UC Berkeley

UC Berkeley Previously Published Works

Title

The CUORE and CUORE-0 experiments at LNGS

Permalink

https://escholarship.org/uc/item/6q23v65z

Journal

Journal of Physics Conference Series, 1056(1)

ISSN

1742-6588

Authors

Alduino, C Alfonso, K Avignone, FT <u>et al.</u>

Publication Date

2018-07-01

DOI

10.1088/1742-6596/1056/1/012009

Peer reviewed

IOP Conf. Series: Journal of Physics: Conf. Series 1056 (2018) 012009 doi:10.1088/1742-6596/1056/1/012009

The CUORE and CUORE-0 experiments at LNGS

C Alduino¹, K Alfonso², F T Avignone III¹, O Azzolini³, G Bari⁴, F Bellini^{5,6}, G C Alduino', K Alfonso', F T Avignone III', O Azzolini', G Bari', F Bellini'', G Benato⁷, A Bersani⁸, M Biassoni⁹, A Branca^{10,11}, C Brofferio^{12,9}, C Bucci¹³, A Camacho³, A Caminata⁸, L Canonica^{14,13}, X G Cao¹⁵, S Capelli^{12,9}, L Cappelli^{7,16,13}, L Cardani⁶, P Carniti^{12,9}, N Casali⁶, L Cassina^{12,9}, D Chiesa^{12,9}, N Chott¹, M Clemenza^{12,9}, S Copello^{17,8}, C Cosmelli^{5,6}, O Cremonesi⁹, R J Creswick¹, J S Cushman¹⁸, A D'Addabbo¹³, D D'Aguanno^{13,19}, I Dafinei⁶, C J Davis¹⁸, S Dell'Oro^{20,13,21}, M M Deninno⁴, S Di Domizio^{17,8}, M L Di Vacri^{13,22}, V Dompè^{5,6}, A Drobizhev^{7,16}, D Q Fang¹⁵, M Faverzani^{12,9}, E Ferri⁹, F Ferroni^{5,6}, E Fiorini^{9,12}, M A Franceschi²³, S J Freedman^{16,7,a}, B K Fujikawa¹⁶, A Giachero^{12,9}, L Gironi^{12,9}, A Giuliani²⁴, L Gladstone¹⁴, P Gorla¹³, C Gotti^{12,9}, T D Gutierrez²⁵, K Han²⁶, K M Heeger¹⁸, R Hennings-Yeomans^{7,16}, H Z Huang², G Keppel³, Yu G Kolomensky^{7,16}, A Leder¹⁴, C Ligi²³, K E Lim¹⁸, Y G Ma¹⁵, L Marini^{17,8}, M Martinez^{5,6,27}, R H Maruyama¹⁸, Y Mei¹⁶, N Moggi^{28,4}, S Morganti⁶, S S Nagorny^{13,21}, T Napolitano²³, M Nastasi^{12,9}, C Nones²⁹, E B Norman^{30,31}, V Novati²⁴, A Nucciotti^{12,9}, I Nutini^{13,21}, T O'Donnell²⁰, J L Ouellet¹⁴, C E Pagliarone^{13,19}, M Pallavicini^{17,8}, V Palmieri³, L Pattavina¹³, M Pavan^{12,9}, G Pagharone^{15,0}, M Pallavicini^{15,0}, V Palmieri¹, L Pattavina^{15,0}, M Pavan^{15,0}, G Pessina⁹, C Pira³, S Pirro¹³, S Pozzi^{12,9}, E Previtali⁹, F Reindl⁶, C Rosenfeld¹, C Rusconi^{1,13}, M Sakai², S Sangiorgio³⁰, D Santone^{13,22}, B Schmidt¹⁶, J Schmidt², N D Scielzo³⁰, V Singh⁷, M Sisti^{12,9}, L Taffarello¹⁰, F Terranova^{12,9}, C Tomei⁶, M Vignati⁶, S L Wagaarachchi^{7,16}, B S Wang^{30,31}, H W Wang¹⁵, B Welliver¹⁶, J Wilson¹, K Wilson¹, L A Winslow¹⁴, T Wise^{18,32}, L Zanotti^{12,9}, G Q Zhang¹⁵, S Zimmermann³³ and S Zucchelli^{28,4}

