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Denoising Surface Waves Extracted From Ambient Noise Recorded by 1-D Linear Array Using Three-Station Interferometry of Direct Waves

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1 2 3	Denoising surface waves extracted from ambient noise recorded by 1-D linear array using three-station interferometry of direct waves
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14	Key points:
15 16	• Surface waves from ambient noise cross correlations are significantly enhanced at high frequencies using three-station interferometry
17 18	• Phase travel times are extracted reliably between 0.3-1.6 s for a 1.6-km-long linear array and are used to perform surface wave tomography
19 20	• Phase velocity models of Rayleigh and Love waves derived via eikonal tomography reveal high-resolution fault zone images

#### 21 Abstract

22 We develop an automatic workflow for enhancing surface wave signals in ambient noise 23 cross correlations (ANCs) calculated for a 1-D linear array. The proposed array-based 24 method is applied to a 1.6-km-long dense linear nodal array crossing surface traces of the 25 San Jacinto fault near Anza, California. Fundamental and higher modes of surface waves 26 are observed in ANCs of the nodal array. After attenuating the surface wave overtones by 27 applying a frequency-dependent tapering window to the ANCs, signals dominated by the 28 fundamental mode surface wave are then enhanced through a denoising process based on 29 three-station interferometry of direct waves. The signal to noise ratio is significantly 30 increased at high frequencies (> 2 Hz) after denoising. Phase travel times are extracted 31 reliably in the frequency domain for the period ranges of 0.3-1.2 s and 0.3-1.6 s for 32 Rayleigh and Love waves, respectively. The corresponding period-dependent phase 33 velocity profiles derived from the eikonal equation reveal high-resolution details of fault 34 zone internal structures beneath the array. A broad (500-1000 m) low-velocity zone that 35 narrows with increasing period is observed, illuminating a flower-shaped structure of the 36 San Jacinto fault damage zone.

## 37 Plain Language Summary

38 Properties of fault damage zone (width of 100-1000's meters), such as its geometry and 39 velocity reduction compared to the surrounding host rock, can have a profound impact on 40 our understandings of earthquake ruptures and the long-term behavior of the fault. 41 Several dense nodal arrays with 10-100 m spacing and aperture of a few kilometers were 42 deployed crossing surface traces of major faults, to provide high-resolution images of the 43 fault zone internal structures. Surface waves travel at frequency-dependent speeds 44 between every two sensors are observed in ambient noise cross correlations. We can infer 45 structures at different depth using surface wave, as the velocity at higher frequency is 46 more sensitive to shallower structures. However, surface waves extracted from ambient 47 noise at high frequencies (> 1 Hz), that are essential to image fault zone in the top 100's 48 meters, are often very noisy. Here, we develop a denoising method that utilizes three-49 station interferometry to effectively suppress non-surface wave signals in a linear 1-D 50 array. The quality of surface waves is significantly improved after the denoising,

especially at high frequencies (> 2 Hz). Reliable measurements at high frequencies
provide better constraints on fault zone internal structures at shallow depth.

#### 53 **1. Introduction**

54 Noise-based surface wave tomography has been widely used to resolve crustal 55 structures at various scales (e.g., Lin et al., 2009; Qiu et al., 2019; Wang et al., 2019; 56 Zigone et al., 2019). In an effort to study high-resolution internal structures of major 57 faults in southern California (SC), several dense arrays with station spacing less than 100 58 m were deployed crossing surface traces of the San Jacinto fault (e.g., Qin et al., 2021) 59 and rupture zone of the 2019 Ridgecrest earthquake sequence (Catchings et al., 2020) for 60 about one month. Analysis of high frequency (e.g., > 1 Hz) surface waves extracted from 61 ambient noise cross correlations (ANCs) of these dense arrays can provide crucial 62 information on the shallow (top 10s to 100s of meters) materials near faults with unprecedented spatial resolution (e.g., Wang et al., 2019) and thus improves our 63 64 understanding of the local seismic hazard.

65 The quality of surface waves reconstructed from ANC depends on the duration of the 66 continuous data and noise source distribution. Previous studies have shown that proper 67 preprocessing steps (e.g., Bensen et al., 2007) can improve signal to noise ratio (SNR) of 68 surface waves extracted from ANC. However, in contrast to high-quality signals at long 69 periods (e.g., > 1 s), extraction of surface waves from ANCs calculated at high 70 frequencies (e.g., > 1 Hz) for these linear arrays remains a challenging topic due to its 71 low SNR (e.g., Wang et al., 2019), even if proper preprocessing procedures (e.g., Bensen 72 et al., 2007) were implemented. This is likely due to the short recording time (e.g., one 73 month) of these dense arrays and complicated pattern of noise sources at high frequencies 74 and near faults (e.g., Hillers et al., 2013, 2014).

To achieve better reconstruction of surface waves from ambient seismic noise with inhomogeneous source distribution, Stehly et al. (2008) used correlations of the coda of the ANC (C3) calculated for a 2-D seismic network. This is achieved by stacking higherorder correlations for triplets of stations with two common virtual receivers. Froment et al. (2011) then investigated Green's function reconstructed by correlations of the coda of C3 and concluded that the C3 method is helpful in suppressing source effects associated

81 with non-isotropic source distribution. Sheng et al. (2018) later proposed an alternative 82 prestack procedure that constructs C3 from the coda of each daily ANC first and then 83 stack over the entire recording period. They showed that the prestack C3 reveals faster 84 convergence and better recovery of arrivals at higher frequencies. The C3 method has 85 also been applied to data recorded by asynchronous seismic networks to enhance the 86 spatial resolution of noise-based surface wave tomography (e.g., Spica et al., 2016).

87 Different from C3, we use correlations of the entire waveform of the ANCs (later 88 referred to as "three-station interferometry of direct waves") in this study to enhance 89 signals of surface waves. Zhang et al. (2020) compared surface waves extracted from C3 90 and three-station interferometry of direct waves using data from the EarthScope 91 USArray. They found that surface waves retrieved from three-station interferometry of 92 direct waves show considerably higher SNR and broader bandwidth but yields small 93 biases in dispersion measurements. Such bias arises from the geometry of the stations 94 used to compute C3, and becomes zero or negligible when these stations align along a 95 nearly straight line (i.e., 1-D linear array), even if noise sources are not evenly distributed 96 (Lin et al. 2008). Therefore, we adopt the idea of using three-station interferometry of 97 direct waves (hereinafter "three-station interferometry" for simplicity) to denoise the 98 fundamental mode surface waves extracted from ANCs for 1-D linear arrays, which has 99 not been done before.

100 We note that the configuration of 1-D linear array is commonly used in many dense 101 deployments, such as linear arrays across major faults in SC (e.g., Catchings et al., 2020; 102 Oin et al., 2021) and fiber optic cables (e.g., Cheng et al., 2021). The deployment periods 103 of these 1-D linear arrays are in the order of a few tens of days to a few months, which 104 are much shorter than those of 2-D broadband arrays that are used in ambient noise 105 tomographic studies (e.g., Qiu et al., 2019). We apply this method to data recorded by a 106 dense linear array deployed at the Ramona Reservation (RR) site across surface traces of 107 the San Jacinto fault, near Anza (Fig. 1), California. Seismic waveforms from the RR 108 array have been analyzed for fault zone internal structures in Qin et al. (2021). ANCs 109 were computed for each station pair of the RR array and the corresponding Rayleigh 110 wave phase velocities for periods from 0.3 s to 0.8 s were derived from double 111 beamforming tomography in Wang et al. (2019).

