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Authors

Falk, Nicholas M
Igel, Adele L
Igel, Matthew R

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1 **The Relative Impact of Ice Fall Speeds and Microphysics**

2 **Parameterization Complexity on Supercell Evolution**

3 Nicholas M. Falk¹, Adele L. Igel¹, Matthew R. Igel¹

4 ¹Department of Land Air and Water Resources, University of California Davis, Davis, 95616,

5 USA

6 *Correspondence to:* Nicholas Falk (Nick.Falk@colostate.edu)

7

1 **Abstract.** The use of bin or bulk microphysics schemes in model simulations frequently
2 produces large changes in the simulated storm and precipitation characteristics, but it is still
3 unclear which aspects of these schemes give rise to these changes. In this study, supercell
4 simulations using either a bin or a double-moment bulk microphysics scheme are conducted with
5 the Regional Atmospheric Modeling System (RAMS). The two simulations produce very
6 different storm morphologies. An additional simulation is run for each scheme in which the
7 diameter-fall speed relationships for ice hydrometeors are modified to be similar to those used by
8 the other scheme. When fall speed relationships are homogenized, the two parameterization
9 schemes simulate similar storm morphology. Therefore, despite the use of largely dissimilar
10 approaches to parameterizing microphysics, the difference in storm morphology is found to be
11 related to the choice of diameter-fall speed relationships for ice hydrometeors. This result is
12 investigated further to understand why. Higher fall speeds lead to higher mixing ratios of
13 hydrometeors at low levels and thus more melting. Consequently, stronger downdrafts and cold
14 pools exist in the high fall speed storms, and these stronger cold pools lead to storm splitting and
15 the intensification of a left-mover. The results point to the importance of hydrometeor fall speed
16 on the evolution of supercells. It is also suggested that caution be used when comparing the
17 response of a cloud model to different classes of microphysics schemes since the assumptions
18 made by the schemes may be more important than the scheme class itself.

19

20

1 **1. Introduction**

2 Despite large advances in atmospheric modelling having been made in the past decades,
3 there remain many unanswered questions, open debates, and opportunities for improvement in
4 atmospheric modeling. Increasing computing power is allowing modelers to increase resolution
5 and move toward more widespread use of cloud-resolving simulations; this is making the
6 parameterization of cloud microphysics increasingly important to understand and to perform
7 accurately.

8 Cloud microphysics schemes must simulate complex processes, and the representation of
9 these processes can differ greatly between models. There are currently two common, but
10 fundamentally different approaches to parameterizing cloud microphysics. These two types of
11 microphysics schemes are typically called bin schemes and bulk schemes. A full discussion of
12 these scheme types is given by Khain et al. (2015). We briefly summarize them here. Bulk
13 schemes assume a functional form for the size distribution of hydrometeors; they predict the
14 mass and/or number mixing ratio of each hydrometeor species. Bulk schemes which predict both
15 properties (double-moment schemes) are increasingly used since they tend to perform better than
16 bulk schemes which predict only one property (single-moment schemes) (Igel et al. 2015). Some
17 triple-moment bulk schemes additionally predict a third property, the radar reflectivity
18 (Milbrandt and Yau 2005; Szyrmer et al. 2005; Shipway and Hill 2012; Loftus et al. 2014). Bin
19 microphysics schemes avoid the need to predict moments of a distribution. Instead, each
20 hydrometeor size distribution is discretized into bins that together span the size range of that
21 species. The number and/or mass mixing ratio is predicted for each bin. This allows for bin
22 schemes to predict size distributions that may not be consistent with the assumed distribution
23 shape in bulk schemes. Additionally, prediction of discrete sizes allows bin schemes to simulate

1 many phenomena more realistically than bulk models, for example collision-coalescence. The
2 drawback of bin schemes is their high computational cost compared to bulk schemes. This cost
3 results in bin schemes being used less frequently than bulk schemes.

4 Many studies have shown that the choice of a bin or bulk microphysics scheme can have
5 a large impact on the simulation of cloud and storm properties (Seifert and Beheng 2006;
6 Morrison and Grabowski 2007; Khain et al. 2009; Li et al. 2009; Fan et al. 2012; Iguchi et al.
7 2012; Fan et al. 2015; Wang et al. 2013). However, because bin and bulk schemes vary in so
8 many ways, it is often difficult to understand why the schemes result in such different
9 simulations. It is unclear if the differences in simulations arise primarily due to the different
10 fundamental construction of bulk and bin schemes, or if they arise primarily due to different
11 assumptions that the schemes make when parameterizing the same process.

12 Certainly, the assumptions made in microphysics schemes of either type can have large
13 impacts on simulated cloud properties (Johnson et al. 2015; Posselt 2016). Several studies in the
14 past have focused on understanding the sensitivity of convective storms to assumptions about
15 parameter values related to the size distribution or physical properties of hydrometeors within
16 bulk microphysics schemes. Gilmore et al. (2004) varied the intercept parameter (and therefore
17 indirectly the mean size and fall speed) of graupel and hail and found that simulations with
18 larger, faster falling graupel produced stronger cold pools. Morrison and Milbrandt (2011), van
19 Weverberg et al. (2012), and Bryan and Morrison (2012) all also found cold pools appeared
20 earlier and were stronger when they switched from less dense, slow falling graupel to more
21 dense, fast falling hail. Other studies disagree. Smaller hydrometeors have larger surface area to
22 volume ratios, so they melt faster and are associated with more latent cooling leading to stronger
23 cold pools (van den Heever and Cotton 2004; Cohen and McCaul 2006; Xue et al. 2017). This

1 increase in melting, along with the subsequent dynamical effects, results in storms with smaller
2 hydrometeors having less total precipitation. However, storms with smaller or slower falling
3 hydrometeors can have larger stratiform regions and wider areas of precipitation, since
4 hydrometeors are slower to fall out and there is more horizontal advection of hydrometeors. This
5 response of the stratiform region is consistent among studies even if the cold pool response is not
6 (Adams-Selin et al. 2013; Bryan and Morrison, 2012; van Weverberg et al. 2012).

7 All these studies change size distribution parameters that indirectly impact the average
8 fall speeds of hydrometeors, but the parameters in fall speed-diameter relationships can be just as
9 important for determining the average fall speed as the size distribution parameter (McFarquhar
10 and Black 2004). Few previous studies directly vary fall speed-diameter relationships. Bennetts
11 and Rawlins (1981) simulated later onset of heavy precipitation and lower total precipitation
12 when they lowered the fall speed of hail to that of rain in their early model. Adams-Selin et al.
13 (2013) found that faster falling graupel led to more intense cold pools.

14 These results taken together indicate that the choice of hydrometeor diameter-fall speed
15 relationship could have large direct impacts on storms and suggest that the sensitivity to
16 parameters that impact the hydrometeor size distribution may in part be a result of indirectly
17 changing the average fall speeds.

18 In this study, we describe a case in which the differences between simulations with a bin
19 and a bulk scheme seem to arise primarily from different, reasonable assumptions that are made
20 about the fall velocity of ice particles. The results suggest that the fundamentally different design
21 of bulk and bin schemes may be less important than the shared choices that the schemes must
22 make about hydrometeor properties, at least in some cases. This conclusion is in agreement with
23 Adams-Selin et al. (2013) who noted that changing hail sizes (and consequent settling rates) had

1 larger impacts on their bow echo simulations than the choice of bulk microphysics scheme. They
2 also found changing graupel fall speeds independently of size had as large of an impact as
3 changing graupel size and fall speed in concert. This is also in agreement with Xue et al. (2017)
4 who found that assumptions about hydrometeor properties could make large changes in storm
5 dynamics and thermodynamics. They also conclude that caution must be taken when comparing
6 bin and bulk schemes. We also discuss the physical processes that cause such a large sensitivity
7 to the fall velocity of ice particles. The simulations are described in Section 2, and the results are
8 discussed in Section 3.

9 **2. Experimental Design**

10 Four supercell simulations were run using the Regional Atmospheric Modeling System
11 (RAMS) (Cotton et al. 2003) to test the relative sensitivity of these storms to changes in
12 microphysical parameterizations. The domain size was 300 km east to west and 250 km north to
13 south. Grid spacing was 1 km in the horizontal, and variable in the vertical with a minimum
14 spacing of 50 m at the surface, a stretching factor of 1.1, and a maximum of 1 km. The
15 simulations were run for 3 hours with a 4 s time step. We used an open radiative boundary
16 condition for outflow and a zero-gradient condition for inflow. The Coriolis force was included
17 using a latitude of 37°N. Radiation was parameterized using the Harrington (1997) radiation
18 scheme and updated every 10 minutes. Sub-grid scale turbulence was parameterized following
19 Smagorinsky (1963) with modifications by Lilly (1962) and Hill (1974). The initial
20 thermodynamic conditions and convective initialization are the same as those used by Grant and
21 van den Heever (2014). Temperature and moisture profiles were initially horizontally uniform
22 and used the analytical sounding of Weisman and Klemp (1982) with 13 g/kg of moisture in the

1 boundary layer. The wind shear follows an idealized half-circle profile with speeds increasing to
2 32 m/s at 5km (Weisman and Klemp 1984). There is a small amount of shear between 5 and 7
3 km and above 7 km the winds are constant with height. Convection was initiated using a square 2
4 K warm bubble 10 km on a side.