- ¹ Dept of Physics and Astronomy, Univ of South Carolina, Columbia, SC 29208, USA
- ² Dept of Physics and Astronomy, Univ of California, Los Angeles, CA 90095, USA
- ³ INFN Laboratori Nazionali di Legnaro, Legnaro (Padova) I-35020, Italy
- ⁴ INFN Sezione di Bologna, Bologna I-40127, Italy
- ⁵ Dipartimento di Fisica, Sapienza Università di Roma, Roma I-00185, Italy
- ⁶ INFN Sezione di Roma, Roma I-00185, Italy
- ⁷ Department of Physics, University of California, Berkeley, CA 94720, USA
- ⁸ INFN Sezione di Genova, Genova I-16146, Italy
- ⁹ INFN Sezione di Milano Bicocca, Milano I-20126, Italy
- ¹⁰ INFN Sezione di Padova, Padova I-35131, Italy
- ¹¹ Dipartimento di Fisica e Astronomia, Università di Padova, I-35131 Padova, Italy
- ¹² Dipartimento di Fisica, Università di Milano-Bicocca, Milano I-20126, Italy
- ¹³ INFN Laboratori Nazionali del Gran Sasso, Assergi (L'Aquila) I-67100. Italy
- ¹⁴ Massachusetts Institute of Technology, Cambridge, MA 02139, USA
- ¹⁵ Shanghai Institute of Applied Physics, Chinese Academy of Sciences, Shanghai 201800, China
- ¹⁶ Nuclear Science Division, Lawrence Berkeley Nat Lab, Berkeley, CA 94720, USA ¹⁷ Dipartimento di Fisica, Università di Genova, Genova I-16146, Italy

IOP Conf. Series: Journal of Physics: Conf. Series 1056 (2018) 012009 doi:10.1088/1742-6596/1056/1/012009

¹⁸ Wright Laboratory, Department of Physics, Yale University, New Haven, CT 06520, USA

¹⁹ Dipartimento di Ingegneria Civile e Meccanica, Università degli Studi di Cassino e del Lazio Meridionale, Cassino I-03043, Italy

²⁰ Center for Neutrino Physics, Virginia Polytechnic Institute and State University, Blacksburg, Virginia 24061, USA

²¹ INFN – Gran Sasso Science Institute, L'Aquila I-67100, Italy

²² Dip di Scienze Fisiche e Chimiche, Università dell'Aquila, L'Aquila I-67100, Italy

²³ INFN – Laboratori Nazionali di Frascati, Frascati (Roma) I-00044, Italy

²⁴ CSNSM, Univ. Paris-Sud, CNRS/IN2P3, Univ. Paris-Saclay, 91405 Orsay, France
 ²⁵ Physics Dept, California Polytechnic State Univ, San Luis Obispo, CA 93407, USA

²⁶ INPAC and School of Physics and Astronomy, Shanghai Jiao Tong University;

Shanghai Laboratory for Particle Physics and Cosmology, Shanghai 200240, China

²⁷ Lab de Fisica Nuclear y Astroparticulas, Univ de Zaragoza, Zaragoza 50009, Spain

²⁸ Dipartimento di Fisica e Astronomia, Università di Bologna, Bologna I-40127, Italy

²⁹ Service de Physique des Particules, CEA / Saclay, 91191 Gif-sur-Yvette, France

³⁰ Lawrence Livermore National Laboratory, Livermore, CA 94550, USA

³¹ Dept of Nuclear Engineering, University of California, Berkeley, CA 94720, USA

³² Department of Physics, University of Wisconsin, Madison, WI 53706, USA

³³ Engineering Division, Lawrence Berkeley National Lab, Berkeley, CA 94720, USA

^a Deceased

E-mail: chiara.brofferio@unimib.it

Abstract. The Cryogenic Underground Observatory for Rare Events (CUORE) is the first bolometric experiment searching for neutrinoless double beta decay that has been able to reach the 1-ton scale. The detector consists of an array of 988 TeO₂ crystals arranged in a cylindrical compact structure of 19 towers. The construction of the experiment and, in particular, the installation of all towers in the cryostat was completed in August 2016 and commissioning started in fall 2016. The experiment has completed the pre-operation phase and is currently in data taking. We present here the achievements of CUORE during the commissioning phase and the limit on the ¹³⁰Te half-life for the neutrinoless double beta decay that has been released after the first 3 weeks of collected data. Physics results from CUORE-0 will also be updated.

