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Spatio-temporal evolution of the L → I → H transition

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We investigate the dynamics of the low(L) → high(H) transition using a time-dependent, one dimensional (in radius) model which self-consistently describes the time evolution of zonal flows (ZF) and mean flows (MF), poloidal spin-up, and density and pressure profiles. The model represents competition. The I-phase appears as a nonlinear transition wave originating at the edge boundary and propagates inward. Locally, I-phase exhibits the characteristics of a limit-cycle oscillation. All these observations are consistent with recent experimental results. We examine the trigger of the L → H transition, by defining a ratio of the rate of energy transfer from the turbulence to the zonal flow to the rate of energy input into the turbulence. When the ratio exceeds order unity, ZF shear gains energy, and a net decay of the turbulence is possible, thus triggering the L → H transition. Numerical calculations indicate that the L → H transition is triggered by this peak of the normalized ZF shearing. Zonal flows act as “reservoir,” in which to store increasing fluctuation energy without increasing transport, thus allowing the mean flow shear to increase and lock in the transition. A counterpart of the L → I → H transition, i.e., an L → H transition without I-phase, is obtained in a fast power ramp, for which I-phase is compressed into a single burst of ZF, which triggers the transition. Effects of neutral charge exchange on the L → H transition are studied by varying ZF damping and neoclassical viscosity. Results show that the predicted L → H transition power increases when either ZF damping or viscosity increase, suggesting a link between recycling, ZF damping, and the L → H threshold. Studies of fueling effects on the transition and pedestal structure with an emphasis on the particle pinch are reported. © 2012 American Institute of Physics.

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I. INTRODUCTION

Understanding of L → H transition physics is crucial to a successful ITER. The L → H transition to a state of the higher confinement was discovered at ASDEX.1 Understanding of the power requirement for accessing the H-mode is so essential that intensive experimental surveys of threshold power in various physical parameters have been carried out.2,3 The power threshold tends to have different behavior in lower and higher density regions. However, a successful theoretical framework within which to understand the physics of the L → H transition and a theoretical prediction of the threshold power have not yet been realized, due to uncertainty concerning the trigger of the transition. The trigger of the transition has not been specified yet. To determine the threshold power scalings, a physical understanding of the trigger is necessary.

In experiments, prior to the L → H transition, as power slowly increases, several kHz quasi-periodic evolution of the $E_r$ (radial electric field) was first observed in ASDEX-Upgrade, and referred to as a dithering cycle.4 Instead, with a fast heat power ramp rate, only a few such cycles appear, while with a slow ramp rate, or at steady power, an extended series of limit-cycle oscillations is observed. This quasi-periodic behavior prior to the L → H transition is also seen in JET,5 DIII-D,6 etc.

Theoretically, the L → H transition is regarded as a transport bifurcation, which is related to $E \times B$ flow shear suppression by multiple types of flows: zonal flow (ZF) and mean flow (MF). Here, mean $E \times B$ flow (MF) $\langle V_E \rangle$ is generated mainly by diamagnetic flow due to global pressure and density profiles, while zonal flow $V_{ZF}$ is fluctuation driven and has a mesoscale radial structure. The mesoscale spectral range is $\Delta \ll l_{meso} \ll L_p$, where $\Delta \sim \rho_i$ is the turbulence correlation length and $L_p$ is a system size characterized by pressure scale length. The mesoscale range may be typified by the geometric mean of the micro and macro scales, $l_{meso} \sim \sqrt{\Delta L_p}$. However, we emphasise that “mesoscale” refers to a range of the scales and not a single scale! Therefore, ZF and MF shearing should be distinguished.
Shear enhanced decorrelation of turbulent fluctuations has been proposed as a mechanism for confinement improvement and turbulence suppression.\textsuperscript{9,10} A mean field predator-prey model originally describes the interplay between turbulence fluctuation and the diamagnetic-driven $E \times B$ mean flow, exhibiting self-consistency, i.e., the turbulence fluctuation and the mean flow amplitude are closely related to, and are determined by, one another.\textsuperscript{11} Reference 12 extended the self-consistent predator-prey system, coupling turbulence with the ZF, using the conservation of the wave action with the flow shear.

Reference 13 proposed a predator-prey model describing interacting ZF and MF shear suppression of turbulence, a simplified coupling of the mean flow shear and pressure gradient by radial force balance, and MF-ZF competition. The model recovers an L-H transition triggered by ZF shearing and a pre-transition limit-cycle oscillation (LCO) due to interplay among turbulence, ZF, and MF. The interplay occurs on a slow zonal flow damping time scale and manifests a phase delay between turbulence and shearings. Here, this model study finds that ZF shearing is an important ingredient in the L $\rightarrow$ H transition, because ZF mediates the transition, reduces turbulence level, enhances MF evolution, and thus regulates the power threshold. When the importance of the ZF to transition was noticed, it was realized that since zonal flow is fluctuation driven, ZF can trigger the transition but cannot sustain it. Thus, the transition intrinsically must be a two predator (ZF and MF) and one prey problem. Mean flow shear affects the fluctuation-driven Reynolds correlation, as well as fluctuation intensities. In this two predator-one prey model, the zonal flow triggers the transition, while the mean flow “locks in” to the H-mode state.

TJ-II has identified the physical mechanism behind the L $\rightarrow$ H transition from the limit-cycle oscillation of $E_i$, interacting with turbulent fluctuations.\textsuperscript{14} The experimental results clearly show limit-cycles in the $E_i$ and turbulence fluctuations in phase space, suggesting a strong similarity to the model of Ref. 13. Other experimental devices also have studied the pre-transition limit-cycle oscillation, which is referred as to intermediate (I)-phase in NSTX, ASDEX-Upgrade, DIII-D, and EAST.\textsuperscript{15–18}

Note that a dynamical system analysis of the model of Ref. 13 indicates that the LCO occurs around a saddle (structurally unstable) fixed point, referred as to the transient state.\textsuperscript{19} This should be distinguished from L $\rightarrow$ H $\rightarrow$ L periodic transition, as is previously discussed in Ref. 4. This is because the transient state LCO is defined by the feedback loop of turbulence and ZF shearing, and thus the ZF damping determines the time scale of the oscillation.

With regard to the pre-transition LCO, spatio-temporal structure has been also elucidated by simultaneous measurement at two radial locations, with a two-channel Doppler reflectometer in TJ-II.\textsuperscript{20} This study identified two-way propagation of $E_i$ oscillations depending on the measured line density, i.e., some cases of lower line density $(2 - 25) \times 10^{19}$ m$^{-3}$ show outward propagation, and some particular cases of higher line density $3 \times 10^{19}$ m$^{-3}$ exhibit inward propagation. The outward propagation speed is the largest at the innermost radial position and gradually decreases as the oscillation reaches the edge $E_i$ shear position. This phenomenon may be linked to the radial spreading of the turbulence from the plasma core to the edge barrier.\textsuperscript{21}

To relate these findings to the two predator-one prey hypothesis of ZF trigger and mediation, at least a one space dimension version of the multi-predator-prey model\textsuperscript{13} is necessary here. Such a one-dimensional model can predict the spatio-temporal evolution of the pedestal through the L $\rightarrow$ H transition, as well as the spatial structure of the LCO in I-phase. Note that since no “first principles” simulation has ever successfully recovered the L $\rightarrow$ H transition, reduced models are the only option. Also, in the event that useful “first principles” simulations become available in the future, reduced models will still be necessary to extract the essence of the transition physics—i.e., to distill the lesson learned.

In this paper, we present novel theoretical results on the spatio-temporal dynamics of L $\rightarrow$ H transition, with special emphasis on the role of ZFs in the trigger process. We present comparison to several recent experimental results. We extend the earlier transition models\textsuperscript{13} to develop a 5-field reduced mesoscale model which evolves turbulence intensity, zonal flow shear, pressure and density profiles, and mean poloidal mass flow in both radius and time. The mean $E \times B$ velocity shear $\langle V_E \rangle$ is calculated via radial force balance using density and pressure profiles and poloidal flow. We present evidence that the ZF can trigger the transition.

To support the hypothesis that the ZF is fundamental to the transition, we explore the sensitivity of the L $\rightarrow$ H threshold to the ZF damping. As ZF shearing can trigger the L $\rightarrow$ H transition and downshift the power threshold, larger ZF damping should weaken the ZF shearing and turbulence suppression and upshift the power threshold. Higher ZF damping can result from higher neutral charge exchange (CX). An increase of the neutral CX friction can be caused by wall saturation, increased re-cycling, etc., leading to increased ZF damping and increased neoclassical poloidal flow viscosity. This damping acts in addition to that originating from ion-ion collisions. This indicates that the wall physics and recycling can alter the L $\rightarrow$ H power threshold and that high edge neutral density is unfavorable to transition.\textsuperscript{22} Based on this hypothesis, we examine how ZF mediates the L $\rightarrow$ H transition in higher neutral CX cases.

We also study how the early stages of pedestal formation depend upon particle fueling. As particle fueling originates at the edge boundary, a particle pinch effect must be considered, so as to allow build-up of the density profile. However, the effects of the particle pinch on L $\rightarrow$ I $\rightarrow$ H transition and pedestal formation have not been discussed. Related to the particle pinch, the deposition layer width is another comparable factor. The deposition layer width is
really set by the fueling method, i.e., pellet injection or gas puffing. Thus, we here investigate two issues regarding fueling: (i) how the particle pinch affects $L \rightarrow H$ transition dynamics and (ii) how the fueling deposition width affects the pedestal structure.

