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A SEARCH FOR NEUTRINO-INDUCED REACTIONS

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May 22, 1963

A Search for Neutrino-Induced Reactions*

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The object of this experiment was to see if an upper limit could be placed on the value of the cross section for the reaction

$$\bar{\nu}_{\mu} + p \rightarrow \mu^{+} + n.$$

Two observations, reported by Cowan in experiments with cosmic-ray neutrinos, may be used to estimate this cross section.¹ A diurnal effect is reported in the counting rate, suggesting that the earth is acting as an absorber per bound proton for the neutrino flux. This would demand a cross section of $\sim 10^{-33} \text{ cm}^2$. A value of $\sim 10^{-28} \text{ cm}^2$ is given by an attenuation measurement in which the rate at the surface of the earth is compared with the rate a known distance below. If there exists such an enhancement over the normal order of magnitude of the cross section for neutrino-induced events, it could be evidence for the appearance of the resonance postulated by Kinoshita² from the intermediate-boson hypothesis of Tanikawa and Watanabe.³ Preliminary results from the experiment of Cowan suggested the possibility of such a resonance at incoming neutrino energies of $\sim 200 \text{ MeV}$.

The 184-Inch Cyclotron should be a reasonable source of 200-MeV neutrinos (decay of a 350-MeV pion gives a 204-MeV neutrino in the forward direction). A rough estimate of the neutrino flux indicates that limits as high as 10^{-28} cm^2 for the cross section for neutrino-induced events can readily be checked with a manageable block of detector placed in the meson cave.

In the present experiment the detector was a 7-by-7-by 3-in. sodium

iodide crystal. The signature for the reaction under investigation was the appearance in the detector of a muon which registered through a delayed coincidence with its decay electron. The position of the crystal in the meson cave is shown in Fig. 1. In addition to the usual shielding in the cave, three feet of steel was placed between the machine and the crystal, which was surrounded on the other five sides by a 1-ft layer of steel. At first no anti-coincidence counter was used to veto events due to cosmic rays, since, in a preliminary run away from the machine, the delayed coincidence rate from cosmic-ray muons stopping in the crystal was about one per minute. It was believed that by using the machine on pulsed operation (64 beam pulses per sec, each ~ 300 μ sec in width) and gating the detector, a factor of ~ 50 could be gained over the cosmic-ray background. However, the rate of accidentals proved to be so great that we had to abandon this method of approach and use the auxiliary dee to improve the duty cycle. Under these conditions, some 50% of the intensity appears in a "beam spike" (~ 300 μ sec), but the remaining intensity is spread out uniformly in time over a "beam stretch" (~ 10 msec). There is a second but comparatively weak "beam spike" at the end of the "beam stretch."

The sodium iodide crystal was viewed by nine (6655) phototubes. Essentially, if two pulses occur in the crystal within 5 μ sec, then an event is registered and their time separation measured from a photograph of the oscilloscope trace. The trace shows two electron pulses separated by a fixed delay (~ 6 μ sec) with the muon pulse between them. A plot of the accumulated data from cosmic rays is shown in Fig. 2a, illustrating the capability of the apparatus to detect muons. It should be noted that π^+ mesons would be registered as muons.

During the experiment an anticoincidence counter was placed over the sodium iodide crystal in an effort to reduce the cosmic-ray background. However, because of the arrangement of the heavy steel shielding, it could not be placed in an optimum position, but some reduction in rate was observed.

EXPERIMENTAL RESULTS

Runs With No Anticoincidence Counter

Rates of events for runs with no anticoincidence counter are:

Events rate with a 10-msec gate on the stretched beam $0.56 \pm 0.05/\text{min}$

Events rate from cosmic rays (beam off) $0.85 \pm 0.03/\text{min}$

Calculated background rate due to cosmic-ray events

during the machine run = $64 \times 10^{-2} \times 0.85$ $0.54 \pm 0.02/\text{min}$

Events rate with an 8-msec gate on the stretched beam $0.38 \pm 0.04/\text{min}$

Calculated background rate from cosmic rays $0.43 \pm 0.01/\text{min}$

The time distribution of the delayed coincidences collected during these machine runs is shown in Fig. 2b. It is evident that the distribution has a lifetime very close to that of the positive muon, but the rate of events is consistent with that of cosmic rays viewed with the duty cycle of the cyclotron. The gate width was reduced from 10 to 8 msec to avoid the second "spike" in the intensity distribution of the stretched beam.

Runs With an Anticoincidence Counter

Rates of events for runs with an anticoincidence counter are:

Events rate with an 8-msec gate on the stretched beam $0.20 \pm 0.02/\text{min}$

Events rate from cosmic rays (beam off) $0.30 \pm 0.02/\text{min}$

Calculated background rate from cosmic rays $0.15 \pm 0.01/\text{min}$

The time distribution is shown in Fig. 2c. The events appear to be

muons, and their rate is really consistent with that expected from cosmic rays. The errors quoted in all cases are from counting statistics only and are therefore somewhat optimistic.

Interpretation of Results

Suppose we assume that some events are produced by the machine with a rate equal to $(0.2 \text{ minus } 0.15) = 0.05 \pm 0.02/\text{min}$. These events could be accidentals, π^+ mesons produced by neutrons, or muons produced by neutrinos. The flux of neutrons in the sodium iodide crystal has not been estimated, but estimates of the accidental rate and the neutrino flux are given below.