1. Introduction

Neutrinoless Double Beta Decay (0vDBD) searches [1][2][3] represent a unique tool to assess the Dirac/Majorana nature of neutrino and to check the Lepton Number Conservation law. The experimental approach is usually based on the observation of a large number of nuclei, stable in normal beta decay but for which a double weak interaction process, changing the nuclear charge by two units, is energetically favorable. The decay involving the simultaneous emission of two electrons and two neutrinos (2vDBD) is allowed by the Standard Model of Electroweak Interactions and has been observed for several nuclides [4]. However, if neutrinos have a Majorana nature (and mass), a process without the emission of any neutrino is possible. A 0vDBD is an unambiguous signal of a Majorana mass and of a violation of the lepton number conservation. Moreover, the study of such decay can put a constraint on the absolute neutrino mass scale, which is at the moment unknown [5].

Most sensitive experiments are presently based on the homogeneous approach in which the nuclei under observation are part of the detector itself. The sensitivity of such experiments scales with the square root of the exposure (the live time times the detector mass) and the inverse of the observed background rate, since it is supposed that this also scales with the detector mass.

Conference on Neutrino and Nuclear Physics (CNNP2017)IOP PublishingIOP Conf. Series: Journal of Physics: Conf. Series 1056 (2018) 012009doi:10.1088/1742-6596/1056/1/012009

Low temperature detectors (LTD) are naturally suitable for this approach. These detectors are simply based on the possibility to register a temperature increase in a crystal, acting as absorber, when a particle interacts with it. In fact, the heat capacity of any cold diamagnetic and dielectric crystal is proportional to the cube of the ratio between the operating and Debye temperatures and can therefore become so small that even the tiny energy released by a particle in form of heat can be revealed by the temperature increase of the crystal by means of a suitable thermal sensor. Many interesting compounds have already been studied as candidates for this approach [6] with excellent results both on intrinsic radio-purity and on energy resolution, which is fundamental to reduce the number of counts due to 2vDBD at the Q-value of the decay in the 2-electrons energy sum spectrum, where the events from 0vDBD are expected. This background is in fact the only unavoidable one that will always be present.

2. The TeO₂ choice and CUORE-0

¹³⁰Te is an optimal candidate to search for 0vDBD [7] due to its high transition energy (2527.5 keV), and large isotopic abundance (34.2 %), which allows performing a sensitive experiment with natural tellurium. TeO₂ has demonstrated to be an optimal choice for LTD crystals to be used in 0vDBD searches, thanks to its high enough Debye temperature [8], good crystal properties, very low intrinsic radioactivity [9] and favorable Te mass percentage in the compound (27%). TeO₂ LTDs have been pioneered by the Milano group in a series of constantly increasing mass (from 6 g to 6.8 kg) experiments carried out at Laboratori Nazionali del Gran Sasso (LNGS) [10], which opened the way to CUORE [11], a 1-ton LTD experiment for ¹³⁰Te 0vDBD.

Suggested by Ettore Fiorini in 1998, CUORE physics potential attracted many international collaborators, that convinced the community (and the funding agencies) of the feasibility of the project with a smaller (41 kg) experiment, CUORICINO [12]. This experiment was in fact able to show not only the technical possibility of scaling in mass by using arrays of LTD towers, but also that the technique was mature to compete in background reduction (0.169 c/keV kg y) and 0vDBD half-life limits (2.8×10^{24} y at 90% C.L.) [13]. Once funded, CUORE detector was designed and optimized under all the possible aspects [14][15], to push the sensitivity as far as possible [16]. The demonstrator of all these efforts has been the single tower CUORE prototype, CUORE-0.

