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Continental flood basalts drive Phanerozoic extinctions

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Refinements of the geological timescale driven by the increasing precision and accuracy of radiometric dating have revealed an apparent correlation between large igneous provinces (LIPs) and intervals of Phanerozoic faunal turnover that has been much discussed at a qualitative level. However, the extent to which such correlations are likely to occur by chance has yet to be quantitatively tested, and other kill mechanisms have been suggested for many mass extinctions. Here, we show that the degree of temporal correlation between continental LIPs and faunal turnover in the Phanerozoic is unlikely to occur by chance, suggesting a causal relationship linking extinctions and continental flood basalts. The relationship is stronger for LIPs with higher estimated eruptive rates and for stage boundaries with higher extinction magnitudes. This suggests LIP magma degassing as a primary kill mechanism for mass extinctions and other intervals of faunal turnover, which may be related to CO2, SO2, Cl, and F release. Our results suggest continental LIPs as a major, direct driver of extinctions throughout the Phanerozoic.

volcanology | mass extinctions | carbon cycle | paleontology

Large igneous provinces (LIPs), comprising $>10^5$ km^3 of typically mafic magma with durations generally less than 5 million years, have long been recognized as a distinct category of voluminous volcanic eruptions with the potential to rapidly perturb the environment (1–3). Throughout the Phanerozoic, LIPs have erupted frequently, both on the continents and in the oceans; both U–Pb and Ar–Ar geochronology have been used to date periods of eruptive activity at several such provinces, revealing most of their immense volume to have been emplaced over submillion-year time periods (4).

Fed by extensive dike systems and composed of volatile-bearing volcanics brought up from the deep mantle in plumes, LIPs—especially in their mafic guise as flood basalt provinces—have had pronounced effects on the climate history of Earth (5). Ocean anoxic events (OAEs), global warming (via $CO₂$ and CH₄), and short-term global cooling (via SO2) have all been attributed to flood basalt volcanism (3, 6, 7). Most remarkably, at least four of the five mass extinctions of the Phanerozoic coincide with flood basalt eruptions: the Deccan Traps at the Cretaceous–Paleogene (K-Pg) boundary, the Central Atlantic Magmatic Province (CAMP) at the Triassic–Jurassic boundary, the Siberian Traps at the Permo–Triassic boundary, and, more tentatively, the Viluy or Kola-Dnieper Provinces at the Frasnian–Famennian boundary (3, 5, 6); mafic magmatism is speculated at the Late Ordovician mass extinction as well, but has not been correlated with a specific known LIP (3, 8). Other drivers have been proposed for each of these mass extinctions, most notably the Chicxulub impactor at the K-Pg (9), but the temporal coincidence with flood basalts has, nonetheless, attracted significant interest (5) and is difficult to dismiss. Notably, in each such case, the corresponding flood basalt province is continental, not oceanic, despite the fact that many of the largest LIPs (including the two largest known to date) are oceanic (10).

Quantitative Assessment of Temporal Coincidence

While a temporal correlation between flood basalt eruption and biotic extinction has long been qualitatively noted (5), the likelihood of such a correlation arising by chance has not been rigorously quantified across the Phanerozoic. To address this problem, we draw upon three well-established, independent datasets: the Geologic Time Scale (11), the Paleobiology Database (12) [which is the basis for the extinction magnitude calculations of Muscente et al. (13) that are used in this analysis], and the LIP compilation of Ernst (14), updated with the most accurate available geochronology (*Materials and Methods*; *[SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)*[, Tables S1 and S2\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental).

While the number of Phanerozoic LIPs identified by Ernst (14) far exceeds the number of catastrophic mass extinctions, not every episode of faunal turnover rises to the magnitude of a mass extinction. Indeed, each Phanerozoic stage and period boundary of the Geologic Time Scale is defined on the basis of faunal assemblages, so each of these biological transitions can be taken as a time of potential faunal turnover, much of

Significance

Although the causes of the five largest mass extinctions remain controversial, geochronological improvements have revealed an apparent correlation between large igneous provinces (LIPs) and periods of Phanerozoic faunal turnover. This paper establishes that this relationship is unlikely to occur by chance and defines an eruptive rate threshold, above which known continental LIPs correlate with large extinctions. Continental LIPs also have an approximately linear relationship between their eruptive rate and extinction magnitude. It is difficult to attribute the causality of any one extreme event like an extinction with certainty, but there is an overall correlation between continental LIPs and extinction events that warrants consideration.

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which occurs as a short pulse near the end of the stage (11, 15, 16). We quantify the intensity of this faunal turnover using stage-level genus mean extinction rates calculated by Muscente et al. (13) with data from the Paleobiology Database (12) on August 12, 2018. Although LIPs have been identified throughout Earth history from the Archean to the Phanerozoic (3), the biostratigraphic record underlying our estimates of the timing and severity of faunal turnover is largely restricted to the Phanerozoic, after the advent of skeletal biomineralization. Consequently, our quantitative analysis focuses on the Phanerozoic.

