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1 **Response of the East Antarctic Ice Sheet to orbital forcing during the Pliocene and Early**
2 **Pleistocene**

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28 **The absence of a strong ~20-kyr precession signal in Early-Pleistocene to Pliocene**
29 **geological reconstructions of global ice volume and sea-level challenges our fundamental**
30 **understanding of how ice sheets respond to orbital forcing. Here, we present an orbital-**
31 **scale time-series of ice-berg rafted debris from a well-dated sediment core (Integrated**
32 **Ocean Drilling Program (IODP) site U1361) adjacent to the Wilkes Land margin of the**
33 **East Antarctic Ice Sheet (EAIS). Our data reveal that under a warmer climate state, mean**
34 **annual insolation, paced by the 41-kyr obliquity cycle, had more influence on Antarctic ice**
35 **volume than summer insolation intensity modulated by the precession cycle. A transition to**
36 **precession dominance after ~3.5 Ma reflects a declining influence of oceanic forcing as the**
37 **high latitude southern ocean cooled and a perennial summer sea-ice field developed. Our**
38 **data show that precession-paced variability of the EAIS marine margin occurs during cold**
39 **climate states, even when the obliquity signal dominates globally-integrated proxy records.**

40

41 Geological reconstructions of global ice volume¹ and sea-level² during the Pliocene and
42 Early Pleistocene (5 to 2 Ma) display regular glacial-interglacial cycles occurring every 41-kyrs,
43 paced by variations in Earth's axial tilt (obliquity). The absence of a strong ~20-kyr precession
44 signal prior is confounding, because precession should impart the greatest influence on high-
45 latitude summer insolation intensity, and therefore polar ice volume³. Indeed, precession is a
46 major control on the orbital pacing of Late Pleistocene cycles¹, and while a number of

47 hypotheses have been proposed to explain its absence prior to this time^{4, 5, 6}, reconciliation of this
48 conundrum remains hampered by direct observational evidence of orbital-scale Plio-Pleistocene
49 variance of ice sheets, in particular from Antarctica^{7, 8, 9}.

50 IODP core U1361 was recovered from ~3000 m water depth on the continental rise
51 adjacent to the Wilkes Land sector of Antarctica, one of the largest marine based sectors of the
52 EAIS, where recent geophysical surveys show the presence of landward deepening reverse slope
53 troughs penetrating beneath the EAIS reaching depths of up to 2 km below sea-level¹¹, heightening
54 the vulnerability of this sector of the EAIS to marine ice sheet instabilities (Fig. 1). U1361
55 provides a well-dated and continuous geological archive of Pliocene and Early Pleistocene
56 orbital scale variability of the marine margin of the EAIS. Sediment deposition at this site is
57 controlled by the interplay between: (i) downslope marine sediment gravity flows triggered by
58 the buildup of sediment on the edge of the continental shelf during glacial advance; (ii) the
59 rainout of biogenic detritus from surface water plankton; (iii) iceberg rafting of terrigenous
60 sediments; and (iv) low energy bottom currents (Supplementary Information).

61 The core consists of eighteen sedimentary cycles spanning an age range of 4.3-2.2 Ma, and
62 comprising alternating terrigenous massive to laminated muds and diatom-rich/bearing silty-mud
63 units (cycles 1-18 Fig. 2; Supplementary Figs. S2 and S3). In places the muds contain packages
64 of well-defined laminae and are consistent with established models of non-erosive overbank
65 hemipelagic deposition onto a channel levee setting via turbidites on the lowermost Antarctic
66 continental rise¹⁰. Steeply dipping, seaward prograding wedge sediments are evident in seismic
67 reflection profiles across the continental shelf and extend onto the upper continental slope above
68 seismic unconformity WL-U8 (~4.2 Ma; Supplementary Fig. S1)^{12, 13}. The geometry of these
69 strata is characteristic of grounding zone deposition by repetitive advances of a marine-based ice