5 Simulations were run with either the RAMS double-moment bulk microphysics scheme
6 (Saleeby and Cotton 2004; Saleeby and van den Heever 2013), or the Hebrew University
7 Spectral Bin Model (SBM) (Khain et al. 2004) which is a recently added option in RAMS (Igel
8 and van den Heever 2017). The RAMS bulk scheme has eight hydrometeor categories: cloud
9 droplets, drizzle, rain, pristine ice, snow, aggregated snow, graupel, and hail. Pristine ice
10 corresponds to smaller ice crystals and snow corresponds to larger sizes. The SBM on the other
11 hand uses seven hydrometeor categories: liquid drops, three categories of ice crystals (plates,
12 columns, and dendrites), aggregated snow, graupel, and hail.

13 There are many different assumptions made by the two schemes. One particularly
14 striking difference is the diameter-fall speed relationships of aggregated snow, graupel, and hail.
15 These three species will be hereafter referred to as “big ice” and the remaining ice categories will
16 be referred to as “small ice”. We ran one RAMS bulk simulation with its standard fall speeds and
17 one SBM simulation with its standard fall speeds. The bulk scheme simulation with its standard
18 fall speeds will be referred to as BULK-H and the SBM simulation with the standard SBM fall
19 speeds will be referred to as BIN-L (H for high fall speed and L for low fall speed). To test
20 whether these different diameter-fall speed relationships could explain differences between the
21 RAMS bulk and SBM simulations (see next section), we ran two additional simulations in which
22 the diameter-fall speed relationships in each scheme were altered. A second bulk scheme
23 simulation was run with the standard bulk scheme fall speeds for aggregated snow, graupel and

1 hail reduced to half in order to effectively mimic the bin scheme fall speeds. This simulation
2 with the SBM-like fall speeds will be referred to as BULK-L. Finally, a second SBM simulation
3 was run using the bulk scheme fall speeds for big ice and will be referred to as BIN-H. Only the
4 diameter-fall speed relationships of big ice are changed; the diameter-fall speed relationships of
5 all other hydrometers are left unchanged. Furthermore, note that due to the use of precomputed
6 collection kernels, the changes in fall speeds do not impact the collection rates in either BULK or
7 BIN. Due to the structure of the code, the fall speeds impact the ventilation coefficients used for
8 condensation and evaporation in BIN, but not in BULK.

9 In the RAMS bulk scheme, piecewise power law relationships are based on data
10 presented in Mitchell (1996). In the SBM, power law relationships for aggregated snow and hail
11 are based on Pruppacher and Klett (1997) and for graupel on Khain et al. (2001). The first
12 column of Figure 1 shows each relationship for 1000 mb. The diameter-fall speed relationships
13 used by the two schemes differ by 50% or more for most sizes across all three categories, with
14 the RAMS bulk scheme fall speeds being greater than those in SBM in all cases. These
15 differences are not related to the choice of mass-diameter relationship which is shown in the
16 second column of Figure 1. For graupel and hail, the mass-diameter relationships are similar and
17 the same, respectively. For aggregated snow, the mass-diameter relationships differ substantially;
18 however, the SBM diagnoses *higher* mass and *lower* fall speed for the same diameter. Therefore,
19 the mass-diameter relationships do not explain why the fall speed relationships disagree. Instead,
20 the different relationships appear to simply be different choices made by the scheme developers
21 and represent some degree of uncertainty present in empirical data. We will reiterate; both sets of
22 relationships are reasonable and based on published literature. Furthermore, we make no
23 judgment here about which set of fall speed relationships is more realistic or appropriate, and in

1 the results that follow, no judgment about which simulation is considered to be the most
2 “correct”. We consider all simulations to be equally likely representations of reality.

3 **3. Results**

4 *a. Supercell Evolution*

5 Figure 2 shows the column and time (from output every 10 minutes; shaded) maximum
6 velocity and accumulated precipitation (white contours) for all four simulations. The
7 immediately notable difference in Figure 2 is that both high fall speed simulations (i.e. “H”
8 simulations) split with distinct right- and left-movers, while neither of the low fall speed
9 simulations (i.e. “L”) exhibit a classic split. BULK-L remains a single cell storm and BIN-L
10 develops some secondary convection that remains near to the main storm. At 60 minutes,
11 splitting in both BULK-H and BIN-H is apparent (Figure 2a and 2c) and after 100 minutes, BIN-
12 H’s left-mover intensifies (Figure 2c). Despite these differences, Figure 2 also indicates that the
13 right mover updrafts are all quite similar. BULK-H and BIN-H have more total precipitation than
14 BULK-L and BIN-L respectively, but both bulk simulations have more precipitation than either
15 bin simulation. Because the storms split, the precipitation fields of the high fall speed simulations
16 cover much more area than their low fall speed counterparts. BULK-H and BIN-H have
17 precipitation fields of approximately the same area; and despite the disorganized secondary
18 convection that forms in BIN-L, BULK-L’s precipitation field is roughly the same area (Figure
19 2). Nearly all precipitation reaches the surface as rain except for a negligible amount of hail early
20 in the bulk simulations. The conclusion we draw from Fig. 2 is that the behavior of simulated
21 supercells depends more strongly on fall speed parameterization than on microphysics scheme

1 type in our simulations. This conclusion is significant as it indicates how critical the fall speeds
2 are to the evolution of supercell thunderstorms.

3 Most of the differences between our simulations can be attributed to splitting. Thus, in
4 the remainder of this work, we further investigate the differences among the simulations and
5 discuss why H simulations split and L simulations do not. To suggest why the left-mover does or
6 does not form, we will first investigate the microphysical behavior of our storms and then link
7 the microphysics back to storm dynamics.

8 *b. Microphysics*

9 The direct result of differing fall speeds should be on the vertical structure of
10 hydrometeors in the storms. Higher fall speeds will logically allow big ice to precipitate faster
11 out of the anvil and core where it can begin to melt and contribute to increased rain near the
12 surface. This effect is observed in the storms beginning at about 30 minutes (not shown). At 60
13 minutes, we see increased rain water content in both H simulations compared to the L
14 simulations (Fig. 3a). At this same time, we see higher ice water contents in H compared to L at
15 lower elevations for all three species – hail, graupel, and aggregated snow (Fig. 3b-d). There are
16 some clear shifts between the BULK and BIN simulations. BULK has more hail and less graupel
17 than BIN, but the summation of the two species is about the same. Additionally, the BIN
18 simulations have a more prominent downward shift of the ice mass, particularly for hail and
19 aggregates. This may be because for small hydrometeor sizes the difference in fall speeds is
20 greater between H and L for the BIN simulations (Fig. 1 a, c, e). Finally, the simulations could
21 develop different mean hydrometeor size that would also affect the fall speeds, but this seems to
22 be a secondary effect on the profiles shown in Figure 3. BIN-H has smaller big ice at low levels

1 than BIN-L which partly offsets the higher fall speeds in BIN-H; however, BULK-H and BULK-
2 L have more similar sizes (not shown).

3 One immediate impact of these differences in hydrometeor mass content is the rate of
4 melting. Figure 4a indicates that melting is sensitive to the fall speed relationship from the start
5 with BULK-H having the most melting, followed by BULK-L, BIN-H, and BIN-L. An increase
6 in evaporation of rain follows from the greater melting of precipitating ice (Figure 4d). BULK-L
7 has more melting than BIN-H, but BIN-H has more evaporation than BULK-L. Due to the
8 differing ice nucleation schemes and riming and aggregation schemes, BIN-H has smaller mean
9 raindrop diameters (not shown). Thus, evaporation occurs more rapidly in BIN-H despite less
10 melting. The high fall speed storms consistently evaporate more rain than their low fall speed
11 counterparts.

12 The additional melting in H simulations compared to their L counterparts is not driven by
13 additional freezing. Freezing primarily occurs in the updrafts, and total freezing is closely linked
14 to updraft mass flux. As previously shown, the updrafts in the right-movers are all quite similar
15 (Figure 2). Furthermore, since we did not alter any of the properties of the warm phase
16 hydrometeors, we expect the total freezing in the initial development of the storms to be similar.
17 As expected, Figure 4b shows the freezing in the storms depends more on the microphysics
18 scheme than the fall speed relationship until splitting occurs. For BULK-H, its freezing is nearly
19 identical to BULK-L until 60 minutes, when its left-mover begins to intensify. For BIN-H, its
20 freezing is very similar to BIN-L, at some points even less, until after 100 minutes when its left-
21 mover intensifies.

22 Figure 4c also shows the ratio of melting to freezing. This metric is introduced as a
23 simple way of examining the upward convective flux of liquid to the downward settling of ice. In

1 all cases, this ratio indicates that the storms freeze more water than they melt. The high fall speed
2 simulations melt a much higher fraction of the mass frozen than do the low fall speed
3 simulations, which is logically consistent with our previous findings. After 40 minutes, BIN-H
4 surpasses BULK-L so the ratio of melting to freezing is determined by the fall speed
5 relationships first and the parameterization scheme second (i.e. BULK-H, BIN-H, BULK-L,
6 BIN-L).

7 *c. Downdrafts and Cold Pools*

8 The consequence of the increase in melting and evaporation in the high fall speed
9 simulations is an increase in latent cooling (Figure 4). The increase in latent cooling coupled
10 with an increase in condensate loading (implied in Figure 3) leads to the formation of stronger
11 downdrafts at low levels (Figure 5). For the first 30 minutes of the simulation, all simulations
12 have almost identical downdrafts below 5 kilometers. From 30 minutes to the split becoming
13 apparent at 60 minutes, BULK-H and BIN-H have discernably stronger and more widespread
14 downdrafts (Figure 5a, c) than BULK-L and BIN-L (Figure 5b, d) in the lowest 5 kilometers.
15 Additionally, the structure of the downdrafts is markedly different. The H simulations have much
16 larger and more intense downdrafts on the northern (forward) flank of their right-movers than the
17 L simulations in the lowest 5 kilometers.