112 Since signals of surface wave overtones were observed in ANCs of the RR array 113 (Wang et al., 2019), in addition to the description of station configuration (Figs. 1a-b) and 114 ANC data (Figs. 1c-d), we first analyze and attenuate signals of higher-mode surface 115 waves for both components (Fig. 2) in section 2. Following the flow chart illustrated in 116 Figure 3a, we present the theoretical formulation for denoising the fundamental mode 117 surface waves using three-station interferometry with a 1-D linear array in section 3.1. 118 The denoising process is demonstrated for an example station pair (Fig. 4) in section 3.2, 119 and the comparison between surface wave signals before and after denoising is illustrated 120 for the linear segment of the RR array (Fig. 5) in section 3.3. In section 4, following the 121 flow chart in Figure 3b, surface wave phase travel times are first extracted from the 122 denoised wavefield and then inverted for phase velocity dispersion models via the eikonal 123 equation (Figs. 6-7). Discussion of the performance of the denoising method and 124 comparison between the resulting phase velocity profiles and fault zone images from 125 previous studies (Wang et al., 2019) are presented in section 5.

#### 126 2. Data and Preprocessing

127 The RR array (red triangle in Fig. 1b) is located at north of Anza (blue square in Fig. 128 1b), California, and crosses surface traces of the Clark segment of the San Jacinto fault 129 (Fig. 1a). The array consists of 94 three-component 5-Hz Fairfield geophones (balloons 130 in Fig. 1a) that were set to record continuously for a month with a sampling rate of 500 131 Hz. ANC is obtained by first computing cross correlations of ambient noise data in 5-min 132 windows, and then stacking them over the entire recording period for each station pair 133 (Wang et al., 2019). The positive and negative time lags of the monthlong stacked ANC 134 are fold and averaged to reduce the effects of the asymmetric noise source distribution. 135 We use ANCs of a sub-array RR01-RR47 (yellow, blue, and red balloons in Fig. 1a) to 136 demonstrate the surface wave denoising process (Fig. 3a) developed in this study. The 137 sub-array has 47 stations with an average station spacing of ~30 m and an aperture of 138 ~1.6 km.

We project stations in the sub-array to the straight line connecting RR01 and RR47 (cyan dashed line in Fig. 1a) and compare interstation distances calculated using station locations before and after the projection. The comparison yields negligible differences (<

1%) suggesting that the sub-array RR01-RR47 is in a 1-D linear configuration (later 142 143 referred to as "the linear RR array"). In Wang et al. (2019), a period-dependent velocity 144 threshold was applied to taper off the contamination of body waves or potential higher-145 mode surface waves. In this study, however, we first only apply a tapering window, with 146 a maximum and minimum moveout velocities of 2 km/s and 0.1 km/s (black dashed lines 147 in Figs. 1c-d), to the raw ANCs for both Transverse-Transverse (TT) and Vertical-148 Vertical (ZZ) components. The tapered ANCs are then filtered between 0.2 Hz and 10 Hz 149 and depicted as colormaps in Figures 1c-d.

150 The aim of this study is to demonstrate that the proposed denoising procedure (Fig. 151 3a; Section 3.1) is effective in denoising surface waves, when only one mode (i.e., 152 fundamental mode) of surface waves is present in the ANCs. Thus, we first analyze the 153 surface wave overtones in the ANCs of the linear RR array by resolving the array-mean 154 group and phase dispersion images (Fig. 2). Details of the analysis are described in 155 supplementary materials (Text S1 and Fig. S1). Figures 2a-b show the array-mean group 156 dispersion images for Love and Rayleigh waves, respectively. The blue dashed curves, 157 connecting points with the largest values of the image obtained at different periods, 158 denote the average group velocity dispersion of the fundamental mode signal. This 159 suggests that the fundamental mode surface wave is the dominating signal in the ANCs 160 filtered between 0.3 s and 1.6 s. This is in a good agreement with the observation of only 161 one peak per period in the array-mean phase dispersion diagrams (MASW; Park et al., 162 1999) at long periods (> 0.6 s in Fig. 2c and > 0.7 s in Fig. 2d).

163 At short periods, however, higher-mode signals are clearly observed with much 164 higher phase velocities in the array-mean phase dispersion diagrams (< 0.6 s in Fig. 2c 165 and < 0.7 s in Fig. 2d). This is consistent with the weaker energy (above white dashed 166 lines) observed at these short periods in the group dispersion images (Figs. 2a-b) that 167 travels at group speeds much higher than that of the fundamental mode (blue dashed 168 curves). The absence of higher mode signals at long periods is likely the result of low 169 amplitude due to subsurface structures and excitation pattern of the ambient noise field. 170 Another possibility is that the group velocity of higher modes at long periods is faster 171 than 2 km/s. To suppress the surface wave overtones observed at short periods, we first 172 highpass filter the ANCs (< 0.6 s for TT and < 0.7 s for ZZ; black dashed lines in Fig. 2)

and apply a second tapering window to the filtered data that removes signals with group speeds above 0.55 km/s before performing surface wave denoising (Fig. 3a). The attenuation of higher-mode signals at short periods is effective, as the signatures of surface wave overtones that are clearly observed in Figures 2c-d are almost missing in the updated phase dispersion diagrams computed for the ANCs after the second tapering (Figs. 2e-f).

# 179 **3. Surface Wave Denoising**

180 Let  $G_{i,j}(t)$  be the positive lag of ANC for the station pair of *i*-th (virtual source) and *j*-

181 th (virtual receiver) sensors in the linear RR array (yellow, green, and red triangles in Fig.

182 1a), we can expand it as

$$G_{i_{j}}(t) = S_{i_{j}}(t) + B_{i_{j}}(t) + N_{i_{j}}(t),$$
(1)

183 where  $S_{i_j}(t)$  and  $B_{i_j}(t)$  represent signals traveling on the surface (i.e., surface waves) and 184 at depth (i.e., diving *P* or *S* body waves) between the source *i* and receiver *j*, respectively. 185  $N_{i_j}(t)$  is the residual (later referred to as "background noise"). This section aims to 186 develop a denoising process that preserves  $S_{i_j}(t)$  while suppressing  $B_{i_j}(t)$  and  $N_{i_j}(t)$  in 187 equation 1.

#### 188 **3.1 Three-station interferometry for a 1-D linear Array**

189 Since surface waves are dispersive, let  $\tilde{G}_{i_j}(\omega)$  be the Fourier transform of  $G_{i_j}(t)$  at 190 the angular frequency  $\omega$ , we can rewrite equation 1 in the frequency domain

$$\tilde{G}_{i_{j}}(\omega) = A_{G_{i}ij} \cdot e^{i \cdot \varphi_{G_{i}ij}} = \tilde{S}_{i_{j}}(\omega) + \tilde{B}_{i_{j}j}(\omega) + \tilde{N}_{i_{j}j}(\omega)$$

$$= \sum_{S} A_{S_{i}ij} \cdot e^{-i\left(\omega \cdot T_{ij}^{S} + \varphi_{S}\right)} + \sum_{B} A_{B_{i}ij} \cdot e^{-i\left(\omega \cdot T_{ij}^{B} + \varphi_{B}\right)} + \tilde{N}_{i_{j}j}(\omega),$$
(2a)

191 where  $A_{S\_ij}$  and  $T_{ij}^S$  are amplitude spectrum and phase travel time of surface wave signals 192 in ANC at the angular frequency  $\omega$  that propagate between the *i*-th and *j*-th stations, 193 while  $A_{B\_ij}$  and  $T_{ij}^B$  represent those of body waves that travel at depth.  $\varphi_s$  and  $\varphi_B$  are 194 initial phases of surface- and body-wave signals in the ANC, respectively, and dependent 195 on the distribution of ambient noise sources (e.g., Lin et al., 2008). Assuming the higher-196 mode surface waves are negligible or have already been removed from ANC (Section 2), 197 we, therefore, can simplify equation 2a as:

$$\tilde{G}_{i_{j}}(\omega) = \tilde{F}_{i_{j}}(\omega) + \tilde{O}_{i_{j}}(\omega) = A_{F_{i}j} \cdot e^{-i\left(\omega \cdot T_{i_{j}}^{F} + \varphi_{F}\right)} + \tilde{O}_{i_{j}}(\omega),$$
(2b)

where the symbol or subscript *F* stands for the fundamental mode surface wave.  $\tilde{O}_{i_j}(\omega) = \tilde{B}_{i_j}(\omega) + \tilde{N}_{i_j}(\omega)$ , that consists of signals from body waves traveling at depth and background noise, is the term we want to suppress in the denoising process. It is interesting to note that  $\varphi_F = \pi/4$  for an azimuthally homogenous ambient noise source distribution (Snieder, 2004), whereas  $\varphi_F = 0$  when noise sources are only present in line with the station pair *i* and *j* (Lin et al., 2008).