70 sheet to the shelf edge during glacial periods¹³. Sediment overloading near the shelf break at
71 submarine canyons heads, in turn triggers turbidity currents down slope channels and leads to
72 overbank deposition at the core site. Low density turbidity currents in overbank “distal” channel
73 levee environments on the Antarctic continental rise are typically non-erosive¹⁰ (Supplementary
74 Information). Thus, turbidite units are associated with periods of glacial advance to the Wilkes
75 Land continental shelf edge, whereas bioturbated, diatom-rich/bearing facies represent warm
76 interglacial periods of relatively ice-free ocean and increased primary productivity when the
77 grounding line had migrated landward away from the shelf edge. Increased productivity during
78 interglacial warm climates may be associated with enhanced upwelling of nutrient-rich
79 Circumpolar Deepwater (CDW)¹⁴ (Supplementary Information), which has been linked to
80 southward expansion of the westerly wind field in response to a reduced pole-equator
81 temperature gradient during past warm periods¹⁵. Presently this relatively warm nutrient-rich
82 CDW upwells to the surface north of the Southern Boundary Front of the Antarctic Circumpolar
83 Current and is marked by areas of enhanced productivity to the north of Site U1361 (Fig. 1).

84 We have developed a high-resolution record (~3-4 kyr sample spacing) of ice-berg rafted
85 debris (IBRD) mass accumulation rates (MAR) for the U1361 core (Supplementary Table S1;
86 Supplementary Fig. S4). Our age model is based on biostratigraphy used to constrain the
87 interpretation of a magnetic polarity stratigraphy¹² (Supplementary Fig. S5), and is not orbitally-
88 tuned. Thus, the age model for our primary spectral analyses (Fig. 3) assumes, long-term
89 (millennial-scale) sedimentation rates between polarity reversal tie points (Supplementary
90 Information).

91 In general the highest intensity of IBRD occurs during transitions from glacial terrigenous
92 mud facies to interglacial diatom-rich/bearing muds up-core until ~47 mbsf, with most IBRD

93 peaks immediately preceding diatom-rich facies and Ba/Al peaks (Fig. 2). Isotopic Nd and Sr
94 provenance indicators suggest that the terrigenous components in these diatom-rich/bearing
95 muds are associated with periods of deglacial retreat of the ice margin back into the Wilkes Land
96 subglacial basin during the Early Pliocene¹⁶. The Antarctic ice sheet loses the majority of its
97 mass via iceberg calving and sub-ice shelf melting¹⁷. However, these processes are intimately
98 linked as enhanced ice shelf melt leads to reduced buttressing, that in turn acts to enhance flow
99 of glacial tributaries in turn enhancing iceberg discharge¹⁸. Ice sheet reconstructions based on
100 geological data¹⁹ and modeling⁸ suggest that the above processes dominated Antarctic mass-loss
101 even during the warmest climates of the last 5 Ma. Thus, we interpret the maxima in IBRD
102 MAR to be the consequence of accelerated calving during glacial retreat from marine
103 terminating outlet glaciers along the Wilkes Land coastline as well as a contribution from EAIS
104 outlet glaciers entering the western Ross Sea (Supplementary Information). This interpretation is
105 consistent with models and paleo-data, which imply the most rapid mass loss of the EAIS margin
106 during the last glacial termination occurred between 17-7 ka, and was primarily the consequence
107 of oceanic warming^{20,21}.

