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THE EVOLUITON OF A SYSTEM WITH A RANDOM HAMIITONTAN*

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(To be submitted to Physics or Fluids)

July 12, 1967

## ABSTTRA.CI

An exact equation of evolution, with the structure of the diffusion equation, is derived for the phase-space probability density of a system whose Hamiltonian is random. No assumption need be made about the magnitude of the fluctuations nor about their time scale. In the limit of short correlation time, the equation reduces to the equation or Birmingham, Northrop, and Fälthammar. The reduction of the Fokker-Planck equation to diffusion form is also demonstrated.

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## -1.

Certain problems of physical interest can be represented as a system evolving under a random Hamiltonian, whose ensemble is independent of the state of the system. The simplest example is that of a charged particle acted upon by a random electromagnetic field. This example has been studied by Birmingham, Northrop, and Falthammar, ${ }^{I}$ for the case of particle diffusion in $(\alpha, \beta)$ - space, for fields conserving the magnetic moment and longitudinal action. They derived a diffusion equation for the evolution of the density in $(\alpha, \beta)$ - space, upon the assumption of small fluctuations of the fields. The present note drops this assumption, extending the work of $3 N P$ to arbitrary fluctuations and to a general Hamiltonian.

A system with Hamiltonian $H(\Gamma, t)$ and phase-orbit $\Gamma_{S}(t)$ has the phase-space density

$$
\begin{equation*}
\rho(\Gamma, t)=\delta\left[\Gamma-\Gamma_{S}(t)\right] \tag{1}
\end{equation*}
$$

which evolves as

$$
\begin{equation*}
\frac{\partial \rho}{\partial t}=-L \rho, \tag{2}
\end{equation*}
$$

where

$$
\begin{equation*}
L \equiv\{, \mathrm{H}\} \equiv \dot{\Gamma} \partial_{\Gamma} \tag{3}
\end{equation*}
$$

Is the Liouville operator. ${ }^{2}$ The symbol $\dot{\Gamma} \partial_{\Gamma}$ represents a scalar product over the phase space:

$$
\begin{equation*}
\dot{\Gamma} \partial_{\Gamma} \equiv \Sigma\left(\dot{q} \partial_{q}+\dot{p} \partial_{p}\right) \tag{4}
\end{equation*}
$$

with summation over the $f$ degrees of freedom of the system. A randomness of $H(t)$ produces a randomness of $\Gamma_{S}(t)$ and correspondingly or $\rho(\Gamma, t)$. The randomness or $\Gamma_{S}(t)$ is (partially) characterized by the probability distribution in phase-space $P\left(\Gamma_{S} ; t\right)$. This quantity is identical to the statistical mean of $\rho(\Gamma, t)$, since

$$
\langle p\rangle(\Gamma, t) \equiv \int d \Gamma_{S} \delta\left(\Gamma-\Gamma_{S}\right) P\left(\Gamma_{S} ; t\right)=P(\Gamma ; t)
$$

We shall derive the equation of evolution for $P(\Gamma ; t)$.
Express 0 and $L$ in terms of their means and their rluctuations about their respective means:

$$
\begin{align*}
& \rho \equiv P+\delta \rho  \tag{5}\\
& L \equiv\langle\mathrm{~L}\rangle+\delta \mathrm{I}
\end{align*}
$$

Substitute (5) into (2), and then take the mean of (2). The result is

$$
\begin{equation*}
\mathrm{dF} / \mathrm{dt}=-\langle\mathrm{SI} \mathrm{SO}\rangle, \tag{6}
\end{equation*}
$$

where

$$
\begin{equation*}
d / d t \equiv(\partial / \partial t) \div\langle L\rangle \tag{7}
\end{equation*}
$$

is the convective derivative under the mean Hamiltonian. The difference between Eqs. (2) and (6) is

$$
\begin{equation*}
[(\mathrm{d} / \mathrm{dt})+\delta \mathrm{L}] \delta \rho=-\delta L P+\langle\delta L \delta \rho\rangle \tag{8}
\end{equation*}
$$

This equation was solved by BNF to first order in the fluctuations, but, since it is linear in $\delta \rho$, it can be solved exactly for $\delta \rho$ and thus for the quantity $\langle\delta L \delta \rho\rangle$ required in (6).