18 The next connection is from increased downdrafts and melting to stronger cold pools;
19 more melting cools the air and stronger downdrafts force it to the surface where it spreads into
20 the cold pool and can excite new convection. Figure 6 shows these effects occur in our
21 simulations. At 100 minutes, BULK-H has a cold pool with an area of 4639 km² (cold pool area
22 defined as grid boxes at the surface with a temperature deviation of more than -0.5 K) and a
23 maximum temperature deviation of -8 K from the background (Figure 6a). BIN-H and BULK-L

1 both have smaller cold pools – 3046 and 3521 km² respectively – with maximum deviations of -4
2 K (Figure 6b and 6c respectively). BIN-L has the smallest cold pool, 1639 km², and a maximum
3 deviation of only -2 K (Figure 6d). The dynamic effect of these cold pools is also clear. In the H
4 simulations, and to a lesser extent in BULK-L, the cold pools are pushing air outward, and then
5 upward (as implied by convergence) at their edge. This is consistent with past research (Rotunno
6 et al. 1988; Weisman and Rotunno 2004). In the H simulations, left-mover updrafts are situated
7 on the edge of their respective cold pools (Figure 6).

8 As discussed in the introduction, there have been conflicting conclusions drawn on the
9 impact of hydrometeor size and fall speed on cold pool intensity in the past. Our results are in
10 better agreement with those studies which associate higher fall speeds with colder cold pools
11 (Bryan and Morrison 2012; Van Weverberg et al. 2012; Adams-Selin et al. 2013). However, we
12 note that in our simulations, there is nearly complete melting of hail and graupel in all
13 simulations. This was not the case in van den Heever and Cotton (2004), Cohen and McCaul
14 (2006), or Xue et al. (2017). If a significant fraction of our hail and graupel had remained frozen,
15 then we may have expected results that were more in line with these studies.

16 *d. “Splitting”*

17 As we have shown, the H simulations have stronger cold pools due indirectly to higher
18 fall speeds of big ice. These stronger cold pools lead to the formation of additional convective
19 cells in the H simulations. In many respects, these new convective cells resemble left movers
20 that have split from their parent supercell. They have anticyclonic updraft vorticity, they move
21 at an acute angle to their parent cell, and they are generally weaker than their parent supercell.
22 However, the formation and maintenance of these cells appears to be largely driven by cold pool

1 dynamics (see next). So, by “split” we simply mean the generation of a leftward moving
2 secondary storm or storms by any mechanism.

3 Figure 7 shows the evolution of updrafts at 1.4 km for BULK-H and BULK-L. BIN-H
4 and BIN-L behave similarly to BULK-H and BULK-L, respectively, and are not shown. Both
5 BULK-H and BULK-L have a main updraft and a smaller, distinct updraft developing on the
6 north side of the main updraft at 50 minutes (Figure 7a, d). This smaller updraft would seem to
7 be the beginning of the left-mover. However, 10 minutes later (Figure 7b, e) these traditional
8 left-movers have dissipated while a third updraft is forming farther to the north along the edge of
9 the cold pool in BULK-H. By 75 minutes (Figure 7c, f), the traditional left-mover has all but
10 dissipated while the “true” left mover in BULK-H is still on the cold pool edge and intensifying.
11 Bulk-H's left mover is even beginning to contribute to the cold pool.

12 Traditional supercell splitting in high shear environments is driven by a top-down process
13 where heavy precipitation and downdrafts act to separate regions of oppositely-signed updraft
14 vorticity. It is certainly reasonable to imagine that the changes to the precipitation properties we
15 have shown above could result in favorable conditions for such processes. Figures 8 and 9 show
16 height vs meridional distance from both BULK and BIN simulations of zonal maximum updraft
17 speed at several times before and during the initial generation of new convection in the H
18 simulations. In all simulations, the main updraft is adjacent to an area of enhanced vertical
19 velocity just to its north (seen as a spreading of contours to the right of in the figure; the
20 “traditional” left mover). With time, this adjacent velocity region descends slightly. In the H
21 simulations, this adjacent velocity max appears to excite new surface-based convection at the
22 leading edge of the cold pool (the “true” left mover). This new convection grows as it moves
23 north of the main updraft. Figure 6 includes an outline of the region whose surface-to-500m

1 vorticity matches the magnitude of the environmental shear in that same layer. Canonically,
2 RKW theory (Weisman and Klemp 1984, 1982) predicts that maintenance of new convection is
3 favored in regions whose low-level vorticity balances the shear. These conditions only appear to
4 be met at the northern flank of the cold pool by H simulation storms. So, Figs. 6, 8, and 9
5 suggest that splitting in the H simulations is due either to enhanced excitation of new surface-
6 based convection by a subtly descending northern downdraft (Figures 8 and 9) and/or due to
7 better convective maintenance mechanisms provided by the stronger cold pools (Fig. 6). Either
8 way, the splits appear to be similar to the mechanisms proposed to result in a supercell flanking a
9 squall (Klemp and Wilhelmson 1978; Wilhelmson and Klemp 1978). Although, we reiterate that
10 it is not precisely our intention to prove why sustained split convection exists, but rather, it is to
11 point out the more novel result that simulating splitting can depend on fall speed choices in at
12 least some cases.

13 **4. Conclusions**

14 Model simulations of a supercell were conducted with the Regional Atmospheric
15 Modeling System (RAMS), which has a double-moment bulk microphysics scheme as well as a
16 bin microphysics scheme. Before alterations were made, the bulk scheme simulated a splitting
17 supercell whereas the bin scheme did not. These schemes have differing built-in diameter-fall
18 speed relationships for aggregated snow, graupel, and hail. Additional simulations were run in
19 which the diameter-fall speed relationships in each scheme were modified to be more like the
20 other. Both high fall speed simulations developed defined left-movers while neither low fall
21 speed simulation did. We conclude for this case that the cause of the starkly different bin and
22 bulk simulations was a result of their choice of diameter-fall speed relationships for ice

1 hydrometeors, and not the scheme design itself. Such a result suggests that the fundamentally
2 different design of bulk and bin schemes may be less important than the assumptions about
3 hydrometeor properties that they make.

4 We investigated why the diameter-fall speed relationships had such a large impact on the
5 supercell development. We suggest that the following chain of events caused the high fall speed
6 simulations to split. Higher fall speeds led to more hydrometeor mass at low levels, which led to
7 greater melting and evaporation. Along with consequently larger condensate loading, this latent
8 cooling led to stronger downdrafts, which led to larger and colder cold pools. These stronger
9 cold pools intensified the left-movers in simulations with high fall speeds. The splitting of these
10 storms was described.

11 These results lead us to make several broad suggestions. First, in agreement with
12 recommendations by Fan et al. (2016), more empirical data on ice diameter-fall speed
13 relationships should be collected. Many efforts in this regard are currently underway (Cheng et
14 al. 2015; Heymsfield et al. 2018; Westbrook and Sephton 2017). Second, while in our
15 simulations the mass-diameter and velocity-diameter relationships were independent, we know
16 that they should instead be linked and both depend on ice density. Ice density is usually
17 prescribed, but more microphysics schemes may need to include ways to diagnose or predict ice
18 density in order to have better relationships between hydrometeor size, mass, and velocity, such
19 is currently done by the P3 bulk microphysics scheme (Morrison and Milbrandt 2015). Finally,
20 care should be taken when comparing microphysics schemes, particularly bin and bulk schemes.
21 Here we have shown that the primary differences in the storm evolution did not come from the
22 large differences associated with the scheme type, but rather to one assumption that both
23 schemes needed to make.

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6 NMF wrote the original draft and ALI and MRI reviewed and edited the manuscript. The current
7 version of RAMS is publicly available at vandenheever.atmos.colostate.edu/~vdhpage/rams.php.
8 Model simulations are available for download at farm.cse.ucdavis.edu/~aigel. Plot data is
9 available in the Dash FAIR-aligned data repository at dash.ucdavis.edu. DOI: 10.25338/B8RC84

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1 **References**

- 2 Adams-Selin, R. D., S. C. van den Heever, and R. H. Johnson, 2013: Sensitivity of Bow-Echo
3 Simulation to Microphysical Parameterizations. *Mon. Weather Rev.*, **28**, 1188–1209,
4 doi:10.1175/WAF-D-12-00108.1. [http://journals.ametsoc.org/doi/abs/10.1175/WAF-D-12-](http://journals.ametsoc.org/doi/abs/10.1175/WAF-D-12-00108.1)
5 00108.1.
- 6 Bennetts, D. A., and F. Rawlins, 1981: Parameterization of the ice-phase in a model of mid-
7 latitude cumulonimbus convection and its influence on the simulation of cloud
8 development. *Q. J. R. Meteorol. Soc.*, **107**.
- 9 Bryan, G. H., and H. Morrison, 2012: Sensitivity of a Simulated Squall Line to Horizontal
10 Resolution and Parameterization of Microphysics. *Mon. Weather Rev.*, doi:10.1175/MWR-
11 D-11-00046.1.
- 12 Cheng, K.-Y., P. K. Wang, and T. Hashino, 2015: A Numerical Study on the Attitudes and
13 Aerodynamics of Freely Falling Hexagonal Ice Plates. *J. Atmos. Sci.*, **72**, 3685–3698,
14 doi:10.1175/jas-d-15-0059.1.
- 15 Cohen, C., and E. W. McCaul, 2006: The Sensitivity of Simulated Convective Storms to
16 Variations in Prescribed Single-Moment Microphysics Parameters that Describe Particle
17 Distributions, Sizes, and Numbers. *Mon. Weather Rev.*, doi:10.1175/MWR3195.1.
- 18 Cotton, W. R., and Coauthors, 2003: RAMS 2001: Current status and future directions.
19 *Meteorol. Atmos. Phys.*, **82**, 5–29, doi:10.1007/s00703-001-0584-9.
20 <http://link.springer.com/10.1007/s00703-001-0584-9>.
- 21 Fan, J., L. R. Leung, Z. Li, H. Morrison, H. Chen, Y. Zhou, Y. Qian, and Y. Wang, 2012:

1 Aerosol impacts on clouds and precipitation in eastern China: Results from bin and bulk
2 microphysics. *J. Geophys. Res. Atmos.*, doi:10.1029/2011JD016537.