108 Spectral analysis of the un-tuned IBRD MAR time-series displays a dominant period of
109 ~40-kyr between ~4.2-3.5 Ma that transitions to strong and significant variance at ~100, and 20-
110 kyr periods after ~3.5 Ma, with a corresponding decrease in power of the ~40-kyr cycle (Figs. 3,
111 4). On the basis of the strong orbital relationship displayed in frequency spectra of the un-tuned
112 IBRD MAR time series, the paleomagnetic age model, and the apparent near continuous long-
113 term sedimentation (Supplementary Fig. S5), we established a nominal one-to-one correlation
114 between cycles in ice margin variability expressed by our IBRD data and orbitally-paced
115 climatic time series (Fig. 2). Between 4.3-3.4 Ma a strong visual correlation can be observed

116 between 40-kyr cycles in IBRD, mean annual insolation and the benthic $\delta^{18}\text{O}$ global ice volume
117 record¹, whereas between 3.3-2.2 Ma, IBRD cycles correlate with the ~20-kyr cycles of summer
118 insolation at 65°S (Fig. 2). We acknowledge that our visual correlations, although constrained by
119 7 precisely-dated paleomagnetic reversals, do not represent a unique solution but, as noted
120 above, they are entirely consistent with the variance in orbital frequencies implied by our
121 spectral estimations.

122 Using these correlative relationships (Fig. 2) and our evolutionary spectral analyses on un-
123 tuned IBRD data (Fig. 3), we then explore the role of longer-period orbital influences on the
124 pattern of iceberg calving. The top of the ~40-kyr-dominated interval is marked by an ~200 kyr-
125 long condensed section between ~3.5-3.3 Ma (Supplementary Fig. S5) and corresponds to a
126 long-term +1‰ glacial $\delta^{18}\text{O}$ excursion punctuated by smaller interglacials spanning Marine
127 Isotope Stage (MIS) MG5 and MIS M2 (indicated by arrow in Fig. 2). The stratigraphic
128 condensation, or possible erosion, at Site U1361 is associated with this glacial excursion. Indeed,
129 this glacial excursion has also been associated with southern high-latitude climate cooling and
130 the re-establishment of grounded ice on the middle to outer continental shelf in the Ross Sea
131 following a ~200-kyr period of warm open ocean conditions^{7,9}. Previous studies of older
132 Oligocene and Miocene $\delta^{18}\text{O}$ glacial excursions have proposed a relationship between intervals
133 of increased glacial amplitude in the $\delta^{18}\text{O}$ record with a coincidence of 1.2 Ma nodes in obliquity
134 and 400-kyr minima in long period eccentricity^{22,23}. This orbital configuration, which favors
135 extended periods of cold summers and low seasonality is considered optimal for Antarctic ice
136 sheet expansion, and occurs at ~3.3 Ma (Fig. 3c) - the time of the transition from obliquity to
137 precession dominance in the IBRD MAR time series from U1361 (Figs. 2, 3a).

138 Observed ~20-kyr-duration IBRD cycles correlate with summer insolation calculated for
139 65°S for the interval of the core between 3.3-2.2 Ma (Fig. 2; Fig. 3a) and are embedded within
140 100-kyr-duration IBRD cycles with an additional peak at 400 kyr nearing the 90% confidence
141 level (Fig 3d). Broad peaks of IBRD maxima are associated with transitions between laminated
142 mudstones to diatom-rich/bearing muds (Fig. 2). A dramatic decrease in the amplitude of ~20-
143 kyr IBRD peaks, and a change to lithofacies associated with non-erosive low-energy bottom
144 currents, is observed at the core site above ~2.5 Ma, broadly coincident with southern high
145 latitude cooling⁹ and the onset of major Northern Hemisphere glaciations²⁴. We attribute the
146 progressive reduction in calving intensity to cooling and a relative stabilization of the EAIS ice
147 margin (discussed below). Homogenization of the turbidite sediments during glacial maxima by
148 enhanced bioturbation and bottom current activity is observed and likely reflects increased
149 Antarctic sea ice and polynya-style mixing at this time producing oxygenated high salinity shelf
150 water⁹ that is transferred downslope over Site U1361 to form Antarctic Bottom Water²⁵
151 (Supplementary Information).