Consider a single geological timescale boundary i occurring at a time constrained by a Riemann-integrable likelihood function $p_i(t)$, reflecting, for example, the Gaussian analytical uncertainty derived from radioisotopic dating of a boundary horizon. For a given time t and time increment δt , the likelihood that boundary *i* occurs between t and $t + \delta t$ is, by definition,

$$
\int_{t}^{t+\delta t} p_i(t)dt = P_i|_{t}^{t+\delta t}, \qquad [1]
$$

where P_i is the cumulative distribution function corresponding to the likelihood function p_i . We may likewise consider an analogous likelihood function $q_i(t)$ for the eruption of a given increment of magma from LIP j (c.f. *[SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)*[, Fig. S1\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental). The likelihood of the two events coinciding within a given interval t to $t + \delta t$ if the distributions p and q are independent is then

$$
\xi_{ij}|_{t}^{t+\delta t} = P_i|_{t}^{t+\delta t} Q_j|_{t}^{t+\delta t}, \tag{2}
$$

with a total coincidence likelihood summed over all such intervals within the Phanerozoic (zero through t_{Phan} , with intervals spaced apart by δt), then given by a sum of the form

$$
\xi_{ij} = \sum_{n=1}^{t_{Phan}/\delta t} P_i|_{t_n}^{t_n + \delta t} Q_j|_{t_n}^{t_n + \delta t}.
$$
 [3]

Considering that $P_i|_t^{t+\delta t}$ is equal to the mean of p over the interval from t to $t + \delta t$ times δt —that is, $\bar{p}_i \vert_t^{t + \delta t} \delta t$ —we may rewrite Eq. **3** as

$$
\xi_{ij} = \sum_{n=1}^{t_{Phan}} \bar{p}_i \vert_{t_n}^{t_n + \delta t} \delta t \bar{q}_j \vert_{t_n}^{t_n + \delta t} \delta t, \tag{4}
$$

which, however, depends on the choice of δt . To remove this dependency, we consider, instead, a scale-invariant version of this coincidence sum

$$
\psi_{ij} = \frac{\xi_{ij}}{\delta t} = \sum_{n=1}^{t_{Plan}/\delta t} \bar{p}_i \vert_{t_n}^{t_n + \delta t} \bar{q}_j \vert_{t_n}^{t_n + \delta t} \delta t, \tag{5}
$$

which in the limit as δt approaches zero is simply

$$
\psi_{ij} = \int_0^{t_{Phan}} p_i(t) q_j(t) dt, \tag{6}
$$

that is, a middle Riemann sum representing the integral of the product of the two likelihood functions p_i and q_j . We define this quantity as the coincidence product ψ_{ij} for boundary i and LIP j. Considering pair-wise all such boundaries i and LIPs j , the total coincidence product for the Phanerozoic (Ψ) is then

$$
\Psi = \sum_{i} \sum_{j} \psi_{ij}.
$$
 [7]

Using Ψ , then, as our metric, we conduct a simple Monte Carlo simulation, generating \sim 10⁸ lists of uniformly distributed stage boundaries and calculating the coincidence products for each (*Materials and Methods*). As shown in Fig. 1, we find that the level of correlation between LIPs and periods of faunal turnover far exceeds that expected by chance—occurring in 1 out of every 1.64×10^4 simulations ($P = 6.09 \times 10^{-5}$) when considering the entire list of Phanerozoic stage boundaries and in 1 out of 2.06×10^6 simulations ($P = 4.84 \times 10^{-7}$) when considering only the stage boundaries associated with the five mass extinctions of the Phanerozoic. This relationship is also robust across various subsets of the dataset. The Siberian Traps is the largest LIP associated with a significant extinction, but when it is excluded, the observed correlation between LIPs and stage boundaries still occurs in only 1 out of every 1.60×10^4 simulations ($P = 6.24 \times$ 10^{-5}). Even when the five LIPs that are temporally correlated with mass extinctions (the Deccan Traps, CAMP, Siberian Traps, Viluy Province, and Kola-Dnieper Province) are not considered, the observed correlation between LIPs and stage boundary faunal turnover occurs in 1 out of every 9.84×10^3 simulations (P = 1.02×10^{-4}). These correlations are still greater than chance and are similar to the P value calculated when the entire LIP list is considered. This demonstrates that the observed correlation between LIPs and Phanerozoic faunal turnover extends to smaller extinction intervals and is not just driven by the mass extinctions. Consistent with previous expectations (5, 17–21), these results provide a first test and confirmation of the statistical significance of the temporal coincidence between LIPs and extinctions.

By contrast, although several Phanerozoic extinctions have been attributed to large bolide impacts (18, 22), we do not find a similarly robust coincidence between Phanerozoic stage boundaries or mass extinction boundaries and the radiometric ages of large impact craters. Fig. 1 shows that much of the correlation observed between impacts and stage-level extinction is driven by the precise coincidence between the Chicxulub impactor and the K-Pg mass extinction, with all P values becoming statistically insignificant when it is not considered. *[SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)*[, Fig. S2](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) shows the same effect when impacts are compared only to mass extinction boundaries. While the coincidence products may change with improved geochronology, Fig. 1 demonstrates that LIP eruptions have a correlation with Phanerozoic extinctions that is more robust than the correlation with bolide impacts and extinctions. These results are consistent with a broader trend that becomes clear from a historical examination of the literature; specifically, apart from the K-Pg, proposed impact-extinction correlations have not been confirmed by improved geochronology. For example, while the ∼100-km Manicouagan structure (23) was once considered as a potential cause for the end-Triassic mass extinction, it was later found to predate that extinction by about 14 million y (Myr) (24, 25). Where Manicouagan ejecta have been directly located in stratigraphic section, no significant extinction seems to be associated with the event (25, 26). Similarly, the Chesapeake and Popigai impactors—the two largest of the Cenozoic—resolvably predate the Eocene–Oligocene boundary and do not appear to be associated with significant levels of faunal turnover (27–29). Correlation of the Woodleigh impact structure with the Late Devonian mass extinction is likewise doubtful (30). It has been shown previously that smaller impacts, especially those much smaller than Chicxulub, are not associated with significant extinction (22, 31). This analysis supports and extends that conclusion by showing no significant correlation between extinction and impacts smaller than Chicxulub (Fig. 1 and *[SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)*[,](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) [Fig. S2\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental).