The remainder of this paper is organized as follows. In Sec. I, we introduce the reduced 5-field mesoscale model, the related physics, and the initial conditions and parameters. In Sec. II, we present basic numerical results of the $L \rightarrow H$ transition and comparison to experimental results. In Sec. III, we discuss numerical results regarding the power threshold study and pedestal profiles. Here, we also suggest implications for experiments based on the results. In Sec. IV, we conclude and discuss remaining issues.

II. STRUCTURE OF THE MODEL

To better understand $L \rightarrow I \rightarrow H$ transition dynamics, we propose a theoretical model as an extension of the two predator-one prey system\(^7\) to treat dynamics and evolution in the radial dimension $r$. The model describes space-time evolutions of turbulence intensity ($I$), zonal flow shear energy ($E_0 = V^2_{ZF}$), pressure ($p$) and density ($n$) profiles, and mean poloidal mass flow ($\langle iv_0 \rangle$). These quantities are averaged over fine scales and fast times. Thus, \textit{a priori}, the spatial scale is longer than $\rho_i$ and the evolution scale is slower than drift wave time $\omega_c^{-1} \sim (c_s/a)^{-1}$, in order to satisfy the scale separation requirement for wave kinetics, which is the foundation of our reduced model.\(^7\) The basis is the following predator-prey model with one prey (turbulence intensity) and two predators ($ZF$ and MF shearings):

$$\partial_t I = I(\gamma_L - \Delta \omega I - x_0 E_0 - x_v E_V) + \chi_0 \partial_r (I \partial_r I),$$

$$\partial_t E_0 = \frac{x_0 E_0}{1 + \omega_c E_V} - \gamma_{damp} \partial_r E_0.$$  \hspace{1cm} (1)

Here, we assume that turbulence originates from the ion channel; the first term on the r.h.s. of Eq. (1) represents turbulence generation by ion temperature gradient (ITG) mode via linear instability, where $\gamma_L = \gamma_{Lo}(c_s/R)\sqrt{(R/L_T) - (R/L_T)_{crit}}$ is the local growth rate of turbulence intensity, with a critical temperature gradient parameter $(R/L_T)_{crit}$, $c_s = \sqrt{T_0/m_i}$ is the ion sound speed, $T_0$ is a reference ion temperature, $m_i$ is the ion mass, $R$ and $a$ are major and minor radii, respectively, and $L_T = [d \ln T/dr]^{-1}$ is the scale length of the ion temperature gradient. The temperature profile is calculated from $T = p/n$. The model is flexibly adaptable to other instabilities, by changing the definition of $\gamma_L$, to that for the trapped electron modes, etc. The second term on the r.h.s. of Eq. (1) represents the nonlinear damping by self-saturation of turbulence due to local spectrum broadening, the increment parameter for which is given by $\Delta \omega$. The third and fourth terms represent turbulence suppression due to $ZF$ and MF shearing, respectively. There $x_0 \sim \tau_{se}$ is a shearing coupling coefficient proportional to the auto-correlation time between turbulence and zonal flow group propagation,\(^11\)\(^23\) and $x_v$ is the shearing coupling coefficient between turbulence and MF shear. The detailed derivation of the shearing coupling parameters is discussed in Appendix A. The fifth term of the r.h.s. of Eq. (1) represents the nonlocal turbulence spreading,\(^21\) where the turbulent thermal diffusivity is $\chi_N \sim \chi_{Lo}(\nu_c c_s^2/a)$, and we assume turbulence spreading is diffusive with $D \sim \chi_N \sim D_{CM}$.\(^24\)

The first term on the r.h.s. of Eq. (2) represents turbulent Reynolds drive, with enhancement of decorrelation of drift wave propagation by a MF shear, which enters the cross phase in the Reynolds stress. $\xi_0$ represents the inhibition of zonal flow growth by MF shear due to weakening of the response of drift wave spectrum to a seed $ZF$.\(^11\) The derivation of the effect on $ZF$ shearing by MF shear is summarized in Appendix B. The second term on the r.h.s. of Eq. (2) represents damping of $ZF$ shear, where $\gamma_{damp}$ is the $ZF$ damping rate originating from ion-ion collisionality and also neutral CX. $\gamma_{damp} = \gamma_{damp,0} \nu_{ii}/R$ and $\gamma_{damp,0} = 1 + \nu_{CX}/\nu_{ii}$, where $\nu_{ii}$ is the ion-ion collision frequency and $\nu_{CX}$ is neutral CX friction. Here, we neglect the dependence of the neutral CX friction on the density profile, for simplicity. The sensitivity of $ZF$ damping to ion collisionality has been elucidated by direct numerical simulation (DNS) using the gyrokinetic equation, e.g., see Ref. 25. $E_V = \langle V^2_n \rangle$ is MF shear energy. The MF shear is obtained from the radial force balance equation. To separate the evolution of mean flow from that of zonal flow, we assume that pressure and density profiles should evolve on time scale much slower than zonal flows. This is due to the fact that $ZF$ is distributed throughout the mesoscale spectrum (i.e., $I \sim \sqrt{\Delta I_{zf} p}, \tau \sim \sqrt{\Delta \Gamma_{zf} p/c_s < \Gamma_{zf}/c_s}$), while MF are macroscopic (i.e., $I \sim \Gamma_{zf}$).

We obtain $\gamma_L$ and $E_V$ from the global profiles, $p$ and $n$. The evolution of the pressure and density profiles are given by the following one-dimensional transport equations with external sources:

$$\partial_t p + (1/r)\partial_r (r \Gamma_p) = \partial_t H,$$

$$\partial_t n + (1/r)\partial_r (r \Gamma_n) = \partial_t S.$$  \hspace{1cm} (4)

Here, $H$ and $S$ are the external heat and particle source flux profiles, respectively, given by

$$\partial_t H = Q_a \exp \left( -\frac{\gamma^2}{2 \Gamma_{h,dep}^2} \right),$$

$$\partial_t S = \Gamma_a \frac{a - r + d_a}{L_{dep}} \exp \left[ -\frac{(a - r + d_a)^2}{2 \Gamma_{dep}^2} \right].$$  \hspace{1cm} (5)

for which typical profiles are illustrated in Fig. 1. Note that source is flux-driven both for particles and heat. The heat flux $H$ is located in the core with the amplitude $Q_a$, while the particle flux $S$ is fueled in the edge region with amplitude $\Gamma_a$, the deposition width $L_{dep}$, and the shift of a gaussian peak from the edge, $d_a$.

$\Gamma_a$ and $\Gamma_p$ are given by the following equations, consisting of diffusive and convective (pinch) terms:

$$\Gamma_n = -(D_{nt} + D_0) \partial_r n + V_p n.$$  \hspace{1cm} (7)
the local heat diffusivity, we have

\[ \frac{\partial (v_0)}{\partial t} = B \frac{\mu_{\text{neo}}}{m_i} \left( \nabla \cdot (\bar{\varepsilon}_0 \overline{\Pi_{\text{turb}}}) \right) + B \mu_{\text{neo}} \left( (v_0) - \langle v_0^{\text{neo}} \rangle \right), \]

(11)

where \( B \) is a toroidal magnetic field. The first term in the r.h.s. represents poloidal spin-up driven by turbulence through the stress tensor, which may be replaced with radial divergence of the Reynolds stress through the Taylor identity. \( \mu_{\text{neo}} = \mu_0^{(\text{neo})} \nu_i q(r)^2 \) is the neoclassical poloidal viscosity, \( q(r) \) is the safety factor, \( \mu_0^{(\text{neo})} = \mu_{\text{gyro}} (1 + v_{\text{CX}}/v_i) \), and \( \mu_0^{(\text{neo})} \) is calculated from the energy weighted momentum equation. \(^3^3\) Here, we assume the neutral CX friction can be added to the ion-ion collisionality in the neoclassical poloidal viscosity \( \mu_{\text{neo}} \), as well as the ZF damping. \( \langle v_0^{\text{neo}} \rangle = - (\mu_0^{(\text{neo})} L \langle \psi \rangle B_\theta) \approx -1.17 \nabla T \), given by Ref. 33, is the neoclassical poloidal flow velocity. Assuming slab geometry and constant \( B \), we simplify Eq. (11) to

\[ \frac{\partial (v_\theta)}{\partial t} = - \frac{\gamma_L}{\omega_n} c_s^2 \frac{\partial I}{\partial \psi} - \frac{\mu_0^{(\text{neo})} \nu_i q \ell R^2}{\rho_i L_T} \left( \langle v_0 \rangle - 1.17 c_s \rho_i L_T \right), \]

(12)

We here use \( \gamma_L \) denoted in Ref. 32. This is the same as used in Eq. (1) of this model and \( \omega_n \sim c_i/L_n \sim c_i/a \) is fixed. Note that the contribution from the zonal flow shearing to the Reynolds drive is neglected here, due to the scale separation assumption.