For surface waves of a certain (e.g., fundamental) mode traveling between three stations i < j < k in a 1-D linear array, the travel times satisfy the following relation

$$T_{ik}^S = T_{ij}^S + T_{jk}^S, (3a)$$

206 whereas

$$T^B_{ik} < T^B_{ij} + T^B_{jk},\tag{3b}$$

for body waves traveling at depth. Therefore, we introduce a third station *k* and perform
three-station interferometry following Zhang et al. (2020):

$$\tilde{I}_{i_{j}}(\omega;k) = \begin{cases} \tilde{G}_{i_{k}}^{*}(\omega) \cdot \tilde{G}_{j_{k}}(\omega), & k < i \\ \tilde{G}_{i_{k}}(\omega) \cdot \tilde{G}_{j_{k}}(\omega), & i < k < j. \\ \tilde{G}_{i_{k}}(\omega) \cdot \tilde{G}_{j_{k}}^{*}(\omega), & k > j \end{cases}$$
(4a)

209 In equation 4a, we cross correlate  $G_{i,k}(t)$  and  $G_{i,k}(t)$  in the time domain, when k < i or k > i210 *j* (later referred to as "outer-source zone"). The interferometry becomes equivalent to the convolution of  $G_{i_k}(t)$  and  $G_{j_k}(t)$  in the time domain for station k located within the two 211 virtual sources (i.e., i < k < j; later referred to as "inter-source zone"). For the case k = i or 212 *j*, we define  $\tilde{I}_{i_j}(\omega; k) = A_{G_ij}^2 \cdot e^{i \cdot \varphi_{G_ij}}$  that approximates the convolution of  $G_{i_j}(t)$  and 213  $G_{i_i}(t)$  or  $G_{j_i}(t)$ , by assuming the amplitude spectrum of the auto-correlation  $G_{i_i}(t)$  or 214  $G_{j,j}(t)$  is similar to that of  $G_{i,j}(t)$ , i.e.,  $A_{G \ ii} \approx A_{G \ jj} \approx A_{G \ ij}$ . 215 216 Combining equations 2b, 3a, and 4a, if the fundamental mode surface wave is the

216 Combining equations 2b, 3a, and 4a, if the fundamental mode surface wave is the 217 dominant signal in ANC (i.e.,  $\tilde{O}_{i_j}$  in Equation 2b is negligible), the phase term of the 218 interferogram  $\tilde{I}_{i_j}(\omega; k) = A_{ij_k} \cdot e^{i \cdot \varphi_{ij_k}}$  is given by

$$\varphi_{ij\_k}(\omega) = \begin{cases} -\omega \cdot T_{ij}^F - 2\varphi_F, & i < k < j \\ -\omega \cdot T_{ij}^F - \varphi_F, & k = i \text{ or } j. \\ -\omega \cdot T_{ij}^F, & k < i \text{ or } k > j \end{cases}$$
(4b)

 $T_{ij}^F$  and  $\varphi_F$  denote the phase travel time and initial phase of the fundamental mode surface 219 wave signal (Equation 2b) extracted from the ANC of station pair *i* and *j*. Considering the 220 221 amplitude spectrum,  $A_{G_{ij}}$ , of the original ANC usually peaks at certain frequencies (e.g., 222 microseism frequency band), the three-station interferometry also acts as a bandpass filter 223 that amplifies signals around those spectral peaks in  $A_{G_{ij}}$ , as the amplitude term of the interferogram  $\tilde{I}_{i_j}(\omega; k)$  is approximately given by  $A^2_{G_ij}$ . Therefore, we take the square 224 root of the amplitude term  $A_{ij,k}$  while preserving the phase term  $\varphi_{ij,k}$  of the original 225 interferogram  $\tilde{I}_{i_j}(\omega; k)$  to suppress the effect of source spectra multiplication introduced 226 in the three-station interferometry (Equation 4a), i.e.,  $\tilde{I}_{i_{-}j}^{c}(\omega; k) = \sqrt{A_{ij_{-}k}} \cdot e^{i \cdot \varphi_{ij_{-}k}}$ . This is 227 228 based on the assumption that amplitude spectra of the input ANCs are similar for all station pairs, i.e.,  $A_{ij_k} = A_{G_ik} \cdot A_{G_jk} \approx \hat{A}_G^2$ . 229

230 Equation 4b suggests that the interferograms within either the inter- or outer-source 231 zones share the same phase, whereas interferograms from different zones are only aligned in phase when  $\varphi_F$  is zero. In cases when the term  $\tilde{O}_{ij}$  is significant, we can divide the 232 interferogram  $\tilde{I}_{i,j}^{c}(\omega;k)$  into two components:  $\tilde{I}_{i,j}^{cF}(\omega;k)$  and  $\tilde{I}_{i,j}^{cO}(\omega;k)$ .  $\tilde{I}_{i,j}^{cF}(\omega;k)$ 233 represents the interferogram that only involves the fundamental mode surface wave 234 signal, i.e., when  $\tilde{O}_{i_j}$  is set to zero in equation 2b. The phase of  $\tilde{I}_{i_j}^{cF}(\omega; k)$ , given by 235 equation 4b, is independent of k when  $\varphi_F$  is zero. Therefore, we can simply stack the 236 interferograms  $\tilde{I}_{i,j}^c(\omega; k)$  over all available station k to enhance the fundamental mode 237 238 surface wave and the denoised waveform is given by:

$$\tilde{C}_{i_{j}}^{3}(\omega) = \sum_{k=1}^{N} \tilde{I}_{i_{j}}^{c}(\omega;k) / N \approx \sum_{k=1}^{N} e^{i \cdot \varphi_{i_{j},k}} \sqrt{A_{i_{j},k}} / N.$$
(5)

Here, *N* is the number of stations in the 1-D linear array, and we assume  $\varphi_F = 0$ . We note that the other component,  $\tilde{I}_{i,j}^{c0}(\omega; k)$ , is suppressed through the stacking. This is because the phase term of  $\tilde{I}_{i,j}^{c0}(\omega; k)$  varies significantly with *k*, as it involves contributions from diving body waves and background noise that do not satisfy equation 3a and thus equation 4b. If  $\varphi_F \neq 0$ , we need to correct the interferogram  $\tilde{I}_{i,j}^c(\omega; k)$  for station *k* within the inter-source zone following equation 4b, i.e.,  $\tilde{I}_{i,j}^c(\omega; k) \cdot e^{i \cdot 2\varphi_F}$ , before stacking.

### 246 **3.2 Surface wave denoising of the station pair RR10 and RR40**

247 As described in section 2, we observe higher-mode Love and Rayleigh waves only at 248 short periods that travel at group speeds higher than 0.55 km/s (Fig. 2) in ANCs of the 249 linear RR array (Figs. 1c-d). Thus, the denoising process (Fig. 3a) is performed directly 250 on the lowpass filtered ANCs for each component, as the surface wave overtones are 251 negligible at low frequencies (> 0.6 s at TT and > 0.7 s at ZZ; Fig. 2). For ANCs at high 252 frequencies, a second tapering window that effectively attenuates higher-mode signals is 253 applied to the highpass filtered data before denoising. All results in later discussions are 254 thus associated with the fundamental mode surface wave.

