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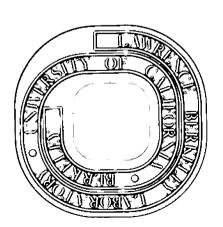
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### RELAXATION EFFECTS ON AUGER ENERGIES

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Abstract:

A hitherto-neglected relaxation-energy effect in the Auger process was found, and an approximate theory for calculating it was developed. The effect ranges from about 6 eV in neon to about 150 eV in fermium. Theoretical  $\mathrm{KL}_{\mathrm{I}}\mathrm{L}_{\mathrm{I}}$  energies show excellent agreement with experiment. The relaxation term is large enough to affect seriously the analysis of unknown Auger spectra. It appears to have very wide application to Auger spectroscopy.

In the course of a systematic study of the theory of Auger electron energies, it has become clear that an important effect—outer-shell relaxation in the final state—has been generally neglected until now. The use of three—or four parameter semi-empirical theories to fit existing data has obscured the significance of outer-shell relaxation, because these theories were flexible enough to give good fits. However, the agreement deteriorates drastically if one attempts to use calculated, rather than empirically adjusted, Slater integrals. In this Letter a new theoretical approach that includes

relaxation is derived. It relates Auger energies to measured one-electron binding energies (as did the semi-empirical theory) plus two-electron integrals from Hartree-Fock calculations in the ground states of neutral atoms. One adjustable parameter is required if the integrals employed are based on nonrelativistic wavefunctions; none is required if the wavefunctions are relativistic. Comparison of the predictions of this theory with experimental  $\mathrm{KL}_{\mathrm{LL}}$  Auger energies of the elements shows generally excellent agreement with experiment. The theory may be readily applied to estimate Auger energies throughout the Periodic Table. An application to energies of MNN groups in xenon is given, to indicate the wide applicability of the relaxation term.

For brevity let us consider the  ${^{K\!L}}_{\rm I}{^{L}}_{\rm I}$  Auger transition explicitly. In an element of atomic number Z this process may be denoted by

$$Z(\overline{K})^+ \rightarrow Z(\overline{L}_I^2)^{2+} + e^-$$
 (1)

where the atom is assumed to have its full complement of electrons excepting those denoted by the "hole-state" notation  $\overline{K}$ ,  $\overline{L}_{\underline{I}}$ , etc. An approximate estimate of the kinetic energy of the outgoing electron is given by

$$E(KL_{I}L_{I}) \cong E_{B}(K) - 2E_{B}(L_{I}) , \qquad (2)$$

i.e., by a combination of one-electron binding energies. In deriving Eq. (2) one implicitly considers Eq. (1) to be obtained as a combination of the processes

$$Z \rightarrow Z(\overline{K})^{+} + e^{-}$$
 (3a)

$$Z \rightarrow Z(\overline{L}_{I})^{+} + e^{-}$$
 (3b)

$$Z(\overline{L}_{I})^{+} \rightarrow Z(\overline{L}_{I}^{2})^{2+} + e^{-}$$
, (3c)

and the energy of step (3c) is assumed to be identical to that of (3b) in the derivation of Eq. (2). Although this is a very poor assumption, it is the <u>only</u> assumption, and efforts to improve theoretical estimates of KL<sub>TL</sub> Auger energies can be focused on Eq. (3c).

At this point we note that empirical binding energies will be used throughout this Letter in estimating Auger energies. Accurate theoretical K and L<sub>I</sub> binding energies are available for relatively few elements. Even in these cases the empirical binding energies are preferred because in combining them to form an Auger transition energy there is a partial cancellation of such effects as extra-atomic relaxation and the work function in condensed phases; in addition, relativity and electron correlation effects are of course included in the empirical binding energies.