3 ———, and Coauthors, 2015: Improving representation of convective transport for scale-aware
4 parameterization: 1. Convection and cloud properties simulated with spectral bin and bulk
5 microphysics. *J. Geophys. Res. Atmos.*, **120**, 3485–3509, doi:10.1002/2014jd022142.

6 ———, Y. Wang, D. Rosenfeld, and X. Liu, 2016: Review of Aerosol–Cloud Interactions:
7 Mechanisms, Significance, and Challenges. *J. Atmos. Sci.*, doi:10.1175/JAS-D-16-0037.1.

8 Gilmore, M. S., J. M. Straka, and E. N. Rasmussen, 2004: Precipitation Uncertainty Due to
9 Variations in Precipitation Particle Parameters within a Simple Microphysics Scheme. *Mon.*
10 *Weather Rev.*, **132**, 2610–2627, doi:10.1175/MWR2810.1.
11 <http://journals.ametsoc.org/doi/abs/10.1175/MWR2810.1>.

12 Grant, L. D., and S. C. van den Heever, 2014: Microphysical and Dynamical Characteristics of
13 Low-Precipitation and Classic Supercells. *J. Atmos. Sci.*, doi:10.1175/JAS-D-13-0261.1.

14 Harrington, J. Y., 1997: The Effects of Radiative and Microphysical processes on simulated
15 warm and transition season arctic stratus. 289 pp.

16 van den Heever, S. C., and W. R. Cotton, 2004: The Impact of Hail Size on Simulated Supercell
17 Storms. *J. Atmos. Sci.*, **61**, 1596–1609, doi:10.1175/1520-
18 0469(2004)061<1596:TIOHSO>2.0.CO;2.
19 <http://journals.ametsoc.org/doi/abs/10.1175/1520->
20 [0469%282004%29061%3C1596%3ATIOHSO%3E2.0.CO%3B2](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%282004%29061%3C1596%3ATIOHSO%3E2.0.CO%3B2).

21 Heymsfield, A., M. Szakáll, A. Jost, I. Giammanco, and R. Wright, 2018: A Comprehensive

1 Observational Study of Graupel and Hail Terminal Velocity, Mass Flux, and Kinetic
2 Energy. *J. Atmos. Sci.*, **75**, 3861–3885, doi:10.1175/jas-d-18-0035.1.

3 Hill, G. E., 1974: Factors Controlling the Size and Spacing of Cumulus Clouds as Revealed by
4 Numerical Experiments. *J. Atmos. Sci.*, **31**, 646–673, doi:10.1175/1520-
5 0469(1974)031<0646:FCTSAS>2.0.CO;2.
6 <http://journals.ametsoc.org/doi/abs/10.1175/1520->
7 [0469%281974%29031%3C0646%3AFCTSAS%3E2.0.CO%3B2](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%281974%29031%3C0646%3AFCTSAS%3E2.0.CO%3B2).

8 Igel, A. L., and S. C. van den Heever, 2017: The Importance of the Shape of Cloud Droplet Size
9 Distributions in Shallow Cumulus Clouds. Part I: Bin Microphysics Simulations. *J. Atmos.*
10 *Sci.*, **74**, 249–258, doi:10.1175/JAS-D-15-0382.1.
11 <http://journals.ametsoc.org/doi/10.1175/JAS-D-15-0382.1>.

12 ———, M. R. Igel, and S. C. van den Heever, 2015: Make It a Double? Sobering Results from
13 Simulations Using Single-Moment Microphysics Schemes. *J. Atmos. Sci.*, **72**, 910–925,
14 doi:10.1175/JAS-D-14-0107.1. <http://journals.ametsoc.org/doi/10.1175/JAS-D-14-0107.1>.

15 Iguchi, T., T. Nakajima, A. P. Khain, K. Saito, T. Takemura, H. Okamoto, T. Nishizawa, and
16 W.-K. Tao, 2012: Evaluation of Cloud Microphysics in JMA-NHM Simulations Using Bin
17 or Bulk Microphysical Schemes through Comparison with Cloud Radar Observations. *J.*
18 *Atmos. Sci.*, doi:10.1175/JAS-D-11-0213.1.

19 Johnson, J. S., Z. Cui, L. A. Lee, J. P. Gosling, A. M. Blyth, and K. S. Carslaw, 2015: Evaluating
20 uncertainty in convective cloud microphysics using statistical emulation. *J. Adv. Model.*
21 *Earth Syst.*, doi:10.1002/2014MS000383.

- 1 Khain, A., M. B. Pinsky, and M. Shapiro, 2001: Graupel-drop collision efficiencies. *J. Atmos.*
2 *Sci.*, **58**, 2571–2595.
- 3 ———, A. Pokrovsky, M. Pinsky, A. Seifert, and V. Phillips, 2004: Simulation of Effects of
4 Atmospheric Aerosols on Deep Turbulent Convective Clouds Using a Spectral
5 Microphysics Mixed-Phase Cumulus Cloud Model. Part I: Model Description and Possible
6 Applications. *J. Atmos. Sci.*, **61**, 2963–2982, doi:10.1175/JAS-3350.1.
- 7 Khain, A. P., L. R. Leung, B. Lynn, and S. Ghan, 2009: Effects of aerosols on the dynamics and
8 microphysics of squall lines simulated by spectral bin and bulk parameterization schemes. *J.*
9 *Geophys. Res. Atmos.*, doi:10.1029/2009JD011902.
- 10 ———, and Coauthors, 2015: Representation of microphysical processes in cloud-resolving
11 models: Spectral (bin) microphysics versus bulk parameterization. *Rev. Geophys.*,
12 doi:10.1002/2014RG000468.
- 13 Klemp, J. B., and R. B. Wilhelmson, 1978: Simulations of Right- and Left-Moving Storms
14 Produced Through Storm Splitting. *J. Atmos. Sci.*, **35**, 1097–1110, doi:10.1175/1520-
15 0469(1978)035<1097:SORALM>2.0.CO;2.
16 [http://journals.ametsoc.org/doi/abs/10.1175/1520-
17 0469%281978%29035%3C1097%3ASORALM%3E2.0.CO%3B2](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%281978%29035%3C1097%3ASORALM%3E2.0.CO%3B2).
- 18 Li, X., W.-K. Tao, A. P. Khain, J. Simpson, and D. E. Johnson, 2009: Sensitivity of a Cloud-
19 Resolving Model to Bulk and Explicit Bin Microphysical Schemes. Part I: Comparisons. *J.*
20 *Atmos. Sci.*, doi:10.1175/2008JAS2646.1.
- 21 Lilly, D. K., 1962: On the numerical simulation of buoyant convection. *Tellus*, **14**, 148–172,

1 doi:10.3402/tellusa.v14i2.9537.
2 <https://www.tandfonline.com/doi/full/10.3402/tellusa.v14i2.9537>.

3 Loftus, A. M., W. R. Cotton, and G. G. Carrió, 2014: A triple-moment hail bulk microphysics
4 scheme. Part I: Description and initial evaluation. *Atmos. Res.*,
5 doi:10.1016/j.atmosres.2014.05.013.

6 McFarquhar, G. M., and R. A. Black, 2004: Observations of Particle Size and Phase in Tropical
7 Cyclones: Implications for Mesoscale Modeling of Microphysical Processes. *J. Atmos. Sci.*,
8 doi:10.1175/1520-0469(2004)061<0422:OOPSAP>2.0.CO;2.

9 Milbrandt, J. A., and M. K. Yau, 2005: A Multimoment Bulk Microphysics Parameterization.
10 Part II: A Proposed Three-Moment Closure and Scheme Description. *J. Atmos. Sci.*,
11 doi:10.1175/JAS3535.1.

12 Mitchell, D. L., 1996: Use of Mass- and Area-Dimensional Power Laws for Determining
13 Precipitation Particle Terminal Velocities. *J. Atmos. Sci.*, **53**, 1710–1723, doi:10.1175/1520-
14 0469(1996)053<1710:UOMAAD>2.0.CO;2.
15 <http://journals.ametsoc.org/doi/abs/10.1175/1520->
16 [0469%281996%29053%3C1710%3AUOMAAD%3E2.0.CO%3B2](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%281996%29053%3C1710%3AUOMAAD%3E2.0.CO%3B2).