Fig. 1. Relationship between observed and expected coincidence products. The observed coincidence products between all Phanerozoic stage boundaries and all LIPs (*A*), along with the corresponding stochastic distribution (which would result if timescale boundaries were spread randomly throughout the Phanerozoic following a uniform distribution), on a logarithmic *y* scale. The probability that a uniform distribution has a higher coincidence product (Ψ) than observed is given by *P*. Note that the vertical log scale visually distorts the shaded regions compared to the reported *P* values. *B* shows the results for LIPs and stage boundaries when the Siberian Traps, the largest LIP correlated with a severe extinction, is excluded. *C* shows the coincidence products for all LIPs and the five Phanerozoic mass-extinction boundaries. *D* shows the results for LIPs and stage boundaries when all of the LIPs temporally correlated with mass extinctions (the Deccan Traps, CAMP, Siberian Traps, Viluy Province, and Kola-Dnieper Province) are not considered. Even without these large, well-correlated events in *B* and *D*, there is still a statistically significant relationship between LIPs and Phanerozoic extinctions. Below are the observed coincidence products and stochastic distributions of coincidence products for impact events. *E* shows the results for all impacts with diameters ≥20 km and all stage boundaries, while *F* excludes the Chicxulub impactor from consideration. Likewise, *G* shows the results for all impacts with diameters ≥40 km and all stage boundaries, while *H* excludes the Chicxulub. The coincidences between impact events and Phanerozoic extinction are statistically significant only when the Chicxulub impactor is included, indicating that its precise coincidence with the K-Pg mass extinction is primarily responsible for the observed relationship between impacts and extinctions. Since the LIP-extinction coincidence is robust to similar exclusions, this supports a significant temporal relationship between LIPs and faunal turnover over the Phanerozoic.

Fig. 2. Greater normalized total coincidence products for larger extinctions and faster bulk eruptive rates. Observed LIP-extinction coincidence products for subsets of the record with varying of extinction severities and eruptive rates, normalized by the maximum possible coincidence product for a subset of that size. The subsets include all LIPs with bulk eruptive rates greater than or equal to the stated bulk eruption rate and all stages with extinction magnitudes greater than or equal to the stated extinction magnitude. For subsets of the record that include only the most severe extinctions and LIPs with the greatest eruptive rates, the observed coincidence products approach the maximum possible values. *[SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)*[, Fig. S3](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) shows the same trend when the subsets are nonoverlapping.

If the observed degree of temporal coincidence between LIPs and faunal turnover is the result of a causal relationship, we might further expect larger LIPs to be associated with more severe extinctions. However, such a correlation, at least with absolute LIP magnitude, has not been previously observed (3, 32–34). Rather than absolute magnitude, we consider, instead, "bulk eruption rate"—that is, the total volume of the LIP divided by the best estimate of the total duration over which that bulk was erupted as a critical parameter. While more difficult to quantify than magnitudes alone (requiring accurate geochronology), rates are highly relevant in the context of critical thresholds in climatic and ecological stability (35–37).

To explore this parameter space, Fig. 2 examines the relative strength of temporal correlation between LIPs and extinctions as a function of both extinction severity and bulk eruption rate. In particular, the maximum possible Ψ diminishes as the number of considered LIPs decreases with an increasingly strict minimum size or rate threshold. This is because the total time span considered remains the same, while the number of events that can potentially be correlated decreases, limiting the observed coincidence product for a given list of events. With that in mind, we consider here as a metric the normalized total coincidence product

$$
\Psi_{norm} = \Psi / \Psi_{max},
$$
 [8]

that is, the Ψ for a given set of LIPs and boundaries divided by the maximum possible Ψ of a set of equivalent size, where each LIP coincides perfectly with an extinction. At higher extinction percents and higher bulk eruptive rates, the normalized total coincidence products increase (Fig. 2). The relatively low value of the highest normalized total coincidence product (0.11; Fig. 2) will likely increase with improved geochronology, but it does indicate that there may be other important factors to consider in the relationship between LIPs and extinctions.

Bulk Eruptive Rate and Extinction Severity

To better understand the potential role of threshold processes in the relationship between bulk eruption rate, LIP duration, and extinction severity, we consider as an analog the framework of

Rothman (37), who determined the critical stability threshold of the Phanerozoic biogeochemical carbon cycle in terms of the nondimensionalized mass and nondimensionalized duration of a given carbon-cycle perturbation. While Rothman determined nondimensionalized masses and durations on the basis of carbon isotope excursions, we can easily adapt this framework to our purposes by considering, instead, the equivalently nondimensionalized durations and CO_2 emissions of LIPs, with a degassed CO_2 concentration of 0.5 wt.% (38). This is taken as a reasonable value for $CO₂$ in basaltic magmas after ref. 38, but a range from 0.2 to 0.5 wt.% is also considered (39) to reflect the potential variation in $CO₂$ concentrations between different magmatic systems. Most of that $CO₂$ should be degassed by continental eruptions, but these values may overestimate $CO₂$ degassing from oceanic LIPs, as more volatiles are trapped in the rocks when they erupt under higher oceanic pressures (40–42). We also adapt this framework to reflect the potentially intermittent nature of volcanic activity by including critical thresholds that represent carbon emissions over 100%, 10%, 1%, and 0.1% of the total event duration (*Materials and Methods*). Hereafter, we consider only nonoceanic LIPs, also known as continental flood basalts (CFBs), because of the potential for decreased volatile degassing in submarine eruptions (40–42) and the lack of preservation of older oceanic LIPs that makes it difficult to extrapolate their effects across the Phanerozoic.