To improve on Eq. (2), Asaad and Burhop<sup>2</sup> noted that Step (3c) differs from Step (3b) in one important way; when the second electron is ejected from the  $L_{\rm I}$  shell, there is no electron remaining in this shell. The binding energy will therefore be increased by the two-electron interaction energy. Neglecting the relatively small electron-correlation term, the two-electron energy for this case is given by the Slater integral  $F_0(2s,2s)$ . Thus Eq. (2) becomes, according to Asaad and Barhop,

$$E(KL_{I}L_{I}) = E_{B}(K) - 2E_{B}(L_{I}) - F_{O}(2s, 2s)$$
 (4)

These authors gave similar, but longer expressions for the other KLL energies. In most cases the required Slater integrals were not available, so empirical expressions of the form

$$F_0(2s,2s) \cong A(Z - Z_s) (1 + \alpha Z^n)$$
, (5)

were used, with the parameters A,  $Z_s$ ,  $\alpha$ , and n determined by comparison with existing Auger energies. In this way semi-empirical KLL Auger energies were estimated and tabulated for the elements by several authors.  $^{3-5}$ 

More recently, Asaad has shown that the relativity factor  $1 + \alpha Z^n$  should have n = 2; whereas the value n = 3 was used in the semi-empirical tables. By itself this would not be too serious, because  $\alpha$  could be readjusted to give a reasonably good fit of the data. However, in light of the availability now of accurate Slater integrals from the Hartree-Fock equations by J. B. Mann, it is no longer desirable to regard A and  $Z_s$  as adjustable parameters. But if the calculated values of  $F_0(2s,2s)$  are used, the calculated KL<sub>I</sub>L<sub>I</sub> energies disagree seriously with experiment, showing that Eq. (4) is not really viable.

$$E(KL_IL_I) = E_B(K) - 2E_B(L_I) + R - F_0(2s,2s)$$
 (16)

The energy R may be estimated by using an "equivalent cores" approach  $^8$  that has proved successful recently in calculating the effects of relaxation on atomic binding energies. The essence of this approach is that a one-electron outer orbital of an atom of element Z which has a hole in an inner orbital may be rather accurately represented by the corresponding outer orbital of a neutral atom of element Z + 1. It follows  $^8$  that

$$F_{0}(2s,nl; Z(\overline{2s})) \cong F_{0}(2s,nl; Z+1)$$

$$G_{\ell}(2s,nl; Z(\overline{2s})) \cong G_{\ell}(2s,nl; Z+1)$$

$$(7)$$

where n > 2 and  $Z(\overline{2s})$  denotes an atom of element Z with a hole in the 2s (or  $L_I$ ) shell. This approach neglects inner-shell (n = 1) relaxation, which is negligible. Intrashell effects are also small. They cannot be estimated by this method, but the results of Hedin and Johansson can be used to estimate the intrashell relaxation energy in light elements. Straightforward calculation gives

$$R \cong \sum_{\ell} \{N(n\ell) \Delta F_0(2s,n\ell) - \Delta G_{\ell}(2s,n\ell)\}, \qquad (8)$$

where N(nl) is the occupation number of the nl subshell and

$$\Delta F_{0}(2s,nl) \equiv F_{0}(2s,nl; Z + 1) - F_{0}(2s,nl; Z)$$

$$\Delta G_{\ell}(2s,nl) \equiv G_{\ell}(2s,nl; Z + 1) - G_{\ell}(2s,nl; Z) . \tag{9}$$

In the actual evaluation of R it is not advisable to simply calculate R for each element, especially in the transition series, because a change in N(nl) from Z to Z + 1 precludes the use of Eq. (7). Good estimates of R for all the elements above Z = 10 were obtained by calculating R for closed shells and subshells, using Eq. (8), and interpolating. For neon the intrashell energy given for Na<sup>+</sup> by Hedin and Johansson was used. No intrashell estimates were made for argon and above, because the R value for argon estimated using Eq. (8) accounted for all the relaxation energy.

Predicted  $\mathrm{KL}_{\mathrm{I}}\mathrm{L}_{\mathrm{I}}$  Auger energies of the elements are set out in Table I, together with experimental values, where available. The predicted values were calculated using empirical binding energies from Ref. 4. Outer-shell relaxation R was calculated using Mann's table, and  $F_0(2s,2s)$  was also taken from Mann's tables. Because these  $F_0(2s,2s)$  values were calculated with a nonrelativistic theory, a relativity correction of the form given by Asaad was applied. Specifically,  $F_0(2s,2s)$  was taken to have the form

$$F_0(2s,2s) = (1 + \alpha Z^2) F_0(2s,2s)_{Nonrelativistic}, \qquad (10)$$

and the single parameter  $\alpha$  was adjusted. The value  $\alpha$  = 4.2  $\times$  10  $^{-5}$  was adopted.