17 Morrison, H., and W. W. Grabowski, 2007: Comparison of Bulk and Bin Warm-Rain
18 Microphysics Models Using a Kinematic Framework. *J. Atmos. Sci.*, doi:10.1175/JAS3980.

19 ———, and J. Milbrandt, 2011: Comparison of Two-Moment Bulk Microphysics Schemes in
20 Idealized Supercell Thunderstorm Simulations. *Mon. Weather Rev.*,
21 doi:10.1175/2010MWR3433.1.

- 1 Morrison, H., and J. A. Milbrandt, 2015: Parameterization of Cloud Microphysics Based on the
2 Prediction of Bulk Ice Particle Properties. Part I: Scheme Description and Idealized Tests. *J.*
3 *Atmos. Sci.*, **72**, 287–311, doi:<https://doi.org/10.1175/JAS-D-14-0065.1>.
- 4 Posselt, D. J., 2016: A Bayesian Examination of Deep Convective Squall-Line Sensitivity to
5 Changes in Cloud Microphysical Parameters. *J. Atmos. Sci.*, doi:[10.1175/JAS-D-15-0159.1](https://doi.org/10.1175/JAS-D-15-0159.1).
- 6 Pruppacher, H., and J. D. Klett, 1997: *Microphysics of Clouds and Precipitation: With an*
7 *Introduction to Cloud Chemistry and Cloud Electricity*. 914 pp.
- 8 Rotunno, R., J. B. Klemp, and M. L. Weisman, 1988: A Theory for Strong, Long-Lived Squall
9 Lines. *J. Atmos. Sci.*, **45**, 463–485.
- 10 Saleeby, S. M., and W. R. Cotton, 2004: A large-droplet mode and prognostic number
11 concentration of cloud droplets in the Colorado State University Regional Atmospheric
12 Modeling System (RAMS). Part I: Module descriptions and supercell test simulations. *J.*
13 *Appl. Meteorol.*, **43**, 182–195, doi:[10.1175/1520-0450\(2004\)043,0182:ALMAPN.2.0.CO;2](https://doi.org/10.1175/1520-0450(2004)043,0182:ALMAPN.2.0.CO;2).
- 14 ———, and S. C. van den Heever, 2013: Developments in the CSU-RAMS aerosol model:
15 Emissions, nucleation, regeneration, deposition, and radiation. *J. Appl. Meteorol. Climatol.*,
16 **52**, 2601–2622, doi:[10.1175/JAMC-D-12-0312.1](https://doi.org/10.1175/JAMC-D-12-0312.1).
- 17 Seifert, A., and K. D. Beheng, 2006: A two-moment cloud microphysics parameterization for
18 mixed-phase clouds. Part 1: Model description. *Meteorol. Atmos. Phys.*,
19 doi:[10.1007/s00703-005-0112-4](https://doi.org/10.1007/s00703-005-0112-4).
- 20 Shipway, B. J., and A. A. Hill, 2012: Diagnosis of systematic differences between multiple
21 parametrizations of warm rain microphysics using a kinematic framework. *Q. J. R.*

1 *Meteorol. Soc.*, doi:10.1002/qj.1913.

2 Smagorinsky, J., 1963: General Circulation Experiments with the Primitive Equations, 1. the
3 basic experiment. *Mon. Weather Rev.*, doi:10.1175/1520-
4 0493(1963)091<0099:GCEWTP>2.3.CO;2.

5 Szyrmer, W., S. Laroche, and I. Zawadzki, 2005: A Microphysical Bulk Formulation Based on
6 Scaling Normalization of the Particle Size Distribution. Part I: Description. *J. Atmos. Sci.*,
7 **62**, 4206–4221, doi:10.1175/JAS3621.1.
8 <http://journals.ametsoc.org/doi/abs/10.1175/JAS3621.1>.

9 Wang, Y., J. Fan, R. Zhang, L. R. Leung, and C. Franklin, 2013: Improving bulk microphysics
10 parameterizations in simulations of aerosol effects. *J. Geophys. Res. Atmos.*, **118**, 5361–
11 5379, doi:10.1002/jgrd.50432.

12 Weisman, M. L., and J. B. Klemp, 1982: The Dependence of Numerically Simulated Convective
13 Storms on Vertical Wind Shear and Buoyancy. *Mon. Weather Rev.*, **110**, 504–520.

14 ———, and ———, 1984: The Structure and Classification of Numerically Simulated Convective
15 Storms in Directionally Varying Wind Shears. *Mon. Weather Rev.*, **112**, 2479–2498.

16 ———, and R. Rotunno, 2004: “A Theory for Strong Long-Lived Squall Lines” Revisited. *J.*
17 *Atmos. Sci.*, **61**, 361–382, doi:10.1175/1520-0469(2004)061<0361:ATFSLS>2.0.CO;2.

18 Westbrook, C. D., and E. K. Sephton, 2017: Using 3-D-printed analogues to investigate the fall
19 speeds and orientations of complex ice particles. *Geophys. Res. Lett.*,
20 doi:10.1002/2017GL074130.

21 Van Weverberg, K., A. M. Vogelmann, H. Morrison, and J. A. Milbrandt, 2012: Sensitivity of

1 Idealized Squall-Line Simulations to the Level of Complexity Used in Two-Moment Bulk
2 Microphysics Schemes. *Mon. Weather Rev.*, doi:10.1175/MWR-D-11-00120.1.

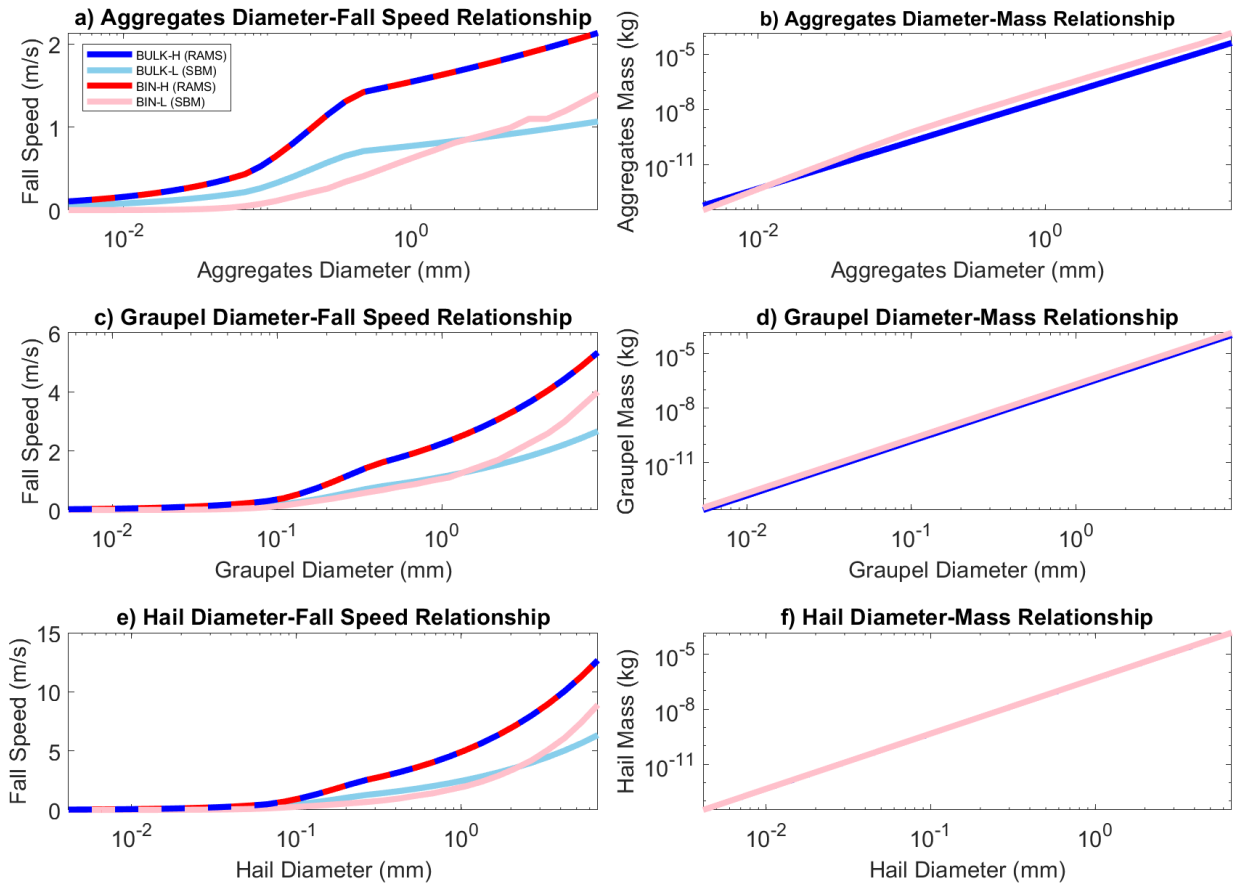
3 Wilhelmson, R. B., and J. B. Klemp, 1978: A Numerical Study of Storm Splitting that Leads to
4 Long-Lived Storms. *Mon. Weather Rev.*, **35**, 1974–1986.

5 Xue, L., and Coauthors, 2017: Idealized Simulations of a Squall Line from the MC3E Field
6 Campaign Applying Three Bin Microphysics Schemes : Dynamic and Thermodynamic
7 Structure. *Mon. Weather Rev.*, 4789–4812, doi:10.1175/MWR-D-16-0385.1.