The exact equation of evolution is thus found to be

$$
\begin{equation*}
d \mathrm{P} / \mathrm{dt}=(1-\langle\delta L G\rangle)^{-1}\langle\delta L G \delta I\rangle P, \tag{9}
\end{equation*}
$$

where $G \equiv[(d / d t)+\delta L]^{-1}$. With the initial condition $\delta \rho(t=0)=0$, G takes the explicit form:
$G A(t)=\int_{0}^{t} d t^{\prime} A\left(t^{\prime}\right)-\int_{0}^{t} d t^{\prime} \delta L\left(t^{\prime}\right) \int_{0}^{t^{\prime}} d t^{\prime \prime} A\left(t^{\prime \prime}\right)$

$$
\begin{equation*}
+\int_{0}^{t} d t^{\prime} \delta L\left(t^{\prime}\right) \int_{0}^{t^{\prime}} d t^{\prime \prime} \delta L\left(t^{\prime \prime}\right) \int_{0}^{t^{\prime \prime}} d t^{\prime \prime \prime} A\left(t^{\prime \prime \prime}\right)-\cdots ; \tag{10}
\end{equation*}
$$

the time-integrals are convective, corresponding to (7). Formally, (10) may be written as

$$
\begin{equation*}
G=G_{0}\left(1+\delta L G_{0}\right)^{-1} \tag{11}
\end{equation*}
$$

where $G_{0} \equiv(d / d t)^{-1}$; i.e., $G_{0} A(t)=\int_{0}^{t} d t^{\prime} A\left(t^{\prime}\right)$.

Using Eq. (11), and $\delta L \equiv \dot{\Gamma} \partial_{\Gamma} \equiv \partial_{\Gamma} \dot{\Gamma}$ (the latter jedentity following from Liouville's theorem), we may manipulate Eq. (9) into the form

$$
\begin{equation*}
\frac{d P}{d t}=\partial_{\Gamma} \quad E \partial_{\Gamma} P \tag{12}
\end{equation*}
$$

which has the formal structure of the standard diffusion equation. However, the "diffusivity" $E$ is here still an integro (in time) differential (in phase-space) bperator:

$$
\begin{equation*}
\bar{\omega} \equiv\left[1-M\left(1+\partial_{\Gamma} M\right)^{-1} \partial_{\Gamma}\right] D, \tag{13}
\end{equation*}
$$

with

$$
\begin{equation*}
D \equiv\left\langle\delta \dot{\Gamma}^{\prime} G \delta \dot{\Gamma}\right\rangle \quad \therefore \tag{14}
\end{equation*}
$$

and

$$
\begin{equation*}
M \equiv\left\langle S \dot{\Gamma} G_{O} \delta L G\right\rangle, \tag{15}
\end{equation*}
$$

also operators. Note that $D$ and $E$ are second-rank tensors in the phase-space, while $M$ and $\partial_{\Gamma}$ are vectors.

The exact equation (12) becomes an actual difiusion equation, if one simultaneously makes two approximations: (a), one replaces $P(t-\tau)$, in the integrand or the time-integrals represented by $G$ and $G_{0}$, by $P(t)$; (b), one keeps only the lowest-order (in $S \dot{\Gamma}$ ) contribution to $L^{-}$:

$$
\begin{equation*}
\dot{N} \rightarrow D_{0} \equiv\left\langle\varepsilon \dot{\Gamma} G_{0} \dot{\delta} \dot{\Gamma}\right\rangle, \tag{16}
\end{equation*}
$$

where, because of the first approximation, $G_{0}$ now operates only on - $8 \dot{\Gamma}$ and not on $P_{0}$ This is the result obtained by ENF, for the case of one degree of freedom:

$$
\begin{equation*}
\frac{d P}{d t}=\partial_{\Gamma} D_{0} \partial_{\Gamma} P \tag{17}
\end{equation*}
$$