152 In summary, our correlations of IBRD variations with the benthic $\delta^{18}\text{O}$ stack and orbital
153 time series identify up to seventeen ~40-kyr-duration cycles (orange lines in Fig. 2) within six
154 major lithological cycles (cycles 13-18, Fig. 2) during the Early Pliocene (4.3-3.5 Ma). This is
155 followed by forty ~20-kyr-duration cycles (blue lines in Fig. 2) modulated by eccentricity, that
156 occur within twelve lithological cycles (cycles 1-12 Fig. 2).

157 Although the marine sediment core recovered by the ANDRILL Program from the Ross
158 Sea region provided the first direct evidence that advance and retreat of the WAIS margin across
159 the continental shelf was paced by obliquity during the Pliocene prior to ~3 Ma⁷, sub-glacial
160 erosion surfaces in the ANDRILL core associated with ice advance have raised the possibility of

161 missing cycles, particularly after 3.1 Ma - e.g., the response of WAIS to orbital forcing is more
162 ambiguous after this time. The U1361 and ANDRILL records together confirm the dynamic
163 response of both the WAIS and EAIS marine margins to obliquity forcing during the warm
164 Pliocene prior to the onset of southern high-latitude cooling at 3.3 Ma⁹.

165 Geological records^{7, 9, 16} and model simulations⁸ of past warm climates both highlight the
166 sensitivity of the marine-based portions of the Antarctic ice sheets to ocean warming. However,
167 the mechanism by which the coastal ocean warms and destabilizes marine grounding lines,
168 particularly in response to obliquity forcing, remains elusive. It has been proposed that changes
169 in the intensity and the meridional distribution of mean annual insolation controlled by obliquity
170 may have a profound influence on the position and strength of the Southern Hemisphere zonal
171 westerly winds, with implications for marine ice sheet instability⁷. Indeed, an aerosolic dust
172 record from the Southern Ocean is dominated by ~40-kyr cycles in iron and leaf-wax biomarkers
173 prior to ~0.8 Ma²⁶. Moreover, prior to ~3.3 Ma the southward expansion of the westerly wind-
174 field over the Antarctic circumpolar convergence zone under a reduced meridional temperature
175 gradient has been associated with a reduced sea-ice field⁹. This may have caused the upwelling
176 of warm, CO₂-rich Circumpolar Deep Water (CDW) onto the continental shelf with
177 consequences for the migration of marine grounding-lines^{15, 27, 28}.

178 The dominance of precession-paced variability and the corresponding reduction in
179 obliquity influence revealed by our data after ~3.5 Ma is interpreted to reflect a declining
180 sensitivity of the EAIS to oceanic forcing, as the southern high latitudes cooled. Both model and
181 geological reconstructions imply that past Antarctic ice sheet expansion is closely linked with
182 development of the sea-ice field²⁹ potentially resulting in northward migration of westerly winds
183 and Southern Ocean fronts⁹. In addition, sea ice expansion after 3.5 Ma⁹ likely restricted

184 upwelling and ventilation of warm CO₂ rich CDW at the Antarctic margin acting to further
185 enhance climate cooling. Under such a scenario, a warmer climate state during the Early to mid-
186 Pliocene with higher atmospheric CO₂ concentration^{30, 31}, the increased duration and intensity of
187 austral summer surface warming, produces a pattern approaching mean annual insolation
188 regulated by obliquity (Fig. 2d), rather than precession that cancels out over the course of a
189 seasonal cycle⁵. Late Pliocene cooling raised the melt threshold such that the duration of the melt
190 season was restricted to times of austral summer insolation maxima controlled by precession
191 (Fig. 2c), and conversely extends the winter sea ice growth season³². We propose this decrease in
192 radiative forcing resulted in extensive sea-ice cover extending into much of the summer season
193 limiting the influence of upwelling CDW on marine grounding line stability.