Among all CFBs with quantifiable bulk eruption rates, only the three that coincide with major mass extinction events (the Siberian Traps with the Permo–Triassic, the CAMP with the Triassic– Jurassic, and the Deccan Traps with the K-Pg) plot near the 100% threshold (Fig. 3).These all plot above the 10% critical rate threshold, along with the Emeishan Province that is associated with the severe Capitanian–Wuchiapingian extinction in the late Permian. Other CFBs near, but below, the 10% threshold are generally associated with lesser extinction events or other environmental perturbations like OAEs. For instance, the second phase of the North Atlantic Igneous Province is potentially correlable with the Paleocene–Eocene Thermal Maximum and associated ocean anoxia, while the Parana–Etendeka has been potentially associated with the Valanginian Weissert OAE—yet both lack significant extinctions (5, 13, 43). The lack of extinction associated with these volumetrically significant CFBs has long been a point of discussion (44–46); if the currently available duration (and thus rate) estimates for these CFBs are accurate, this may be simply a consequence of falling below the critical rate threshold, meaning that the rates of $CO₂$ degassing fall within the range that can be effectively buffered by silicate weathering. However, future geochronological constraints may well result in a reevaluation of these bulk eruption rates. The Karoo–Ferrar LIP, which has been linked to the early Toarcian OAE and a biotic crisis, also falls just below the critical line, but is already well-dated enough that it is unlikely to move above that threshold (47).

The apparent pattern of higher extinction magnitudes for CFBs that plot farther above the 10% critical threshold line in Fig. 3 (i.e., those with greater estimated $CO₂$ degassing rates) suggests the further possibility of a direct correlation between bulk eruption rate and extinction severity. While CFBs with larger bulk eruptive rates do not necessarily have the largest total volumes in the LIP record, they appear to have much stronger environmental impacts than those that slowly erupt for long periods of time. Indeed, plotting bulk eruption rate vs. extinction magnitude of the largest coincident stage boundary for each of the well-dated $(1\sigma < 0.2\%)$ CFBs of the Phanerozoic reveals a clear, positive, approximately linear correlation (Fig. 4)—in sharp contrast to the relative lack of correlation observed when considering total

Fig. 3. Carbon-cycle perturbations from continental LIPs relative to the Rothman threshold. The dimensionless carbon-cycle mass perturbation *M* = |Δ*m*|/*m*∗, plotted as a function of the dimensionless timescale of CFB eruption, adapting the nondimensionalizations from Rothman's (37) analysis of C-isotope excursions to flood basalt eruptions. Log scale. The leftmost diagonal (identity) line represents the critical rate threshold $\phi = \phi_c = 0.23 \pm 0.07 = f_{org}$ representing the maximum possible normalized flux perturbation in a mass-conserving carbon cycle with no anomalous sinks or sources (37). The other diagonal lines apply that same critical threshold to events with intermittent CO₂ degassing, occurring over 10%, 1%, or 0.1% of the total duration. CFB provinces that are associated with the highest extinction magnitudes fall closest to the continuous degassing (100%) critical rate threshold and plot above the 10% critical rate threshold. Horizontal error bars reflect 1σ age and φ*^c* uncertainties with the exception of the lower uncertainty on the Emeishan Province, which assumes a minimum τ*env* of 10,000 y because the small nominal duration of that province causes its start and end age uncertainties to overlap. The 0.5 wt.% CO₂ is taken as a reasonable value after ref. 38, but vertical error bars show the range of potential dimensionless CO₂ masses that result from varying the estimated wt.% CO₂ in the erupted material from 0.2 to 0.5 (39) (*Materials and Methods*).

LIP volume alone (3, 32–34). Moreover, the strength of this correlation appears to suggest that extinction is driven in large part by factors directly proportional to bulk magma eruption or intrusion rate, and not exclusively dependent on, e.g., the type of basin in which a CFB is emplaced or the configuration of the Earth system at the time of the eruption.

While these results suggest a correlation between total magma volume and total volatile emissions at the scale of an entire CFB, such a relationship should not be assumed to hold on smaller spatial scales or timescales. On the contrary, eruption rates per se should not be assumed to act as a proxy for volatile fluxes, especially for volatiles such as $CO₂$ with high vapor pressures under magmatic conditions. For example, the apparent highest CO_2 flux (and coincident global warming) associated with the Deccan Traps began ∼300 kilo annum (ka) prior to the K-Pg extinction and does not coincide with the highest volumetric eruption rate of Deccan lavas (48, 49). This appears to be a case of pre-eruptive (or passive) degassing (50). The decline of this warming began about 150 ka before the extinction and may have been driven by weathering-related drawdown of $CO₂$ and/or sulfate aerosol emission from Deccan magmas (48). A decoupling of C vs. S degassing is consistent with exsolution of $CO₂$ preceding that of S, in keeping with the higher solubility of S than of $CO₂$ in basaltic magmas (41). Indeed, $CO₂/SO₂$ fractionation as a function of degassing pressure is sufficiently systematic to be used as a measure of magmatic degassing depth (42).

Burgess et al. (51) inferred that, in the case of the Siberian Traps, the main trigger for the Permo–Triassic extinction was CO_2 sourced from supracrustal coal beds that were thermally degassed by interaction with basaltic sills. This represents another version of pre-eruptive degassing, in that erupted magma itself was not the dominant source of $CO₂$. This differs from the Deccan situation

in that there are abundant coal beds known in the substrate of the Siberian Traps, but not in the Deccan. These abundant sedimentary basin-sourced volatiles degassed by the Siberian Traps may explain why the extinction magnitude at the Permo–Triassic boundary is far above the other extinctions, though it is consistent with the trendline that includes CFBs without such nonmagma volatile inputs (Fig. 4). Conversely, the volatile-poor eolian sandstones intruded by the Parana–Etendeka may help explain that eruption's limited environmental impact (44). However, although peaks may be decoupled at finer timescales, high eruption rates must reflect high heat and mass flux on the scale of the lifetime of a CFB. Sedimentary basin-sourced volatiles and pre-eruptive degassing likely play a role in the extinction severity associated with a CFB (32), but does not invalidate the implication that high bulk eruptive rates are mechanistically correlated with significant climate perturbations, leading to major extinctions. Magmatic degassing is the main factor in the critical rate threshold (Fig. 3) and eruptive rate-extinction magnitude linear correlation (Fig. 4) shown here, but passive degassing can explain some of the variability within those relationships.