255 Figure 4 shows results of the three-station interferometry applied to ANCs of TT 256 component for an example station pair RR10 (i = 10) and RR40 (j = 40). As surface waves are dispersive, we compare the input ANCs and the interferograms,  $\tilde{I}_{i_{-}j}^{c}(\omega; k)$ , 257 after narrow bandpass filtering at two example periods, 0.8 s (Figs. 4a-c) and 0.3 s (Figs. 258 259 4d-f), to better demonstrate the performance of the denoising process in the low and high 260 frequency bands, respectively. The bandpass filter is generated following section 3.1 of 261 Qiu et al. (2019). Figures 4a and 4b show the ANCs filtered at 0.8 s with RR10 and RR40 262 as the virtual source,  $G_{i,k}(t)$  and  $G_{i,k}(t)$ , respectively. Black and blue waveforms denote the filtered ANCs with the virtual receiver station k inside the outer-source (k < 10 or k > 10263 264 40) and inter-source (10 < k < 40) zones (i.e., k is the y-axis of Fig. 4), respectively. The 265 filtered ANC of the example station pair RR10 and RR40 is depicted in red.

Figure 4c demonstrates in colors the resulting interferograms  $\tilde{I}_{i,i}^{c}(\omega; k)$  transformed 266 to the time domain through inverse Fourier transform. The interferograms show coherent 267 268 surface waves filtered at 0.8 s that are well aligned in phase for all k values. This is 269 consistent with our derivations in section 3.1 and the observation of high-quality surface 270 wave signals in the filtered ANCs (Figs. 4a-b). The fact that the interferograms  $\tilde{I}_{i}^{c}(\omega; k)$ 271 are well aligned between the inter- (blue vertical arrow) and outer-source (black vertical 272 arrow) zones suggests the initial phase  $\varphi_F \approx 0$  (Equation 4b) for ANCs filtered at 0.8 s. 273 Therefore, we obtain the denoised waveform for the example station pair RR10 and 274 RR40 filtered at 0.8 s (black waveform; Equation 5a) by stacking all the interferograms 275 in Figure 4c. The surface wave signals (between the dashed lines) are almost identical between waveforms before (in gray) and after (in black) denoising, with the denoisedwaveform showing smaller coda waves.

278 Coherent surface waves are also observed in Figure 4f for interferograms narrow 279 bandpass filtered at 0.3 s. Again, signals of surface wave in the interferograms filtered at 280 high frequency are also well aligned between the inter- and outer-source zones, indicating 281 the assumption of  $\varphi_F \approx 0$  is valid. We also verified that the initial phase  $\varphi_F \approx 0$  for all 282 periods (from 0.3 s to 1.6 s) and components (ZZ and TT) analyzed in this study, by 283 estimating the systematic phase difference between interferograms in the inter- and outer-284 source zones computed for the example station pair RR10 and RR40 (not shown here). 285 The SNR of surface waves in interferograms at high frequency is much lower than those 286 at low frequency, as incoherent arrivals with large amplitudes that vary significantly with 287 the choice of station k are seen (Fig. 4f). This is because of the lower SNR for surface waves (i.e.,  $\tilde{O}_{i,i}$  in Equation 2b is non-negligible) in the input ANCs filtered at high 288 frequency (Figs. 4d-e). The SNR of surface wave is increased after denoising (in black), 289 290 as the large coda waves present in the filtered ANC (in gray) are significantly suppressed.

291 Similar observations are obtained for Rayleigh waves extracted from ANCs at ZZ 292 component (Fig. S2). Although the surface wave signals are generally aligned well between different interferograms filtered at both the low and high frequencies, small 293 294 fluctuations in the travel times are still observed in Figures 4c and 4f. This is likely due to 295 variations in SNR, i.e., amplitude ratio between surface and coda waves, of the 296 interferograms. In addition, although a narrow bandpass filter is applied, the peak 297 frequency of the filtered interferograms can still deviate from the center frequency of the 298 filter. Thus, variations in peak frequency can also lead to visible changes in surface wave 299 travel times, particularly for long interstation distances, as surface waves are dispersive 300 (e.g., blue dashed curves in Fig. 2). We note that, based on our derivations in section 3.1, 301 the denoising method is applicable to any arrivals if (1) the wave propagation satisfies equation 3a (e.g., teleseismic arrivals, body waves traveling along the surface or refracted 302 303 from a subsurface impedance contrast) and (2) only one such arrival is present in the 304 input wavefield  $G_{i}(t)$ .

# 305 3.3 Surface wave denoising of the linear RR Array

306 In section 3.2, we demonstrate the workflow and effectiveness of the three-station-307 interferometry-based denoising process for an example station pair RR10 and RR40. As 308 illustrated in the flow chart (Fig. 3a), we can further enhance surface waves extracted 309 from ANCs of the entire linear array by performing the denoising process for multiples 310 times: first self-normalize the output wavefield of the current iteration, and then use the 311 normalized wavefield as the input for the next iteration. The number of iterations is 312 determined so that the difference between input and output wavefields of the last iteration is negligible. As surface waves are dispersive, we use symbol  $C_{i,i}^{2+n}(t;T_c)$  to represent the 313 waveform of station pair i and j, after first applying  $n \geq 1$  iterations of the denoising 314 process and then narrow bandpass filtering at period  $T_c$  for better illustration of the 315 316 denoised results.

317 Figure 5a shows the comparison between the TT component ANC (black) of the 318 example station pair RR10 and RR40 filtered at 0.8 s and the corresponding denoised waveforms (in red) of the first  $C_{i,i}^3(t; 0.8)$  and second  $C_{i,i}^4(t; 0.8)$  iteration. Since the 319 waveforms  $C^3$  and  $C^4$  are almost identical, this suggests that only two iterations are 320 needed for the denoising results to converge at low frequencies (> 0.6 s for TT and > 0.7321 322 s for ZZ; Fig. 2). On the other hand, Figure 5e suggests that four iterations (red 323 waveforms) are needed to ensure a convergence of the denoising process at high 324 frequencies, as the SNR of surface wave is much lower in the filtered ANC (black 325 waveform) after the attenuation of higher-mode signals (Section 2). Although the surface 326 wave SNR gradually increases with the number of denoising iterations (from bottom to 327 top; Figs. 5a and 5e), the surface wave signals are always coherent and aligned in phase.

328 The ANC data  $G_{i,i}(t)$  of TT component filtered at 0.8 s and the corresponding denoised waveforms  $C_{i,j}^4(t; 0.8)$  are illustrated in Figures 5b-c, respectively, for all 329 330 station pairs. Although waveforms before and after denoising filtered at low frequency 331 are similar, the difference is still noticeable (Fig. 5c). For instance, the background 332 fluctuations with irregular arrival patterns in the coda waves are reduced; the arrival prior 333 to surface wave with a phase velocity of  $\sim 2$  km/s at long interstation distances (> 1.2 km), 334 which are likely related to the tapering window (with an upper limit velocity of 2 km/s; 335 Section 2), are suppressed after the denoising. Figure 5d compares the array-mean 336 amplitude spectra averaged over all station pairs for data before (in black) and after (in

red) the denoising. The similarity between the two average amplitude spectra is
consistent with the high SNR of surface waves in the raw ANCs filtered at low
frequency.

340 The difference between wavefields before and after denoising is much larger at high 341 frequency (0.3 s; Figs. 5f-g). Although coherent surface wave signals are seen 342 propagating at a group speed slightly slower than 0.5 km/s in the filtered ANC data (Fig. 343 5f), wavelets with large amplitudes are observed in coda waves (e.g., black waveform in 344 Fig. 5f). These wavelets sometimes have arrival times similar to those of surface waves 345 and thus can interfere with and bias the surface wave dispersion measurements. After 346 four iterations of denoising, the background noise is greatly suppressed in the wavefield 347  $C^{6}$  (Fig. 5g). Figure 5h shows the array-mean amplitude spectra, with the one averaged 348 over data after denoising being smoother (in red). This is likely due to the interference 349 between the surface and coda waves that contributes to the complicated array-mean 350 amplitude spectrum calculated for the data before denoising (in black).