8

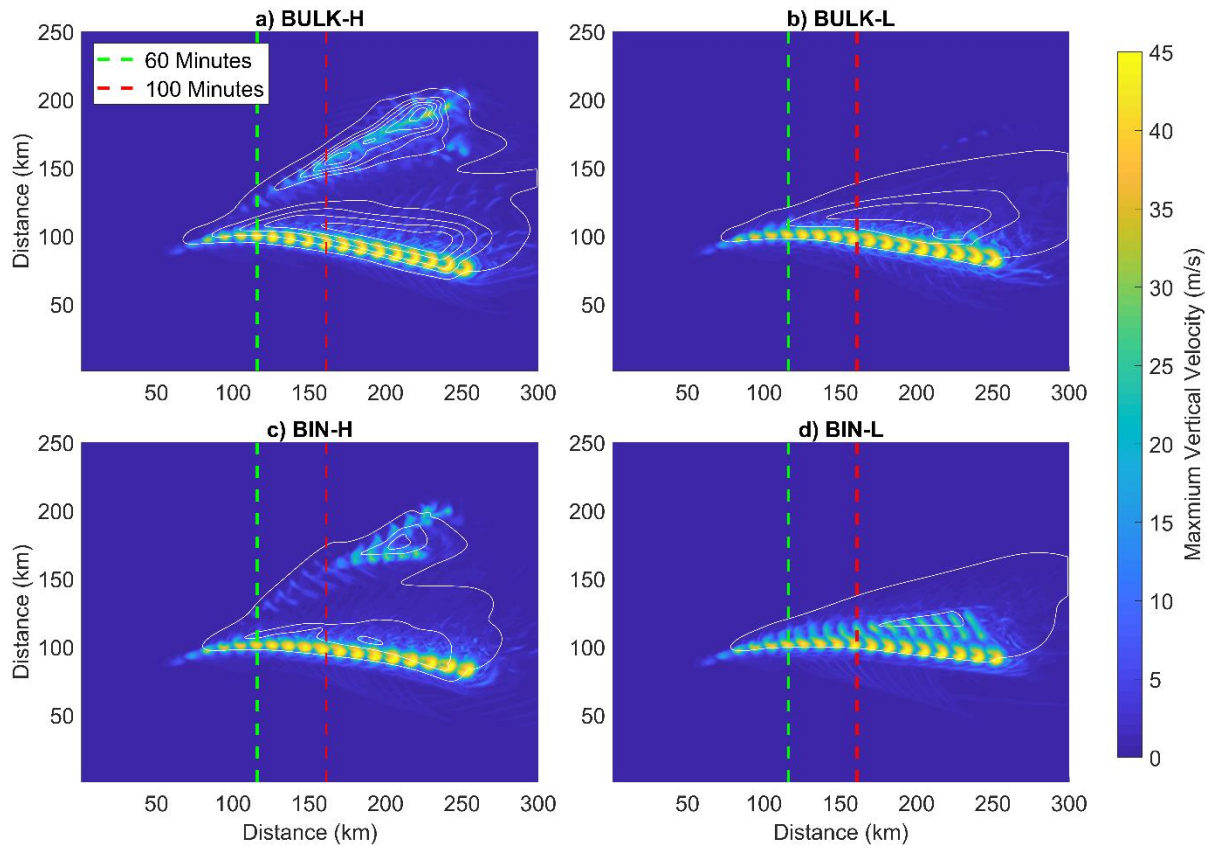
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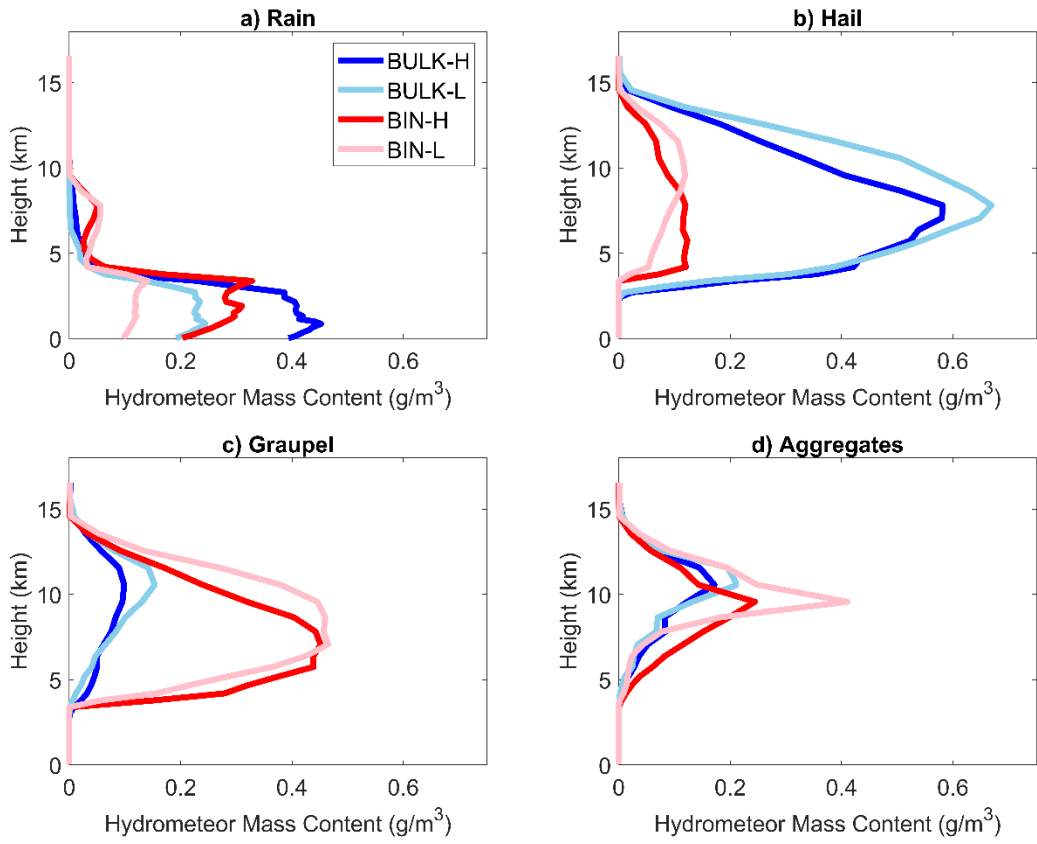
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3 Figure 1: Fall speed-diameter (left) and mass-diameter relationships (right). Blue and red
4 indicate the standard RAMS relationships used in BULK-H and BIN-H, pink indicates the
5 standard SBM relationship used in BIN-L, and light blue indicates the modified relationship used
6 in BULK-L. Diameter-fall speed relationship for (a) aggregates, (c) graupel, and (e) hail.
7 Diameter-mass relationship for (b) aggregates, (d) graupel, and (f) hail.



1

2 Figure 2: Accumulated precipitation (white contours) and time and column maximum vertical
 3 velocity (shaded). Precipitation contours are drawn in white every 3 mm with the first at 0.1 mm.
 4 Vertical velocity data are taken every 2 outputs (10 minutes). The green and red lines indicate
 5 the location of the maximum vertical velocity at 60 minutes and 100 minutes, respectively, for
 6 each storm.

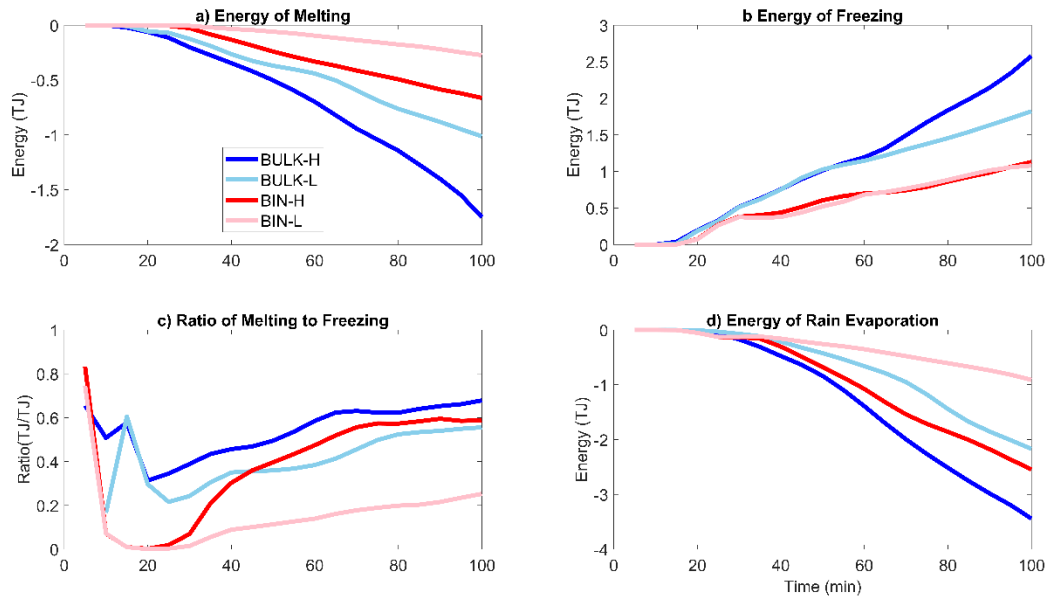


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2 Figure 3: Vertical profiles of average hydrometeor content (g/m³) at 60 minutes, for a) rain, b)

3 hail, c) graupel, and d) aggregates.