194 The significant variance at ~20 kyr precession and coincident increase in the ~100-kyr
195 eccentricity frequency bands between 3.5-2.5 Ma is intriguing (Figs. 3a, d, and Fig. 4) and is
196 similar to the pattern of orbital response expressed in ~100-kyr-duration glacial-interglacial
197 cycles in the benthic δ¹⁸O and greenhouse gas ice core records of the last 600-kyr^{33, 34}. Thus the
198 relative dominance of eccentricity in EAIS at the Wilkes Land marine margin variability after
199 3.5 Ma may reflect a threshold-response of the ice sheet to orbital forcing under a colder climate
200 regime, as has been proposed to explain the relative dominance of ~100-kyr cycles in global Late
201 Pleistocene ice volume, albeit with different mass balance forcings^{35, 36, 32}. Here we propose that
202 the precession response of perennial sea-ice after 3.3 Ma was modulated by eccentricity (Fig. 4),
203 such that aerial extent of sea ice was significantly reduced every ~100 kyr enhancing the
204 influence of oceanic warming at those times.

205 Our data are also consistent with a reduction of obliquity variance in the benthic δ¹⁸O^{1, 37} at
206 ~3.5 Ma (Fig 3b), coincident with an obliquity node in the astronomical solution³⁸. The gradual

207 re-emergence of a 40-kyr signal at ~3 Ma in benthic $\delta^{18}\text{O}$ record³⁷ (Figs. 3b and 4c) most likely
208 reflects a similar reemergence of strong obliquity forcing in the orbital records and possibly also
209 a direct response of the developing northern hemisphere continental ice sheets to obliquity
210 forcing. It is also possible that precession-driven, anti-phase oscillations in both hemispheric ice
211 volume histories may have cancelled out in globally integrated proxy records after 3-2.8 Ma (e.g.
212 ref. 4).

213 A recent reconstruction of Pliocene Arctic climate³⁹, ice sheet models⁴⁰, and records of
214 iceberg rafted debris from the north Pacific and north Atlantic oceans²⁴ preclude a significant
215 northern hemisphere contribution to the $\delta^{18}\text{O}$ signal of global ice volume prior to 2.8 Ma. Thus, it
216 is unlikely that an anti-phase argument can be used to explain the dominance of the obliquity
217 signal seen in our record prior to 3.0 Ma. The strong obliquity signal seen in our IBRD record
218 also suggests that the intensity of summer insolation was a not a direct control on the mass
219 balance of the EAIS prior to 3.0 Ma. The geometry of strata on the Wilkes Land continental shelf
220 indicate that the EAIS in this region periodically expanded towards the continental shelf edge
221 during glacial maxima in the Pliocene¹² and suggests most Antarctic ice volume variance at this
222 time was growth and retreat of the marine-based ice sheets (Supplementary Fig. S1). During the
223 Early Pliocene, when the sea ice field was reduced and the Wilkes Land margin of the EAIS was
224 in more direct contact with oceanic influences, iceberg calving occurred more regularly within
225 both glacial and interglacial facies. Based on the significant decrease in IBRD after 2.5 Ma (Fig.
226 2; 3) and Southern Ocean records inferring decreased SSTs^{9,41}, we also infer the marine margins
227 of the EAIS became less sensitive to ocean-induced melting compared to the WAIS²⁸. After this
228 time, the fully-glaciated East Antarctic ice volume (excluding drainage into marine embayments

229 of the Ross and Weddell Sea) likely fluctuated at a magnitude similar to that of Late Pleistocene
230 cycles, with a minimal contribution to the $\delta^{18}\text{O}$ signal^{8,20}.