351 In addition to comparisons between the input ANCs and denoised wavefield at low 352 (0.8 s) and high (0.3 s) frequencies in Figure 5, we also compute the array-mean group 353 and phase dispersion images for the denoised wavefield at TT (Figs. S4a-b) and ZZ (Figs. 354 S5a-b) components following the procedures described in Text S1. The array-mean group 355 and phase velocity dispersion curves (white markers in Figs. S4-S5) are determined as the 356 period-dependent velocity that yields the largest amplitude of the curve extracted from 357 the image at each corresponding period. The dispersion relations obtained for data before 358 (blue dashed curve) and after (white markers) denoising are compared in Figure 2. 359 Differences between results before and after denoising are generally smaller than 5% for 360 array-mean group velocities but much larger ( $\sim 10\%$ ) for phase velocities. This is because 361 array-mean phase velocities obtained from MASW (e.g., Figs. 2c-f) are inferred in the 362 frequency domain (Text S1) and thus very sensitive to background noise.

363

## 4. Surface Wave Tomography

In this section, we use denoised waveforms of TT and ZZ components (e.g., Figs. 5c, 5g, S3c, and S3g) to infer phase velocity structures of Love and Rayleigh waves beneath the array, respectively. Following the flow chart shown in Fig. 3b, we first determine 367 cycle-skipped phase travel times of surface waves propagating between all available 368 station pairs at each period (e.g., Fig. 6a) in the frequency domain, which is much simpler 369 than measuring in the time domain but requires high SNR (Section 4.1). Second, we infer 370 phase velocity structures beneath the linear RR array, using travel time measurements 371 after cycle-skipping correction from section 4.1, via the eikonal equation in section 4.2 372 (e.g., Fig. 6b). The aim of this section is to demonstrate that robust surface wave phase 373 velocity models can be resolved from the denoised waveforms.

# **4.1 Determination of phase travel time**

375 Frequency time analysis (FTAN) is widely used in previous studies to determine 376 phase travel time of surface wave signal in ANC (e.g., Bensen et al., 2007; Lin et al., 377 2008; Qiu et al., 2019). First, Gaussian narrow bandpass filters centered on a series of 378 consecutive frequencies are applied to the ANC, then the phase travel time dispersion is 379 measured using the envelope and phase functions of the filtered ANC in the time domain. 380 The advantage of FTAN is that reliable phase travel times can still be extracted when 381 SNR is low at high frequencies. However, ad hoc criteria and thresholds are required to 382 automate the FTAN. Additional details on the FTAN method can be found in section 3 of 383 Qiu et al. (2019). Since our goal is to verify that the signals after denoising are 384 representative of surface waves and high SNR is achieved for all frequencies, we thus 385 measure phase travel times from the denoised waveforms in the frequency domain, which 386 is much simpler than the FTAN method and described in detail below.

387 Although surface wave SNR is high in the denoised waveform, we still observe 388 waves with very small amplitudes before and after the surface wave (e.g., black 389 waveforms in Figs. 5c and 5g). This is because we can only attenuate rather than remove 390 completely signals that are not surface waves. Here, we apply a frequency-dependent 391 tapering window (e.g., black dashed lines in Figs. S4d-S5d) centered on the surface wave 392 to further remove these background fluctuations. Width of the tapering window is set to 393 six times the dominant period of the array-mean amplitude spectrum (e.g., red curve in 394 Fig. 5d), whereas the center is determined by the array-mean phase and group velocities 395 of the array at the target period (e.g., white markers in Fig. 2). We note that this tapering 396 process is to ensure the accuracy of phase travel times measured in the frequency domain, 397 which is unnecessary if FTAN is implemented.

Since  $\varphi_F \approx 0$  (e.g., Fig. 4; Section 3.2) and the term  $O_{i_j}$  in equation 2b is negligible after denoising (e.g., Fig. 5) and tapering (e.g., Fig. S4d-S5d), we have

$$\tilde{C}_{i_{j}}^{DT}(\omega; f_{c}) = A_{F_{i}j}(\omega; f_{c})e^{-i\cdot\omega T_{i_{j}}^{F}(\omega)},$$
(6)

where  $\tilde{C}_{i_{-}i_{-}}^{DT}(\omega; f_c)$  is the spectrum of the denoised and tapered waveform filtered at the 400 center frequency  $f_c$  for station pair *i* and *j* (e.g., blue waveform in Fig. 6a). Equation 6 401 402 suggests that we can extract cycle-skipped phase travel time from the phase spectrum of the tapered waveform. It is important to note that the peak frequency  $f_{max}$  of the tapered 403 waveform, i.e., the peak of  $A_{F ij}$ , may deviate from  $f_c$ , the center frequency of the filter. 404 Therefore, we measure the wrapped phase (i.e., between  $-2\pi$  and 0) of the spectrum 405  $\tilde{C}_{i\,i}^{DT}(\omega; f_c)$  at  $\bar{f}_{max}$ , the peak frequency of the amplitude spectrum averaged over all 406 station pairs, where the array-mean SNR of the surface wave is the highest. Then, the 407 cycle-skipped phase travel time is computed as the wrapped phase divided by  $-2\pi \bar{f}_{max}$ . 408

Figure 6a shows the cycle-skipped phase travel times (white circles) measured at  $\bar{f}_{max} = 3.14$  Hz for surface waves filtered and tapered at 0.3 s from station pairs associated with a common virtual source RR10 (y-axis of 0 km). To obtain the actual phase travel time, we perform a simple cycle-skipping correction as follows:

413 (1) As illustrated in Figure 6a, we first extract all the cycle-skipped phase travel times for
414 surface waves of a virtual shot gather and arrange them as a function of the location
415 to the virtual source.

416 (2) We perform cycle-skipping correction for surface waves traveling NE (toward RR47)
417 and SW (toward RR01) separately.

418 (3) For surface waves traveling in the same direction, the principle of the cycle-skipping
419 correction is to ensure that the travel time of any virtual receiver is larger than those
420 of receivers that are closer to the virtual source after the correction.

421 (4) In practice, we examine measurements  $T_i$  and  $T_{i+1}$  of every two adjacent virtual 422 receivers with the *i*-th station being closer to the virtual source. If  $T_i \ge T_{i+1}$ , we use  $T_i$ 423 as the reference and add  $N/\bar{f}_{max}$  (*N* is an integer) to  $T_{i+1}$  so that  $T_{i+1} + N/\bar{f}_{max} > T_i \ge$ 424  $T_{i+1} + (N-1)/\bar{f}_{max}$ . The correction is performed for closer-to-source pairs first.

425 Travel times, for the virtual shot gather of RR10, after the correction are illustrated as red426 stars in Figure 6a. We note that a more sophisticated cycle-skipping correction (e.g.,

using phase velocity structure inferred at a longer period as the reference) is needed whenstation spacing is larger than one wavelength.

#### 429 **4.2 1-D eikonal tomography**

We use the eikonal equation to derive phase velocity structures using travel time measurements of all station pairs in the linear RR array (Section 4.1). First, we project all stations to the straight line connecting RR01 and RR47 (cyan dashed line in Fig. 1a). Second, travel time measurements associated with each virtual source *i* at the target frequency  $\bar{f}_{max}$  are extracted and interpolated (e.g., black curve in Fig. 6a) with a regular grid size of  $\Delta$ =50 m. Since variations in topography (Fig. 2b of Qin et al., 2021) have a negligible effect (< 0.5%) on the results, the eikonal tomography can be simplified as:

$$\tilde{v}_i(x; \bar{f}_{\max}) = 2 \cdot \Delta / [T_i(x + \Delta; \bar{f}_{\max}) - T_i(x - \Delta; \bar{f}_{\max})],$$
(7)

437 where  $\tilde{v}_i(x; \bar{f}_{max})$  and  $T_i(x; \bar{f}_{max})$  are the local phase velocity and interpolated phase 438 travel time, respectively, of the grid cell at location *x*. Since the local phase velocity  $\tilde{v}_i$ 439 only varies with the grid cell location, it is independent of virtual source *i*. Thus, we can 440 average the 1-D phase velocity profiles resolved from all available virtual sources at the 441 same frequency  $f_{max}$  to achieve a more reliable phase velocity model:

$$\bar{v}(x; f_{\max}) = \sum_{i=1}^{N_x} \tilde{v}_i(x; \bar{f}_{\max}) / N_x, \qquad (8a)$$

442 and estimate the corresponding uncertainty as the standard deviation:

$$\delta(x; \bar{f}_{\max}) = \sqrt{\sum_{i=1}^{N_x} \left[ \tilde{v}_i(x; \bar{f}_{\max}) - \bar{v}(x; \bar{f}_{\max}) \right]^2 / N_x}, \tag{8b}$$

443 where  $N_x$  is the number of virtual sources available for stacking at location x.