The \( E \times B \) mean flow shear \( \langle V_E \rangle \) is related to the density and pressure gradients, the second derivative of the pressure profile, and the mean poloidal mass flow \( \langle v_0 \rangle \) by the radial force balance equation,

\[ \langle V_E \rangle \equiv \frac{1}{eB} \left[ \frac{- n' p'}{n^2} + \frac{\rho''}{n} \right] + a \left( \frac{v_0}{q R} \right)' - \langle v_0 \rangle', \]

(13)

which can be rewritten as

\[ \langle V_E \rangle' = \rho_i c_s L_p^{-1} (-L_n^{-1} + L_p^{-1}) - \langle v_0 \rangle', \]

(14)

where \( L_p = (d \ln f / dr)^{-1} \) for \( f = n, p, p' \). Here, the first term in Eq. (13) corresponds to diamagnetic shearing, proportional to pressure and density gradient. The second term is the pressure profile curvature, proportional to the second derivative of pressure profile. The pressure profile curvature is often dropped because of difficulty in the analyses (e.g., Ref. 26), but some references discuss effects of the pressure

\[ \Gamma_p = -(\chi_{\text{turb}} + \chi_0) \partial_j p. \]

(8)

Here, \( D_{\text{n.t.}} \) and \( \chi_{\text{n.t.}} \) are non-turbulent particle and heat diffusivities in H-mode, respectively, corresponding to the neo-classical transport. \( D_0 \) and \( \chi_0 \) correspond to turbulent particle and thermal diffusivities of the electrostatic turbulence, respectively given by

\[ D_0 = \chi_0 = \frac{\tau \ell^2 j}{1 + \chi_0 (V_E)^2}. \]

(9)

Note that \( \tau \) is the correlation time of turbulence and \( \chi_0 \) is a cross-phase modification by mean flow, normalizing the MF shear \( (V_E)^2 \). Here, we assume the particle and the heat fluxes have the same basic correlation time scale, i.e., we neglect possible differences in cross-phases of the diffusivities, for simplicity. The mean flow shear suppression factor \( \chi_0 \) also appears in Refs. 24, 26, and 9. As low \( \beta \) is assumed, there is no magnetohydrodynamics (MHD) activity—edge localized modes (ELMs), etc., after the development of the pedestal. Describing such evolution is beyond the scope of this paper. We thus consider only electrostatic turbulence here.

\[ V_\theta = V_{\text{poloidal}} (D_{\text{n.t.}} + D_0) \left( \frac{2}{R} + \frac{1}{L_T} \right), \]

(10)

where the first term of the latter set of parenthesis on the R.H.S of Eq. (10) represents a turbulent equipartition (TEP) pinch \(^2^8,9^9,3^0\) and the second term is a thermodiffusive pinch. \(^3^1\) Note that both pinch velocities are inward for ITG turbulence. Note also the total particle flux could be negative, at least transiently, until the density gradient steepens in H-mode. This is because a large heat flux can drive a strong inward density pinch via the ITG mode. The particle pinch consists of the TEP and thermodiffusive pinch, \( V_\theta = V_{\text{TEP}} + V_{\text{Th}} \). The thermodiffusive pinch is proportional to the turbulence intensity, as given by \( V_{\text{Th}} \sim |V_\psi| / \chi \). Since \( |\nabla T| \sim Q / \chi \) follows, where \( Q \) is the heat flux and \( \chi \) is the local heat diffusivity, we have \( V_{\text{Th}} \sim Q \). Thus, \( \Gamma_n < 0 \) is possible if sufficient heat flux is carried. This has been observed, i.e., probe studies of particle flux show cross phase of \( V_\psi \) and \( n \) such that total flux is inward. \(^3^1\) Note that this observation applies to the H-mode pedestal and not to the L-mode preceding the transition. Of course, the particle pinch causes density profile peaking, so roughly \( \nabla n / n \sim V / D \). We do not include a heat pinch in the pressure evolution, because the heat source is assumed to be applied in the core, and we are not concerned with global temperature profile structure.

By coupling toroidal and parallel force balance equations, we obtain the time evolution of poloidal mass flow, \(^3^2\)

\[ - \frac{\partial (v_\theta)}{\partial t} = B \frac{\mu_{\text{neo}}}{m_i} \left( \nabla \cdot (\bar{\varepsilon}_0 \overline{\Pi_{\text{turb}}}) \right) + B \mu_{\text{neo}} \left( (v_0) - \langle v_0^{\text{neo}} \rangle \right), \]
profile curvature on the transition.\textsuperscript{34,35} The profile curvature term becomes important especially when there is a corner in the profile. At the top of the pedestal in the H-mode profile, there necessarily must be a corner, and thus the profile curvature is large. How the pressure curvature evolves through $L \rightarrow H$ transition is not yet clear. Therefore, we here keep this term and investigate its evolution. The forth term corresponds to a contribution from the mean poloidal momentum. Here, we neglect the toroidal momentum of the third term in Eq. (13), for simplicity. This effect likely is essential for internal transport barrier (ITB) formation, however.\textsuperscript{36}

Note that this model can be reduced to the local predator-prey model of Ref. 13 with the following simplification. In Eqs. (1) and (2), neglecting the nonlocal effect ($\chi_N = 0$), and letting $\gamma_L = N'$, where $N' \propto (-dp/dr)$, we obtain the evolution equations for turbulence intensity and zonal flow shearing. These correspond to Eqs. (6) and (7) in Ref. 13. Using the simplified radial force balance of Eq. (14), assuming a constant temperature profile, neglecting the second derivative and toroidal and poloidal momentum, we obtain $(V_F)^2 = dN^2$. Reference 13 described the evolution of $N'$ by the expression $\partial_t N = Q - (c_1 + c_2)N'$ in Eq. (3). There $Q$ is the heat source and $c_1$ and $c_2$ are the collisional/residual and turbulent transport coefficients, respectively.

A. Parameters and boundary conditions

Finally, we obtain the 5-field ($I, E_0, p, n, \langle v_0 \rangle$) equations (Eqs. (1)–(4) and (12)) and the radial force balance equation, Eq. (13). We numerically solve these equations, using a finite difference method in radial space, and an implicit finite difference method in time for the time integration, with the following conditions and parameters.

The simulation region is a cylindrical, one-dimensional space between $r = 0$ at core and $r = a$ at the edge. Boundary conditions are the following, simplest set. For pressure and density profiles, we take

\begin{align*}
p(r = a) &= 0.01, \quad n(r = a) = 0.1, \\
P'(r = 0) &= n'(r = 0) = 0, \\
P''(r = 0, a) &= n''(r = 0, a) = 0.
\end{align*}

Here, $p$ and $n$ are normalized by the reference quantities $n_0 = 10^{20} [m^{-3}]$ and $p_0 = T_0 n_0 = 1 [keV] \times 10^{20} [m^{-3}]$, respectively. We apply free boundaries at the cores of $p$ and $n$ but fixed boundaries at the edge. For the second derivatives of $p$ and $n$ at the edge, we impose the free boundary conditions, thus the third derivatives are also set to zero. We assume no interaction between the edge and scrape-off-layer (SOL) region and use a simple boundary condition. We believe that the free-forced boundary condition on the second derivative is the simplest applicable one. The simple boundary condition is consistent with experimental results, since quantities at the last closed flux surface (LCFS) are mostly constant during the $L \rightarrow H$ transition in experiments. Here, we neglect any explicit modeling of SOL region ($r > a$). If we go beyond the free boundary condition, a separate SOL-edge model is necessary to match to. This is beyond the scope of this paper and will be addressed in future work.

Note that we need to deal with the poloidal asymmetry and X-point structure in a complete model. This requires modelling of the SOL flows and the effect of the edge boundary condition. In the present model, we simply assume a simulation box inside the LCFS and a fixed boundary at the edge. We present further discussion of this issue in Sec. IV. The study of detailed models will be pursued in the future.

For turbulence intensity, $Z_F$, and mean poloidal flow, we impose

\begin{align*}
I'(r = 0) &= I'(r = a) = E_0'(r = 0) = E_0'(r = a) = 0 \Rightarrow \langle v_0 \rangle'(r = 0) = \langle v_0 \rangle'(r = a) = 0.
\end{align*}

The other parameters we used in this numerical simulation are the following: The amplitude of particle flux source $Q_a$ is held constant at $Q_a = 10^{-4}$. The strength of heat flux power $Q_T$ ramps linearly from $1.0 \times 10^{-4}$ to $3.2 \times 10^{-2}$, after reaching initial equilibrium from arbitrary initial conditions. Here, $Q_a$ and $Q_T$ are normalized by $(c_i n_0)$ and $(c_i p_0)$, respectively.

The ion-ion collision frequency, $\nu_{ii}$, normalized by the ion cyclotron frequency $\omega_{ci} = eB/m_i$, is

\begin{align*}
\frac{\nu_{ii}}{\omega_{ci}} &= \frac{n_0 Z^2 e^4}{3^{1/2} 6\pi \epsilon_0 m_i^{1/2} T_{i0}^{3/2}} \left( \frac{eB}{m_i} \right)^{-1}.
\end{align*}

Here, we assume a single hydrogen ion, $Z = 1$, $A = 1$, $e = 1.6 \times 10^{-19} [C]$, $\Lambda = 20$, $B = 3.5 [T]$, $m_i = 1.67 \times 10^{-27} [kg]$, $\epsilon_0$ is the dielectric constant of vacuum, and $T_{i0} = 1 [keV]$ is a reference ion temperature.