The critical rate threshold (Fig. 3) and the approximately linear correlation between extinction severity and bulk eruptive rate (Fig. 4) also suggest that any changes in ecosystem stability over time (e.g., refs. 52–54) have not been able to sufficiently buffer the rapid perturbations caused by high-rate CFBs. The lack of severe extinctions associated with more recent eruptive events, such as the North Atlantic Igneous Province, can, given the currently available data, be explained by lower total eruptive rates without invoking a state change in ecosystem stability. Background climate state and boundary conditions, like the $CaCO₃$ saturation state of the oceans (55–57), changes in surface oxygenation (58), and the presence of volatile-rich sediments in the magma ascent path (59),

Fig. 4. CFB bulk eruptive rates correlate strongly with extinction severity. Observed correlation between CFBs and extinction magnitudes results in a logistic regression (r² = 0.653; black) of the form $y = \frac{100 - y_0}{1 + e^{-k(\ln(x) - x_0)}} + y_0$ with parameters $k = 2.42 \pm 1.24$ (1 SE), $x_0 = 0.92 \pm 0.17$ (1 SE), and $y_0 =$ 17.71 \pm 3.66 (1 SE). For the most precisely dated CFBs (1 σ < 0.2%; red), the observed correlation with extinction magnitude can be approximated by a regression line ($r^2 = 0.956$; red) with a slope of 21.02 \pm 2.02%/(km³/y) (1 SE; $t = 10.40$; $P = 0.0001$) and an intercept of 4.73 \pm 3.68% (1 SE; $t =$ 1.29; $P = 0.2555$). The intercepts for both regressions are nonzero, potentially because a background extinction rate was not subtracted from the extinction magnitudes from Muscente et al. (13) used here. Moderately well-dated CFBs (1 σ < 0.85%; purple) broadly fall along the same linear trend: 1, Parana-Etendeka; 2, North Atlantic Igneous Province; 3, Tarim; 4, Qiangtang; 5, Jutland (Skagerrak); and 6, Chon Aike. All other less precisely dated flood basalts are shown in blue. To avoid plotting duplicate correlations for less precisely dated CFBs, boundaries that are correlated with multiple eruptions are shown only with the highest eruptive rate CFB, and CFBs that overlap multiple boundaries are plotted only with the highest extinction magnitude stage. Horizontal error bars reflect 1 σ age uncertainties only, not those related to volume. The Emeishan, Parana-Etendeka, and Jutland (Skagerrak) provinces have sufficiently large duration uncertainties to produce negative eruption rates, so those 1 σ errors are, instead, set equal to the bulk eruptive rate. Vertical error bars show the range from minimum to maximum extinction magnitude at a given boundary, as calculated by Muscente et al. (13). Exact values are given in *[SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)*[, Table S5.](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)

may be important for determining climate impacts and play a role in the decreased severity of recent extinctions. However, according to our analyses, these factors are subordinate to eruption rate in determining the severity of extinctions across the Phanerozoic.

Degassing Efficiency and Submarine Flood Basalts

While the results thus far appear to explain most of the Phanerozoic mass extinctions, one major discrepancy remains: In contrast to the CFBs, oceanic LIPs do not appear to obey the same relationship between bulk eruption rate and extinction severity. In particular, the two largest known oceanic flood basalts (the Ontong Java and Kerguelen plateaus), while long in duration, have total volumes well above those of any CFB (10). If oceanic flood basalts fully degassed the same concentration of $CO₂$ as their continental counterparts, these LIPs would fall well above the 10% critical rate threshold (*[SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)*[, Fig. S4\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental), yet are associated with lower extinction magnitudes than CFB eruptions above that threshold. The Ontong Java Province even falls above the 100% critical threshold.

One possible solution may involve decreased magma degassing in submarine eruptions due to the confining pressure of the overlying water column. Firstly, for volatiles with low vapor pressures (including SO_2 , Cl, and F), exsolution is a strong function of pressure (41, 42). Gaillard et al. (40), for example, estimate six times less SO_2 exsolution for magmas erupted under 1 km of

water than at the surface on the basis of the increased solubility of such volatiles in basaltic magmas at increased pressure. Secondly, even once exsolved, volatiles may be physically trapped in vesicles or other pore spaces. Such physical trapping would affect all volcanic volatiles, including CO_2 , and is dramatically enhanced in submarine magmas to the extent that some seafloor basalts spontaneously and violently fracture upon being raised to the surface due to the expansion of the compressed $CO₂$ and other volatiles trapped in their vesicles (60).

In this context, it is perhaps notable that for the only attested historical analog of a CFB eruption, the ∼12-km³ 1783 to 1784 Laki fissure eruption in Iceland, records suggest marked direct mortality of humans (with an estimated total of some ∼50,000 across Europe), livestock, and both inland and coastal fish populations—all mediated by SO_2 , Cl, and F, not CO_2 (61–63). The ecological effects of these volatiles during the Laki eruption may not directly parallel those expected to produce the genus-level marine extinctions analyzed here. Still, it is a notable, if small-scale, example of how such basaltic eruptions can negatively impact the biosphere, including coastal marine populations. The most significant climatic consequence of these eruptions, an estimated ∼1 ◦C decrease in Northern Hemisphere surface temperatures for 2 to 3 y following the eruption, is likewise a result of stratospheric sulfate aerosols (62) and independent of $CO₂$. Nonetheless, we also observe that the CO_2 -based critical rate threshold (Fig. 3) does effectively separate CFBs associated with mass extinctions from those with more limited environmental impacts—perhaps an unlikely coincidence if CO_2 itself played no significant role in extinction severity. In this context, the potential temporal decoupling of volcanic $CO2₂$ and $SO₂$ emissions by preeruptive degassing may be especially significant. Given the greater volatility and expected earlier (and deeper-sourced) pre-eruptive degassing of $CO₂$, we may expect a general pattern that features early pre-eruptive (CO_2 -driven) warming to precede rapid syneruptive (SO_2 -driven) cooling, followed by a return to warming as the emission of SO_2 aerosols declines. Such climate whiplash may be expected to accentuate biotic stress, where the ecosystem must adapt to rapidly reversing warming and cooling trends, along with both CO_{2} - and SO_{2} -driven acidification (64).