In surface wave studies, phase velocities derived at near-virtual-source grid cells are often excluded to satisfy the far-field approximation (e.g., Bensen et al., 2007). The size of the exclusion zone is usually multiples of the analyzed wavelength (e.g., one wavelength in Wang et al., 2019). Here, however, we set an exclusion zone with a fixed size of 100 m, i.e., discard phase velocities derived at the four grid cells closest to the virtual source, to avoid any potential bias in travel time gradient estimation near the virtual source. Figure 6b shows the 1-D phase velocity profile, in white dots, derived 451 using measurements associated with the virtual source RR10 (black curve in Fig. 6a) for 452 Love waves at  $\bar{f}_{max} = 3.14$  Hz, whereas phase velocity profiles resolved from all virtual 453 sources are illustrated in gray curves and as the colormap. The average phase velocity 454 and uncertainty profiles are calculated via equation 8 and demonstrated as red stars and 455 error bars, respectively, in Figure 6b.

456 Figure 7 shows phase velocity models resolved at periods ranging from 0.3 s to 1.6 s 457 for Love waves (Fig. 7a) and 0.3 s to 1.2 s for Rayleigh waves (Fig. 7b), together with the 458 corresponding uncertainty estimations (Figs. 7c-d). The period range in the plot is 459 determined so that the resolved maximum uncertainty is smaller than 0.1 km/s. In 460 general, the uncertainties are smaller than 0.03 km/s for both Rayleigh and Love waves at 461 all analyzed periods, indicating the resolved phase velocity structures are robust and 462 reliable. The small uncertainty at low frequencies also justifies our choice of an exclusion 463 zone with a 100-m-radius. This is because one wavelength at low frequency (e.g., ~800 m 464 for Rayleigh wave at  $\sim 0.8$  s; white markers in Fig. 2f) is much larger than 100 m. If the 465 one wavelength exclusion zone is necessary, phase velocities calculated at a target grid 466 cell should vary significantly for virtual sources within and outside the exclusion zone, 467 and thus yield large uncertainty values.

468 Phase velocity models of both Love and Rayleigh waves show a ~500- to 1000-m-469 wide low-velocity zone at low frequencies (e.g., > 0.8 s) that gradually narrows with the 470 period. Combined with the fact that phase velocity at lower frequency is more sensitive to 471 structures at greater depth, this observation likely indicates a flower-shaped (i.e., width 472 decreases with depth) fault damage zone beneath the linear RR array. We also see several 473  $\sim$ 100-m-wide narrow zones, that are close to the mapped fault surface traces (black 474 dashed lines in Figs. 7a-b), with extremely low phase velocities (< 500 m/s) at high 475 frequencies (e.g., 0.3-0.6 s). However, the shape and location of these low-velocity zones 476 are different between Figures 7a and 7b. Structure patterns that are inconsistent between 477 models of Love and Rayleigh waves may indicate the existence of radial anisotropy or 478 complicated structures of Vp/Vs ratio.

#### 479 **5. Discussion**

We present a denoising method based on three-station interferometry that effectively enhances surface waves extracted from ANCs of a 1-D linear array. This array-based denoising method complements the existing tools that utilize denoising filters (e.g., Baig et al., 2009; Moreau et al., 2017) or phase weighted stacking (e.g., Schimmel et al., 2011; Ventosa et al., 2017) to improve the quality of ANC calculated for a single station pair. There are three assumptions in the theoretical derivations of this method (Section 3.1):

486 (1) The wave propagation satisfies equation 3a, i.e., the array configuration is 1-D and487 linear.

488 (2) The phase of the target signal is given by  $\omega \cdot T_{ij}^F + \varphi_F$  (Equation 2b) for source *i* and 489 receiver *j*, i.e., the bias  $\varphi_F$  irrelevant to wave propagation is a constant.

490 (3) Only one such arrival is present in the input wavefield.

While the derivations are applicable to other signals (e.g., teleseismic and refracted body waves) and datasets (e.g., earthquake data), we focus on the demonstration of enhancing the fundamental mode surface waves extracted from ANCs of the linear RR array in this paper. Performance of the proposed denoising process is dependent on how well these three assumptions are satisfied using the field data.

496 To verify the first assumption, we estimate the error between the linear RR array and an ideal 1-D linear configuration. The station-configuration error is approximately given 497 498 by the difference (in percentage) between interstation distances calculated using station 499 locations before and after projecting the array to a straight line (green dashed line in Fig. 500 1a). For a rough estimation, the mean and maximum of the station-configuration error for 501 the linear RR array are ~0.1% and 1%, respectively. The uncertainties estimated from 502 eikonal tomography (Figs. 7c-d) suggest ~1% and ~3% for the mean and maximum 503 perturbations in the resolved phase velocities, which is larger than the estimated error in 504 array geometry. Therefore, we conclude that this denoising method is robust when the 505 station-configuration error is less than the allowable uncertainty of the resulting phase 506 velocity model (e.g., mean and maximum of 1% and 3% in this study).

Regarding the second assumption, as shown in Lin et al. (2008), the phase term  $\varphi_F$  is related to the effect of noise source distribution. In section 3.2, we analyze the term  $\varphi_F$ systematically for triplets of stations with two common virtual sources, RR10 and RR40, at various periods between 0.3 and 1.6 s (e.g., Figs. 4c and 4f). The observation that the interferograms calculated for different triplets of stations are well aligned suggests that not only the second assumption is valid but also  $\varphi_F \approx 0$ . It is interesting to note that  $\varphi_F \approx 0$  indicates the noise sources recorded by the linear RR array are not isotopically distributed, as otherwise  $\varphi_F$  would be  $\pi/4$  (Snieder, 2004). Since  $\varphi_F$  can be calculated for any given noise source distribution, measurements of  $\varphi_F$  using ANCs of sub-arrays aligned in a straight-line taken from a 2-D array along different angles may provide a new way of resolving the noise source distribution.

Although clear higher-modes of surface waves are observed in ANCs of the linear RR array at high frequencies (Figs. 2a-d), we find that these higher-mode signals can easily be separated from the fundamental mode surface waves and effectively attenuated through a second tapering process (Section 2) to satisfy the third assumption. However, if two modes (*F* and *M*) of surface waves are present in ANCs, following equation S3c derived in Text S2, we can still apply three-station-interferometry-based denoising to the ANCs by stacking the interferogram  $\tilde{I}_{i_j}(\omega; k)$  defined in Equation 4a:

$$\tilde{C}_{i\_j}^{3}(\omega) = \sum_{k=1}^{N} \tilde{I}_{i\_j}(\omega;k) / N \approx A_F^2 \cdot \left[ e^{-i\omega \cdot T_{ij}^F} + R^2 \cdot e^{-i\omega \cdot T_{ij}^M} \right], \tag{9}$$

where  $T_{ij}^F$  and  $T_{ij}^M$  are the phase travel times of the mode F and M, respectively.  $A_F$ 525 denotes the array-mean amplitude spectrum of the mode F, whereas the constant R526 527 indicates the amplitude ratio between the two modes. As demonstrated by derivations in 528 Text S2 and the synthetic test in Figures S6-S7, waveforms after denoising via equation 9 529 still preserves accurate phase travel times of both modes, as cross terms between the two 530 modes are suppressed through stacking. We use equation 9 in the denoising process when 531 two modes are present in the wavefield, as the application of equation 5 to such data 532 leads to pseudo arrivals (Nakata, 2020; Figs. S8-S9).