The agreement between predicted and experimental values of  $E(KL_{\mathrm{I}}L_{\mathrm{I}})$  is generally excellent, as perusal of Table I will show. It is even slightly better than that of the published semi-empirical values. Thus the necessity for considering the static relaxation energy R has been demonstrated, and a recipe has been given for calculating R with reasonable accuracy.

Further work is in progress. Preliminary results indicate that the R term, together with the intermediate-coupling model of Asaad and Burhop, will accurately predict the energies of all nine lines of the KLL multiplet throughout the Periodic Table. Applications to other Auger transitions are also encouraging. For example, this theory predicts the average energies of the xenon  $M_4N_4$ ,  $5^N4$ ,

Other applications are under study. A more detailed account of the static relaxation energy will be published elsewhere. The assistance of Mr. D. N. Shirley in calculating the Auger energies is gratefully acknowledged.

#### FOOTNOTES AND REFERENCES

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Table I. Predicted and Experimental E(KL,L,) Energies of the Elements (in eV).

		Table I. Predicted and Experimental $E(KL_{\underline{I}}L_{\underline{I}})$ Energies of the Elements (in eV).					
z	R	E(KL <sub>I</sub> L <sub>I</sub> ) This Work	E(KL <sub>I</sub> L <sub>I</sub> ) Experiment <sup>a</sup>	z	R	E(KL <sub>I</sub> L <sub>I</sub> ) This Work	E(KL <sub>I</sub> L <sub>I</sub> ) Experiment <sup>a</sup>
10	6	751	748.0(4)	56	92	25320	
11	. 8	922	922.8(4)	57	93	26239	25251(6)
12	.10	1101	1101.3(4)	58	95	27190	
13	12	1296		59	96	28161	•
14	14	1511	*	60	98	29153	•
15	16	1739		61	99	30160	
16	18	1980		62	101	31188	31175(20) <sup>b</sup>
17	20	2247		63	103	32238	3225#(50)p
18	22	2524		64	104	33305	32224(20)
19	25	2814	2809	65	105	34393	34430(50)
20	28	31.21		66	107	35502	35496(6)b
21	31	3451		67	109	36632	37490(0)
22	33	3793		68	110	37781	
23	36	4163	4159(6)	69	112	38953	38958(25)
24	39	4552	,	. 70	113	40146	
25	42	4953	4962(2) <sup>b</sup>	71	115	41359	40149(4) 41351(10) <sup>b</sup>
26	45	5373	5376 <sup>b</sup>	72	117		41351(10)
27	48	5806	2510	73	118	42588	
23	51	6265		74		43832	15505(15)
19	54	6734	6735(6) <sup>b</sup>	ļ.	120	. 45097	45080(40)
30	57	7216	7220(4)	75	121	46387	46400(25)
31	58	7713	1220(4)	76	122	47694	
32	60	8218	8212(6)	77	123	49026	
33	61	8749	8742(10)	78	124	50382	50370(100)
14	62	9283	0142(10)	79	126	51761	51780(20)b
15	63	9839	9860(10)	80	127	53160	53141(25) <sup>b</sup>
6	65	10411	3000(10)	81	129	54567	54510(100)
7	67	10994	:	82	130	56007	
8	68	11593	11584.4(16)	83	132	57468	57467(30)b
9	70	12211	11304.4(10)	84	133	58939	58920(50)
o .	72	12849	12851.8(15)	85	135	60450	
1	73	13501	12071.0(17)	8 <b>6</b>	136	62006	
2	75	14176	14176.1(13)	87	137	63552	
3	76	14862	141/0.1(13)	88	138	65134	
4	78	15570		89	139	66754	
5	80	16294		90	140	68379	•
,	81			91	141	70056	
7	83	17034		92	142	71748	71738(20) <sup>b</sup>
8	85	17793	17740(60)	93	143	73486	73555(150) <sup>b</sup>
)	85	18563	10050(1)	94	144	75257	75180(15)
, )	95 86	19347	19352(1)	95	145	77116	
, L	87	20149	*	96	146	78928	
· !	88	20968	0100011	97	147	80660	**
: 3	88	21806	21787(10)	98	148	83352	•
<b>3</b> ¥	89	22659	22652(10)	99	149	85294	
•	90	23516 24415	24395(14) <sup>b</sup>	100	150	87286	

arrow Ref. 5. Errors in last place appear parenthetically.

b In cases for which two or more experimental values are available. Either an average was taken or a single value chosen.

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