The Accidental Rate

The events rate during a run with a 0.5-msec gate on the beam "spike" was $0.74 \pm 0.07/\text{min}$. The calculated cosmic-ray background for these conditions was only $\sim 0.03 \text{ events/min}$, so the accidental rate was $\sim 0.7 \text{ events/min}$. If we assume that the intensity of the beam in the "spike" roughly equals that in the "stretch," then in going from a 0.5-msec gate on the "spike" to an 8-msec gate on the "stretch," the accidental rate should be reduced by a factor of $(16)^2$, giving 0.003 events/min .

An Estimate of the Neutrino Flux

The complicated geometry of the internal target of the cyclotron and the effects of the fringe field make a simple estimate of the flux difficult. The approach used finally was to take the known flux of high-energy π^- mesons in the meson cave and assume the following:

- a. The transmission efficiency of the magnet system after the internal quadrupole was unity, i. e., the flux in the cave gave the number of pions of

a given energy entering a known solid angle (60% of the aperture of the internal quadrupole).

b. The percentage contamination of muons in the beam was used to give a measure of the fraction of pions that would give neutrinos into an acceptable solid angle.

The π^- fluxes and muon contaminations used in this work are 310 MeV, 4.5% muons; ⁴374, 417, and 454 MeV, 5% muons; ⁵and 250 MeV, 8% muons. ³The estimated flux of neutrinos with energy greater than 180 MeV entering the solid angle of the sodium iodide crystal is $\sim 4 \times 10^3$ /sec. This estimate is believed to be accurate to about a factor of 2, and is purposely conservative since the conditions for accepting neutrinos are not as stringent as those for muons, where the momentum requirements of the beam must be satisfied.

Estimate of Possible Limits on the Cross Section and Boson Mass

The result of the present experiment may be interpreted in terms of an upper limit on the value of the cross section for neutrino-induced events, or a lower limit on the mass of the intermediate boson. If we assume that the events rate of 0.05 ± 0.02 /min is different from zero and due to neutrino-induced events, then we can estimate these limits.

The number of events expected in the 3-in. -thick sodium iodide crystal is

$$(4.3 \times 10^{23} \times \text{flux/sec} \times \sigma) \text{ per min.}$$

No attempt has been made to correct for inefficiencies in the detection system (for instance, a 5- μ sec gate was used to look for decaying muons), since these errors are negligible compared with that introduced in the neutrino flux. The experiment suggests that σ is less than $4 \times 10^{-32} \text{ cm}^2$ per bound proton.

The limit on the boson mass may be placed by a method of analysis like that used by Kinoshita. Because of the motion of the protons in the nucleus, neutrinos with a broad band of energies can react with the resonant cross section for a given boson mass. Therefore the effective cross section, σ_e , is given by the cross section at resonance, σ_R , reduced by a factor, $\Gamma_R/\Delta E$, where Γ_R is the width of the resonance, and ΔE the width of the band of neutrino energies.

Figure 3 shows a plot of the expected events rate for different masses of the intermediate boson. The end point of the neutrino spectrum from the 184-in. cyclotron is ~ 250 MeV, and neutrinos with this energy in collision with a stationary proton would produce a boson of mass equal to $2270 m_e$. However, with the momentum distribution in the nucleus, higher boson masses may be attained, but only a small fraction of the protons can participate, so the rate of events falls off rapidly.

Because of the low energy of the neutrinos produced at the 184-in. cyclotron, only a rather conservative limit of $2130 m_e$ can be placed on the mass of the intermediate boson.

We would like to thank Professor Luis Alvarez for suggesting this measurement and showing a keen interest in its progress, and also Professor Clyde Cowan for communicating his results before their publication. Our thanks are due Mr. Howard Goldberg, Professor Robert Kenney, and Mr. James Vale and the crew of the cyclotron, without whose full cooperation the run would not have been possible. We are also grateful to Messrs. Philip Beilin, Ned Dairiki, and Robert Shafer for their help in running the experiment.

FOOTNOTES AND REFERENCES

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1. Clyde L. Cowan, Catholic University, Washington, D. C.; invited paper W.2, "Cosmic Neutrinos," Washington American Physical Society Meeting, April 1963, and further private communication.
2. Toichino Kinoshita, Phys. Rev. Letters 4, 378 (1960).
3. T. Tanikawa and S. Watanabe, Phys. Rev. 113, 1344 (1959).
4. Hugo R. Rugge, Lawrence Radiation Laboratory Report UCRL-10564¹⁰²³⁻²
May 20
~~November 15~~, 1962 (unpublished).
5. Richard J. Kurz, Lawrence Radiation Laboratory Report UCRL-10564
November 15, 1962 (unpublished).
6. Howard Goldberg, Lawrence Radiation Laboratory, private communication.

FIGURE LEGENDS

Fig. 1. Geometrical arrangement of sodium iodide crystal with respect to the 184-Inch Cyclotron target.

Fig. 2. Time distributions of events in the sodium iodide crystal.

- (a) Cosmic-ray muon data;
- (b) machine on, no cosmic-ray anticoincidence;
- (c) machine on, cosmic-ray anticoincidence.

Fig. 3. Expected events rate in sodium iodide crystal for different values of the intermediate boson mass. Here $A(0.05 \pm 0.02)$ is the point from the present experiment if (beam-minus-cosmic-ray background) gives the neutrino-induced events. Errors are from counting statistics only.

Fig. 1.

MU-30572

MUB-1843

Fig. 2.

Fig. 3.

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