However, unlike equation 5, the denoising process defined by equation 9 also acts as a bandpass filter ( $A_F$ ) around the peak frequency,  $f_{max}$ , of the input wavefield and thus inherently attenuates signals at frequencies away from  $f_{max}$  (e.g., Fig. S7d). It is interesting to note that the amplitude ratio between the two modes changes to  $R^2$  in the denoised waveform  $\tilde{C}_{i_j}^3$  in equation 9, i.e., the weaker mode in the input wavefield (e.g., the mode *M* if R < 1) is attenuated (by a factor of *R*) after the denoising. Therefore, even if higher-mode signals are present in the field data, the third assumption is still valid if the ratio *R* between the fundamental mode and any higher modes in the input wavefield satisfies  $R^2 \ll 1$  (e.g., R < 0.5; Figs. S6-S7). Although a constant ratio *R* is assumed in the derivations of Text S2 and *R* may vary with many factors (e.g., distance), a good approximation of such constant ratio would be the array-mean of *R* values measured from all station pairs.

545 To further evaluate the surface wave signals in the denoised waveforms, we derive 546 phase velocity dispersion models for Love (0.3-1.6 s; Fig. 7a) and Rayleigh (0.3-1.2 s; 547 Fig. 7b) waves extracted from the denoised wavefield at TT and ZZ component in section 548 4, respectively. Uncertainties of the dispersion model is calculated as the standard 549 deviation of phase velocities derived for the same grid cell from different virtual sources 550 (Equation 8b). The small median uncertainties of ~20 m/s for both Love (Fig. 7c) and 551 Rayleigh (Fig. 7d) waves are consistent with our derivations in section 3.1, as the 552 denoising process aims at enhancing arrivals that have a propagation time between two 553 receivers independent of the source location (Equation 3a). We note that, although the 554 uncertainties are much larger (> 50 m/s) at longer periods (> 1.3 s in Fig. 7c and > 1.1 s 555 in Fig. 7d), the errors in percentage are still small ( $\sim 2-3\%$ ).

556 We compare Rayleigh wave phase velocity models derived from this study and Wang 557 et al. (2019) in the overlapping period (0.3-0.8 s) and spatial (RR01-RR47) ranges (Figs. 8a-b). In their study, the double beamforming technique is applied to the ANCs. The 558 559 local phase velocity is first obtained through grid search for each sub-array (three nearby 560 stations) pair: first sum all nine ANCs of the two sub-arrays through slant-stacking using 561 different slowness values, then determine the local phase velocity of each sub-array based 562 on the maximum amplitude of the envelope function for the stacked waveform. The final 563 phase velocity is given by the average value of phase velocities derived for the same 564 receiver but different source sub-arrays. They did not derive phase velocities for Rayleigh 565 wave at low frequencies (> 0.8 s) as the size of their exclusion zone (one wavelength) is 566 comparable to the array aperture (~1.6 km; Fig. 1a).

567 Wang et al. (2019) used the standard deviation of the median as the uncertainty. This 568 is because their phase velocities obtained from different sources are statistically 569 independent, whereas such redundant information has already been implemented in our 570 denoising process (Equation 5). We find extremely large uncertainty values (> 0.15 km/s) 571 in regions with high phase velocities (left bottom corner) and near fault surface traces 572 (i.e., F1-F3) from their results (Fig. 8d). Since surface waves are dispersive, these large 573 values may be the result of the violation of the assumption required by the double 574 beamforming method that peak frequencies of the applied bandpass filter and waveform 575 after slant stacking are the same. For instance, such deviation between peak frequencies 576 of the applied filter (3.33 Hz) and stacked amplitude spectrum (3.16 Hz) is clearly shown 577 in Figure 5g. Except for regions with errors > 0.15 km/s, the uncertainties of Wang et al. 578 (2019) (~30-40 m/s in Fig. 8d) are comparable with (slightly larger than) those of this 579 study ( $\sim 20$  m/s in Fig. 8c).

580 Similarly, consistent velocity values and structural patterns are seen in both models, 581 such as an ultra-low velocity (< 0.4 km/s) zone on the NE side of the middle fault surface 582 trace (F2) and faster velocities (> 1.2 km/s) at the bottom left of Figures 8a-b, suggesting 583 both phase velocity models are generally reliable. On the other hand, small-scale 584 differences between the two models are clearly observed. For instance, we find several 585 high-velocity anomalies with narrow width (50-100 m) near fault surface traces F2 and 586 F3, outlined by blue dashed lines in Fig. 8b, in the model of Wang et al. (2019), which 587 are missing from our model (Fig. 8a). Considering uncertainties associated with these 588 anomalies resolved from Wang et al. (2019) are extremely large (> 0.15 km/s; blue 589 dashed lines in Fig. 8d), we conclude that these features are likely artifacts, and the phase 590 velocity model derived in this study (Fig. 8a) from the denoised wavefield is a better 591 representation of fault zone structures beneath the linear RR array.

592 Different from Wang et al. (2019), we do not construct a shear wave velocity model 593 for structures beneath the linear RR array using the resolved phase velocity dispersion 594 profiles (Fig. 7). The reasons are twofold: (1) the phase velocity model derived from 595 eikonal tomography assumes that wave amplitudes are varying smoothly in space (e.g., 596 Lin et al., 2009; Lin & Ritzwoller 2011) and (2) the piecewise 1-D Vs inversion scheme 597 adopted in Wang et al. (2019) is only accurate when structures are varying smoothly 598 laterally. Considering the array is deployed crossing fault damage zones that are laterally 599 heterogenous (Figs. 1a and 7) and can significantly amplify seismic motions (Qin et al., 600 2021), we propose performing full waveform tomography (e.g., Zhang et al., 2018) on

the denoised surface waves to infer accurate fault zone structures beneath the array as thesubject of a future study.

603 In the present study, we only show phase velocities resolved at frequencies up to  $\sim 3$ 604 Hz (Fig. 7) from the denoised ANCs by requiring a minimum wavelength of 150 m (i.e., 605 the maximum station spacing) for eikonal tomography (Section 4). However, this 606 denoising method can be applied to ANCs at even higher frequencies. Using high-quality 607 surface waves denoised from ANCs of linear arrays crossing major fault zones for a wide 608 range of frequencies (e.g., 0.6-3 Hz in this study; 2-40 Hz in Zigone et al., 2019), a shear 609 wave velocity model that extends to both shallow (top few tens of meters) and deep (top 610 1-2 km) structures can be derived. Such fault zone velocity model with unprecedented 611 high-resolution will complement the qualitative and semi-quantitative models inferred from traditional fault zone analyses (e.g., Qin et al., 2018, 2021; Qiu et al., 2017, 2020; 612 613 Share et al., 2017, 2019). An integration of both quantitative and qualitative fault zone 614 models can have significant implications for seismic hazard evaluations (e.g., Ben-Zion 615 & Shi, 2005; Spudich & Olsen, 2001) and long-term behavior of the fault (e.g., Thakur et al., 2020). 616

#### 617 **6.** Conclusions

618 We develop a simple workflow that enhances surface waves extracted from ANCs of 619 a 1-D linear array by taking advantage of the redundant information of surface wave 620 propagation along the same straight-line, i.e., the amount of time that surface waves 621 travel between two stations is independent of the source location. We demonstrate the 622 effectiveness and robustness of the three-station-interferometry-based surface wave 623 denoising method in improving SNR of surface waves extracted from ANCs of the linear 624 RR array, particularly at high frequencies (e.g., > 2 Hz). The proposed surface wave 625 denoising method can be applied to a wide range of topics in the future:

- Reduce the minimum duration of ambient noise recording and preprocessing stepsneeded to achieve high-quality surface waves from ANCs.
- 628 2. Provide high-quality surface wave signals both at high (> 2 Hz) and low frequencies
  629 (< 1 Hz) for better constraints of shallow (top 10s to 100s of meters) materials</li>
  630 through full waveform tomography of surface waves.