CFBs Drive Phanerozoic Extinctions

Our results suggest that CFB eruptions are a primary known driver of mass extinction throughout the Phanerozoic. While it is difficult to assign the causal mechanism for any single given extinction on the basis of temporal correlation alone, the broader correlation across the Phanerozoic between faunal turnover and LIP eruptions (Fig. 1) is extraordinarily unlikely to have arisen by chance. In the absence of an external mechanism producing the mantle plumes that drive LIP eruptions and the environmental perturbations that lead to extinctions, this degree of correlation suggests a causal relationship. This observed correlation between extinction and LIP eruption pertains both for mass extinctions and lesser faunal turnovers, but is strongest for boundaries with higher rates of extinction and flood basalts with higher bulk eruptive rates (Fig. 2).

The additional correlation we observe between bulk eruption rate and extinction magnitude (Fig. 4) allows a direct estimation of the degree to which the severity of a given mass extinction corresponds with the expected severity if the extinction were caused by a coincident CFB eruption alone. The linearity of this correlation indicates that the deadly consequences of CFBs are, in general, directly proportional to their total volumetric eruption rate, though intrusive and passive degassing may well play a critical role in the timing of volatile release. Additionally, given the generally gradual nature of flood basalt volatile release on biological timescales, our results reinforce the expectation (35–37) that threshold processes—be they ecological, biogeochemical, or climatic—play a major role in mass extinction. Continued improvements in the high-precision geochronology of flood basalts and continued biostratigraphic study of faunal turnover and extinction in expanded sedimentary sections throughout the Phanerozoic will continue to refine our understanding of these correlations and of the exact mechanisms implicated therein.

Although a correlation between LIPs and mass extinctions is now well established, there is still much resistance to the notion that the Deccan Traps played any significant role in driving species extinction at the K-Pg boundary (65, 66). In this context, it has not escaped our notice that the correlation observed in Fig. 4 between CFB eruptive rates and extinction magnitude implies that the Deccan Traps may have produced a sizeable extinction around the K-Pg boundary, even without the coincidence of the Chicxulub impactor. It has long been noted that the Chicxulub impactor is precisely coincident with and expected to drive extremely rapid extinction at the K-Pg boundary (65, 67, 68), but high-latitude records are compatible with gradual extinction leading up to the boundary, which may be attributable to the longer-term environmental stresses of the Deccan Traps eruptions (69–71). The position of the K-Pg extinction above the regression line in Fig. 4 is consistent with the hypothesis that the Chicxulub impactor increased the severity and brevity of the K-Pg extinction, but the role of the Deccan Traps as a preimpact stressor and extinction mechanism should not be minimized. Following the linear extrapolation in Fig. 4, the 41.9% extinction observed at the K-Pg boundary is within the uncertainty of the $40.3 \pm 5.1\%$ (1σ) extinction predicted by the eruption rate of the Deccan Traps. Moreover, despite the emphasis frequently placed on bolide impact as a driver of the most charismatic extinction of the Phanerozoic, a significant coincidence between impacts and mass extinctions is not borne out beyond the K-Pg.

Further exploration of this correlation between CFBs and extinctions, along with detailed geographical constraints on faunal abundance data, could potentially illustrate a "common cause" for the observed association between faunal turnover and relative sealevel fall (72). The mantle plumes that drive CFB volcanism are known to cause domal uplift of the crust with an amplitude on the order of 1,000 m and wavelength on the order of 1,000 km, persisting for up to several Myr (73). With relative continental uplift (and, thus, relative sea-level fall) from this mantle upwelling and associated thermal anomalies, we may expect CFBs, but not their oceanic equivalents, to be directly correlated with the regression and disconformity often seen during the intervals of faunal turnover. This regression and loss of habitat in the epicratonic marine realm may then compound the direct stresses imposed by volcanic volatile emissions that this work suggests are a primary kill mechanism during Phanerozoic extinction intervals.

Materials and Methods

We used the extinction record calculated by Muscente et al. (13) from Paleobiology Database records (12) to evaluate the faunal turnover at each geologic stage boundary ([SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Table S1\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental). The Geologic Time Scale 2012 (11) was used to define the date and uncertainty of each of these stages. The LIP compilation of Ernst (14) was updated with the most precise available geochronology ([SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Table S2\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) and compared with the Muscente et al. calculations (13) ([SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Table S1\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) to analyze the coincidence between LIPs and periods of faunal turnover.

For each LIP,we define an estimated eruptive duration that includes a majority of the known eruptive products, using the most disparate high-precision-andaccuracy age bounds to define the start and end ages ([SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Table S2\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental). The eruptive rate for each Phanerozoic LIP was then calculated by dividing the best estimate of the total LIP volume by the eruptive duration to get an average rate in terms of km $^3/$ y over the bulk of the eruptive period (*[SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)*[, Table S2\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental). LIPs erupted before the Carboniferous are more heavily eroded and do not have robust volume estimates, so their eruptive rates were not calculated.