For the non-turbulent particle and thermal diffusivities in H-mode, we use the following thermal diffusion coefficients in the banana region:

\begin{align*}
\chi_{Ti} &= \epsilon_i^{-3/2} \left[ q(r) \right]^{2} r^2 \nu_{ii}, \quad (16) \\
D_{ie} &\sim (m_e/m_i)^{1/2} \chi_{Ti}, \quad (17)
\end{align*}

where $m_e$ is the electron mass. We set the safety factor ($q$) profile, $q(r) = 0.95 + 2.0 \times (r/a)^2$. Here, the transport in H-mode should be related to the neoclassical transport, ($\chi_{Ti}$ and $D_{ie}$, respectively), and other residual turbulence transport in H-mode. To model the non-turbulent transport, we apply the enhancement parameters $\chi_{n,t} = 30 \chi_{Ti}$ and $D_{n,t} = 15 \chi_{Ti}$, for the purpose of numerical convenience. With only ITG turbulence, we could not obtain proper initial profiles from the time evolution in the arbitrary initial condition, in the case with lower or purely neoclassical $\chi_{Ti}$ and $D_{ie}$. To obtain correct initial equilibria, some modest residual particle diffusion is necessary to control the density gradient. The actual residual particle transport in the pedestal is not yet understood and remains a topic of active research. This set of the parameters still retains the necessary qualitative properties, since $\chi_{n,t} > D_{n,t}$, $\chi_{0} \gg \chi_{n,t}$, and $D_{0} \gg D_{n,t}$ are satisfied.
Here, we apply a digital filter to the 5-field evolutions in each time step for the purpose of smoothing or compensating irrelevant higher radial wavenumber oscillations.\textsuperscript{38,39} The digital filter can correspond to hyper-viscosities in the field evolutions, which determine a minimum scale size of the mesoscale profile structure. The scale is determined by finite Larmor radius effects or neoclassical polarization shielding, based on gyrokinetic theory. We here avoid a specified estimation of the hyper-viscosities, but retain the essence of the model by using a digital filter with an arbitrary resolution. We set the digital filter so as to maintain the long wavelength structure \(\Delta \approx \rho_i\), and to effectively dissipate the short wavelength structure \(\Delta < \rho_i\). This treatment also ensures the radial scale separation law of the mesoscale envelope from the turbulence scales.

Other parameters are \(\rho_e = (\rho_i/a) = 0.01, \epsilon_e = a/R = 0.25, L_{\phi e} = a = 0.15, L_{d e} = a = 0.10, d_a/a = 0, \tau_e = 1.0(a/c_i), \alpha_e = 1.0(a/c_i)^2, \gamma_e = \Delta\omega = 10^{-2}(c_i/R), \omega_0 = 1.0 \times 10^2/\rho_e, (c_i/a), \alpha_d = 2.0 \times 10^{-2} \sqrt{\rho_e}/(c_i/a), \zeta_0 = 10^2(a/c_i)^2, (R/L_T)_{\text{crit}} = 3.7, \zeta_{\text{damp}} = 1.0\omega_0/(c_i/a), \zeta_N = 0.5\alpha_e/(a/c_i)^2, \nu_0 = 1.0a^{-2}, \alpha_s = 5.0 \times 10^3, \beta_0 = 1.0, \beta_1 = 10^2\). Radial space and time scales are normalized by the minor radius \(a\) and the characteristic time of the drift wave \(\alpha_s = (a/c_i)^2\), respectively. The radial grid size is \(\Delta r/a = 1/400\), and the time step is \(\Delta(a/c_i) = 0.125 - 0.5\).

**III. BASIC STUDIES OF L \xrightarrow[]{} I \xrightarrow[]{} H TRANSITION DYNAMICS**

With the parameters described above, in Fig. 2, we present a numerical result which shows an L \xrightarrow[]{} I \xrightarrow[]{} H transition using a slow power ramp. As can be seen clearly, there are three distinct stages. The early stage, L-mode, is identified as a state with slowly growing turbulence, self-regulated by zonal flows. In the L-mode, as MF shearing is weak, the interaction between ZF and MF shearing is not significant, i.e., little decorrelation of ZF shearing by MF shear appears (see also Fig. 10(c)). As the heat flux through the edge increases with power, the system evolves into the next stage, namely the I-phase. The I-phase, which starts at \(t = 2.16 \times 10^5(a/c_i)\), is characterized by the presence of nonlinear shearing waves, nucleated near the edge boundary and propagating inward while also expanding outward. The nonlinear shearing wave structure is consistent with that observed in DIII-D.\textsuperscript{17} Local phase portrait of the nonlinear waves appears with the LCOs observed in other experiments. These are triggered by an increase of ZF in the edge region (0.90 < \(r/a\) < 1.00). One example of the local phase portrait is shown in Fig. 3, plotting ZF energy \(E_0\) and MF shear energy \(E_V\) in I-phase (\(t = 2.4 - 3.2 \times 10^5(a/c_i)\)) at the edge region, \(r/a = 0.975\). The plots show quasi-periodic behavior with a certain phase relation. During the I-phase, a radially coherent phase shift among local values of turbulence, ZF, and MF can be identified. Once the ZF shears increases, the MF follows the behavior of the ZF shear. At \(t = 3.53 \times 10^5(a/c_i)\), the L \xrightarrow[]{} H transition occurs. At the transition, the I-phase terminates abruptly with the quench of the edge turbulence on a fast time scale and that of ZF in a slower time scale. The time scale that ZF decreases after the transition corresponds to the ZF damping rate. On the other hand, some weak turbulence and ZF persist on the pedestal shoulder, even after the transition. Note that here, the actual transition time scale is very short, even for a slow power ramp! During the transition, the pedestal begins to expand inward, as the MF shear grows rapidly. The transition time is a mixture of turbulence transport and neoclassical transport time scales.\textsuperscript{30} Crudely, it is estimated as \(\sim 1\text{msec}\) for DIII-D parameters. After the L \xrightarrow[]{} H transition, the pedestal width still expands slowly inward and then saturates at \(\Delta_{\text{ped}} \sim 0.10a\) in \(r > 3.6 \times 10^5(a/c_i)\).

Fig. 4 shows a bird’s-eye view of the space-time evolutions of turbulence intensity, ZF, and MF for the same case.

**FIG. 2. Spatio-temporal evolution of turbulence intensity \(I\), (a) (b) ZF shearing energy \(E_0\), and (c) logarithm of MF shearing energy \(\ln(E_V)\) as functions of time \(t\) during a power ramp (2 \(\times 10^5 < t < 4 \times 10^5\)) and as a function of radius (0.5 < \(r/a\) < 1.0). Three distinct stages, L-mode, I-phase (LCO), and H-mode, are evident. In L-mode, turbulence and ZF grow self-consistently from the edge region; In I-phase, extended space and time structure of turbulence intensity, ZF, and MF appears. At the transition, turbulence and ZF drop rapidly. After transition, decay of turbulence and ZF is seen, while mean shear persists.**

**FIG. 3. Plots of ZF energy \(E_0\) versus MF shearing energy \(E_V\). Data points are taken from \(t = 2.4 \times 10^5(a/c_i)\) to \(t = 3.2 \times 10^5(a/c_i)\) with each \(\Delta t = 8 \times 10^5(a/c_i)\). The number of plots are 100.) The plots exhibit the limit-cycle oscillations. The propagating nonlinear waves are locally the limit-cycle oscillations with a phase delay between different radii. The filled circular plot denotes the final point at \(t = 3.2 \times 10^5\). It indicates that the limit cycle rotates counterclockwise.**
as Fig. 2. We see that the LCO is a propagating nonlinear wave of turbulence, ZF, and MF shear fields in the edge layer around \( r/a > 0.85 \). The mean flow shear, representing the profiles, oscillates during I-phase. This is likely related to the oscillation of \( D_a \) signal observed in experiments. The local maximum of turbulence intensity peaks just prior to the transition.

Fig. 5 shows time evolution of turbulence intensity, ZF, and MF energies at various radial locations at the edge region (\( r/a = 0.975, 0.950, 0.925 \)). Inward propagation of MF shear is identified. We notice that as the heat flux approaches criticality, the LCO phase delay between I and E increases from \( \sim \pi/2 \) to \( \sim \pi \), while the nonlinear LCO period increases, i.e., the cycle slows. Interestingly, this tendency is seen in the local model\(^1\) and also in DIII-D.\(^1\) In the DIII-D experiments in L-mode, no correlation of MF shear and turbulence is seen. During the LCO a \( \pi/2 \) phase shift is found, and the phase shift approaches \( \sim \pi \) at final H-mode transition, as equilibrium flow shear quenches the Reynolds stress.

Fig. 6 shows the evolution of MF and ZF shearings through \( L \rightarrow I \rightarrow H \) transition, so as to compare with recent experimental results. Fig. 6(a) shows the temporal evolution of the diamagnetic shearing, i.e., \( \frac{\partial E_x}{\partial B_{dia}} = L_n^{-1} L_p^{-1} \left( \frac{c_i}{\rho_i} \right) \). The diamagnetic shear oscillates with growing amplitude in I-phase, then increases abruptly at \( L \rightarrow H \) transition. This is consistent with experimental results.\(^1\) This indicates that pressure and density gradients oscillate during I-phase, since the diamagnetic shear does. The oscillation of the density gradient must be related to \( D_a \) signal oscillations in experiments.