- 631 3. Investigate the initial phase  $\varphi_F$  (Equation 2b) for 1-D linear arrays extracted from a 632 2-D deployment at different angles, and the possibility to infer the ambient noise 633 source distribution from measurements of  $\varphi_F$ .
- 4. Since higher mode surface wave signals are observed in ANCs of the RR array and
  provide extra constraints on the subsurface structure, the surface wave overtones at
  short periods (e.g., < 0.6 s) can also be enhanced and analyzed by first attenuating the</li>
  fundamental mode signal (i.e., suppress waves traveling at a speed slower than 0.55
  km/s; Figs. 2a-b) and then applying the same denoising process.
- 639

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# 774 **Figure captions**

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776 Figure 1. (a) Google map for the RR array (colored balloons) deployment that crosses 777 surface traces of the San Jacinto fault (colored lines). The stations colored in white are 778 not analyzed in this study, whereas the green balloons denote three sensors closest to 779 each corresponding fault surface trace. Surface wave denoising procedure is 780 demonstrated for an example station pair RR10 and RR40 (red balloons). (b) Zoom out 781 map of the San Jacinto fault zone. The background colors indicate topography. The red 782 star and blue square denote locations of the RR array and the town of Anza. The black 783 lines illustrate surface traces of major faults in this area. EF – Elsinore Fault; SAF – San 784 Andreas Fault; SJF - San Jacinto Fault. (c) Ambient noise cross correlations at TT 785 component of all station pairs for the sub-linear-array RR01-RR47. The cross correlations

are arranged according to interstation distance with red and blue colors representing positive and negative values. All the waveforms are first tapered using a velocity range of 2 km/s and 0.1 km/s (dashed lines), and then bandpass filtered between 0.2 and 10 Hz. (d) Same as (c) for the ZZ component. The black waveforms in (c) and (d) are the TT and ZZ component correlation functions of the station pair RR10 and RR40, respectively.

791 Figure 2. (a)-(b) Array-mean group velocity dispersion images (Text S1) for Love 792 and Rayleigh waves extracted from ambient noise cross correlations (ANCs) at TT and 793 ZZ component, respectively. The white dashed line denotes a group speed of 0.55 km/s 794 that separates the energy of the fundamental mode surface waves from that of the higher-795 mode signals. (c)-(d) Multichannel analysis of surface waves (MASW; Park et al., 1999) 796 for Love and Rayleigh waves, respectively. The black dashed line splits the target period 797 range with the higher-mode surface waves only visible at short periods. (e)-(f) Same as 798 (c)-(d) for ANCs after attenuating signals of higher-mode surface waves (Section 2). The 799 background colors of each panel illustrate the likelihood of the array-mean surface wave 800 velocity dispersion, with blue and red representing values between 0 and 1. The blue 801 dashed curve and white markers illustrate the array-mean surface wave velocity 802 dispersions inferred from data before and after denoising (Figs. S4-S5), respectively.

Figure 3. (a) Flow chart of the surface wave denoising and imaging procedure developed in this study. The dashed box outlines the part of the diagram that performs surface wave tomography. The workflow adopted in this study for surface wave tomography is shown in (b). ANC – Ambient Noise Cross-correlation. The filters applied to the ANCs before denoising are determined by the black dashed line shown in Figure 2. The removal of higher-mode surface waves at high frequencies are described in section 2.

Figure 4. (a) Ambient noise cross correlations (ANCs) of TT component narrow bandpass filtered at 0.8 s associated with the virtual source RR10 (red star). Waveforms are arranged by the station number of the virtual receiver. (b) Same as (a) for virtual source RR40. Waveforms in black and blue represent ANCs of virtual receivers in the outer- and inter-source zones, respectively, while the red waveform denotes the ANCs of the station pair RR10 and RR40 (red balloons in Fig. 1a). (c) Interferograms (colors) calculated via three-station interferometry, i.e.,  $\tilde{I}_{i,i}^{c}(\omega; k)$  in equation 5. The gray

waveform denotes the filtered ANC of the station pair RR10 and RR40, whereas the
linear stack of all the interferograms is shown in black. The black dashed lines (same as
Fig. 5a) outline the surface wave signal. (d)-(f) same as (a)-(c) for F-ANC, i.e., data after
the attenuation of higher-mode signals (Section 2; Fig. 3a), narrow bandpass filtered at
0.3 s.

821 Figure 5. (a) Comparison between the ANC (in black) and denoised waveforms (in 822 red) of the station pair RR10-RR40 narrow bandpass filtered at 0.8 s for TT component. 823 The black dashed lines outline the surface wave signal ( $\pm 2$  periods centered on the envelope peak of  $C^4$ ). (b) ANCs narrow bandpass filtered at 0.8 s for TT component, with 824 red and blue colors representing positive and negative values. The three white dashed 825 826 lines illustrate moveout velocities of 2 km/s, 0.5 km/s, and 0.1 km/s. The black waveform is the same as the bottom black waveform in (a). (c) The denoised  $C^4$  wavefield narrow 827 bandpass filtered at 0.8 s. The white dashed line denotes a moveout velocity of 0.5 km/s. 828 829 The black waveform is the same as the top red waveform in (a). (d) Array-mean 830 amplitude spectra of waveforms shown in (b) and (c) are depicted in black and red, 831 respectively. The peak frequency of the red amplitude spectrum (red dashed line) is 832 labeled in the top right corner of the panel (c). (e)-(h) same as (a)-(c) for F-ANC, i.e., 833 data after the attenuation of higher-mode signals (Section 2; Fig. 3a), narrow bandpass 834 filtered at 0.3 s.

835 Figure 6. (a) Love waves associated with the virtual source RR10 extracted from the 836 wavefield after denoising and tapering. The narrow bandpass filter centered at 0.3 s is 837 applied. Red and blue colors represent positive and negative values, respectively. White 838 circles denote cycle-skipped phase travel time measurements extracted at the peak 839 frequency of the filtered wavefield (3.14 Hz), whereas red stars indicate travel times after 840 cycle skipping correction (Section 4.1). The black curve illustrates the corrected phase 841 travel time after interpolation using a grid size of 50 m. (b) Phase velocity profiles 842 resolved for Love waves at 3.14 Hz. White circles depict the 1-D phase velocity profile 843 derived via eikonal equation for virtual source RR10, i.e., using the black curve in (a). 844 The colormap illustrates phase velocity profiles obtained using different stations as 845 virtual sources (x-axis), with white space illustrating the near-source exclusion zone. The

phase velocity profile averaged over results of all virtual sources are depicted as red stars,
with the error bar representing the corresponding standard deviation. The black vertical
dashed line denotes the array-mean phase velocity estimated at 0.3 s (red star in Fig.
S4b).

Figure 7. Phase velocity dispersion profiles for (a) Love and (b) Rayleigh waves beneath the RR array. The vertical dashed lines denote locations of the mapped fault surface traces (Fig. 1a). Uncertainties of the resolved phase velocity profiles are shown in (c) and (d) for Love and Rayleigh waves, respectively.

Figure 8. Comparison of (a) phase velocity and (b) uncertainty profiles of this study and those, (b) and (d), of Wang et al. (2019) in the overlapping period, 0.3-0.8 s. and spatial, RR01-RR47, ranges. The white space indicates the area not covered by the final model. Figure 1.



Figure 2.



Figure 3.



Figure 4.



Correlation time (s)

Figure 5.





Figure 6.



Figure 7.



Figure 8.