CUORE-0 consisted of 52 5×5×5 cm³ TeO₂ crystals assembled in a tower of 13 floors for a total TeO₂ mass of 39 kg. All the technical details of the detector and its performance can be found in the literature [17]. CUORE-0 was not only the first prototype of a CUORE tower, but a small scale competing experiment, that after a 9.8 kg y exposure was able to improve the background (0.058 c/keV kg y) and 0vDBD half-life limits (2.7×10^{24} y at 90% C.L.) with respect to CUORICINO [18] and to measure with the highest precision ever obtained the half-life of ¹³⁰Te 2vDBD ($T_{2\gamma} = (8.2 \pm 0.2_{stat} \pm 0.6_{syst}) \times 10^{20}$ y) [19]. Many other searches are still under study, like the decays of ¹²⁰Te and of ¹³⁰Te on the first 0⁺ excited state, and will be published soon.

A detailed Monte Carlo simulation that exploits the information about the contaminations of materials obtained through radio-assay screening campaigns and bolometric measurements was built and fine-tuned on CUORE-0 data with excellent results [19]. The information extracted from CUORE-0 background model were then used to tune the general CUORE Monte Carlo simulation [20], that confirmed that the goal of a Background Index (BI) in the Region Of Interest (ROI) of 10^{-2} c/keV kg y was within reach. With CUORE-0 the 5 keV FWHM energy resolution goal was also achieved, on the full 9.8 kg y exposure.

3. CUORE construction and commissioning

The CUORE detector is composed by a closely packed array of 19 towers CUORE-0-like, for a total of 988 TeO₂ bolometers operated at a temperature of \sim 10 mK in a custom-made cryostat installed at Laboratori Nazionali del Gran Sasso (LNGS) in Italy. Each bolometer is equipped with a neutron transmutation doped (NTD) thermal sensor and a silicon heater, for gain stabilization [21][22]. With a

total detector mass around 740 kg of TeO_2 (206 kg of ¹³⁰Te), CUORE is the first ton-scale cryogenic detector ever put into operation.

To cope with the strong requirements for surface (re)contamination, the detector components were produced, handled and cleaned according to specific protocols, the towers were assembled by use of specific glove-boxes always flushed with nitrogen to avoid ²²²Rn or ²¹⁰Pb [23], and their installation in the cryostat was performed using a radon free modular clean room [24]. To shield the detector from gamma and neutron environmental background different layers of copper, lead (70 ton aprox, of which 8 ton inside the cryostat, at low temperature), borated polyethylene and boric acid are used.

The installation of the 19 towers was successfully completed in Summer 2016, obtaining 984 functioning bolometers out of 988. The cryostat interfaces and radiation shields were assembled in the following months. Cool down started in Dec. 2016. The custom-made dilution refrigerator system works without LHe and the cooling at intermediate temperatures (~ 40 K and ~ 4 K) is based on five pulse tube cryo-coolers. The dimensions, experimental volume (~ 1 m³), mass (~17 t), and cooling power (3 μ W at 10 mK) make it the largest and most powerful cryogen-free dilution refrigerator system in operation. Nonetheless, it took almost 1 month to cool to base temperature. A stable temperature of ~ 7 mK was reached on January 27, 2017. We then started the detector pre-operation phase, to optimize the signal readout, the mechanical and electrical noise reduction and the working points of the bolometers. In April 2017 the optimization was not yet completed but the Collaboration decided to start a preliminary science run to extract the very first hints on energy resolution and BI in the ROI.

4. First data-set results

In May 2017, we collected three weeks of physics data bracketed by two calibration periods. During calibration 12 "sausage-like" strings containing ²³²Th sources are temporarily deployed inside the detector region in order to guarantee an approximately uniform γ -ray illumination of the detectors [25]. Six γ -lines (from 239 keV to 2615 keV) are then used to perform the energy calibration of the 984 bolometers. During the physics run we acquired an exposure of 38.1 kg y of TeO₂ (10.6 kg y of ¹³⁰Te), already greater than the total exposure collected by CUORE-0.

For this initial analysis 889 detectors (~90 %) were used, discarding those that would have required more efforts on stability and noise reduction. In our analysis approach, once the calibration is successfully applied, the physics data are blinded. This is obtained by introducing an artificial peak at the Q-value. The model and fitting strategy are then optimized and fixed before unblinding. A detailed description of the full procedure, that was developed for CUORE-0 and then used on this first dataset, is available in the literature [26]. The physics spectrum, see figure 1, is built after applying a series of selection criteria aimed at improving the experimental sensitivity. We therefore remove periods of low quality data (caused by noisy laboratory conditions) and reject pile-up events. Then we select only signals consistent with a proper template waveform (pulse shape analysis) in order to identify real particle events and finally we exclude events that simultaneously trigger more than one crystal, to reduce the background due to events depositing energy in multiple crystals.