![FIG. 4. Three-dimensional color maps of the time evolution of (a) turbulence intensity \( I \), (b) ZF energy \( E_\theta \), and (c) MF shearing energy \( \log(\alpha_r E_V) \) as functions of time \( t \) (during the slow power ramp regime \( 2 \times 10^5 < t < 4 \times 10^5 \)) and radius \( 0.5 < r/a < 1.0 \). These pictures show nonlinear waves propagating inward from the edge layer as the transition develops. What locally appears as a limit cycle is actually a slice of propagating nonlinear wave in the edge layer.](image1)

![FIG. 5. Time evolution of turbulence intensity \( I \) (blue solid line), ZF energy \( E_\theta \) (green solid lines), and mean square MF shear \( E_V \) (red bold lines) at various radial location of (a) \( r/a = 0.975 \), (b) \( r/a = 0.950 \), and (c) \( r/a = 0.925 \). The arrow indicates inward propagation of the mean flow peaks. At constant phase, the innermost radius leads in time, suggesting inward propagation.](image2)

![FIG. 6. Time evolution of (a) diamagnetic shearing (the first term of Eq. (13)), (b) \( \eta \), i.e., the ratio of turbulence energy transfer to the ZF, normalized to the net energy input into the turbulence, and (c) the ratio of MF shear to ZF shear, from the onset of I-phase at \( t = 2.1 \times 10^5 (a/c_s) \) to the L-H transition indicated at \( t = 3.5 \times 10^5 (a/c_s) \), ZF shearing is strongly enhanced, and sufficient to quench the turbulence. Then the mean shear is sufficient to lock in the transition.](image3)
A. Studies of a ZF role in the L $\rightarrow$ H transition: Triggering

Next, we investigate how ZF shearing mediates the L $\rightarrow$ H transition. In order to clarify the parameters which we will use, it is useful to write a schematic set of coupled predator-prey equations \(^1\) for turbulence and ZF, which are a simplified version of Eqs. (1) and (2), neglecting mean flow and nonlocal effects. These are

\[
\begin{align*}
\frac{\partial l}{\partial t} &= \gamma_{\text{eff}} I - \langle \dot{v}_i \dot{v}_j \rangle \frac{\partial v_{ZF}}{\partial r}, \\
\frac{\partial E_0}{\partial t} &= \langle \dot{v}_i \dot{v}_j \rangle \frac{\partial v_{ZF}}{\partial r} - \gamma_{\text{damp}} E_0.
\end{align*}
\]

Here \(\gamma_{\text{eff}} = \gamma_l I - \Delta a I^2\) is a total, effective growth rate, including the gradient drive and nonlinear damping.

\[
P_\perp = \langle \dot{v}_i \dot{v}_j \rangle \frac{\partial v_{ZF}}{\partial r} \sim \alpha_0 I E_0
\]

is the Reynolds work of the fluctuations on the flow. \(\langle \dot{v}_i \dot{v}_j \rangle \sim \pm \partial v_{ZF}/\partial r\) indicates negative (i.e., ZF growth) or positive (i.e., ZF damping) viscosity, respectively.

The obvious criterion for triggering the L $\rightarrow$ H transition is \(\partial l/\partial t < 0\), with a positive phase between \(\langle \dot{v}_i \dot{v}_j \rangle\) and \(\partial v_{ZF}/\partial r\) — i.e., negative viscosity which results in a net decay of the fluctuation energy. In this case, the zonal flow extracts energy from the turbulence faster than the turbulence grows. This requires

\[
\eta \equiv \frac{\langle \dot{v}_i \dot{v}_j \rangle \frac{\partial v_{ZF}}{\partial r}}{\gamma_{\text{eff}} I} = \frac{\alpha_0 E_0}{\gamma_l - \Delta a I} > 1,
\]

which emerges as a natural figure of merit for the collapse of the turbulence and the onset of transition. Note that \(\eta > 1\) does not always guarantee triggering of the L $\rightarrow$ H transition. This is not a sufficient condition but is at least the necessary condition to trigger the L $\rightarrow$ H transition by the zonal flow and mean flow interaction. In Ref. \(^41\), results from “measurements” of \(\eta\) defined as \(P_\perp/\nu_{\text{net}} (\langle \dot{v}_i \rangle^2\rangle\), instead of \(\gamma_{\text{eff}} I\), were reported, where \(\nu_{\text{net}}\) is the effective rate of energy input into the turbulence during periods of weak ZF. Results indicated that when \(\eta\) exceeds order unity, the L $\rightarrow$ H transition is triggered in EAST experiments.\(^{31}\) This shows that the ZF is fundamental to L $\rightarrow$ H transition, at least in those cases. Thus, we here compare the experimental results with numerical results based on the model presented here.

Fig. 6(b) shows the time evolution of \(\eta\), indicating that peaks of the ZF shearing increase significantly close to L $\rightarrow$ H transition, consistent with the analyses from the experiment. This suggests that ZF shearing really becomes dominant just prior to the L $\rightarrow$ H transition. Fig. 6(c) shows a time evolution of MF shearing normalized by the ZF shearing. The MF shear starts to grow after the onset of the I-phase. In accord with the LCO, the normalized MF shear oscillates with growing amplitude. The size of this oscillation is largest just prior to the L $\rightarrow$ H transition. After the final peak of the MF shearing just prior to the transition, the MF shear rapidly drops and then increases, crossing zero. Finally, in the H-mode pedestal, the MF shearing becomes much stronger than the ZF shearing.

B. Studies of the general L $\rightarrow$ H transition with a fast ramp of heat flux

The rate of heat flux increase, related to the rate of increase of power, can be another factor which determines “the type of L $\rightarrow$ H transition.” We know that there are actual cases for which the L $\rightarrow$ H transition occurs without an I-phase. The answer to the question comes from considering the rate of the heat flux increase. Indeed, the I-phase was only identified by carefully creating experiments operating near the power threshold. To this end, we examine the case with a faster power ramp, shown in Figs. 7(a)–7(c). In this case, the LCO is compressed into a single burst of ZF, which triggers L $\rightarrow$ H transition at \(t = 2.7 \times 10^4 (a/c_{s})\). For the single burst scenario to apply the heat flux increase, time scale must be shorter than the time scale of a single limit-cycle. While turbulence is quenched immediately after the L $\rightarrow$ H transition is triggered, the ZF is damped more slowly, in accord with the modest ZF damping time \(\sim \gamma_{ZF}^{-1}\). These features seen in Figs. 7(a)–7(c) resemble those of TJ-II, shown in Fig. 7 of Ref. \(^42\). A small increase of the MF shearing in accord with the burst of ZF in \(0.65 < r/a < 1.0\) and \(2.3 \times 10^4 (a/c_{s}) < t < 2.7 < 2.3 \times 10^4 (a/c_{s})\) is seen in Fig. 7(c). This indicates that the ZF induces MF growth through turbulence suppression.

We discuss the trigger of the L $\rightarrow$ H transition without the I-phase, using the parameter \(\eta\) in Eq. (21). In Figs. 7(d)–7(f), temporal evolutions of turbulence intensity \(I\), ZF energy \(E_0\), MF energy \(E_F\), the product quantities \(P_\perp = \alpha_0 E_0\), and \(\eta = \alpha E_0/\gamma_l - \Delta a I\) are shown at a specific radial location \(r/a = 0.9625\). To make clear the sequence of events...
C. Basic properties of density, pressure, turbulence intensity, zonal flow, and mean flow shear in L-mode, I-phase, and H-mode

In this subsection, we show further analyses obtained from the case with a slow power ramp. The condition is the same as that in Fig. 2. Fig. 8 shows spatial evolutions of density, pressure, and temperature profiles at typical times of the L-mode, I-phase, and H-mode. Pedestal formation in the H-mode is clearly recovered in the edge region, 0.9 < r/a < 1.0. Note that these profiles including pedestals are not empirically specified, but rather evolve given by the time evolution of the self-consistent theoretical model.

In L-mode or I-phase, both heat and particle diffusivities in the edge region are governed by turbulent transport, since the relation $D_{\text{nt}} \approx Z_{\text{nt}} \ll D_{\text{D}}$ is satisfied. On the other hand, in the H-mode pedestal, the diffusivities are governed by the non-turbulent transport instead, due to the quench of turbulence. Once the L → H transition occurs, the diffusivities in the pedestal region drop immediately. As a consequence steep profile gradients are formed, and the MF shear is rapidly excited. On the pedestal shoulder, residual turbulence and ZF persist. They round off the profiles at the top of the pedestal after the L → H transition, thus leading to formation of a convex pedestal corner.

We focus on the profile in the edge region as seen in Fig. 9. The pressure profile in L-phase at $t = 2.8 \times 10^5(a/c_s)$ is not peaked in the edge region, compared to that in L-mode. There is some enhanced turbulence because of the decorrelation of ZF shearing by the MF shear, leading to stronger turbulent transport. Judging from Fig. 5(a), at $t = 2.8 \times 10^5(a/c_s)$ in the edge region $r/a > 0.95$, turbulence is most enhanced and thus the ZF and MF shear exhibit minima. Therefore, the profile behaves like an L-mode, with minimum flow shear. For the contrary phase, at $t = 2.72 \times 10^5(a/c_s)$, turbulence is minimal, and the ZF and MF shear are enhanced. Thus, a steeper gradient than that in...
L-mode is observed. This is how the profile oscillates in I-phase.

Fig. 10 shows the spatial evolution of turbulence intensity, ZF, and MF energies. Turbulence is quenched in the H-mode pedestal. Consequently, in Fig. 10(b), in H-mode, the ZF is eliminated from the pedestal, except at the pedestal shoulder. ZF is enhanced around \( r/a \approx 0.85 \) but decreases in the edge region for \( 0.9 < r/a < 1.0 \), in correspondence to the increase of turbulence, due to decorrelation of ZF shearing by MF shearing. In Fig. 10(c), MF shear exhibits significant structure in the pedestal. The MF shear in the H-mode pedestal is much stronger than the ZF shear. In the I-phase, MF shearing oscillations are seen as part of the LCO propagation. In the L-mode, MF shearing contribution is negligible.