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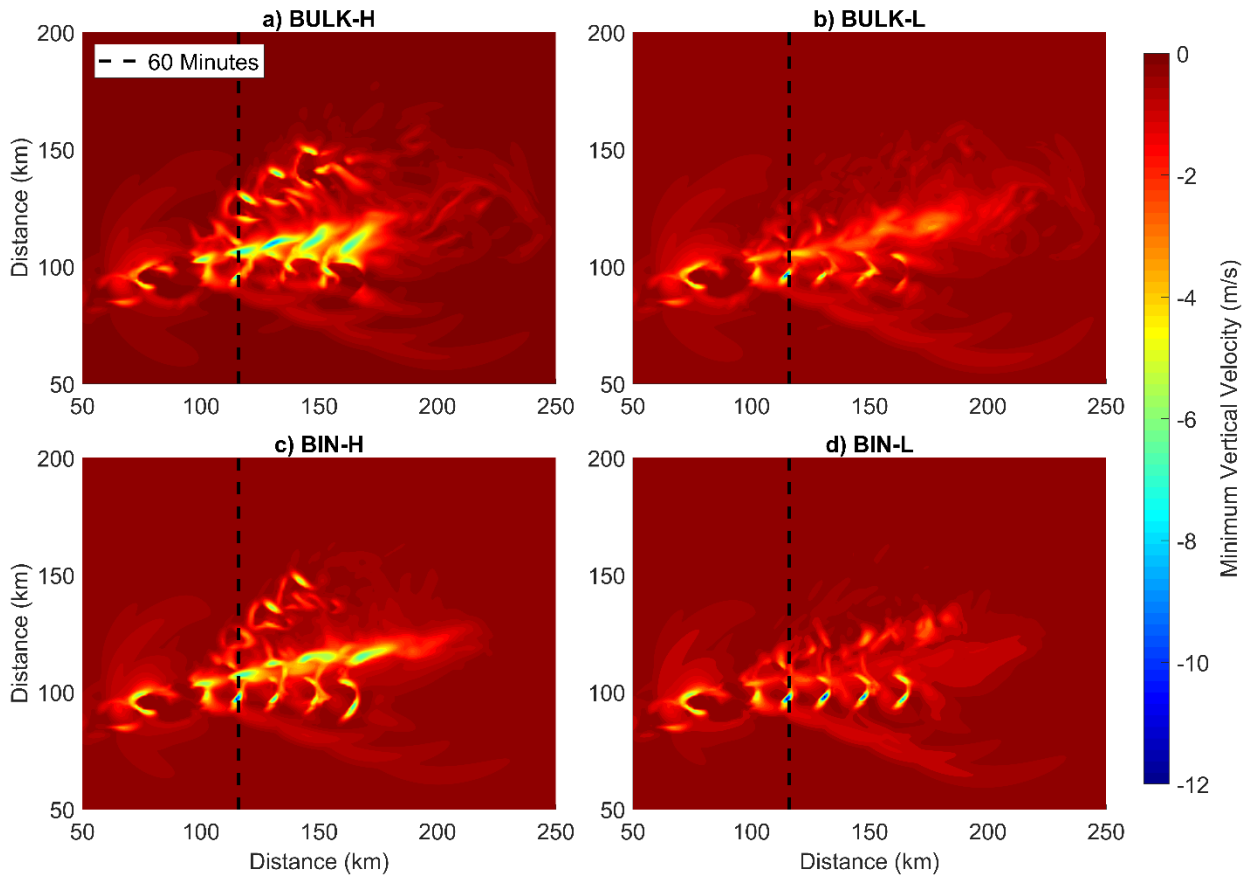


2

3 Figure 4: The latent heating (TJ) associated with a) melting, b) freezing, and d) evaporation. c)

4 The ratio of melting to freezing.

1

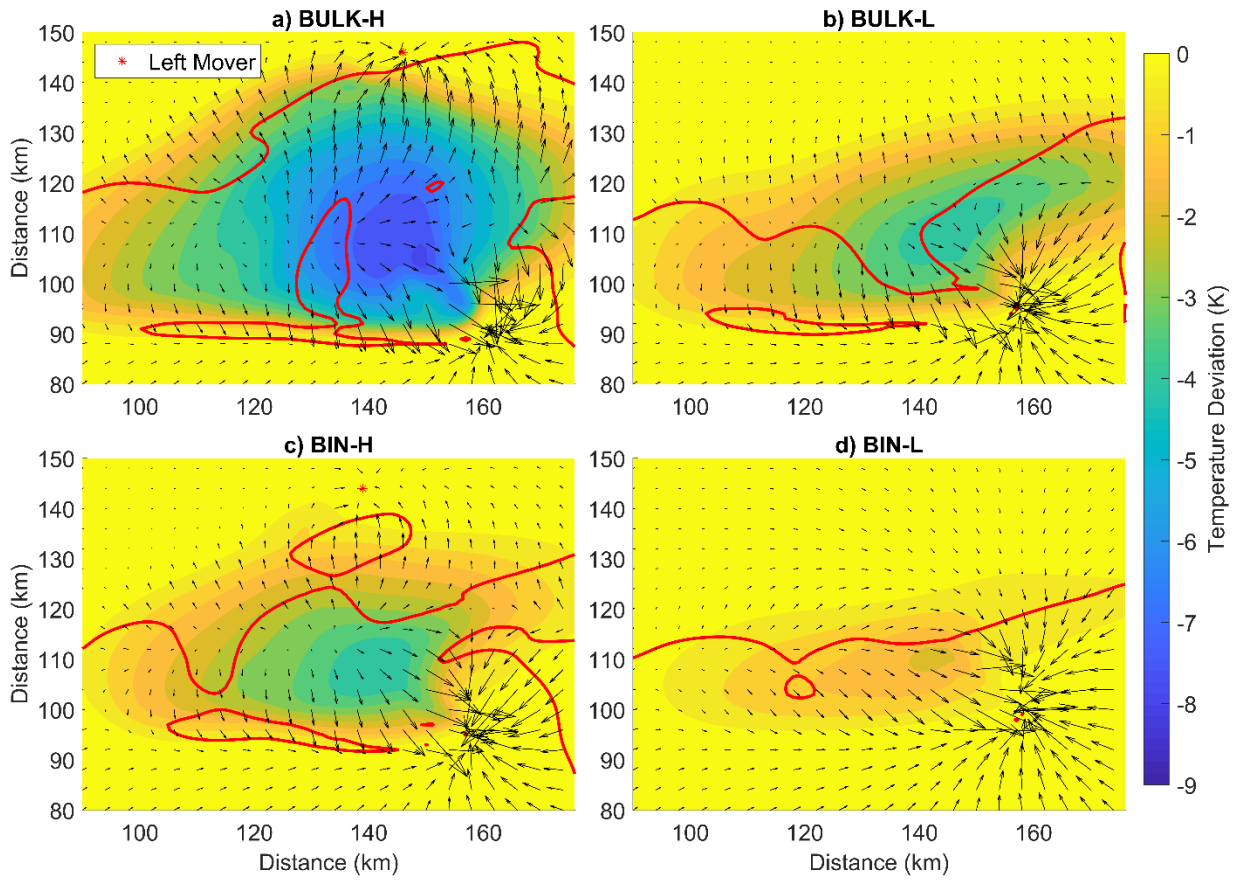


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3 Figure 5: Like Figure 1, but for the strongest downdraft below 5 km. The vertical velocity data is

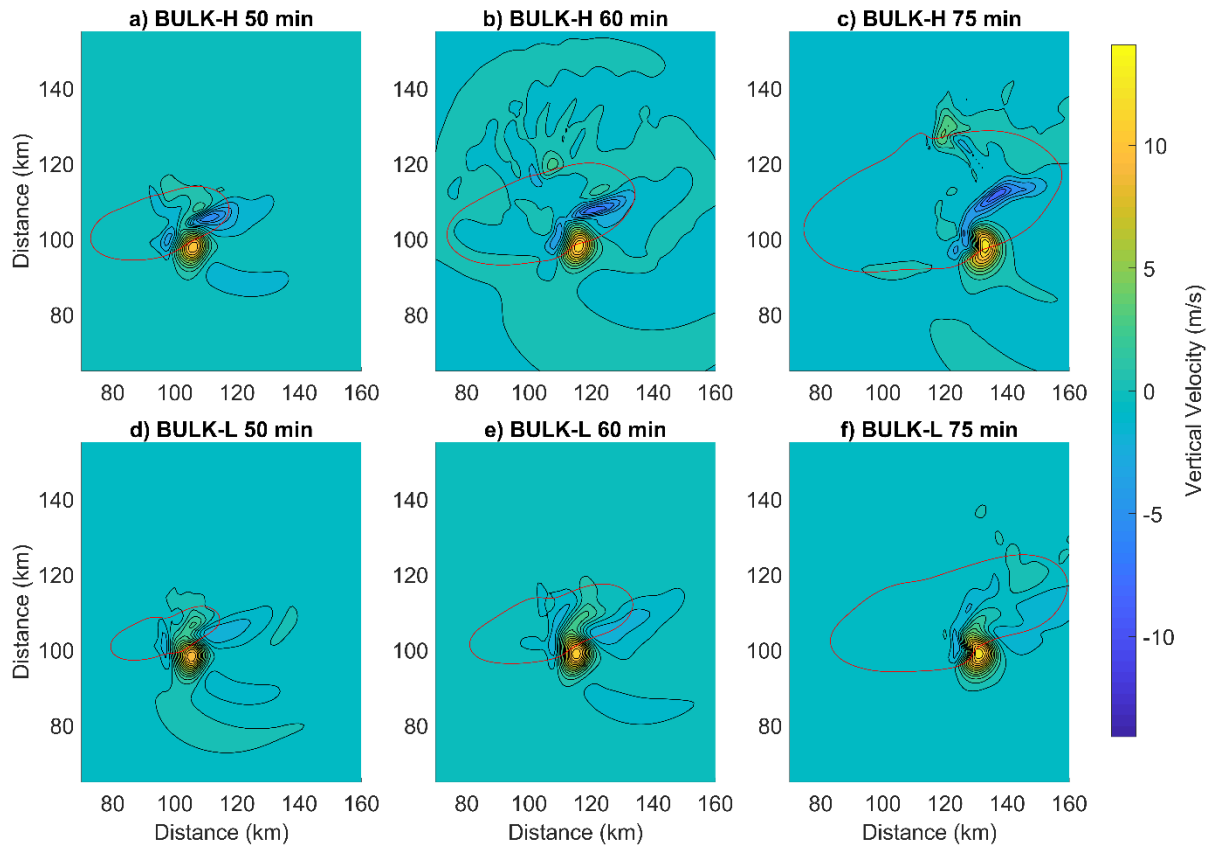
4 taken every 15 minutes and cut off at 100 min.