For this first run we evaluated an overall detection efficiency of $(55.3 \pm 3.0)\%$, which includes a $(88.35 \pm 0.09)\%$ probability that a 0vDBD is fully contained in a single crystal and a $(62.6 \pm 3.4)\%$ probability that a physics event is not discarded when the selection criteria are applied.

We evaluate the detector energy resolution near the ROI by fitting the 2615 keV line in the physics spectrum. The harmonic mean of the detector FWHM resolutions for this preliminary run has been found to be 7.9 ± 0.6 keV.

To estimate the background in the ROI and the 0vDBD rate ($\Gamma_{0\nu}$) we perform an Unbinned Extended Maximum Likelihood fit in the [2465-2575] keV range (the ROI) with the same procedure used for CUORE-0 [26]. The best-fit values are $0.98^{+0.17}_{-0.15} \times 10^{-2}$ counts/(keV kg y) for the background rate, and ($-0.03^{+0.07}_{-0.04}$ (stat) ± 0.01 (syst)) $\times 10^{-24}$ y⁻¹ for $\Gamma_{0\nu}$. We find no evidence for the 0vDBD of ¹³⁰Te and we can only calculate an upper limit of $\Gamma_{0\nu}$ by integrating the profile

IOP Conf. Series: Journal of Physics: Conf. Series 1056 (2018) 012009 doi:10.1088/1742-6596/1056/1/012009

likelihood in the physical region ($\Gamma_{0\nu} \ge 0$). This corresponds to a half-life lower limit of $T_{1/2}^{0\nu} > 4.5 \times 10^{24}$ y (90% C.L.).



Figure 1: Comparison of physics spectra in the gamma region measured with CUORE and CUORE-0, with prominent γ -lines labeled.

Finally, we combine the first results from CUORE with those obtained from CUORE-0 [18] and Cuoricino [13] with 9.8 kg y and 19.8 kg y exposure of ¹³⁰Te, respectively. The half-life lower limit obtained by combining the profile negative-log-likelihood curves of the three experiments is $T_{1/2}^{0\nu} > 6.6 \times 10^{24}$ y (90% C.L.).

The combined half-life limit can be then interpreted as a limit on the effective Majorana neutrino mass $(m_{\beta\beta})$ in the framework of models of 0vDBD mediated by light Majorana neutrino exchange. When using the phase-space factors from [27] and nuclear matrix elements from a broad range of recent calculation models [28][29][30][31][32], with the nucleon axial coupling constant $g_A = 1.27$ we get $m_{\beta\beta} < (210-590)$ meV at 90% C.L., depending on the nuclear matrix element estimates employed. We do not consider other g_A values since there is still quite a lack of certainties on what to expect to be the correct number to be used for 0vDBD. See for instance the discussions in [33] and [34].

5. Another optimization campaign and a new science run

After this preliminary science run another optimization campaign of the detector operating conditions was started. We implemented an active noise cancellation system on the pulse tube cryocoolers using micro-step motor linear drives to control the relative phases of the pulse tube pressure oscillations and fixing the configuration that maximizes the noise cancellation, leveraging the interference between the noise sources [35]. We improved the electrical grounding of the experiment and we performed a temperature scan of the base temperature of the detectors, choosing to work at 15 mK. We optimized the software bandwidth for the pulse amplitude analysis and extended the software trigger window from 5 s to 10 s.

A new science run was then carried out during August 2017. The corresponding dataset as well as the one described in this paper were completely re-processed with a slightly improved analysis procedure and results were presented after the CNNP conference, during CUORE inauguration at LNGS on October, 23rd and will be published on PRL [36].