Fig. 11 shows profiles of diamagnetic shearing (the first term of Eq. (13)), pressure profile curvature (the second term of Eq. (13)), poloidal flow shearing (the forth term of Eq. (13)), and neoclassical poloidal flow shearing (\((\nu_{\parallel}^{(\text{neo})})^2\) in Eq. (12)) for the I-phase and the H-mode. These parameters are the ingredients of the mean flow shearing. As seen in Fig. 11(a), in I-phase, the main contributor to the mean flow shear in the edge region is the poloidal flow, which is mostly neoclassical poloidal flow. Poloidal flow appears to be predominantly neoclassical, except for a very thin layer within the thin pedestal. The pressure curvature is comparable to the diamagnetic shearing, indicating that analyses without the pressure curvature are not valid. The radial structure of the pressure profile curvature appears to be out of phase with the poloidal flow. The fine structure of the MF shearing driven by the pressure curvature and neoclassical poloidal flow may be related to the nonlinear wave propagation discussed previously, since profiles of the diamagnetic shearing without the second derivative of profile stay constant through I-phase.

In H-mode, shown in Fig. 11(b), the diamagnetic shear dominates the mean flow shearing in the pedestal, while the pressure curvature contribution in the pedestal is not significant. Because of the quench of the turbulence in the pedestal, the poloidal flow is mostly governed by the neoclassical contribution. At \( r/a = 0.90 \), pedestal corner forms, so the curvature contribution is necessarily important there. This result is also consistent with the argument in Ref. 35. While the turbulence driven poloidal flow is localized at the thin edge layer, the poloidal flow profile is mostly neoclassical, as seen in Fig. 11(b). This is, however, dependent on the neoclassical poloidal viscosity, proportional to \( \nu_{\parallel} \). Therefore, the situation may be different for lower \( \nu_{\parallel} \).
Here, the ZF damping rates and the neoclassical viscosities are fundamental to the transition. Increasing ZF damping reduces power threshold. This is not surprising, since ZF shearing is leading to an upshift in the transition threshold. We note that ZF shearing and its effect on the suppression of turbulence, thus more MF shear, instead of the ZF shearing, affects the turbulence. 

In Fig. 12, we show cases with various ZF damping rates \( \gamma_{\text{damp},0} = 1, 2, 3 \) and neoclassical viscosities \( \mu_{0}^{(\text{neo})} = 0.5, 1, 1.2 \). (a)-(c) show cases with fixed \( \mu_{0}^{(\text{neo})} \) but increasing \( \gamma_{\text{damp},0} \). Increase in \( \gamma_{\text{damp},0} \) is followed by delays of the transition. (a) and (d) and (e) show cases with fixed \( \gamma_{\text{damp},0} \), but varying \( \mu_{0}^{(\text{neo})} \). As \( \mu_{0}^{(\text{neo})} \) increases, the transition occurs later.

IV. NEUTRAL DENSITY AND FUELING EFFECTS ON TRANSITION AND THE PEDESTAL

In this section, we first show results of numerical studies of tests on the L \( \rightarrow \) H power threshold for various ZF damping and poloidal viscosity. These tests represent variable neutral density effects. As is widely known, high edge neutral density is unfavorable to the transition. While this may be due to higher radiation reducing confinement, increased ZF damping is also a possibility. We here re-visit the issue of neutral related problems in relation to the ZF damping rate. We investigate studies with a higher ZF damping rate, corresponding to cases with higher edge neutral CX friction. Next, we discuss fueling effects on the density profile. By changing the fueling depth with and without the particle pinch effect, we discuss how the pedestal width can vary.

In Fig. 12, we show cases with various ZF damping rates \( \gamma_{\text{damp},0} = 1, 2, 3 \) and neoclassical viscosities \( \mu_{0}^{(\text{neo})} = 0.5, 1, 1.2 \). Here, the ZF damping rates and the neoclassical viscosities are normalized by \( \nu_{ii}/R \) and \( \nu_{ii} a^2 R^2 \), respectively. We here fix the ramp speed slow, which is the same case as Fig. 2. As seen in Figs. 12(a)–12(c), the transition delays as the ZF damping rate increases, with fixed neoclassical viscosity. Thus, the power threshold increases as the ZF damping increases. The period of LCO decreases as the ZF damping increases. This indicates that the period of the LCO depends on the ZF damping rate, a tendency which is consistent with the original model of Ref. 13. A similar tendency to delay transition can be found in comparison with different \( \mu_{0}^{(\text{neo})} \) and the fixed ZF damping rate, as seen in Figs. 12(a), 12(d), and 12(e). Delays of the transition are also observed when the neoclassical viscosity \( \mu_{0}^{(\text{neo})} \) increases. The period of LCO decreases as the neoclassical viscosity \( \mu_{0}^{(\text{neo})} \) decreases. This is because more poloidal velocity shear, driven by turbulence, is excited in the lower \( \mu_{0}^{(\text{neo})} \) and thus more MF shear, instead of the ZF shearing, affects the turbulence intensity evolution. Thus increasing either the ZF damping or the neoclassical viscosity delays transition at constant heat flux power ramp, causing a power threshold upshift.

Fig. 13 shows plots of the power thresholds \( Q_{a,\text{crit}} \) obtained from numerical tests with various ZF damping rate \( \gamma_{\text{damp},0} \) and neoclassical poloidal viscosity \( \mu_{0}^{(\text{neo})} \), for a fast power ramp. It is clear that the power threshold increases as either ZF damping or neoclassical poloidal viscosity increases. Thus, we expect higher neutral CX to increase the power threshold. This is not surprising, since ZF shearing is fundamental to the transition. Increasing ZF damping reduces ZF shearing and its effect on the suppression of turbulence, leading to an upshift in the transition threshold. We note that recent XGC1 simulations with neutrals found results which suggested damping of zonal flows by neutral drag.

We examine the cases without ZF shearing, i.e., \( \gamma_{\text{damp},0} = 0 \), with various neoclassical viscosities. Without the ZF shearing, the L \( \rightarrow \) H transition occurs when \( \mu_{0}^{(\text{neo})} \leq 0.5 \).
$Q_{a\text{, crit}} = 0.0013$ is obtained when $\mu_0^{(\text{neo})} = 0.5$ and $Q_{a\text{, crit}} = 6.7 \times 10^{-3}$ is obtained when $\mu_0^{(\text{neo})} = 0.1$. These power thresholds are asymptotic to those in the cases with higher ZF damping, $\gamma_{\text{damp}} = 10$ with the same $\mu_0^{(\text{neo})}$, plotted in Fig. 13. Cases with higher ZF damping asymptotically approach the case with no ZF shearing. We estimate the power threshold without ZF shearing. In cases with $\mu_0^{(\text{neo})} = 0.1, 0.5$, the power threshold without ZF shearing is found to approach the cases with higher ZF damping, i.e., $\gamma_{\text{damp},0} = 10$. For cases with $\mu_0^{(\text{neo})} > 1.0$ and $\gamma_{\text{damp}} > 5$, the transition does not occur during the power ramp to $Q_a < 0.05$. At least in these cases, the power threshold without ZF shearing increases drastically. This observation supports the hypothesis that the zonal flow is fundamental to the L $\to$ H transition.

An important question here is whether the L $\to$ H transition can occur without ZF shearing. According to the local predator-prey model analyses, without ZF shearing, the L $\to$ H transition can still occur with a sufficiently high level of the heat flux. The QH state exists in any power level of the heat flux and is stabilized above a power which is sufficient to excite the mean flow shear. The L state cannot exist above a certain heat flux, due to strongly excited mean flow. The fundamental role of ZF shearing in the L $\to$ H transition is to reduce the power threshold by reducing the turbulence level. However, with higher ZF damping, there is no significant difference in the turbulence level, with or without ZFs. Therefore, the power threshold with higher ZF damping can be asymptotic to that without ZF shearing.

We study fueling effects on the transition and pedestal structure, with special focus on the particle pinch. In Fig. 14, we plot the calculated power thresholds $Q_{a\text{, crit}}$ with $(V_{\text{th}} = 1.0)$ and without $(V_{\text{th}} = 0.0)$ particle pinch, and with various fuel deposition widths $L_{\text{dep}} = 0.05 - 0.20$. The power thresholds with particle pinch are higher, as compared with those without particle pinch. Either with or without the particle pinch, there is little difference in power threshold for various fuel depositions. To discuss more about the effects of particle pinch, we show profiles of turbulence intensity with and without the particle pinch in Fig. 15. Comparing cases with and without the pinch, there is no qualitative difference in the turbulence profile. However, the turbulence intensities are different because of different $\gamma_L$, as the particle pinch effectively reduces the temperature gradient. This is because the particle pinch tends to steepen the density gradient, and since the pressure gradient is approximately constant, $[L_T]^{-1} = [L_p]^{-1} - [L_n]^{-1}$, the temperature gradient necessarily decreases. As turbulence is thus more weakly excited, ZF

![FIG. 14. Plots of power threshold $Q_{a\text{, crit}}$ for various deposition layer width $L_{\text{dep}}$ with and without particle pinch effects. The power threshold with particle pinch are higher than that without particle pinch. Either with or without the particle pinch, little sensitivity is found in the power threshold for various deposition layer widths.](image)

![FIG. 13. Investigation of power threshold $Q_{a\text{, crit}}$ for various ZF damping rates $\gamma_{\text{damp}}$ as well as for neoclassical poloidal flow viscosities $\mu_0^{(\text{neo})}$. In cases with $\mu_0^{(\text{neo})} \geq 2.0$, no transition occurs for $Q_a < 0.05$. In case of $\mu_0^{(\text{neo})} = 1.5$ and $\gamma_{\text{damp}} > 2.0$, the transition does not occur for $Q_a < 0.05$. In cases with $\mu_0^{(\text{neo})} = 1.0$ and $\gamma_{\text{damp}} > 1.0$, the transition does not occur for $Q_a < 0.05$. In case of $\mu_0^{(\text{neo})} = 0.5$, and $\gamma_{\text{damp}} = \infty$, i.e., without ZF shearing $(x_0 = 0)$, the transition occurs at $Q_a = 0.013$. In case of $\mu_0^{(\text{neo})} = 0.1$, without ZF shearing, the transition occurs at $Q_a = 6.7 \times 10^{-3}$, which power threshold is equivalent to the case of $\mu_0^{(\text{neo})} = 0.1$ and $\gamma_{\text{damp}} = 1.0$.](image)
excitation drops too, so a higher power threshold can be expected. Note that with pinch effects, a much larger core density is also achieved, as compared to the case without a pinch. The higher density may induce higher radiation or ZF damping due to the CX friction. This may further increase the power threshold. Further studies are left to the future, but at least we expect a significant effect of the density pinch on the power threshold.