While the likelihood functions of geological timescale boundaries are generally Gaussian, the separate Gaussian distributions for LIP start and end ages require that the likelihood functions for LIPs be defined in terms of the product of those two normal cumulative distribution functions: one representing the likelihood that LIP eruption has begun and a second representing the likelihood that LIP eruption has ceased (c.f. [SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Fig. S1\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental). Consequently, the integral of the product of p_i and q_i given in Eq. **6** generally cannot be solved analytically. Instead, we integrate this equation numerically over a time interval from 10 SDs before the mean start age to 10 SDs after the mean end age for each LIP.

By generating some 10^n uniform random lists of timescale boundaries (i.e., synthetic boundary lists, where each timescale boundary is distributed as Unif (0, t_{Phan}), with t_{Phan} as the start time of the Phanerozoic) for some sufficiently large n, it is possible to determine numerically what proportion, if any, result in a higher coincidence product than the real list of geologic timescale boundaries and Phanerozoic LIPs (Fig. 1 and *[SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)*[, Table S1\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental). When a one-sample exact Kolmogorov–Smirnov test was applied to the observed distribution of Phanerozoic stage boundaries before the Quaternary ([SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Fig. S5\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental), the 95% confidence outcome could not reject the null hypothesis that the Phanerozoic stage boundaries are drawn from a uniform distribution (two-sided $P = 0.8141$). This supports the use of these uniform random boundary lists for calculating coincidence products. The probability (P) is given by dividing the number of random lists that have a coincidence product larger than that of the real list of timescale boundaries by the number of uniform random lists generated.

To efficiently generate this large number of uniform random boundary lists and calculate their coincidence products, we wrote a scalable Julia program (74) (available at [https://github.com/Theodore-Green/LIP-Extinction-Correlations\)](https://github.com/Theodore-Green/LIP-Extinction-Correlations) (75) using MPI.jl for distributed-memory parallel programming by message passing (76) and VectorizedRNG.jl for random number generation using a vectorized Xoshiro256 $++$ (77) algorithm. We then ran this program on the Dartmouth College linux cluster Discovery, completing 10⁸ simulations for each of the LIP log relative frequency plots in Fig. 1.

A similar approach was used to generate the log relative frequency plots for impacts in Fig. 1. The impact compilation of Schmieder and Kring (23) was combined with estimated impact crater diameters from the Earth Impact Database (78) ([SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Table S3\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) and, like the LIP record, compared with the stage boundary extinction record ([SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Table S1\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) to analyze the coincidence between Phanerozoic impacts and periods of faunal turnover. Since the bolide impacts have only one Gaussian event age instead of the separate start and end ages of LIPs, the integral of the product of p_i and q_i given in Eq. 6 generally can be solved analytically for impact events.Synthetic boundary lists were generated and coincidence products calculated in the same manner as previously described for LIPs. The observed coincidences for impact crater diameters \geq 20 km and \geq 40 km with stage-level and mass extinctions are statistically significant (Fig. 1 and [SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Fig. S2\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental). However, the Chicxulub impactor is obviously coincident with the K-Pg mass extinction, based on the way that boundary is defined (11), so its role in producing the correlations seen in Fig. 1 and *[SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)*[, Fig. S2](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) is explored by excluding it from the analyses in those figures as well. The values for P (the probability that a uniform distribution has a higher coincidence product [Ψ] than observed) indicate that the observed correlation between impacts and faunal turnover likely could be produced by chance without the Chicxulub impactor, indicating that the observed trends are driven primarily by that one event. This shows that other Phanerozoic impactors, all smaller than the Chicxulub, do not show the proposed correlation with extinctions (18, 22).

To test whether the observed LIP-extinction coincidences were likewise being driven by a single event, we repeated the experiments for Fig. 1 after removing the Siberian Traps from the list of considered LIPs. Even without this LIP precisely correlated with the Permo–Triassic mass extinction (51), Fig. 1 and [SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Fig. S6](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) show a relationship between LIPs and extinction that is unlikely to be produced by chance. Similarly, this was repeated for LIPs and stage boundaries when all of the LIPs temporally correlated with mass extinctions (the Deccan Traps, CAMP, Siberian Traps, Viluy Province, and Kola-Dnieper Province) were not considered. The relationship between LIPs and Phanerozoic extinctions remains statistically significant, even without these large, well-correlated events (Fig. 1). This demonstrates that the correlation between LIPS and Phanerozoic extinctions is robust and applicable even to smaller LIPs.

We also repeated the analyses just for CFBs. CFBs, like the whole LIP list, show observed coincidence products that are unlikely to be produced by chance, even when the eruptions most precisely correlated with extinctions are not considered ([SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Fig. S7\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental). This provides confidence that the analyses shown in Figs. 3 and 4, which only consider CFBs, reflect the same trend observed in Fig. 1 with all LIPs.

The trend in Fig.2 was also tested to see if it was driven primarily by the largest events. [SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Fig. S3](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) shows the results from dividing the considered LIP list into nonoverlapping categories of low, low-medium, medium-high, and high eruptive rates and extinction magnitudes. Even without allowing events to appear in multiple categories like Fig. 2, the trend of the observed coincidence product approaching the maximum possible coincidence product for a given list (normalized total coincidence product increasing) at higher eruptive rates and extinction magnitudes holds ([SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Fig. S3\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental).

