal{R}_{\omega} \frac{\overline{H}(\overline{z},-)}{h(\overline{z})} \tag{I}
$$

Long Derivation

The long derivation is essentially an exercise in the use of the integrated form of the divergence relation for the light field vector: 5

> $\bar{P}(s,-) = \alpha v \hat{U}(M)$, (3)

 $\bar{\chi}$

where M is <u>any</u> region of the optical medium, S is its boundary, and $P(S) - I$ is the net inward flux across S into M . $U(M)$ is the radiant energy content of M , $\mathcal V$ is the speed of light in M , and α is the required value of the volume absorption coefficient.

It is interesting to observe that (3) yields a value of α in any homogeneous medium, regardless of the structure of the light field:

$$
\alpha = \frac{\overline{P}(S, -)}{\nu \, \nu(M)} \, . \tag{4}
$$

The numerator of (4) can be obtained by traversing the boundary of M with flat plate collectors. The denominator is obtained by probing the interior of M with a spherical collector (to find $h(\boldsymbol{\mu})$ at each point $\boldsymbol{\mu}$) and integrating the values over the volume of \mathcal{M} .

In the present case, the extreme regularity of the asymptotic light field allows one to estimate $\bigcup (:\mathbb{M})$ knowing only one value of the scalar irradiance at a boundary point of \mathcal{M} . This fact holds also for $P(5,-)$. Specifically, consider a region \mathcal{N} in the form of a vertical column of unit cross section, and bounded by two parallel planes at depth \mathcal{Z}_1 , and \mathcal{Z}_2 , such that $\mathcal{Z}_0 \leq \mathcal{Z}_1 \leq \mathcal{Z}_2$. SIO Ref. 58 - 79 - 6 -

The medium is homogeneous and stratified; hence

$$
\overline{P}(S,-) = \overline{H}(\overline{z}_1,-) + \overline{H}(\overline{z}_2,+).
$$
 (5)

By hypothesis, we have:

$$
H(Z_{z_1} \pm) = H(z_{i_1} \pm) e^{-\oint_{\omega} (z_2 - z_i)}, \qquad (6)
$$

so that

$$
\overline{P}(S,-) = \overline{H}(\overline{z}_{1,-}) \left[1 - e^{-\frac{\hat{R}_{\omega}(z_{2}-z_{1})}{\hat{Z}_{1}}}\right]. \qquad (7)
$$

 $\bar{\beta}$

Furthermore,

 \cdot

$$
U(M) = \int_{\tilde{z}_1}^{\tilde{z}_2} h(\tilde{z}) d\tilde{z}
$$

= $h(\tilde{z}_1) \int_{\tilde{z}_1}^{\tilde{z}_2} e^{-\hat{k}_{\omega}(\tilde{z}-\tilde{z}_1)} d\tilde{z}$
= $\frac{h(\tilde{z}_1)}{\hat{k}_{\omega}} [1 - e^{-\hat{k}_{\omega}(\tilde{z}_2 - \tilde{z}_1)}].$ (8)

 λ

 $\ddot{}$

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Inserting (7) and (8) into the general formula (4) , we have the desired result-

$$
\alpha = \mathcal{R}_{\omega} \quad \frac{\mathcal{H}(\mathcal{Z}_1, -)}{h(\mathcal{Z}_1)} \quad , \quad \mathcal{Z}_1 \geq \mathcal{Z}_{\omega} \tag{1}
$$

DERIVATION OF (II)

The formula (II) can be obtained directly from (I) by writing

$$
H(z,-) = H(z,-) - H(z,+)
$$

h(z) = h(z,-) + h(z,+)

and invoking the definitions of $R(\vec{z}, -)$, and $D(\vec{z}, \pm)$. That is, in general:

$$
\overline{H}(\vec{z},-)=H(\vec{z},-) - R(\vec{z},-) H(\vec{z},-) = H(\vec{z},-) [1 - R(\vec{z},-)]
$$

and

 \mathbf{r}

$$
h(\mathcal{Z}) = D(\mathcal{Z}_{j-})H(\mathcal{Z}_{j-}) + D(\mathcal{Z}_{j+})H(\mathcal{Z}_{j+})
$$

So that when $Z \geq Z_0$, we have

$$
\alpha = \hat{K}_{\infty} \underbrace{\frac{(\mathsf{I} - \mathcal{R}_{\infty})}{D(\mathsf{I} + \mathcal{R}_{\infty} D(\mathsf{I})}}, \qquad (\text{II})
$$

which is the desired alternate formula. We observe in passing that (II) is a limiting form of the exact formula:

$$
\alpha(z) = \frac{K(z,-) - R(z,-) K(z,+)}{D(z,-) + R(z,-) D(z,+)} \qquad (9)
$$

which follows from equation (25) of reference 4. Clearly, as $\mathcal{F} \rightarrow \tilde{\mathcal{F}} \omega$, equation (9) takes the limiting form (II). Furthermore if we assume $D(f) = 2$, as is done in the classical Schuster two-flow theory of the light field, then (II) reduces to the classical relation:

$$
\alpha = \frac{k_{\infty}}{2} \frac{1 - R_{\infty}}{1 + R_{\infty}} \tag{10}
$$

APPLIED NUMEROLOGY: A RULE OF THUMB

Formula III is to be taken as a convenient rule of thumb, and as such, is subject to possible revision whenever specific optical media are under study. Formula III arises from the ... hatural unless following observed regularities in the values of R_{∞} and $D(T)_N$: \mathcal{R} as is usually found to be in the neighborhood of \mathcal{O} . $\mathcal{O}2$, give or take $O. \bigcirc I$. Furthermore $D(\pm)$ appears to be such

that the sum $D(+) + D(-)$ is usually very nearly equal to .; and the ratio $D(t)/D(-)$ is usually very nearly equal to 2, over great ranges of depths and in many media. Solving these two simultaneous equations yields, to two significant figures:

$$
D(-) = 1.3 ,
$$

\n
$$
D(+) = 2.7 ,
$$
 (11)

which agrees very well with experimental results. It follows that, to the nearest rational number with small integers for numerator and denominator, we have from (II):

C

or

$$
\alpha = \frac{2}{4} k_{\infty},
$$
\n(III)\n
$$
\frac{4}{3} \alpha = k_{\infty}.
$$

Any similarity between the appearance of the fraction $4/3$ in (III), and the index of refraction of water is, unfortunately, coincidental.

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