Here, we discuss effects related to the extent of the fueling deposition width. As is discussed above, in both cases with and without particle pinch, the fueling depth has little effect on the power threshold, as seen in Fig. 14. However, pedestal formation in the early stage of H-mode exhibits different features. In Fig. 16, we show pedestal structure of density profiles in H-mode with and without particle pinch and with various fueling deposition widths \( L_{\text{dep}} \). With particle pinch and increasing the fueling depth, there is little difference in pedestal width, as seen in Fig. 16(a). However, without the particle pinch, an increase in the fueling depth causes an increase of pedestal width, as seen in Fig. 16(b). It is interesting to note that the density gradients in the pedestal are similar to those for fixed fueling source strength, while the core density increases as the pedestal width increases. Inward from the pedestal, density profiles are flat without the particle pinch. One reason why there is little dependence of the fueling width on the particle pinch is that the peaking effect of the particle pinch may be overestimated and thus may overwhelm the fueling.

V. CONCLUSION

We have investigated the space-time evolution of the \( L \rightarrow H \) transition, using a time-dependent, one dimensional mesoscale model which self-consistently describes the evolution of turbulence intensity, ZF, density and pressure profiles, and mean poloidal mass flow. The model captures the essential physics of ZF and MF interaction, turbulence suppression by ZF and MF shearing, and poloidal flow evolution, including that driven by turbulence. We here have elucidated how ZF shearing mediates the transition. These findings are in good agreement with findings from several DBS\(^{11,20,42} \) and probe experiments\(^ {18,41} \) and point to the crucial role of ZFs in the transition dynamics. The specific results of this study are as follows:

I. Studies with a slow power ramp have manifested an \( L \rightarrow H \) transition via I-phase, characterized by a series of nonlinear waves, which, locally, are LCOs. The I-phase nucleates near, but not at the LCFS. In the model, the MF shear peak nucleates at the fixed edge boundary. This result, a consequence of boundary conditions, should be compared with results from DIII-D, where the total \( E \times B \) flow velocity negative well nucleates in the edge region and rises to be positive outside of the LCFS.

II. The LCO period increases abruptly at transition, with rapid growth of MF shear. At the transition, pedestals in density and temperature begin to expand inward. Local turbulence intensity peaks just prior to transition. Mean flow shear growth begins after the onset of I-phase.

III. As is seen in DIII-D, the LCO period increases approaching the transition. The I-phase terminates abruptly at transition, consistent with the 0D Kim-Diamond model\(^ {13} \) and also DIII-D results. The diamagnetic shear oscillates with growing amplitude in I-phase, then increases abruptly at the \( L \rightarrow H \) transition, as seen in DIII-D. The growth of the diamagnetic shear amplitude occurs only in I-phase, not in L-mode. The peak of the ZF shear increases just prior to transition. This is consistent with the analyses of EAST experiments, suggesting that ZF shearing is dominant just prior to the \( L \rightarrow H \) transition.

IV. The phase delay between turbulence and zonal flow increases from \( \pi/2 \) to \( \pi \) during the I-phase, consistent with the 0D Kim-Diamond model\(^ {13} \) and also DIII-D results. The diamagnetic shear oscillates with growing amplitude in I-phase, then increases abruptly at the \( L \rightarrow H \) transition, as seen in DIII-D. The growth of the diamagnetic shear amplitude occurs only in I-phase, not in L-mode. The peak of the ZF shear increases just prior to transition. This is consistent with the analyses of EAST experiments, suggesting that ZF shearing is dominant just prior to the \( L \rightarrow H \) transition.

V. At the \( L \rightarrow H \) transition, the MF first grows and then collapses to a small value as \( L_p^{-1} \) drops. The ZF then peaks and extracts the energy from the turbulence. Then, \( L_p^{-1} \) and \( r_n^{-1} \) increase rapidly, as does the MF. This is consistent with the analyses of EAST

**FIG. 16.** Density profiles in H-mode with various fueling deposition layer width \( L_{\text{dep}} \) and (a) with and (b) without particle pinch. In (a), little difference results from changing the fueling deposition layer width, while in (b), the pedestal width increases as the fueling width increases. Flat profiles and much lower core density quantities are obtained in case without the pinch.
experiments, suggesting that ZF shearing is dominant just prior to the L $\to$ H transition.

VI. The actual transition event is abrupt, even if the power ramp and LCO evolve slowly. After the transition, edge pedestals in $n$, $p$, and also $T$ form quickly. This picture is consistent with the experiments, in which the pedestal expands after the L $\to$ H transition.

VII. Numerical studies reveal that two kinds of the transition are possible. With slow power ramp, L $\to$ H transition occurs via I-phase, which clearly manifests a quasi-periodic oscillation. On the other hand, during a fast power ramp, the I-phase is compressed into a single burst of ZF, leading to a transition without I-phase. Here, how fast the ramp must be so as to prevent I-phase depends on the ramp-up rate as compared to the period of the limit-cycle.

Studies of power threshold with various ZF damping and neoclassical poloidal viscosity indicate that larger neutral CX increases the power threshold. More generally, increasing ZF damping increases the power threshold, suggesting that ZF is fundamental to the transition. ZF can act as “reservoir” in which to store large fluctuation energy without increasing transport, thus allowing the mean flow shear to grow. Mean flow shear, however, ultimately is required to “lock in” the state of quenched turbulence. Therefore, the ZF damping must enter the power threshold condition but does not exclusively determine it. More systematic and quantitative studies of the power threshold require modification of mean flow evolution by SOL-edge interaction incorporating up-down asymmetry of X-point location in the single divertor configuration.44 It is interesting to compare our present results to those of another model, which incorporates SOL and core interaction to formulate criteria that the L $\to$ H transition occurs.45 The Fundamenski model does not address the temporal evolution of the L $\to$ H transition and also posits rather simplified core plasma dynamics. Nevertheless, it recovers certain experimental parameter scalings, including a bifurcation of the power threshold in density. Applying such SOL and core interaction conditions to the present model, we will explore estimation of the power threshold in a future study.

These results also suggest implications for future steady state experiments. We find that neutral CX can damp zonal flows in experiments,46 indicating that high edge neutral density is unfavorable to transition. This can be related to the long established experimental lore concerning the power threshold, “dirty machines,” re-cycling, etc. The results have implication for the recovery of H-mode in steady state operation, if it is lost. In such the case of the recovery of H-mode, wall saturation and consequent increased re-cycling can ultimately lead to strong CX damping of ZFs, making it difficult to recover the H-mode should a back-transition occur. This suggest that proper wall conditioning, or reduction of wall impurity saturation, is necessary throughout the long pulse H-mode operation, because the ZF shearing necessary to trigger the transition may be more difficult to excite.

The trend that higher ZF damping increases the power threshold may explain the experimental scaling trend in the higher density region, because $\nu_{\text{damp}} \propto \nu_x \propto n$. As well, increasing $q$ may make a power threshold higher, since the neoclassical viscosity increases with $q$. $B_T$ may change the power threshold through the change of $\rho_i$. We also note that according to Ref. 47, the simple model with Gyro-Bohm turbulence will give the power threshold $P_{\text{thresh}} \sim B_T^2$, $1 < \alpha < 3$. Thus, the model finds that $B_0$ scaling in the power threshold appears from the $\rho_i$ dependence of the coefficients. However, this result is mode dependent and needs further study.

Investigation of the transition and the pedestal structure, including dependence on density pinch and fueling deposition depth, has aimed to promote the understanding of the particle fueling dynamics. What we have found are the following:

I. The density pinch reduces turbulence intensity and ZF shearing excitation, because the density peaking effect (due to the pinch) enhances reduction of the temperature gradient by the larger density gradient. This causes an upshift of the power threshold. With the particle pinch effect, the density profile can be peaked, and the total particle flux can be negative in the presence of a sufficiently steepened temperature gradient, assuming the ITG turbulence model.

II. The fueling deposition depth has virtually no impact on the L $\to$ H transition, with or without particle pinch. Without particle pinch, deeper deposition makes a deeper density pedestal, while with particle pinch, little effects on the pedestal width are observed. This may be because here the peaking effects overwhelm the fueling effects.

For further parameter surveys to such cases where the external particle flux source and the pinch-induced peaking coexist, we need to expand the turbulence model to electron turbulence, such as CTEM. These will give more detailed physical insights for the fueling study and address such phenomena as gas-puffing, pellet injection, and also supersonic molecular beam injection (SMBI).