The validity of these two approximations may be investigated by estimating the relative sizes of the respective leading correction terms. For (a), we obtain the condition

$$
\begin{equation*}
\tau_{P} \gg \tau_{\delta}, \tag{18}
\end{equation*}
$$

where $\tau_{P}$ is the characteristic evolution time for $P$, while $\tau_{\delta}$ is the correlation time for the fluctuating Hamiltonian. This is the usual Markov assumption. From Eqs. (12) and (16), we may estimate

$$
\begin{equation*}
\tau_{P} \sim \sigma_{\Gamma}{ }^{2} D_{0}{ }^{-1} \sim \sigma_{\Gamma}{ }^{2}|\delta \dot{\Gamma}|^{-2} \tau_{5}{ }^{-1}, \tag{19}
\end{equation*}
$$

where $\sigma_{\Gamma}$ is a measure or the spread of $P$ in phase-space. Substituting (19) Into (18), we obtain the condition

$$
\begin{equation*}
\sigma_{\Gamma} \gg|\delta \dot{\Gamma}| \tau_{\delta} \tag{20}
\end{equation*}
$$

as equivalent to (18).
Turning now to (b), we have the condition

$$
\begin{equation*}
1 \gg\left|\delta L G_{0}\right| \sim|\delta \dot{\Gamma}| \tau_{\delta} \sigma_{\Gamma}^{-1} \tag{21}
\end{equation*}
$$

which we see is identical to the previous condition (20). We conclude that the standard diffusion equation is valid for the evolution in phasespace or a system with random Hamjltonian, if only the Markov condition (18) is satisfied. From familiar examples of this equation, we know that $\tau_{P} \sim t$; thus the diffusion equation eventually becomes valid, after a time $t \gg \tau_{\delta}$.

Finally, we demonstrate why the Fokker-Planck equation

$$
\begin{equation*}
\frac{\partial P}{\partial t}=-\partial_{\Gamma} \frac{\langle\Delta \Gamma\rangle}{\Delta t} P+\partial_{\Gamma} \partial_{\Gamma} D_{0} P \tag{22}
\end{equation*}
$$

which is valid under (18) and (21), reduces to the diffusion equation (17) for the present problem. We have

$$
\begin{equation*}
\Delta \Gamma[\Gamma(t), t, \Delta t] \equiv \int_{t}^{t+\Delta t} d t^{\prime} \dot{\Gamma}\left[\Gamma\left(t^{\prime}\right), t^{\prime}\right] \tag{23}
\end{equation*}
$$

In the argument of the integrand, we set $\Gamma\left(t^{\prime}\right) \equiv\langle\Gamma\rangle\left(t^{\prime}\right)+\delta \Gamma\left(t^{\prime}\right)$, where the mean is here conditional on $\Gamma(t)$. Expanding $\dot{\Gamma}$ to first order in $3 \Gamma$, we obtain

$$
\Delta \Gamma=\int_{t}^{t+\Delta t} d t^{\prime}\left\{\dot{\Gamma}\left[\left\langle\Gamma^{\prime}\right\rangle\left(t^{\prime}\right), t^{\prime}\right]+\delta \Gamma\left(t^{\prime}\right) \partial\left\langle\Gamma^{\prime} \dot{\Gamma}\left[\langle\Gamma\rangle\left(t^{\prime}\right), t^{\prime}\right]\right\}\right.
$$

Upon expressing $\delta \Gamma\left(t^{\prime}\right) \equiv \int_{t}^{t} d t^{\prime \prime} 8 \Gamma^{\prime}\left(t^{\prime \prime}\right)$ herein, and then taking the mean, we find

$$
\begin{equation*}
\frac{\Delta \Delta \Gamma}{\Delta t}=\langle\dot{\Gamma}\rangle+\partial_{\Gamma} D_{0}, \tag{24}
\end{equation*}
$$

to lowest order in $\Delta t$ and to second order in $8 \dot{\Gamma}$. Substituting (24) into (22), we obtain

$$
\begin{equation*}
\left(\frac{\partial}{\partial t}+\langle\dot{\Gamma}\rangle \partial_{\Gamma}\right) P=\partial_{\Gamma} D_{0} \partial_{\Gamma} P, \tag{25}
\end{equation*}
$$

which is Eq. (17).

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1. T. J. Birmingham, T. G. Northrop, and C.- G. Fälthammar, Charged Particle Diffusion by Violation of the Third Adiabatic Invariant, submitted to Physics of Fluids. This paper is here denoted by BNF.
2. We use the convention that an operator (like $I$ or $\partial_{\Gamma}$ ) operates on everything to its right.

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