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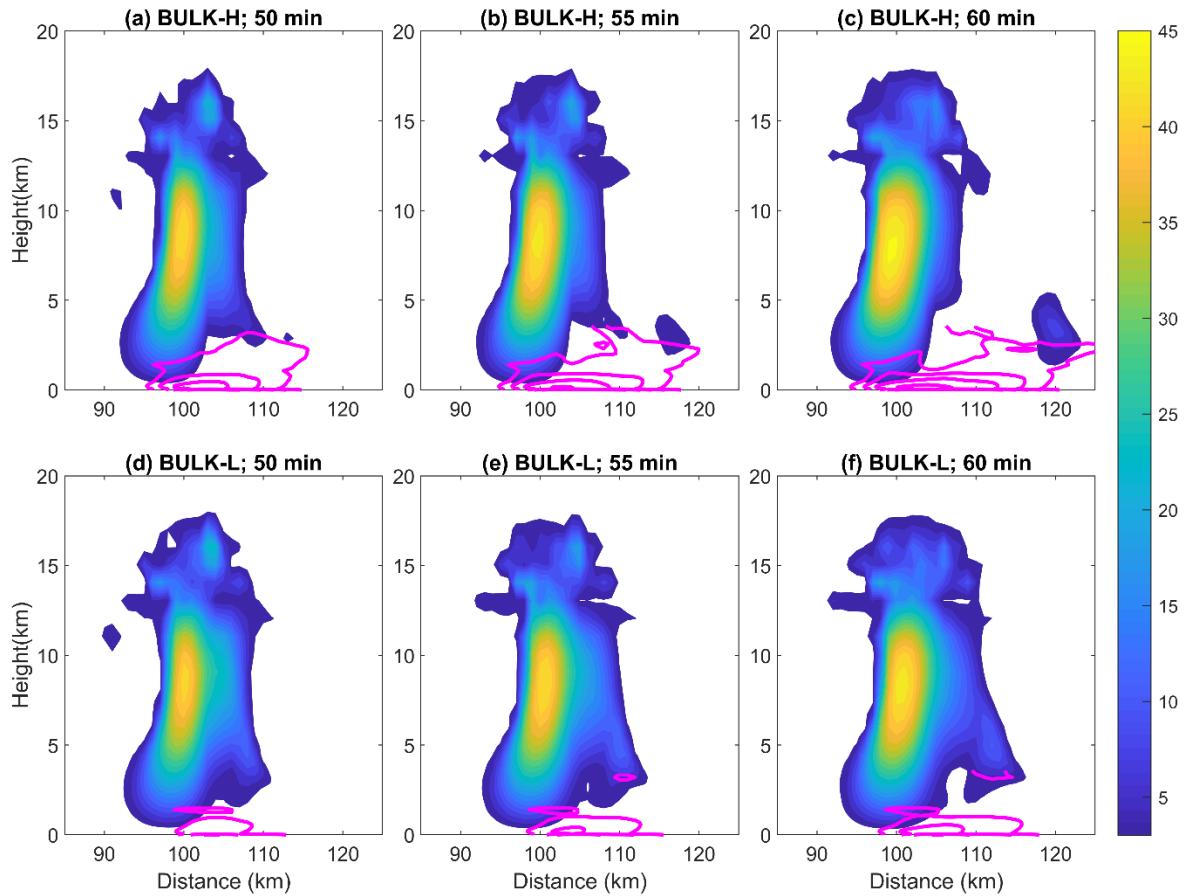
3 Figure 6: Surface air temperature deviation at 100 min with surface winds, left-movers in the
4 high fall speed runs are marked. Red lines show regions where the magnitude of the horizontal
5 vorticity matches or exceeds the the low-level vertical wind shear.



1

2 Figure 7: Vertical velocity at 1.4km and cold pool contour at a deviation of $-0.5K$.

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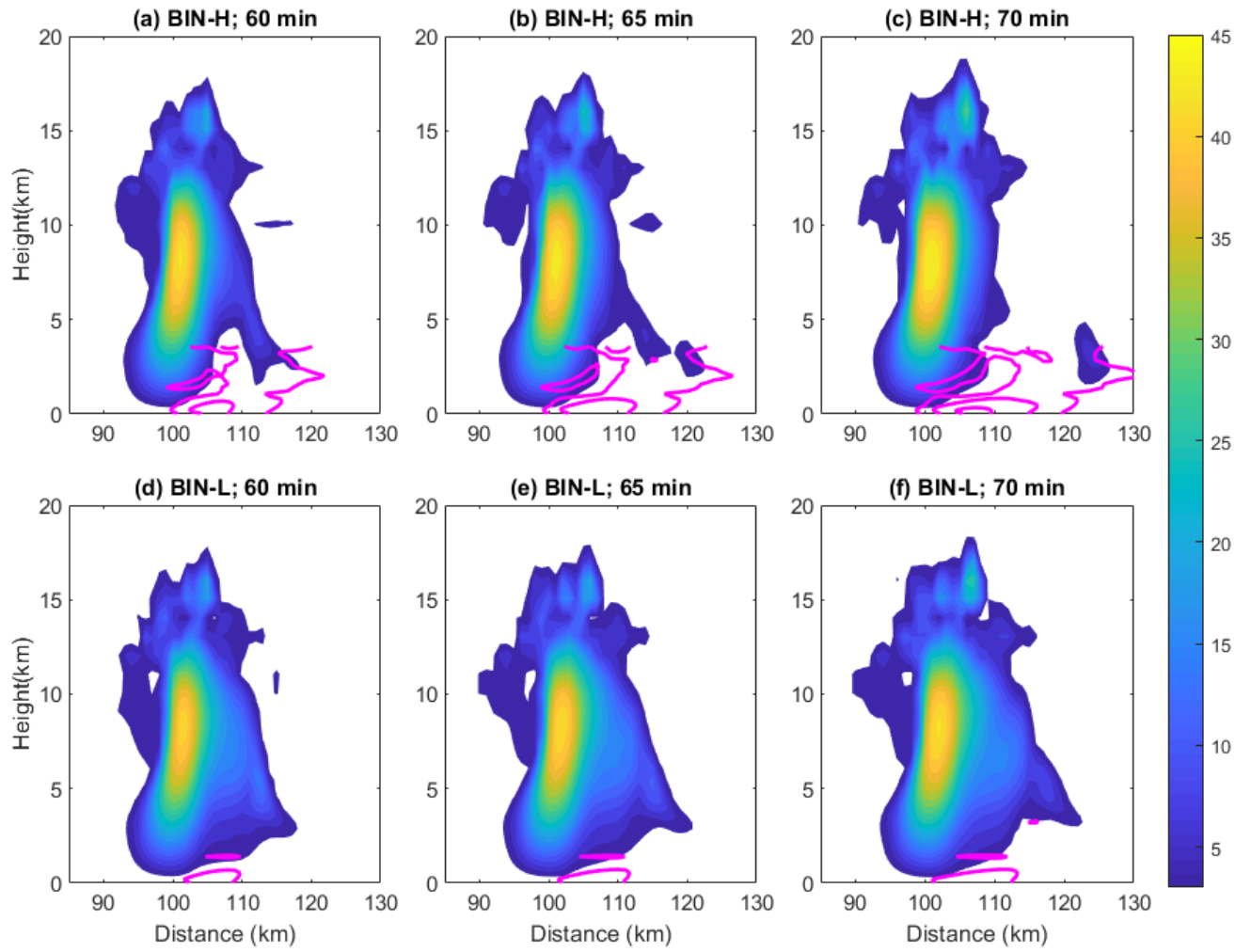


1

2 Figure 8: Height vs meridional distance for the evolution of the zonal maximum vertical wind
 3 speed (colored contours; minimum of 3 m/s) and maximum potential temperature depression
 4 every 1K (magenta contours) in BULK-H and BULK-L.

5

6



1

2 Figure 9: Like Figure 8 but for the BIN-H and BIN-L.