231 Notwithstanding this relative stability of the marine-based sectors of EAIS, model results
232 suggest that ~20-kyr-duration fluctuations in the total Antarctic ice volume (i.e., including
233 WAIS) may have contributed up to 15m of the total amplitude during Late Pleistocene glacial-
234 interglacial cycles⁸ and, given the $\delta^{18}\text{O}$ composition of Antarctic ice, this could have offset a
235 larger out-of-phase precessional change in Northern Hemisphere ice volume (e.g. 20-40m)
236 resulting in an enhanced obliquity signal in globally integrated sea-level and ice volume proxy
237 records after 3.0-2.8 Ma (e.g., ref 4, Fig. 3b, Fig. 4). Alternatively, direct obliquity forcing of the
238 Northern Hemisphere (NH) ice sheet is supported by proxy evidence including ice rafted debris
239 records²⁴ and a recent dust flux record⁴², suggesting that NH ice sheet variability (marine-based
240 margins) and climate primarily responded to obliquity under a relatively warm NH climate state.
241 NH cooling and ice sheet growth across the mid-Pleistocene transition ~0.8 Ma has been
242 implicated in a similar switch to precession and eccentricity-paced glaciations, albeit by different
243 mass balance forcing mechanisms⁴³. In contrast, however, our results imply that Southern Ocean
244 sea-ice feedbacks caused a fundamentally different response of the marine-based sectors of the
245 EAIS under a cooler Late Pliocene/Early Pleistocene climate state, characterized by a dominance
246 of precession-paced variability.

247 We conclude that prior to 3.5 Ma, under a warm climate state, EAIS demonstrates high
248 sensitivity on orbital timescales to a relatively small increase in atmospheric CO_2 concentration
249 and mean global surface temperature. With atmospheric CO_2 concentrations and global surface
250 temperatures projected to remain above 400 ppm and $>+2^\circ\text{C}$ beyond 2100⁴⁴, our results have
251 implications for the equilibrium response of the Antarctic ice sheets, and suggest that the marine

252 margins of EAIS ice sheet as well as the marine-based WAIS, may become increasingly
253 susceptible to ocean warming with the potential for widespread mass loss raising sea-level by
254 meters over the coming centuries to millennia.

255

256 **Methodology**

257 20cc samples were treated with H₂O₂ to remove organic material and 2M NaOH to
258 remove biogenic opal for grain size analysis. The MAR of the coarse sand fraction was then
259 estimated using the following equation:

260

$$261 \quad \text{IBRD MAR} = \text{CS\%} * \text{DBD} * \text{LSR}$$

262

263 where IBRD MAR is the mass accumulation rate (g/cm²/k.y.), CS% is the terrigenous coarse-
264 sand weight percent, DBD is the dry-bulk density of the nearest value (g/cm³) and LSR is the
265 interval average linear sedimentation rate (cm/k.y.). Visual examination of every individual
266 sample for authigenic minerals and volcanic glass was conducted and these were absent,
267 indicating that the IBRD volume percent was directly equivalent to the terrigenous CS%⁴⁵.

268 Following application of the biostratigraphic/magnetostratigraphic age model¹² and the
269 astrochronologic age model derived in this study, we performed time series analysis using “astro:
270 An R package for Astrochronology⁴⁶. Prior to analysis, all data series were conservatively
271 resampled on a 5 kyr grid using piecewise linear interpolation (the median sampling intervals for
272 the IBRD-MAR data and benthic δ¹⁸O stack are 3 kyr and 5 kyr, respectively). Time-frequency
273 analyses utilized the multitaper method (MTM)⁴⁷, with a 500 kyr window (5 kyr step) and three
274 2π prolate tapers; each window was linearly detrended prior to analysis.

275

276 **Author Contributions**

277 MOP, RM, TN designed the study, conducted sedimentological and time series analyses and
278 wrote the paper. SRM carried performed time-frequency analysis. CE and HB led IODP
279 Expedition 318 and provide seismic reflection data interpretation. FJ and CE analysed XRF
280 geochemical data. LT led development of the age model. MR contributed to writing the
281 manuscript. All authors contributed to the interpretations.

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445

446 **Figure 1. Location of Site U1361 and bathymetry offshore of the Wilkes Land margin,**
447 **Antarctica.** Also shown is the location of the Miocene-Plesitocene ANDRILL AND-1B core
448 recovered in the northwestern corner of the Ross Ice Shelf, the southern boundary of the
449 Antarctic Circumpolar Current (ACC), the Mertz Glacier tongue and paleo ice sheet drainage
450 path (white arrows) extending off shore into a slope and rise canyon system. Black lines
451 represent seismic reflection profile tracks represented in Supplementary Figure S1.