Conference on Neutrino and Nuclear Physics (CNNP2017)IOP PublishingIOP Conf. Series: Journal of Physics: Conf. Series 1056 (2018) 012009doi:10.1088/1742-6596/1056/1/012009

Acknowledgments

The CUORE Collaboration thanks the directors and staff of the Laboratori Nazionali del Gran Sasso and the technical staff of our laboratories. This work was supported by the Istituto Nazionale di Fisica Nucleare (INFN); the Director, Office of Science, of the U.S. Department of Energy under Contract Nos. DE-AC02-05CH11231 and DE-AC52-07NA27344; the DOE Office of Nuclear Physics under Contract Nos. DE-FG02-08ER41551 and DEFG03-00ER41138; the National Science Foundation under Grant Nos. NSF-PHY-0605119, NSF-PHY-0500337, NSF-PHY-0855314, NSF-PHY-0902171, and NSF-PHY-0969852; the Alfred P. Sloan Foundation; the University of Wisconsin Foundation; and Yale University. This research used resources of the National Energy Research Scientific Computing Center (NERSC).

References

- [1] Cremonesi O and Pavan M 2014 *Adv. High Energy Phys.* 2014 951432
- [2] Dell'Oro S et al. 2016 Adv. High Energy Phys. 2016 2162659
- [3] Giuliani A and Poves A 2012 Adv. High Energy Phys. 2012 857016
- [4] Patrignani C et al. (Particle Data Group) 2016 Chin. Phys. C 40:100001
- [5] Ge S and Lindner M 2017 *Phys Rev* D 95 033003
- [6] Kim Y H 2017 Proc. 17th Int. Workshop on Low Temperature Detectors Kurume City, Japan, July 17-21. To appear on a special issue of J Low Temp Phys
- Barabash A S 2012 J. Phys. G: Nucl. Part. Phys. 39 085103
 Artusa D R et al. 2014 Eur. Phys. J. C 74 3096
- [8] Barucci M 2001 J. Low Temp. Phys. 123 303
- [9] Arnaboldi C et al. 2010 J. Cryst. Growth 312 2999
- [10] Brofferio C and Dell'Oro S 2018 arXiv: 1801.03580 [hep-ex]
- [11] Artusa D R et al. 2015 Adv. High Energy Phys 2015 879871
- [12] Arnaboldi C et al. 2008 Phys. Rev. C 78 035502
- [13] Andreotti E et al. 2011 Astropart. Phys. **34** 822
- [14] Alessandria F et al. 2012 Astropart. Phys. 35 839
- [15] Alessandria F et al. 2013 Astropart. Phys. 45 13
- [16] Alduino C *et al.* 2017 *Eur. Phys. J.* C **77** 532
- [17] Alduino C *et al.* 2016 *J. Instrum.* **11** P07009
- [18] Alfonso K *et al.* 2015 *Phys. Rev. Lett.* **115** 102502
- [19] Alduino C et al. 2017 Eur. Phys. J. C 77 13
- [20] Alduino C et al. 2017 Eur. Phys. J. C 77 543
- [21] Andreotti E et al. 2012 Nucl. Instrum. Meth. A 664 161–170
- [22] Carniti P et al. 2017 arXiv:1710.05565 [physics.ins-det]
- [23] Buccheri A et al. 2014 Nucl. Instrum. Meth. A 768 130
- [24] Benato G et al 2018 J. Instrum. **13** P01010
- [25] Cushman J S et al. 2017 Nucl. Instrum. Meth. A 844 32–44
- [26] Alduino C et al. 2016 Phys. Rev. C 93 045503
- [27] Kotila J and Iachello F 2012 Phys. Rev. C 85 034316
- [28] Barea J, Kotila J and Iachello F 2015 Phys. Rev. C 91 034304
- [29] Simkovic F, Rodin V, Faessler A and Vogel P 2013 Phys. Rev. C 87 045501
- [30] Hyvarinen J and Suhonen J 2015 Phys. Rev. C 91 024613
- [31] Menendez J, Poves A, Caurier E and Nowacki F 2009 Nucl. Phys. A 818 139
- [32] Rodriguez T R and Martinez-Pinedo G 2010 Phys. Rev. Lett. 105 252503
- [33] Suhonen J 2017 Phys. Rev. C 96 055501
- [34] Suhonen J 2017 Frontiers in Physics 5 55
- [35] D'Addabbo A *et al.* 2017 *arXiv*:1712.02753 [physics.ins-det]
- [36] Alduino C et al. 2018 Phys. Rev. Lett. in the press (arXiv:1710.07988 [nucl-ex])