We remark on the relation of this study to high $\beta$ plasma physics. We here do not consider electromagnetic physics such as kinetic ballooning mode (KBM) or peeling ballooning mode possibly excited in high $\beta$ plasmas. Therefore, in the present model, we do not see any traditional ELM oscillations. Though KBM could play a role in determining the pedestal structure, KBM should not directly impact the L $\to$ H transition itself, since KBM is excited in higher $\beta$ plasmas close to the ideal limit. For actual L $\to$ H transition physics—as opposed to pedestal physics—lower $\beta$ plasmas are more relevant. Furthermore, we have already reproduced most of the detailed transition dynamics without considering high $\beta$ physics.

In future work, we will study back transitions, i.e., H $\to$ L events. We note here some results, in relation to the ramp speed issue. Model studies of slow power ramp downs indicate that the H $\to$ L transition occurs via an I-phase. Instead, fast ramp downs exhibit a rapid burst of ZF activity, but no clear LCO. These features are similar as the phenomena, seen in the L $\to$ H forward transition. Furthermore, we will discuss effects of noise on the L $\to$ H transition in another work.
L → H transition is fundamentally non-deterministic, as the edge layer bombarded by ensemble of core avalanches, producing large variability in the local heat flux. As edge heat flux variability is induced by core avalanches, the edge heat flux frequency spectrum is better taken to be 1/f, i.e., pink noise, than white. 1/f noise is more effective for producing transitions than white noise, because 1/f noise tends to be more coherent. We will discuss these issues in future papers.

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APPENDIX A: THE DERIVATION OF THE FEEDBACK LOOP OF TURBULENCE, ZONAL FLOW, AND MEAN FLOW SHEARING

The derivation of the model to describe feedback loop between turbulence and ZF is shown here, taking into account multiple shearings such as MF. We start from the well-known stability of the model. Thus, here we do not consider this additional effect, since we consider a minimal model of the feedback loop.

As the GAM in the I-phase is observed in ASDEX Upgrade,16 it is interesting to consider a turbulent state regulated by the GAM there. In the possible GAM and ZF coexistence state discussed in Ref. 49, we expect that no GAM will exist. In such cases that the GAM is relevant, the turbulence spreading induced by the GAM may modulate the propagation speed of the LCO.23 Furthermore, GAM and

1/τq,Ω = |Δ(kv − ωk)| = |(v − vgr)Δk|, (A5)

Thus, for quasi-particles with drift wave phase velocity

v = ωk/k = vth resonating with group propagation of the ZF shearing vgr = ∂Ω/∂qr = Ωqr, we obtain

τq,Ω = |(∂Ω/∂qr(k))Ωqr|−1, (A6)

where Δkq is the typical width of the envelope of the ZF and MF wave packet. Rewriting the mean turbulence energy as I, i.e., I ≡ ⟨⟩ = ∫ dκκ(Nk), we find a temporal evolution equation for turbulence intensity with linear growth and nonlinear damping as well as shearing effects as

∂I/∂t = γI − ΔωI − ∑ qr,Ω XqrΩV2ZFqr,Ω, (A7)

where Xqr,Ω ∼ τq,Ω is the coupling parameter between turbulence and ZFs related to the correlation time of the shears.

Here, we retain two different modes of flows, i.e., mesoscale ZF and large scale mean flow. ZF has a shearing scale qr ∼ 1/√Ωqr, while MF has qr ∼ 1/Λr, since MF shear is determined by the profile gradients. We finally obtain

∂I/∂t = γI − ΔωI − xqIE0 − xqIEv. (A8)

This comes to Eq. (1), including the additional nonlocal diffusion term.

Note that we neglect the geodesic curvature −(2c4/R) in the zonal flow evolution Eq. (A2). If the geodesic curvature includes, we may expect the higher frequency eigenmode of ZF, i.e., the geodesic acoustic mode (GAM). A model incorporating the geodesic curvature and also the sound wave propagation is discussed in Ref. 49. As most experimental results show, however, GAM is not relevant in I-phase and H-mode. Therefore, it is sufficient to neglect the geodesic curvature in the model incorporating the mean flow shearing. A screening factor of the zero-frequency ZF, A0 = (1 + 2q2)−1, in the fluid closure may be included in the evolution of ZF, originating from the parallel nonlinearity. However, the screening factor does not qualitatively change the stability of the model. Thus, here we do not consider this additional effect, since we consider a minimal model of the feedback loop.
ZF are more similar than different. Both are $n = 0$ (no transport) secondary modes excited by three wave interactions via modulational instability. Both shear the turbulence, and so extract energy from it. The major difference is that GAMs have $\omega \sim c_s / R$, while ZFs have $\omega \sim 0$. Also, GAM propagates radially. We rather consider the similar basic structure of two predator (GAM and MF or ZF and MF)-one prey (turbulence) systems. Very similar results can be expected, but we expect additional sensitivity of the collisionless damping to the power threshold.

**APPENDIX B: THE DERIVATION OF THE INHIBITION OF ZONAL FLOW SHEARING BY MEAN FLOW SHEARING**

We discuss how ZF shear shrinks in the presence of the MF shear, based on the wavekinetic treatment. Here, we define $(\vec{V}_E)$ as the mean $E \times B$ shear flow and $V_{ZF}$ as zonal flows with the form of $\exp(i_0 x)$. In the presence of the mean shear flow, the linearized wave kinetic equation for the perturbation $\tilde{N}_k$ and mean $\langle N_k \rangle$ are written as

$$\frac{\partial \tilde{N}_k}{\partial t} + i_0 \vec{v}_p \tilde{N}_k - k_0 \langle \vec{V}_E \rangle \frac{\partial \langle N_k \rangle}{\partial k_r} + \gamma \tilde{N}_k = k_0 V_{ZF} \frac{\partial \langle N_k \rangle}{\partial k_r}. \quad (B1)$$

Here, the effect of a mean shear flow $(\vec{V}_E)$ on $\tilde{N}_k$ is explicitly shown in the third term on the l.h.s. of Eq. (B1). We solve Eq. (B1) along a non-perturbed orbit by introducing a total time derivative $D_t$ as

$$D_t = \frac{\partial}{\partial t} - k_0 \langle \vec{V}_E \rangle \frac{\partial}{\partial k_r}. \quad (B2)$$

In this coordinate, the shearing effect by $\langle \vec{V}_E \rangle$ is explicitly reflected in the linear increase of $k_r$ in time as

$$D_t k_r = -k_0 \langle \vec{V}_E \rangle. \quad (B3)$$

Equation (B1) can be integrated along this nonperturbed orbit as

$$\tilde{N}_k(q_r, t) = \int_{-\infty}^{0} d t' \exp \left\{ -\gamma (t - t') - i_0 \int_{t'}^{t} d t'' \vec{v}_{gr}(t'') \right\} \times k_0 V_{ZF}(\rho, t') \frac{\partial \langle N_k(t') \rangle}{\partial k_r(t')}, \quad (B4)$$

where a term depending on the initial condition is dropped. The shearing effect by a mean flow is embedded in the time dependent group velocity $\vec{v}_{gr}$ and equilibrium wave quanta density spectrum $\partial \langle N_k(k_r(t')) \rangle / \partial k_r(t')$. In the limit where the mean shearing occurs on a time scale larger than other dynamical time scales (i.e., $1/\gamma$, $1/\vec{v}_{gr} q_r$, and $1/\Omega$), we can approximate $\partial \langle N_k(k_r(t')) \rangle / \partial k_r(t') \sim \partial \langle N_k(k(t)) \rangle / \partial k_r(t)$. Note that if we take $\vec{v}_{gr}(t')$ and $k_r(t')$ to be constant, the result is the usual modulational instability without the mean flow shear. However, by taking into account the dependence of $\vec{v}_{gr}$ on $\langle \vec{V}_E \rangle$ through $k_r$, we expect the slowdown of the propagation of drift waves due to enhanced inertia via mean shearing. Then, the substitution of the time dependence of $\exp(-i \Omega t)$ for $\tilde{N}_k$ and $V_{ZF}$ simplifies Eq. (B4) to

$$\tilde{N}_k(q_r, \Omega) \sim k_0 V_{ZF} R \frac{\partial \langle N_k \rangle}{\partial k_r}. \quad (B5)$$

Here, the real part of $R$ becomes

$$Re(R) = -1 \left[ \gamma - \frac{\gamma q_{gr}^2}{\gamma^2 + \frac{2}{\gamma^2} \Omega^2} \right] \quad (B6)$$

for $k_1 \rho_n < 1$ and $\gamma > q_{gr} \chi$, where $\omega_s = k_0 \nu_s = k_0 T_e / e B_0 L_a$ is the electron diamagnetic drift frequency, originating from the group velocity. The sign of $Re(R)$ is always positive since $R$ was obtained by treating the effect of $\langle \vec{V}_E \rangle$ as a small perturbation. From Eqs. (A2) and (B6), ZF growth is also obtained as

$$\Omega \sim i q_{gr}^2 \frac{2}{1 + k_r^2 \gamma^2} R \left( \frac{\partial \langle N_k \rangle}{\partial k_r} \right) \quad (B7)$$

From Eq. (B7), it is clearly shown that a mean flow shear suppresses the growth rate of zonal flows. The reduction arises due to the time variation of $\vec{v}_{gr}$, related to the decorrelation of drift wave propagation by a shear flow, weakening the coherent modulation response of the drift wave spectrum. Thus, this effect is put as $\zeta_0$ in Eq. (2).

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