452

453 **Figure 2. Depth series developed for IODP site U1361 sediment core between 4.4-2.2Ma of**
454 **(a) XRF-based Ba/Al, (b) IBRD MAR correlated with time series of (c) January insolation**
455 **and total integrated summer energy (where melt threshold $[t]=400\text{GJm}^{-2}$), (d) mean annual**
456 **insolation and total integrated summer energy (where melt threshold $[t]=250\text{GJm}^{-2}$), (e)**
457 **eccentricity³⁸, and (f) the stacked benthic $\delta^{18}\text{O}$ record¹.** Also shown is the down core
458 distribution of lithofacies, lithological cycles (transitional lithologies are represented by both

459 symbols) and magnetic polarity stratigraphy¹². Maxima in the productivity proxy Ba/Al coincide
460 with bioturbated/diatom-rich mudstone facies. Grey shaded ellipse denotes alignment between a
461 1.2 Ma node in (d) obliquity modulated mean annual insolation and (e) a 400-kyr minimum in
462 eccentricity which favours polar ice sheet growth and corresponds to (f) a 1‰ glacial $\delta^{18}\text{O}$
463 excursion culminating with MIS M2 (arrow). A significant increase in (f) $\delta^{18}\text{O}$ glacial values
464 from 2.7 Ma (arrow) corresponds with a marked decline in the amplitude of (a) IBRD and a
465 100ppm decrease in (g) reconstructed atmospheric CO_2 concentration⁴⁴.

466

467 **Figure 3. 2π MTM time-frequency analysis results (500 kyr window) displaying (a)**
468 **normalized MTM power for the U1361 IBRD-MAR data using the (untuned) age-model¹²**
469 **compared to (b) normalized MTM power for the LR04 $\delta^{18}\text{O}$ stack (1) and (c) MTM**
470 **amplitude for the eccentricity-tilt-precession (ETP) solution³⁸.** Also shown are 2π MTM
471 power spectral estimates for (d, f) the U1361 IBRD-MAR data using the untuned age-model¹²,
472 and (e, g) for the $\delta^{18}\text{O}$ stack¹, with red noise confidence levels using three different approaches.
473 In panels a-c, red indicates large values and blue indicates low values. Milankovitch orbital
474 frequencies are labelled. The results highlight the lack of a strong obliquity signal (40-kyr) in
475 3500-2500 Ma IBRD-MAR data and the relatively strong influence of \sim 400-kyr eccentricity
476 compared with the LR04 $\delta^{18}\text{O}$ stack for the same time interval.

477

478 **Figure 4. Quantitative assessment of the evolution of power in the precession band (0.04-**
479 **0.06 cycles/kyr), obliquity band (0.02-0.035 cycles/kyr), and eccentricity band (0-0.015**
480 **cycles/kyr), using the MTM time-frequency results in Figure 3.** The cumulative power in each
481 frequency band is determined for (a) the untuned IBRD-MAR data, (b) the astronomically-tuned

482 time IBRD-MAR data based on the tielines in Figure 2, and (c) for the LR04 $\delta^{18}\text{O}$ stack¹. These
483 results indicate a generally dominant obliquity signal > 3500 kyr, with a reduction in obliquity
484 and enhancement of eccentricity < 3500 kyr in the IBRD-MAR record. The IBRD-MAR
485 eccentricity power is attributable to the clipping of precession, which transfers precession
486 power to its eccentricity modulator. Clipped-precession power (eccentricity + precession)
487 increases dramatically in all three time series < 3500 ka. Coincident with the intensification of
488 northern hemisphere continental glaciations, obliquity comes to dominate again at ~ 2.8 Ma in the
489 LR04 $\delta^{18}\text{O}$ stack, but not in the IBRD-MAR data. These results imply a local insolation control
490 on EAIS variability < 3500 kyr.

491

492