Particularly at these high eruptive rates that have stronger correlations with extinction events, $CO₂$ degassing from flood basalt magmas would be a significant perturbation to the Earth system that, like the carbon isotope excursions explored by Rothman (37, 79), should result in mass extinction when the rate of environmental change exceeds a critical threshold. Following Rothman (37), the dimensionless inorganic carbon mass perturbation is

$$
M = \frac{|\Delta m|}{m^*}.
$$

where $m^* = 139$, 200 Pg is the steady-state value of inorganic CO₂ in the oceans [from the 38,000-Pg preindustrial reservoir of inorganic $C(80)$ used in Rothman's calculations] and Δm is the mass change. The dimensionless duration is

$$
I = \frac{\phi_c * \tau_{env}}{2\tau_0},
$$
 [10]

where ϕ_c is equal to the organic burial fraction through geologic time of 0.23 \pm 0.07 and the dimensionless, maximum, normalized flux (ϕ) for each event (37). τ_{env} is the duration of the event, and τ_0 is the 140,000-y turnover time for inorganic carbon in the oceans (37). Plotting the dimensionless duration of an event against its corresponding dimensionless carbon mass change shows where each event lies in relation to the critical threshold, defined by Rothman (37) as the identity line where

$$
\phi = \phi_c = 0.23 \pm 0.07. \tag{11}
$$

Those events that lie above the critical threshold are expected to result in mass extinctions and other environmental catastrophes. Rather than investigate carbon isotope events, we adapt this framework to LIP eruptions and associated carbon fluxes. The dimensionless mass is found by using a value of 0.5 wt.% degassed $CO₂$ for basaltic magmas (38). This is taken as a reasonable, but potentially high, value (38), so a range from 0.2 to 0.5 wt.% $CO₂$ is used for error bars on that value (39). That value for pre-eruptive $CO₂$ content is then used in order to calculate the direct CO₂ release from a given LIP eruption as the Δm value in the dimensionless mass term. The same wt.% $CO₂$ is used for oceanic and continental LIPs here, but higher ocean pressures may have decreased the amount of that $CO₂$ released by trapping gas in bubbles within the magma (i.e., popping rocks) (60, 81). The duration of a given LIP ([SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Table S2\)](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) is used for τ_{env} in calculating the dimensionless duration of the event. To account for LIP eruptions that may occur in distinct pulses, rather than assuming constant output over the entire duration, we also define critical thresholds for events that input C

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2. M. Coffin, O. Eldholm, Volcanism and continental break-up: A global compilation of large igneous provinces (Special Publications, Geological Society, London, 1992), vol. 68, pp. 17–30.

in pulses comprising 10%, 1%, and 0.1% of their total durations. Fig. 3 shows the dimensionless duration and dimensionless C mass perturbation for CFBs, while *[SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)*[, Fig. S4](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) shows the results for oceanic and continental LIPs. The LIP points are colored according to the extinction magnitude at the timescale boundary that coincides with that eruption.

[SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Fig. S8](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) applies the same framework to the present oceanic uptake of $CO₂$ and that predicted by 2100 based on the various emissions scenarios (Shared Socio-economic Pathway [SSP] 1.9, 2.6, 4.5, 7.0, and 8.5) defined in the 2021 Intergovernmental Panel on Climate Change report (82). These modern inputs are plotted on the same axes as the CFBs shown in Fig. 3, but the short duration necessitates an adaption of the critical threshold. For events with durations less than 10,000 y, the critical size

$$
M_c = 0.0082 \pm 0.0041,
$$
 [12]

is used to define the critical threshold instead of the rate equation in order to account for the minimum amount of time necessary for the ocean carbon cycle to respond to an influx of material (37). The present oceanic $CO₂$ uptake and that for the low-emission SSP 1.9 scenario fall below the critical size threshold, but the other emission scenarios plot at or above the critical threshold for modern carbon-cycle inputs. This indicates that $CO₂$ inputs since 1850 are comparable to the CFBs that are associated with the highest extinction magnitudes.

For Fig. 4, the ratio of the average LIP start and end times and the uncertainties on those ages is used to define the relative uncertainty (1 σ error) of each eruption. The most precisely dated CFBs have relative uncertainties of less than 0.2% and are included in the linear correlation. Moderately well-dated CFBs are those with 1σ error less than 0.85%. Moderately well-dated CFBs are not included in the linear correlation, but tend to fall near the line. With improved dating, these eruptions may shift closer to the line. The proposed eruptive durations of many less-precisely dated CFBs (1 $\sigma > 0.2\%$) overlap multiple geologic timescale boundaries. For these, the CFB with the highest eruptive rate at a given boundary is considered, and, when a CFB overlaps more than one boundary, only the boundary with the highest extinction magnitude is plotted in Fig. 4. All the CFBs are included in the logistic regression, which constrains the function to the physically possible range of values of the data (0 to 100%). The logistic trend is approximately linear in the considered range. Best estimates of volume are used throughout, but volume remains difficult to constrain for many LIPs. For well-dated and moderately well-dated CFBs, the extinction magnitudes predicted from the linear trend are listed and compared to the extinction magnitude at the corresponding boundary in [SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Table S5.](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)

The precisely dated CFBs are split into two categories (bulk eruptive rates <1 km³/y and \geq 1 km³/y) to plot the two regression lines shown in [SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Fig. S9.](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) Both categories show a positive, linear relationship between bulk eruptive rate and extinction magnitude, though they have different slopes and intercepts. The regression line for the higher rate eruptions more closely matches the overall trend, which suggests that those high-rate CFBs are a strong influence on the observed trend, but the linear trend in each category is still interesting in comparison to the lack of observed correlation between volume and extinction severity (3, 32–34).

Data, Materials, and Software Availability. All datasets used are included as [SI Appendix](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental)[, Tables S1–S3.](https://www.pnas.org/lookup/suppl/doi:10.1073/pnas.2120441119/-/DCSupplemental) The code, datasets, and coincidence product results for this project are available at GitHub [\(https://github.com/](https://github.com/Theodore-Green/LIP-Extinction-Correlations) [Theodore-Green/LIP-Extinction-Correlations\)](https://github.com/Theodore-Green/LIP-Extinction-Correlations) (75). Previously published data were used for this work (11, 13, 14, 23, 78).

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